

13 May 2019 | 1

Herschel & AGN – 10 years after Peter Barthel, Kapteyn Institute, Groningen Herschel Mission Scientist 1997-2017

Take-away message:

AGN (i.e., accreting MBHs) are basic (boring/ important/fascinating ...) ingredients of starforming galaxies, throughout cosmic time.





First of all, let us recall Rowan-Robinson (1995): "an AGN is inevitably accompanied by a starburst but the reverse is not true", and look at the SED of a prototypical starburst galaxy, LIRG M82 (having SFR ~ $10M_{\odot}/yr$).





13 May 2019 | 3

SFR(z) on Main Sequences: SFR~ $M_*^a(1+z)^b$, with a~0.8 and b~2.5





13 May 2019 | 4

Herschel: 90% of the high redshift star-forming galaxies are on the MS, and "merger-starbursts" represent a small fraction.







But what about AGN hosts?

As central black hole mass scales with host bulge mass, starformation and black hole growth must somehow go hand-in-hand.



McConnell & Ma 2013



13 May 2019 | 6

Moreover, MBH accretion over cosmic time, n_{QSO}(z) (colors, scaled up), runs more or less like the cosmic SF history (black line).

Pre-Herschel JCMT/IRAM mmphotometry indeed detects thermal emission in certain distant AGN =>

NEED THE SENSITIVITY, RESOLUTION, AND WAVELENGTH COVERAGE OF HERSCHEL TO STUDY LARGE SAMPLES



Madau + Dickinson 2014



Key question: are AGN in regular star-forming galaxies?



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m R}$ Annu. Rev. Astron. Astrophys. 52:589–660

AGN: efficient vs. inefficient accretion



For instance, the Sombrero galaxy and M87 (Virgo A) – ultra-massive, dust-poor galaxies – are inefficient accretors, starving for fuel, while radio-loud quasars, Cygnus A, and radio-quiet QSOs are dusty, efficient accretors, happy consumers. Obscuring tori in the efficient class are key element in unification scenario's.



13 May 2019 | 8

NGC1068 and Mkn231 are famous AGN-starburst symbiotic cases: what about all the others?









Shao+10 and Rosario+12: X-ray AGN host galaxies provide extra-strength FIR emission.

Using SED composition, the Leipski/Meisenheimer team found the same in very high z SDSS QSOs.





AGN, of both Type-1 and Type-2

4

3

↓₂ N

1

10¹



13 May 2019 | 11

The beauty of spectroscopy: SPIRE FTS line diagnostics (VanderWerf+10), calling for XDRs; plus the wealth of (PACS, SPIRE, HIFI) outflow data









X-ray selected AGN - on MS, with redshift evolution (Mullaney+12, and other studies).







13 May 2019 | 13

Now the all-sky monsters: ultra-powerful (3C) FR2 radio sources, à la Cygnus A, in very massive hosts – GT and OT studies by the Hergé-, Tadhunter-, and Barthel-teams.





Typical 3C radio galaxy and QSR SEDs, indicative of prodigious (200 – 900 $\rm M_{\odot}/yr$) starbursts in 1/3 of the sample.

But … still on the Main Sequence (massive hosts!)





... and the average z>1 3C Herschel SEDs, supporting orientation unification, including the aspect-dependent warm torus emission.





Podigachoski+15



Comparing the N(SFR) distribution of the 3C AGN with non-AGN (in the same host mass range) we see no sign of negative AGN feedback whatsoever. Note also the SFR-distribution of SDSS QSOs (Ma&Yan15, Pitchford+16), going through the ceiling (but watch out!)





13 May 2019 | 17

We observe stronger starbursts in smaller (younger) 3C radio sources: jet-induced star-formation, i.e., positive feedback? TBC.





Podigachoski+15



Further quantification of 3C radio galaxy host properties, using the Pégase-3 SED code, in an IAP collaboration (Podigachoski+16).

Other studies of the YSPs (Young Stellar Populations) are done by the Tadhunterteam. More clever work is needed to constrain the timing/impact of the events.





So, yes, AGN live in regular star-forming/bursting galaxies

Take-away image (C.Carreau)





Outstanding (!) issues

- 1. Quantifying the circumnuclear starbursts, and the related positive/negative feedback: location, size, strength, timing w.r.t. AGN, duration, winds, ..
- 2. AGN outflows as negative feedback: nature, strength
- 3. Torus models/observations



1. The starburst issue

- a. Hergé-team (f.i. Falkendal+18) and Tadhunter/Morganti teams make progress re timing/duration and feedback
- b. Starburst-driven superwinds in distant QSOs are observed as associated absorption in optical-uv, over ~kpc dimensions (à la Mkn231) – Barthel+17
- c. Lutz-team (e.g., Lutz+18) has made clever size estimates (~kpc) of the circumnuclear starbursts by deconvolving the PACS images
- d. ALMA imaging by Barthel-team indeed finds sub-kpc dimensions, and positive feedback (Barthel+18)



Ad b: another take-away image



Ad d: 3C298 (SFR=930, strong CIV absorption) imaged by ALMA at 0.15arcsec resolution

- The 1mm emission is \sim 75% non-thermal and \sim 25% thermal
- The thermal emission resides in an asymmetric plateau
- Its morphology implies jet-ISM interplay, on the sub-kpc scale

See Barthel+18; similar results in 3C318 and 3C454.0 (June 2019 issue of ESO Messenger).



HST: Chiaberge+17



Barthel+18



2. The AGN outflow issue

Bottom line: I am still confused as to the magnitude of pure AGN feedback, that is to say, it's global time-integrated importance – please update/convince me. I also have trouble understanding how it would quench star-formation on the kpc scale: radio jets – yes, but accretion disk radiation or winds – no.



3. The torus models: see f.i. Zhuang+18





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13 May 2019 | 26



Concerning AGN and their hosts:

Herschel explored

ánd

Herschel delivered



13 May 2019 | 27

N.B. To my surprise there is still a big Herschel telescope !





RGO - Greenwich



13 May 2019 28



Finally, quoting Friedrich Wilhelm Herschel:

"They (the stars) are now seen to resemble a luxuriant garden which contains the greatest variety of productions in differing flower beds..." "... every stage through which the plant passes is brought to our view.."

"There are two kinds of happiness or contentment for which we mortals are adapted: the first we experience in thinking and the second in feeling"



13 May 2019 | 29

I FEEL PROUD TO HAVE BEEN PART OF IT!!



Herschel Science Team (2009, a few months before the launch)

n.b. Herschel was #7 in the Time Magazine list of greatest inventions of the year 2009

