

PLANCK 2014

THE MICROWAVE SKY IN
TEMPERATURE AND POLARIZATION





Dick Bond **CITA** on behalf of the Planck collaboration

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Mapping *the Planck 2014 Early Universe*

Phonon Spectrum $\zeta = \ln a|_{\rho} = \ln \rho|_a / 3(1+\langle w \rangle)$

*Acceleration Histories $\varepsilon(Ha) = 3(1+w)/2 = -d \ln \rho / d \ln a / 2$ &
Inflaton Effective Potential $V(\phi)$ Reconstruction*

Maps = (radical) **compressions** of the *time ordered information Tol* onto
a parameterized space: *Linear maps, Quadratic maps (power), cosmic parameter maps*
a *Map* is an *ensemble* = *mean-map* + *fluctuation-maps*, encoding correlated errors



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Mapping the Planck 2014 Early Universe

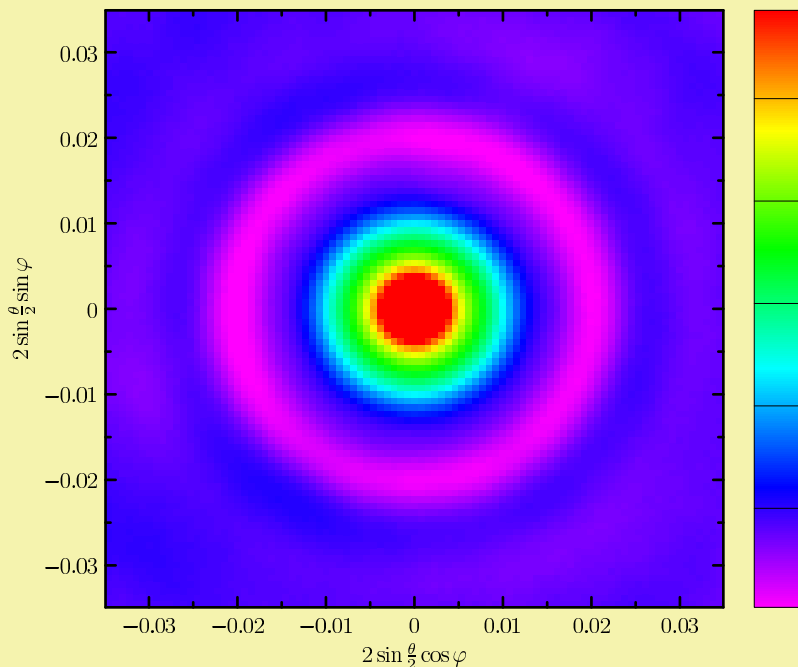
Phonon Spectrum $\zeta = |\ln a|_\rho = |\ln \rho|_a / 3(1 + \langle w \rangle)$

Planck 2014 Isotropy and Statistics paper: extensive stacking of T, Q, U, E, ζ_{dv} on $T, P^2 = Q^2 + U^2, E, \zeta_{dv}$ peaks as in Martinez-Gonzalez, Huang/Frolov talks; & Wandelt

stacking $\Rightarrow \delta \zeta_{dv}$ destructive interference $\Rightarrow$ reveals $\langle \zeta_{dv} | T\text{-peak} \rangle, \langle \zeta_{dv} | \zeta_{dv}\text{-peak} \rangle$

stacked
 $\langle \zeta_{dv} | \zeta_{dv}\text{-pk} \rangle$

20857 patches on ζ maxima, random orientation, threshold $\nu=0$



ζ stacks of
P13 & WMAP9
look v. similar

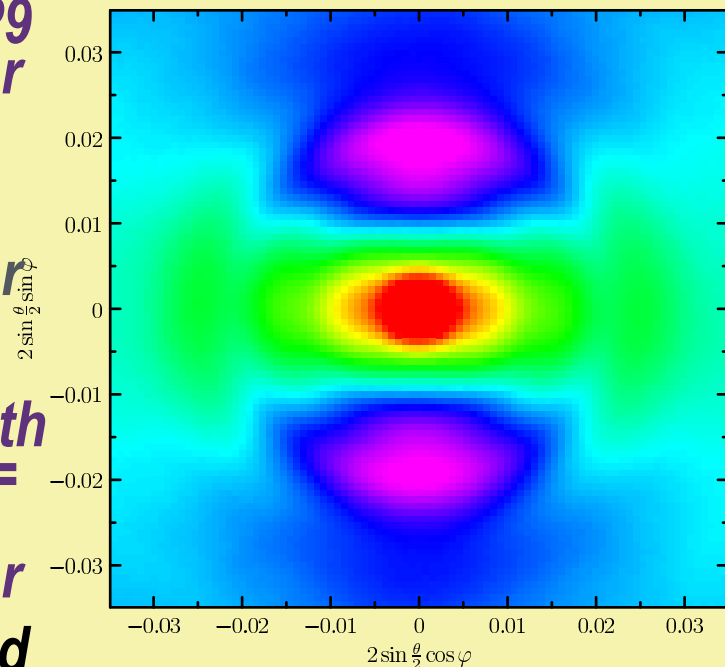
FFP8
simulations
look v. similar

stack one
FFP8 map with
fluctuations =
FFP8 mean
look v. similar

not de-lensed

stacked
 $\langle \zeta_{dv} | \text{oriented } \zeta_{dv}\text{-pk} \rangle$

20854 patches on ζ maxima, oriented, threshold $\nu=0$





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Mapping the Planck 2014 Early Universe

Phonon Spectrum $\zeta = \ln a|_{\rho} = \ln \rho|_a / 3(1+\langle w \rangle)$

phonon $\sim \zeta_{NL} = \ln(\rho a^{3(1+w)}) / 3(1+w) \sim \nabla^{-2}$ scalar curvature @ iso-density

Inflation = phenomenology of **phonons** = energy-density quanta

inflaton = “condensate” of phonon fluctuations, $\langle \rho | k < H a \rangle + \delta \rho$ oscillations
relativistic negative-pressure Equation of State $(1+w)$

phonon = collective mode composed of fundamental scalar fields (many ϕ_b ?)
in linear perturbation theory, the phonon = linear combination of fundamental scalars

Geometrical view: theory of isotropic (volume) strain, Trace (Strain) $\sim$ phonons
& gravity waves Transverse Traceless (Strain)

all that CMB+LSS can deliver is this phonon/strain wave **Inflation** Phenomenology
how does it fit into a UV-complete theory (ultra-high energy to the Planck scale) strings, landscape, ..
& **IR-complete theory** (post-inflation heating $\rightarrow$ quark/gluon plasma)??? TBD



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$\rho(\phi_b, \pi_b, \ln a) \Rightarrow$ coarse-grained $k < H a$ **Hamiltonian-density attractor** $\rho(\phi_b) = 3M_P^2 H^2$
 $d\phi_b / d\ln a = -M_P^2 \nabla_{\phi_b} \ln \rho$, a **gradient / Morse flow** $\Rightarrow$ **Hamilton-Jacobi eqⁿ**,
 “adiabatic” fluctuations along the Morse flow (phonons) **isocurvature directions** $\perp$ the flow

Geometrical view: theory of **isotropic** (volume) **strain**, **Trace (Strain)** $\sim$ **phonons**
 & **gravity waves** **Transverse Traceless (Strain)**

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 how does it fit into a **UV-complete theory** (ultra-high energy to the Planck scale) strings, landscape, ..
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Maps = (radical) **compressions** of the **time ordered information Tol** onto a parameterized space: *Linear maps, Quadratic maps (power), cosmic parameter maps*
a Map is an ensemble = mean-map + fluctuation-maps, encoding correlated errors

Quadratic ζ -Maps aka Radical $\mathcal{P}_\zeta$ Compressions via *expansion in knot-centered modes*

compression onto primordial power spectrum bandpowers
 $\Leftrightarrow$ **compression onto C_L bandpowers** e.g., the nuisance-marginalized Gibbs-sampled grand unified CMB-only bandpowers, bandcentres = knots, bandshapes = modes

Parameters of the $\mathcal{P}_\zeta$ Quadratic Mode Expansion

which function to expand/scan $u = \ln \mathcal{P}_\zeta(k) / \mathcal{P}_{\zeta\text{fid}}(k) = \sum \mathbf{q}_b \mathbf{e}_b(\ln k)$

which modes $\mathbf{e}_b(k)$ (cubic spline, linear, B-splines, Chebyshev, ...)

how many modes aka **knots** (until feature convergence, 12 OK, 16 ~same, 8 too few)

what are the priors/measures on amplitudes $\mathbf{q}_b$ (uniform, independent)

since the C_L^{TT} anomalies of WMAP, we have tried $n_s(\ln k)$, $\epsilon(\ln H a)$, $\ln H(\ln H a)$ as well, $\Rightarrow$ priors change, but we get qualitatively similar features. **Planck 2014 clarifies features & their story.**

fiducial choice $\ln \mathcal{P}_{\zeta\text{fid}}(k) = \ln A_s + (n_s - 1) \ln(k/k_{\text{pivot}})$, $n_s \equiv 0.967$ ($m^2 \phi^2$) (also $\mathcal{P}_{\zeta\text{fid}} = A_s$)

maxL solutions are (quadratic) Wiener-filtered maps! with Fisher/Hessian errors

here we use MCMC, show the mean $\langle \ln \mathcal{P}_\zeta(k) | TT, TE, EE, \dots \rangle$ trajectory &

$\langle \ln \mathcal{P}_\zeta \rangle + \delta \ln \mathcal{P}_\zeta \pm 1\sigma$ trajectories to illustrate ensemble variance & coherence

scan $\mathbf{q}_b$, $\ln A_s$, $r(k_{\text{pivot}, t} = .05/\text{Mpc})$ ($n_t = -r/8$); consistency $\Rightarrow$ reconstruct $\epsilon(\ln H a)$, $V(\phi)$

Quadratic $\ln \mathcal{P}_\zeta(\ln k)$ Maps aka Radical Compressions
 $\Rightarrow$ ultra-early Universe sound/phonon spectrum

Preliminary 12 knots, cubic spline

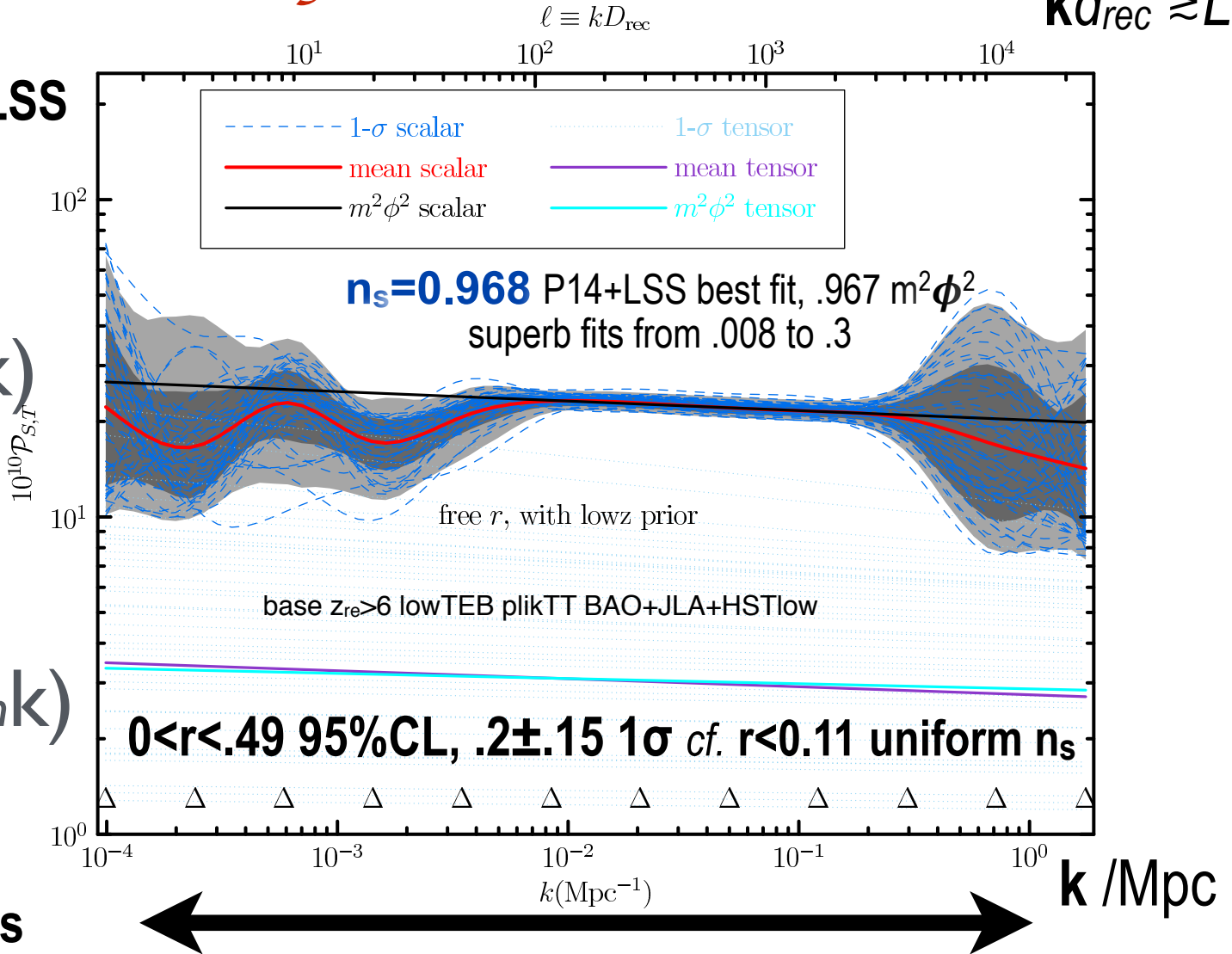
$kd_{\text{rec}} \gtrsim L$

Planck14+LSS

$\ln \mathcal{P}_\zeta(\ln k)$

r - $\mathcal{P}_\zeta$ partial
degeneracy
if r floats

$\ln \mathcal{P}_{\text{GW}}(\ln k)$



Quadratic $\ln \mathcal{P}_\zeta(\ln k)$ Maps aka Radical Compressions
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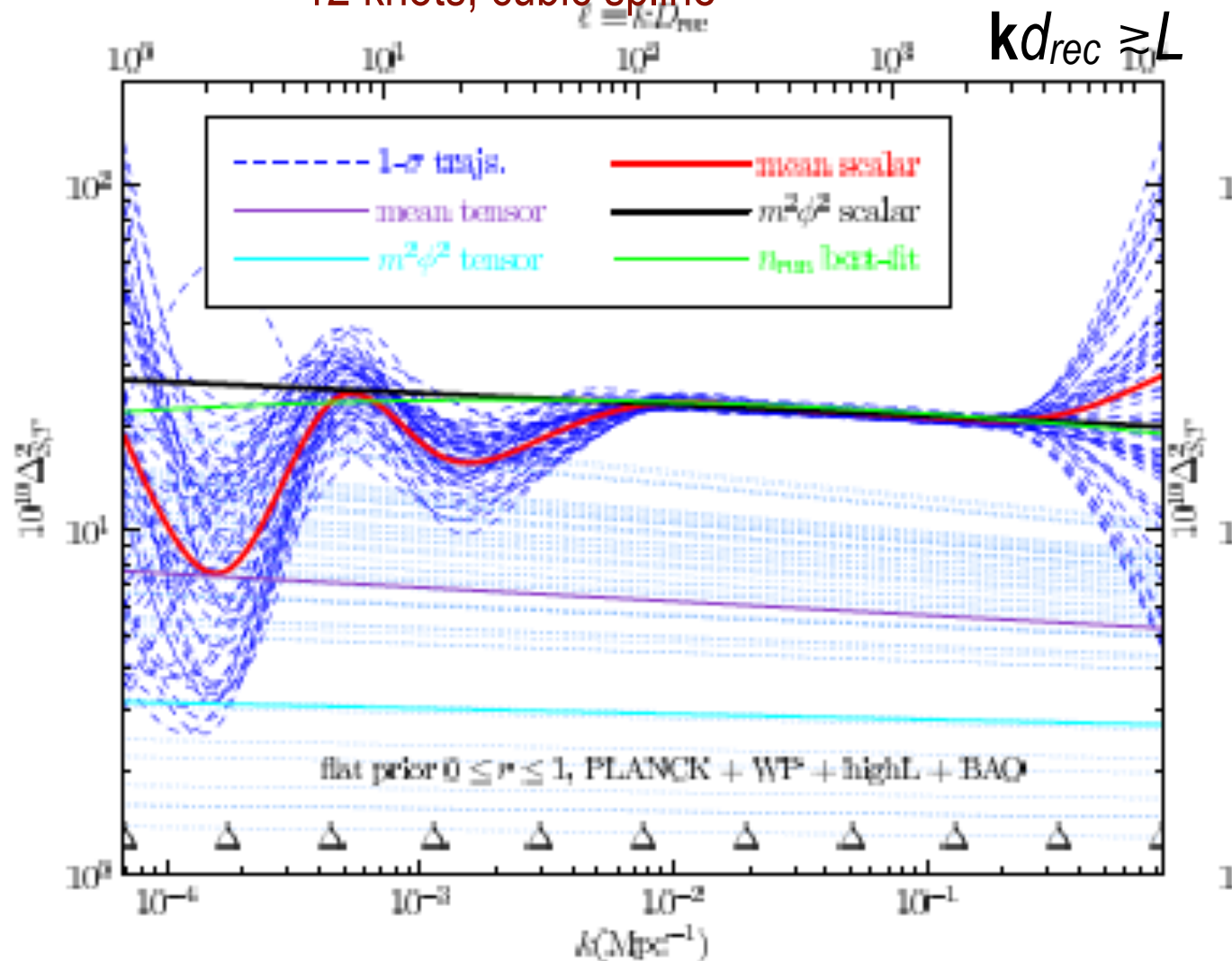
12 knots, cubic spline

cf. Planck13+LSS

$\ln \mathcal{P}_\zeta(\ln k)$

Planck13
& WMAP
 $\Rightarrow$ stable
features

$\ln \mathcal{P}_{\text{GW}}(\ln k)$



9 e-folds

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Preliminary 12 knots, cubic spline

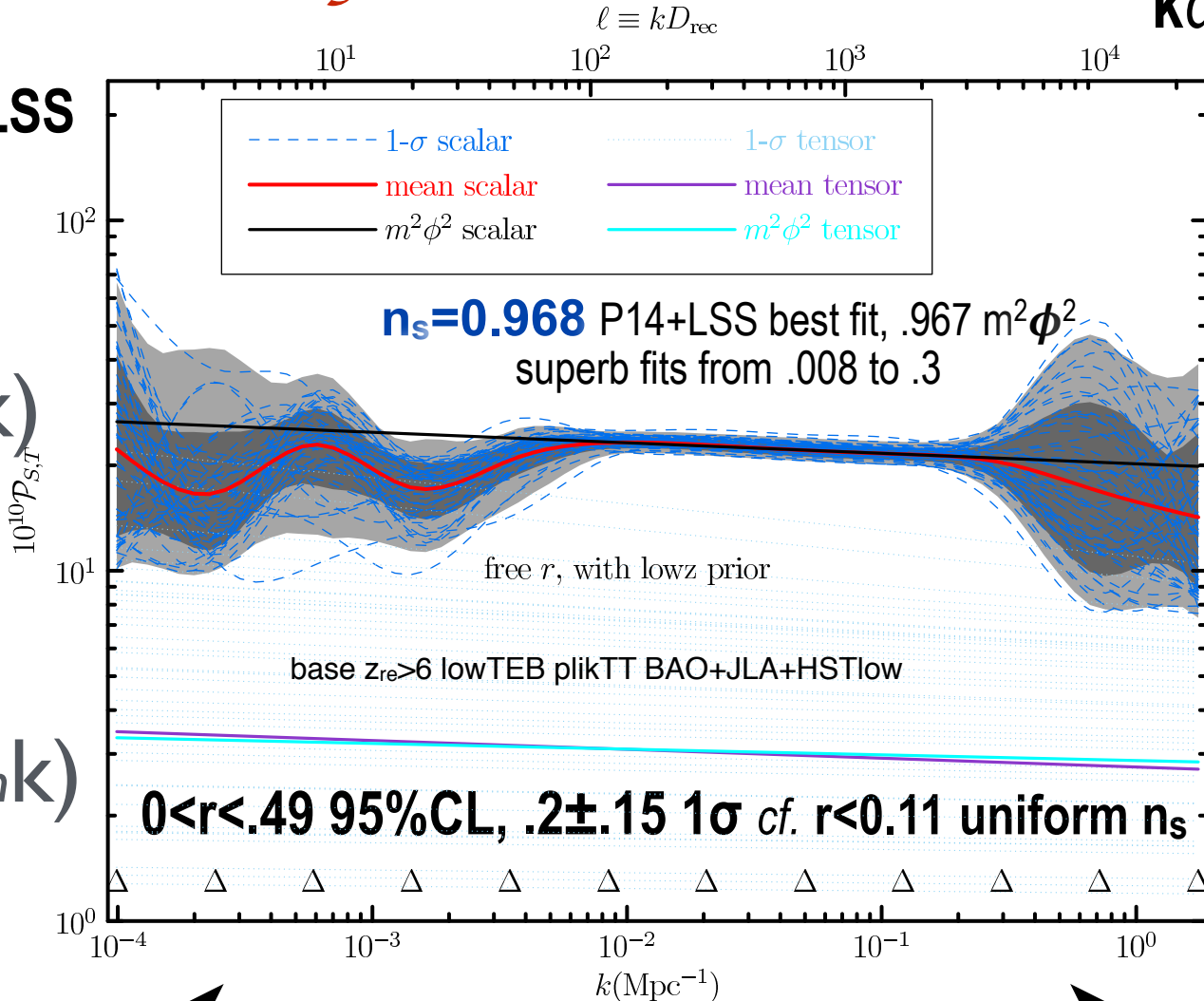
$$kd_{\text{rec}} \approx L$$

Planck14+LSS

$$\ln \mathcal{P}_\zeta(\ln k)$$

Planck13
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 $\Rightarrow$ stable
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Quadratic $\ln \mathcal{P}_\zeta(\ln k)$ Maps aka Radical Compressions
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Preliminary 12 knots, cubic spline

constrain $r \equiv 0.1$

Planck14+LSS

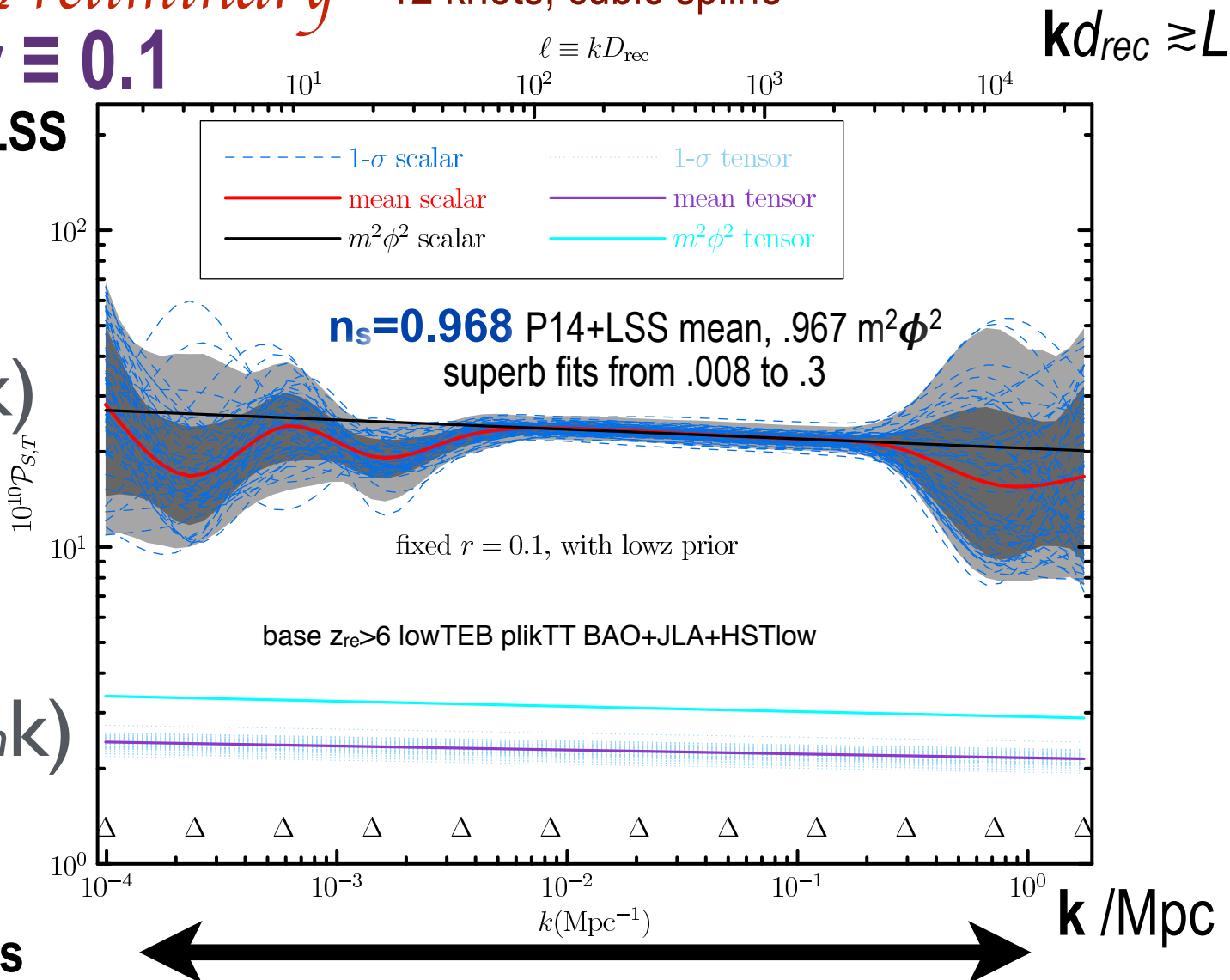
$\ln \mathcal{P}_\zeta(\ln k)$

future $r=0.1$
 constraint
 break

degeneracy $\Rightarrow$
 stable features

$\ln \mathcal{P}_{\text{GW}}(\ln k)$

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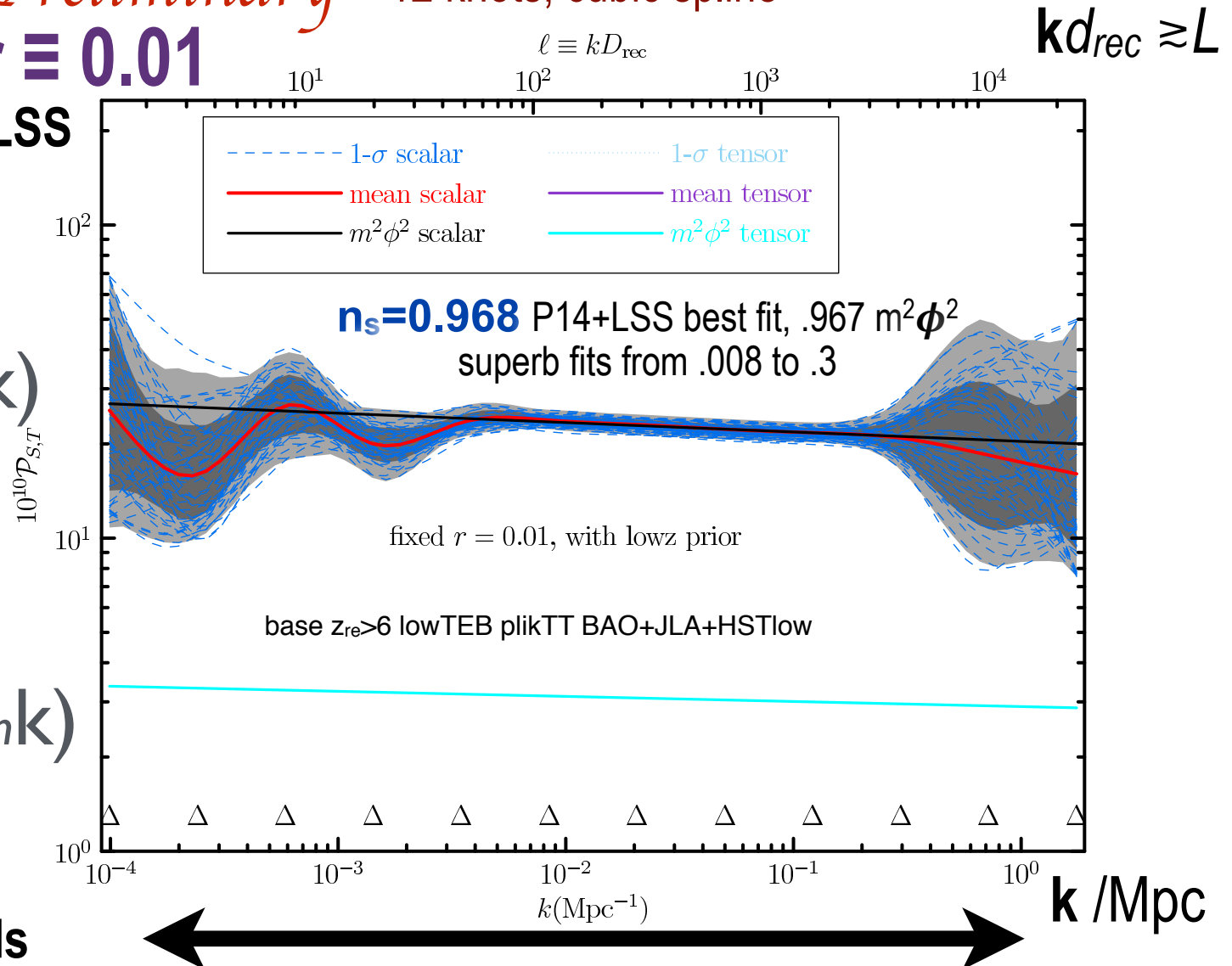
constrain $r \equiv 0.01$

Planck14+LSS

$\ln \mathcal{P}_\zeta(\ln k)$

future $r=0.01$
 constraint
 break
 degeneracy $\Rightarrow$
 stable features

$\ln \mathcal{P}_{\text{GW}}(\ln k)$



Quadratic $\ln \mathcal{P}_\zeta(\ln k)$ Maps aka Radical Compressions
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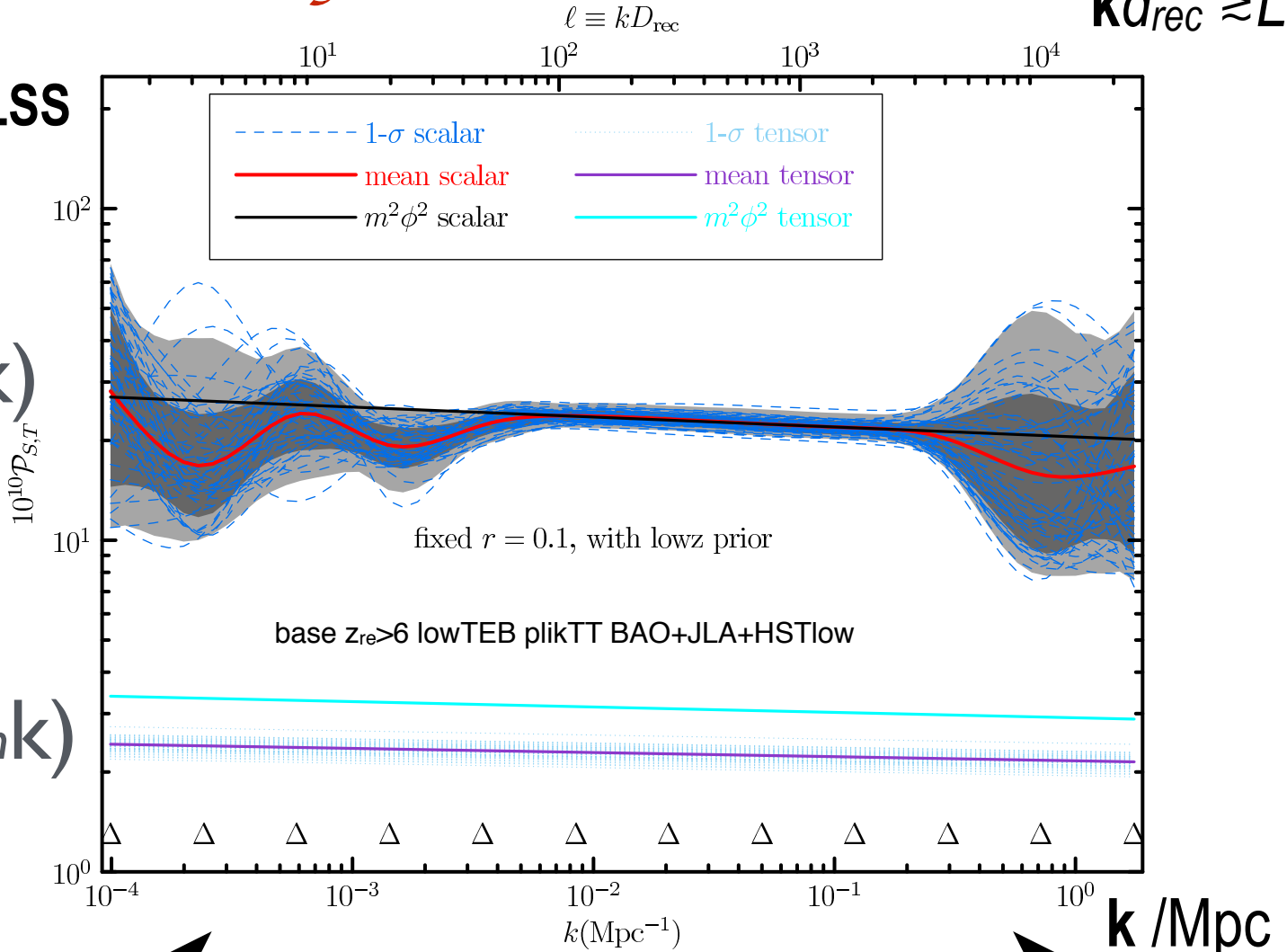
$kd_{\text{rec}} \gtrsim L$

Planck14+LSS

$\ln \mathcal{P}_\zeta(\ln k)$

future $r=0.1$
 constraint
 break
 degeneracy $\Rightarrow$
 stable features

$\ln \mathcal{P}_{\text{GW}}(\ln k)$



Quadratic $\ln \mathcal{P}_\zeta(\ln k)$ Maps aka Radical Compressions $\Rightarrow$ ultra-early Universe sound/phonon spectrum

Preliminary

12 knots, linear

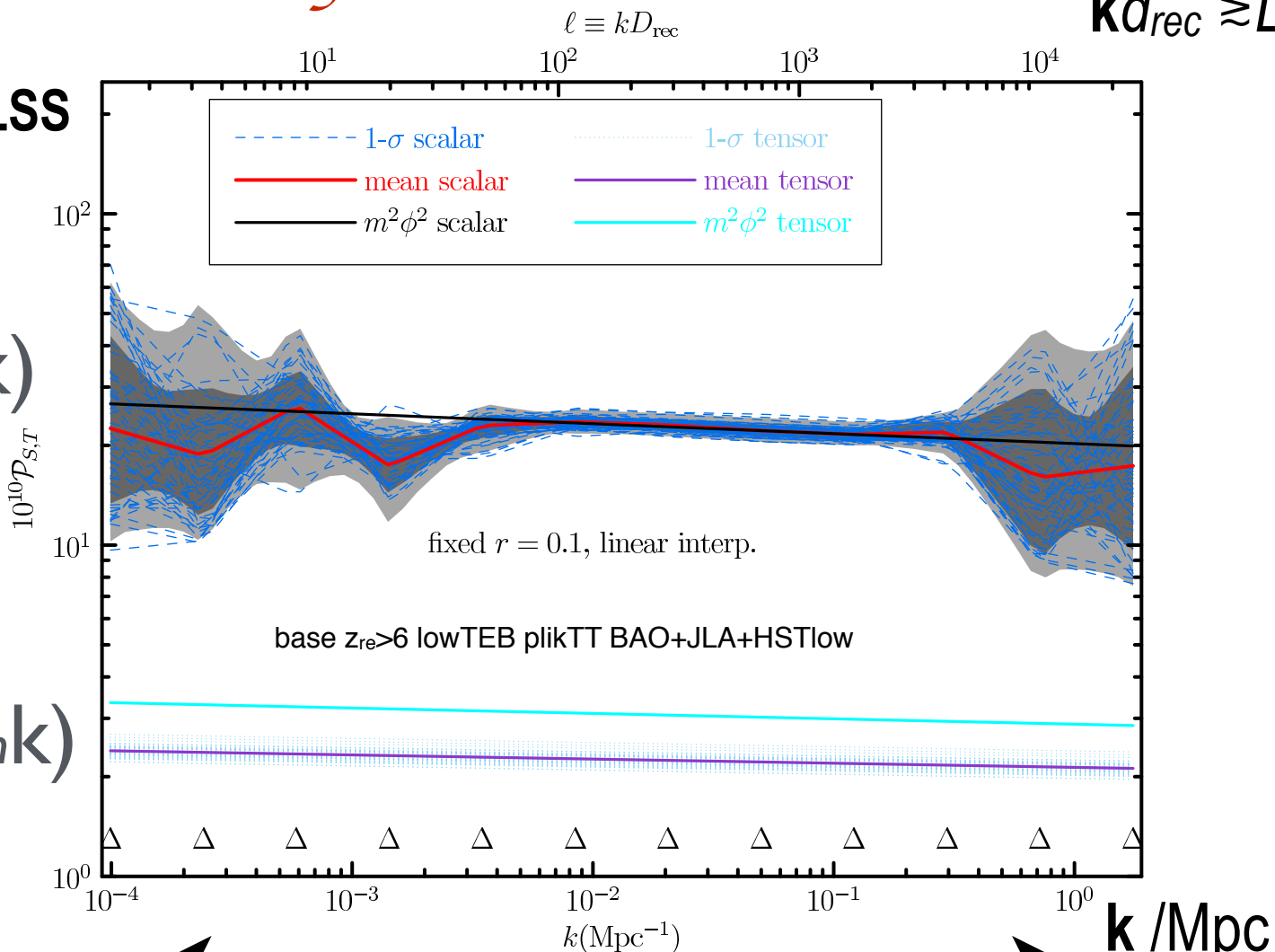
$kd_{\text{rec}} \gtrsim L$

Planck14+LSS

$\ln \mathcal{P}_\zeta(\ln k)$

mode
function
change to
linear $\Rightarrow$
stable
features

$\ln \mathcal{P}_{\text{GW}}(\ln k)$



9 e-folds

Quadratic $\ln \mathcal{P}_\zeta(\ln k)$ Maps aka Radical Compressions
 $\Rightarrow$ ultra-early Universe sound/phonon spectrum

Preliminary 8 knots, cubic spline

$$k d_{\text{rec}} \gtrsim L$$

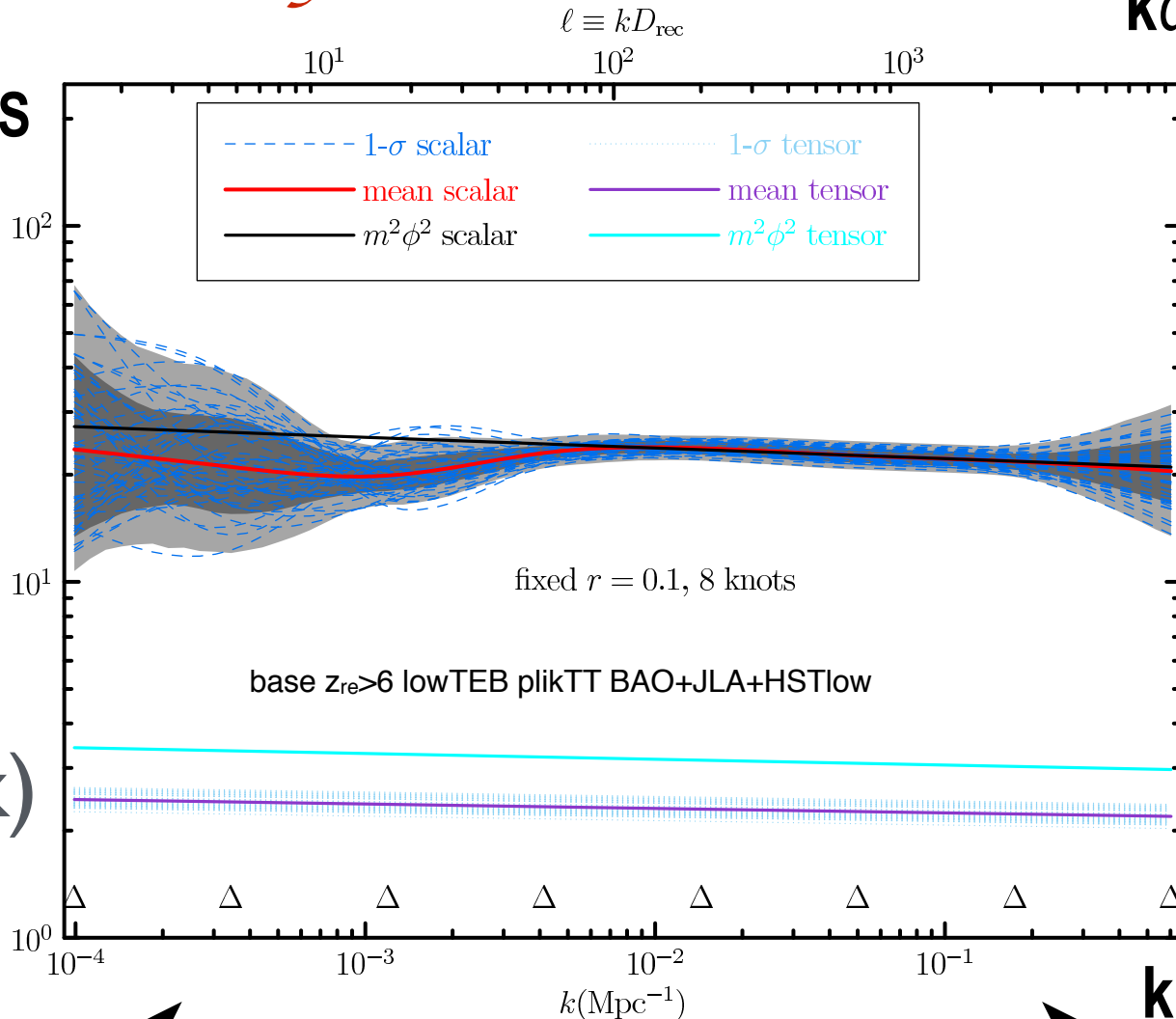
Planck14+LSS

$$\ln \mathcal{P}_\zeta(\ln k)$$

$$10^{10} \mathcal{P}_{S,T}$$

8 knots too
few for stable
features

$$\ln \mathcal{P}_{\text{GW}}(\ln k)$$



9 e-folds

k / Mpc

Quadratic $\ln \mathcal{P}_\zeta(\ln k)$ Maps aka Radical Compressions
 $\Rightarrow$ ultra-early Universe sound/phonon spectrum

Preliminary 12 knots, cubic spline

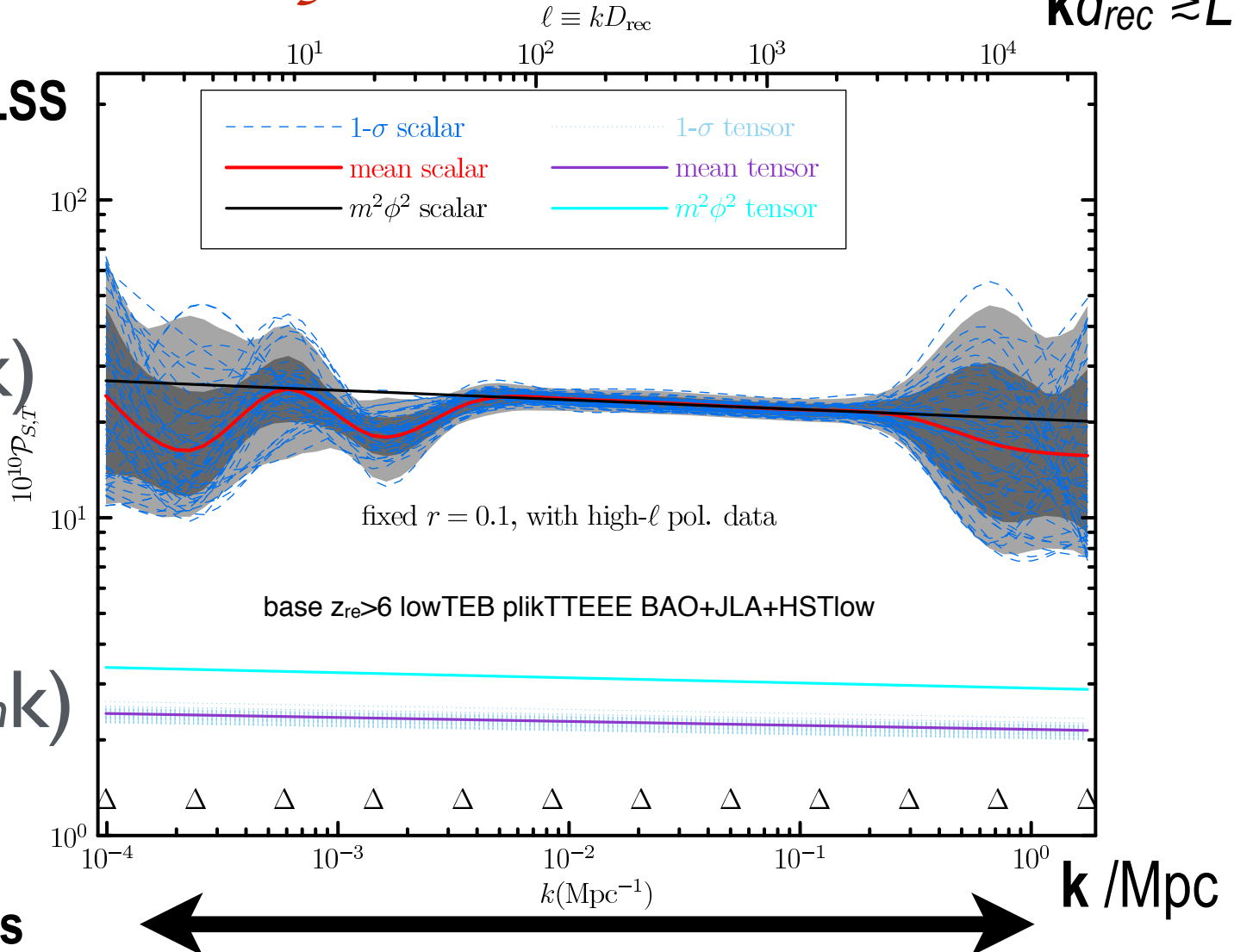
$kd_{\text{rec}} \gtrsim L$

Planck14+LSS

$\ln \mathcal{P}_\zeta(\ln k)$

adding high L
polarization
 $\Rightarrow$ stable
features

$\ln \mathcal{P}_{\text{GW}}(\ln k)$



Quadratic $\ln \mathcal{P}_\zeta(\ln k)$ Maps aka Radical Compressions
 $\Rightarrow$ ultra-early Universe sound/phonon spectrum

Preliminary 12 knots, cubic spline

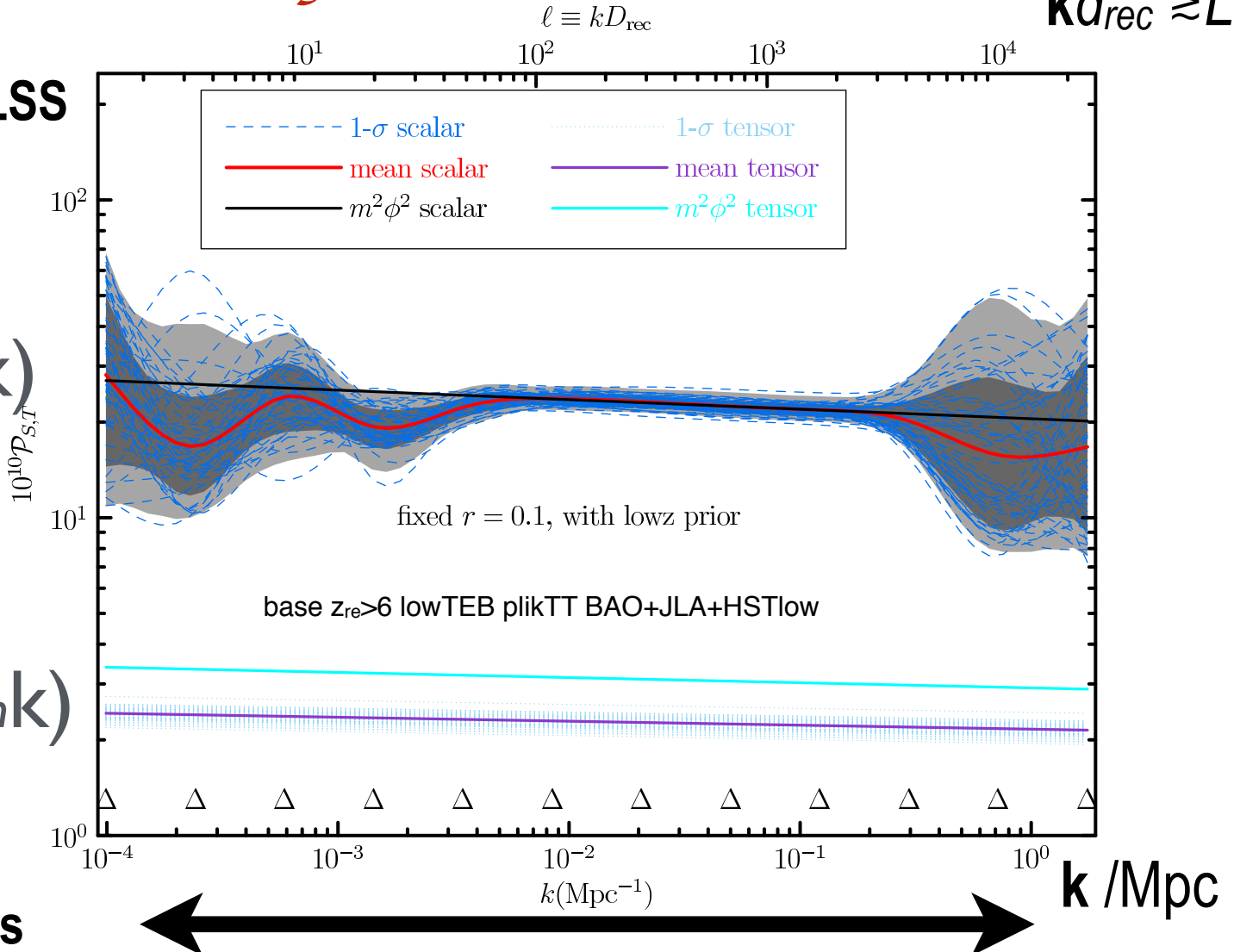
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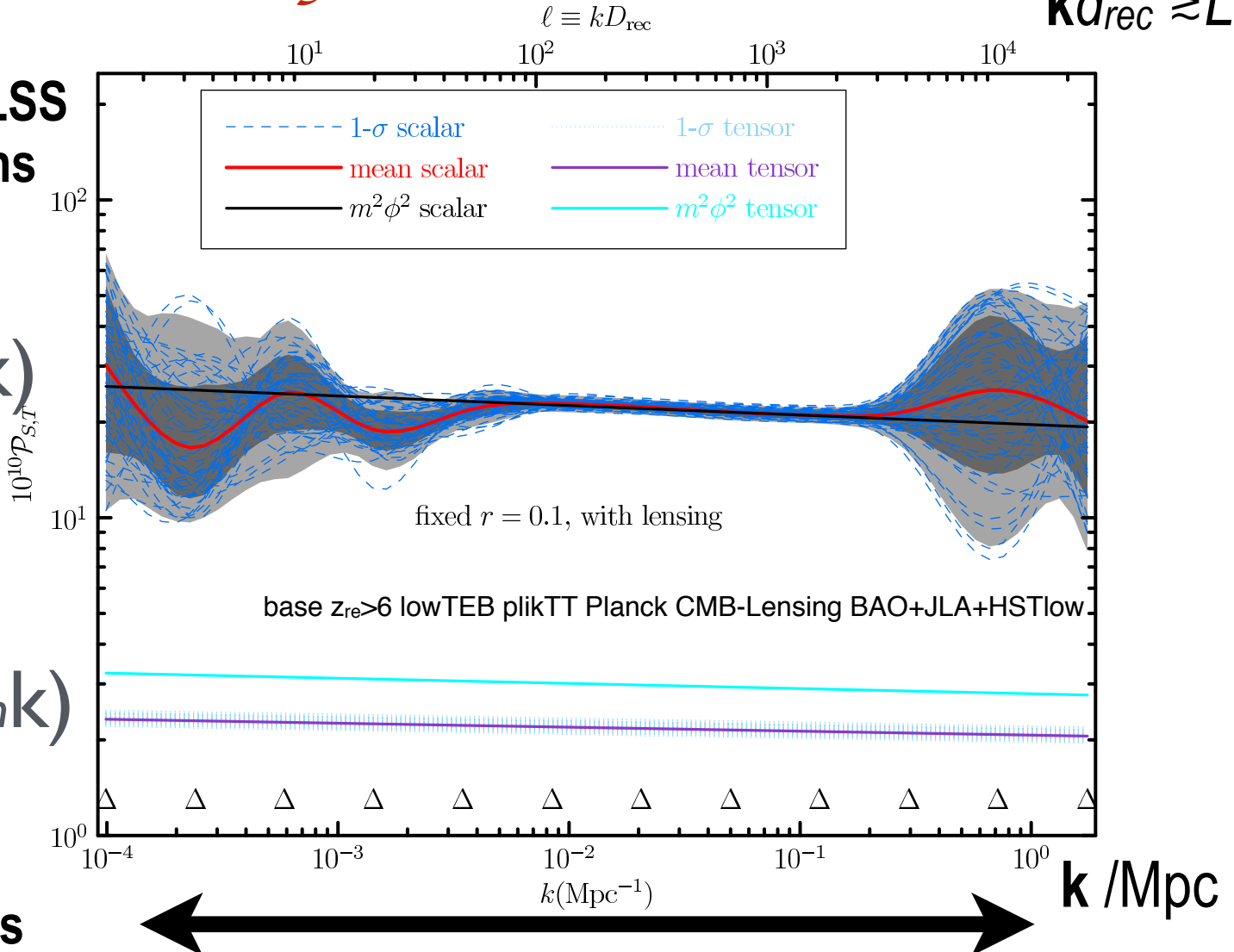
$kd_{\text{rec}} \gtrsim L$

Planck14+LSS
 +CMB Lens

$\ln \mathcal{P}_\zeta(\ln k)$

adding CMB
 lensing $\Rightarrow$
 stable
 features

$\ln \mathcal{P}_{\text{GW}}(\ln k)$



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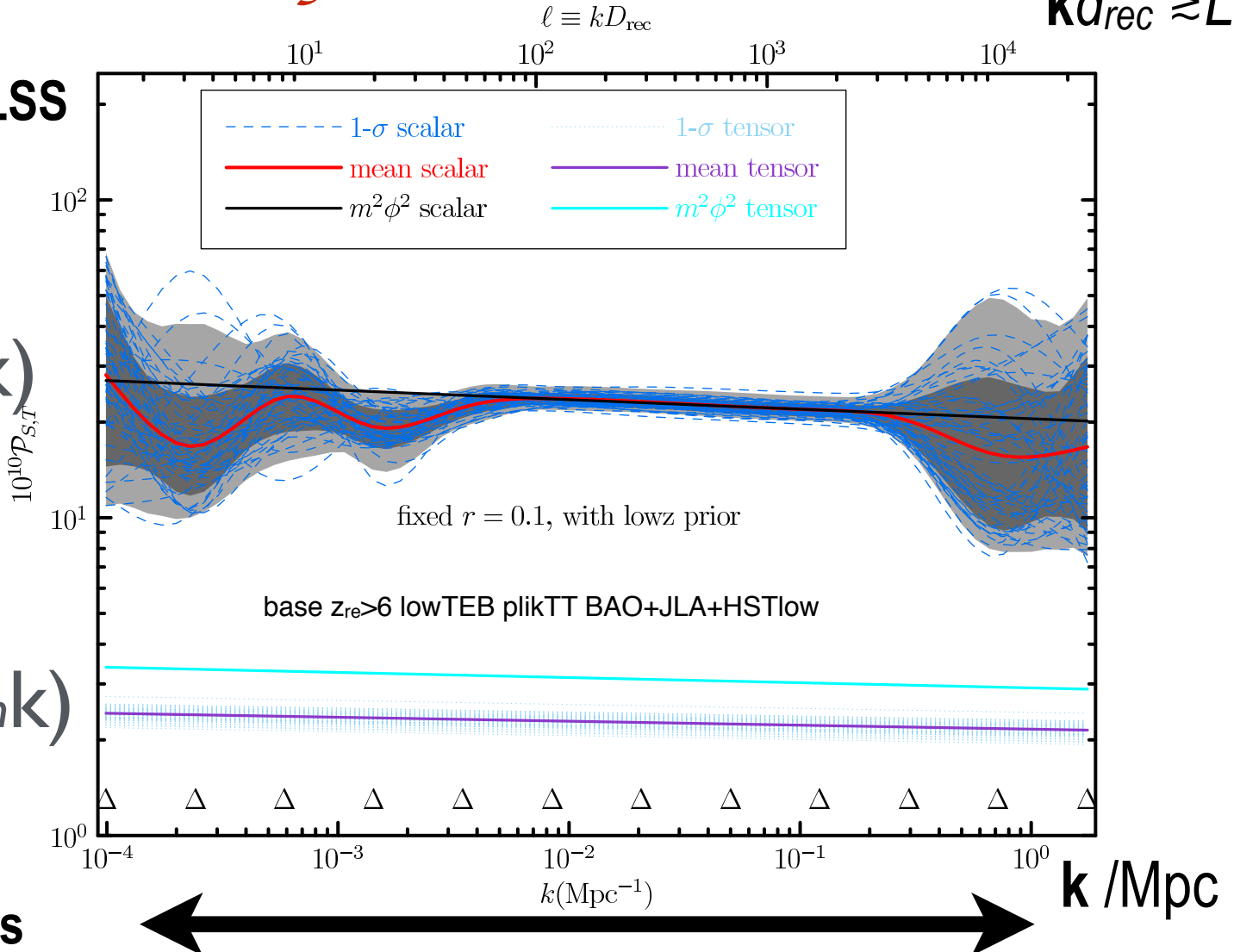
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Planck14+LSS

$\ln \mathcal{P}_\zeta(\ln k)$

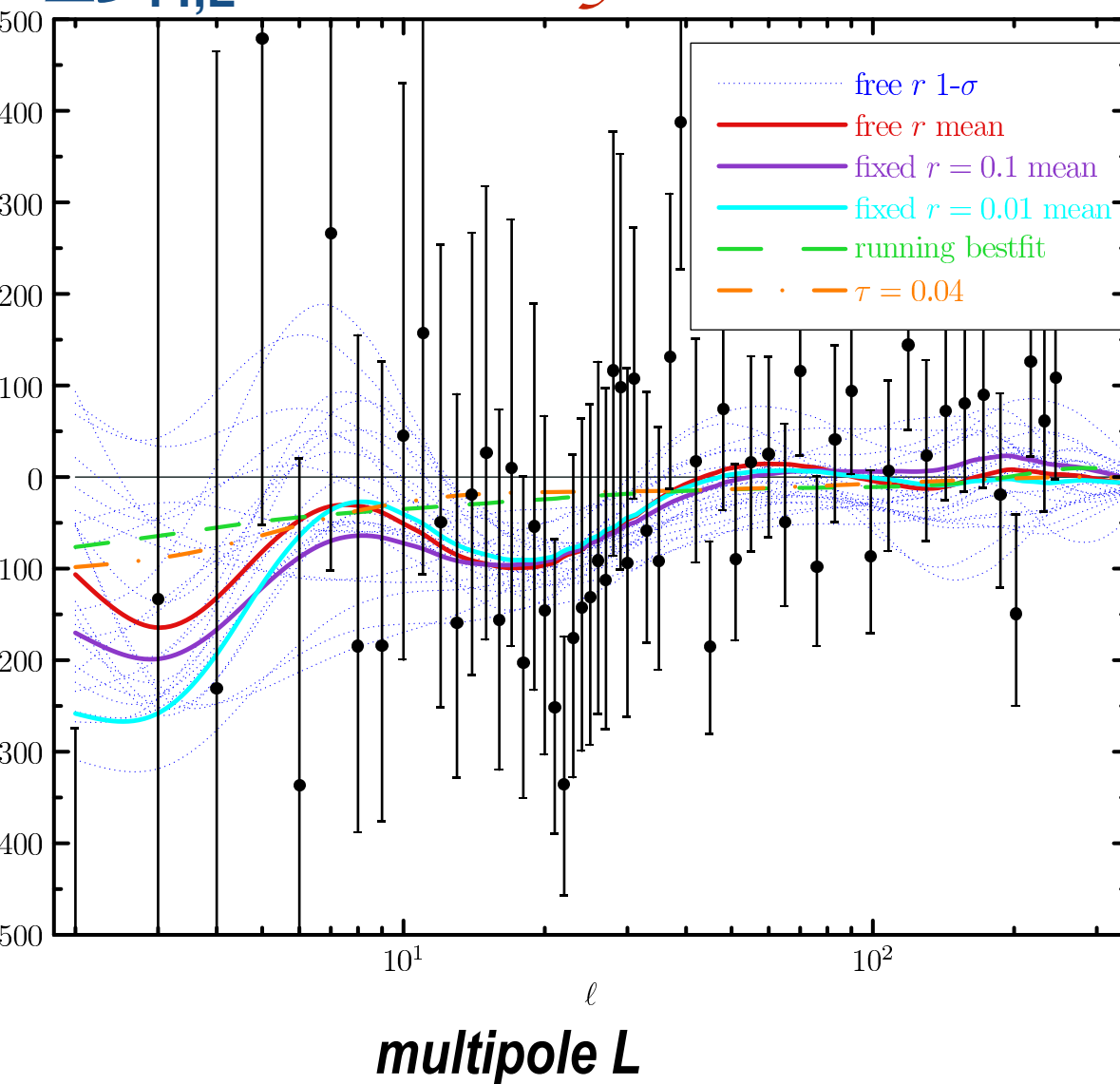
adding CMB
lensing $\Rightarrow$
stable
features

$\ln \mathcal{P}_{\text{GW}}(\ln k)$



trajectories of $\mathcal{D}_{\text{TT},L}$ cf. Planck 2014 Commander Low L spectrum with Blackwell-Rao errors

$\Delta\mathcal{D}_{\text{TT},L}$ Preliminary 12 knots, cubic spline



running of $\mathcal{P}_\zeta$
 $\equiv$ 3 Chebyshev modes
 $\Rightarrow$ very stiff

$\Rightarrow$ not what the data wants
Lower $\tau \Rightarrow$ shape similar to
running at low L

similar response on $\mathcal{D}_{\text{TT},L}$
for constrained & free r
modified by τ freedom

Quadratic $\ln \mathcal{P}_\zeta(\ln k)$ Maps aka Radical Compressions
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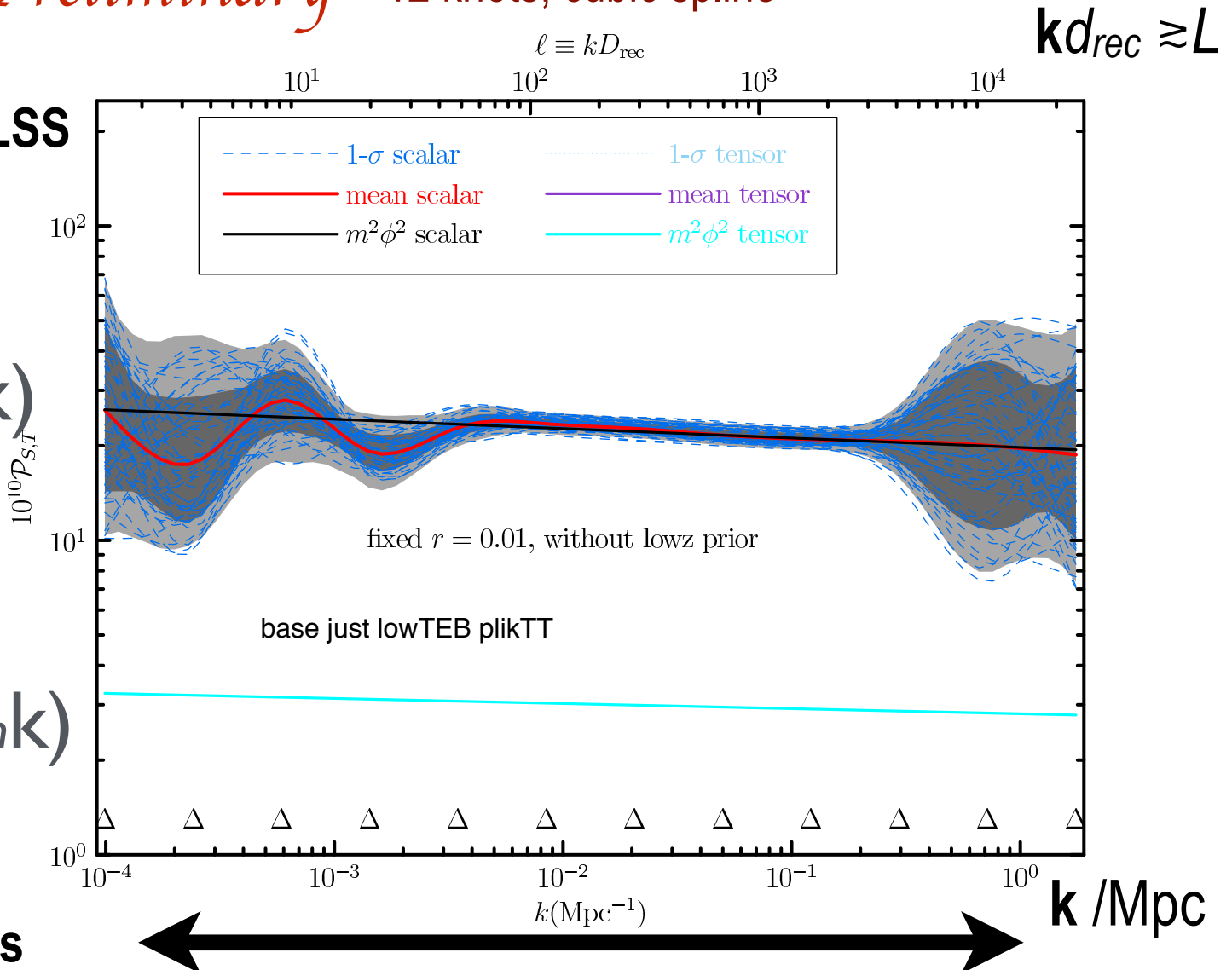
Preliminary 12 knots, cubic spline

Planck14 no LSS

$\ln \mathcal{P}_\zeta(\ln k)$

drop $z_{\text{re}} > 6$ &
LSS, stable
features

$\ln \mathcal{P}_{\text{GW}}(\ln k)$



Quadratic $\ln \mathcal{P}_\zeta(\ln k)$ Maps aka Radical Compressions $\Rightarrow$ ultra-early Universe sound/phonon spectrum

Preliminary 12 knots, cubic spline

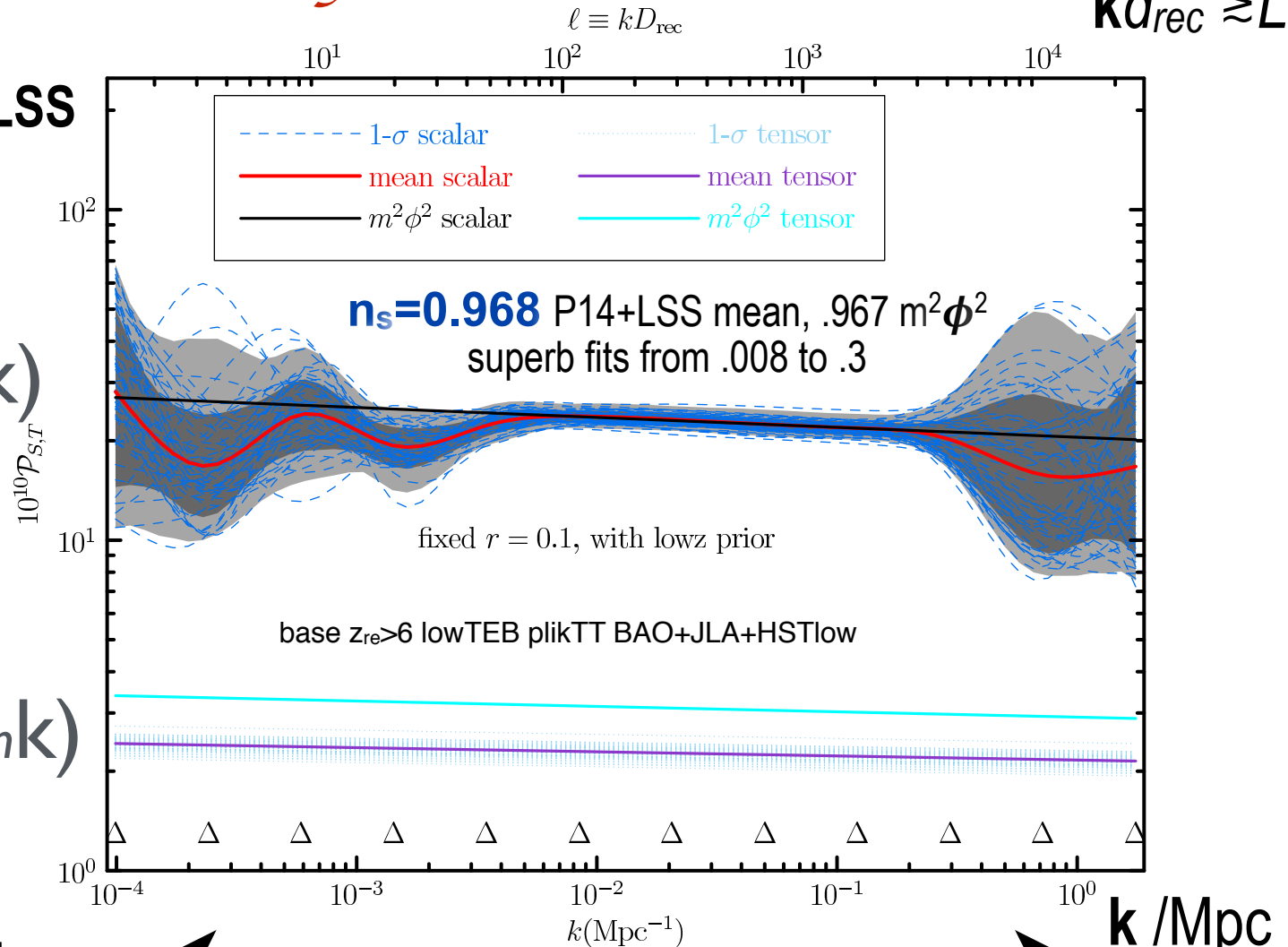
$kd_{\text{rec}} \gtrsim L$

Planck14+LSS

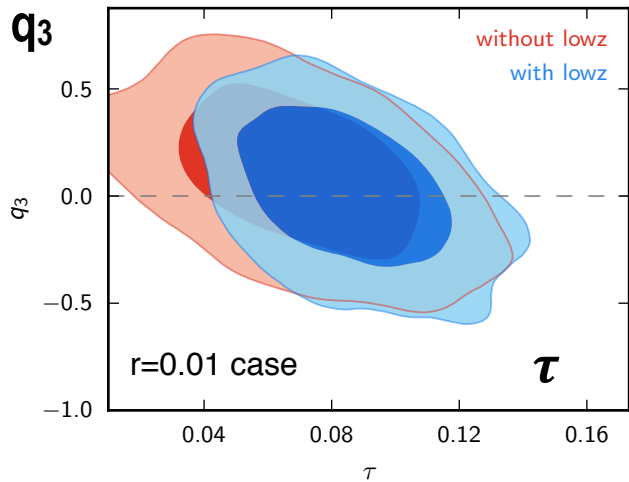
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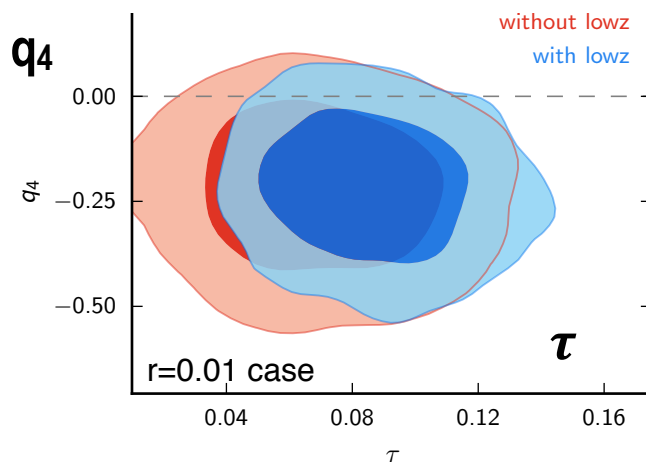
9 e-folds



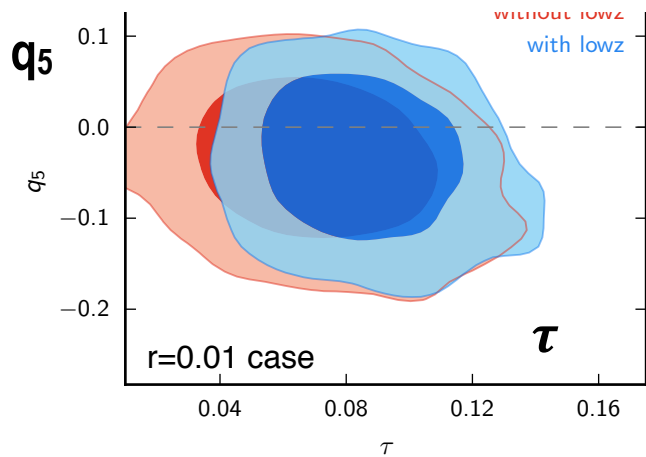
$$q_b = \ln[\mathcal{P}_\zeta(k_b) / A_s (k/k_p)^{(n_s-1)}] \text{ at knot } k_b$$

τ -dependence
creates the slight
low L bump

Preliminary



L~20-30 downturn
creates the q_4 dip
no τ -dependence
beyond since
before / beyond
reionization



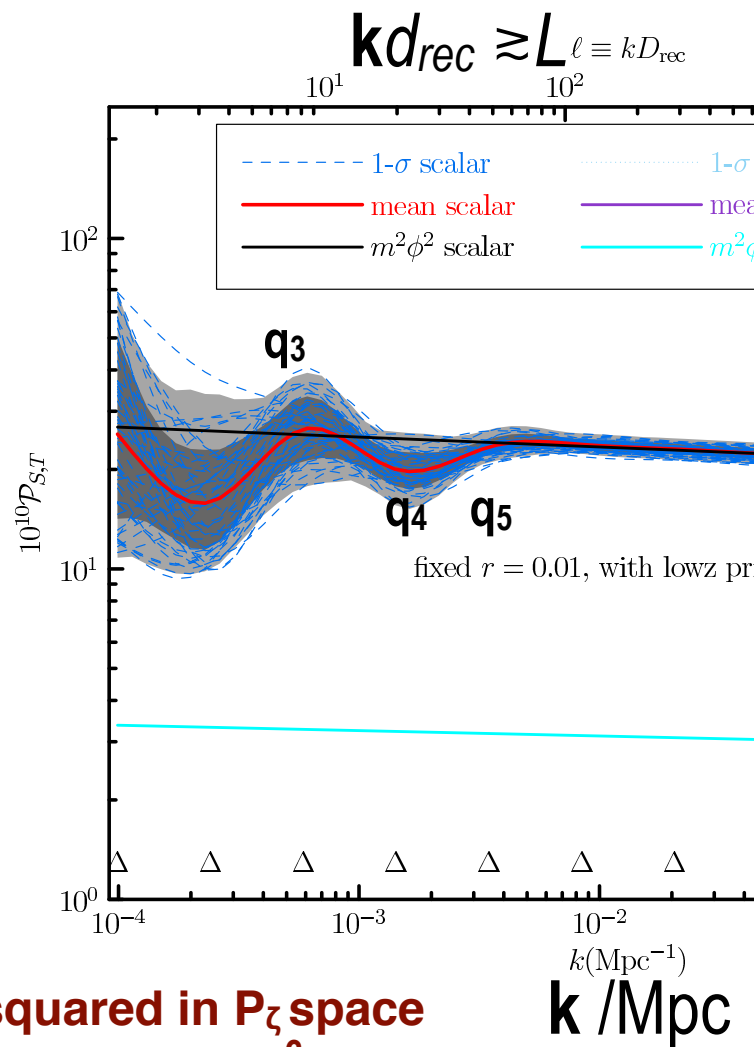
deviation from
 Λ CDM mended
by q_5

p-values from chi-squared in P_ζ space

q_1 to q_5 $L < 50$ p-value = 0.42 uniform $n_s \chi^2/\text{dof} = 1.00$

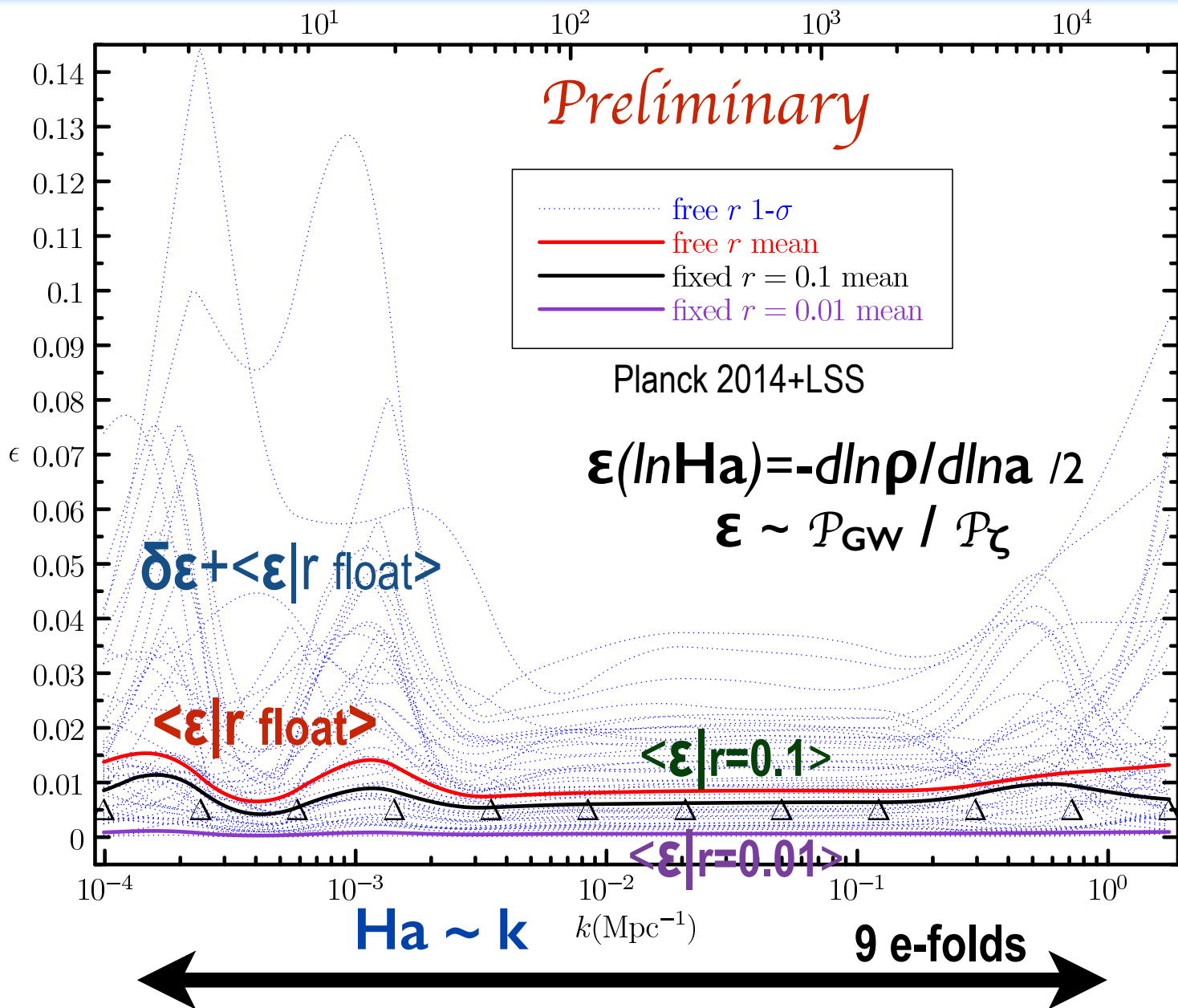
q_6 to q_{12} $L > 50$ P-value = 0.99 uniform $n_s \chi^2/\text{dof} = 0.11$

trajectory coherence => correlated errors



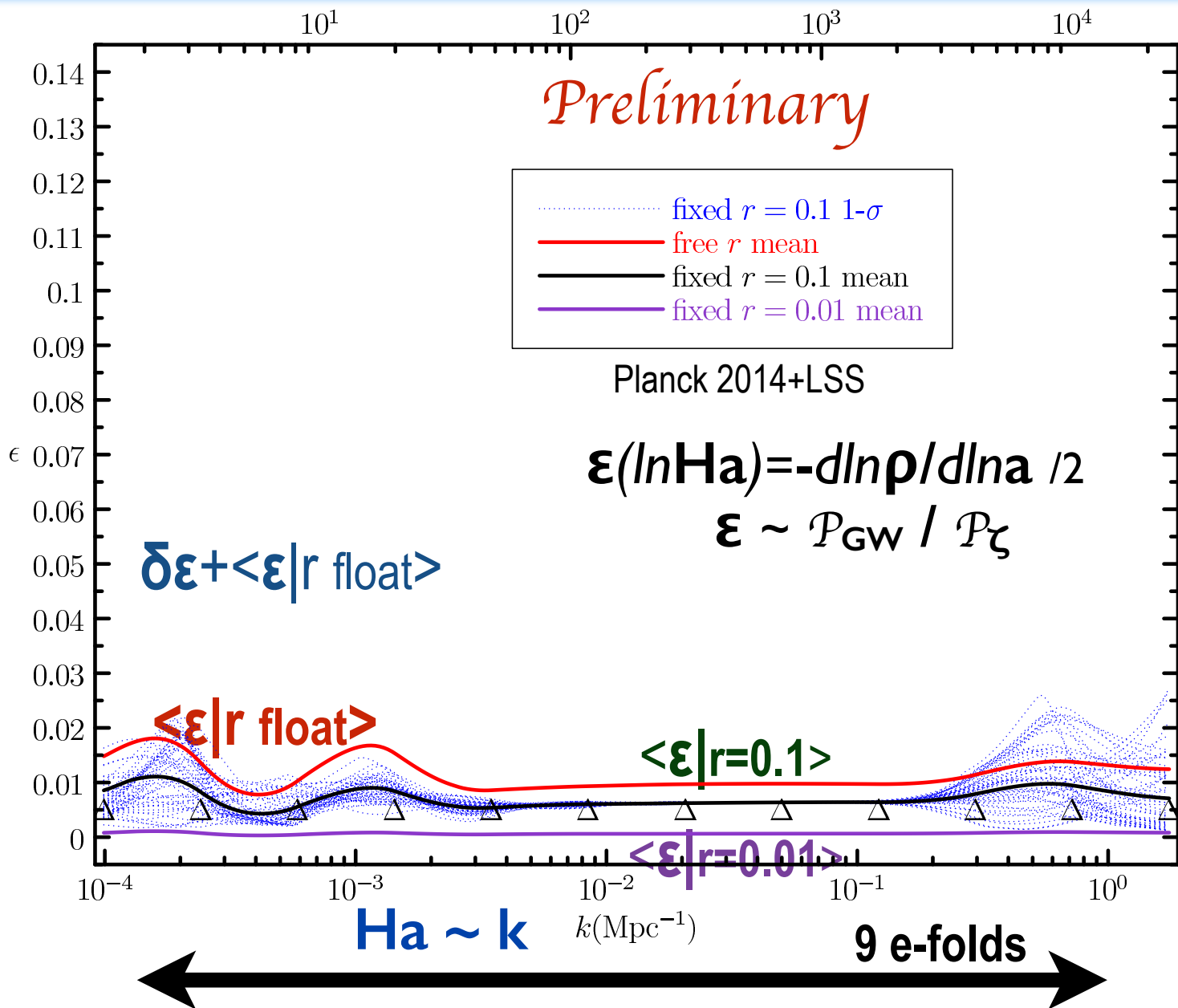
$$\mathcal{E} = 3(1+w)/2$$

$$\approx r(k)/16$$

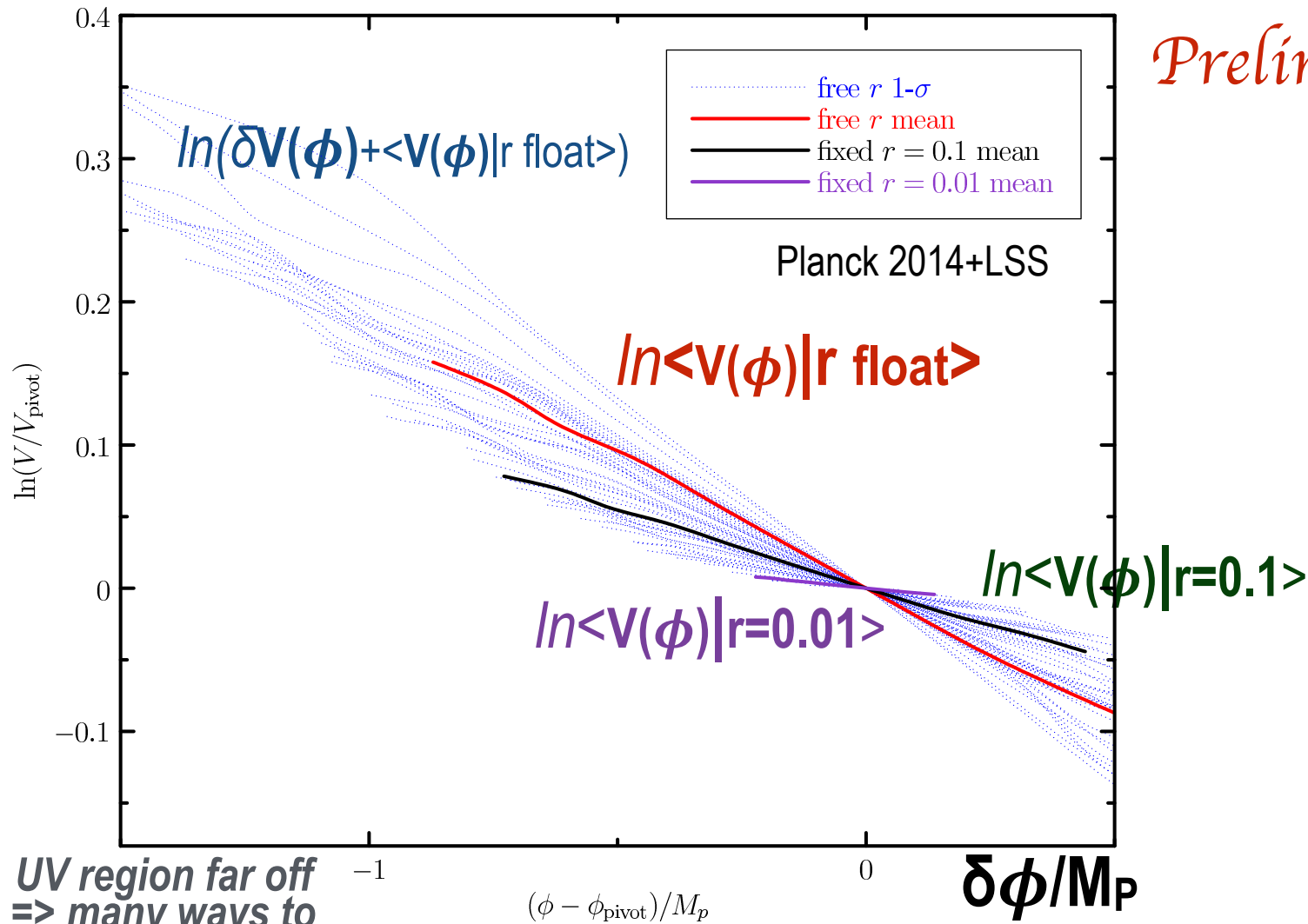


$$\mathcal{E} = 3(1+w)/2$$

$$\approx r(k)/16$$



inflaton $V(\phi)$ -maps $= 3M_P^2 H^2 (1-\epsilon/3)$ HJ eqn, $d\phi/M_P/d\ln a = \pm \sqrt{2\epsilon}$
 along the gradient / Morse flow



Preliminary

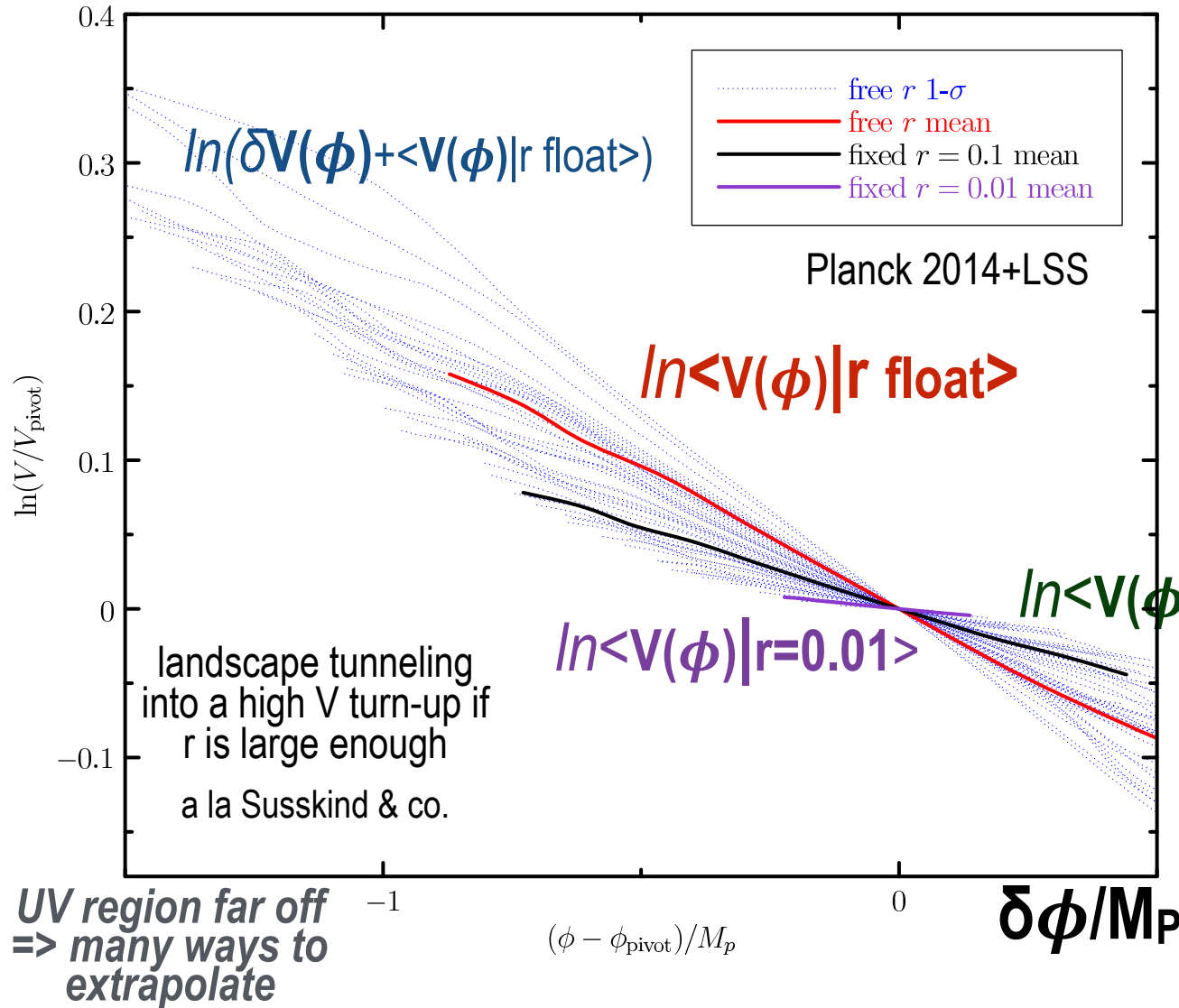
IR heating
 region is far off
 $\Rightarrow$ many ways to
 extrapolate

UV region far off
 $\Rightarrow$ many ways to
 extrapolate

r to ± 0.02 Spider forecast

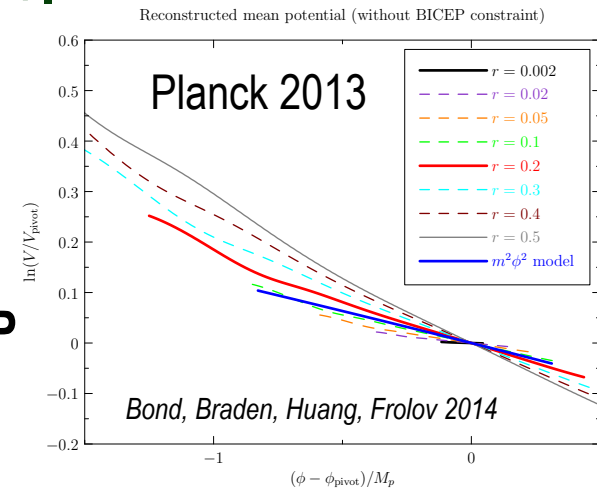
r to ± 0.003 AdvACTpol forecast w/ fgnds

inflaton $V(\phi)$ -maps $= 3M_P^2 H^2 (1-\epsilon/3) HJ \text{ eq}^n$, $d\phi/M_P/d\ln a = \pm \sqrt{2\epsilon}$
 along the gradient / Morse flow



Preliminary

*IR heating region is far off
 $\Rightarrow$ many ways to extrapolate*



r to ± 0.02 Spider forecast

r to ± 0.003 AdvACTpol forecast w/ fgnds

conclusions

convergence testing by increasing the number of knots, changing the knot mode functions, changing the fiducial to scale invariant => stable features & statistics

12 knots is Good for low L features; too few knots (8) => too stiff to respond to the CL data

degeneracies in $\mathcal{P}_\zeta$ cf. $\mathcal{P}_{GW}$ unless r is constrained/measured => $r=0.01, 0.1$ examples
=> same stable features. mild degeneracy with τ for lowest k -bands explain details of $L < 10$ features

2 other $\mathcal{P}_\zeta$ reconstructions in Planck 2014 Inflation paper. e.g., Will Handley poster uses moving linear knots: the stable features & conclusions agree. Planck 2014 Inflation also reconstructs V directly 2 ways

simple uniform n_s triumphs at high k from 0.3 to 0.008/Mpc, OK (r) at low k
 $\sim 10,000,000$ T/E modes = Λ CDM $L_k > 50$ p-value .98 (r free), .99 ($r=0.01$), .99 ($r=0.1$)

$\lesssim 1000$ T/E modes hint of uniform- n_s deviation, $\lesssim 100$ T/E probe reionization history
no statistical evidence of oscillation patterns, cutoffs, at this level of coarseness/stiffness;
 $\exists$ a mean-power change on large $L < 50$ scales exists which is not well-fit by running
the mean is statistically beaten by coherent power fluctuations: NO ANOMALY beyond 2 sigma
statistically insignificant deviation: low- k $L_k < 50$ p-value .40 (r free), .42 ($r=0.01$), .14 ($r=0.1$)
all our anomaly hints are at low L , quadratic & linear: we are victims of cosmic variance

inflaton EOS aka ϵ trajectories => V trajectories: higher r , bigger $\delta\phi/M_P$ & steeper V , upturn?

the $L_k=20-30$ TT-driven mean downturn was/is a phenomenology (less in P14 cf. P13), no matter r
fit into a UV-complete theory (ultra-high energy to the Planck scale) strings, landscape, ..
& IR-complete theory (post-inflation heating -> quark/gluon plasma)??? TBD

The scientific results that we present today are a product of the Planck Collaboration, including individuals from more than 100 scientific institutes in Europe, the USA and Canada.



Planck is a project of the European Space Agency, with instruments provided by two scientific Consortia funded by ESA member states (in particular the lead countries: France and Italy) with contributions from NASA (USA), and telescope reflectors provided in a collaboration between ESA and a scientific Consortium led and funded by Denmark.