



## **CESAR Science Case – Calculation guide**

## Calculate the heliographic coordinates of a sunspot

 $R_{m} = \sqrt{X^{2} + Y^{2}}$   $P_{m} = \operatorname{arc} tg(X/Y)$   $\rho = \operatorname{arc} \sin (R_{m}/R) - \alpha/2(R_{m}/R)$   $\operatorname{sen}(B) = \operatorname{sen}(B_{0}) \cos(\rho) + \cos(B_{0}) \operatorname{sen}(\rho) \cos(P - P_{m})$   $\operatorname{sen}(L - L_{0}) = \operatorname{sen}(P - P_{m}) \operatorname{sen}(\rho) \cos(B)$ 

Where X and Y are the coordinates of a sunspot measured in pixels on the image and L, B the heliographic longitude and latitude.

Use the value of  $0.5244^{\circ}$  for summer,  $0.5422^{\circ}$  for winter and 0.5333 for autumn/spring in the North hemisphere.

 $L_0$ ,  $B_0$ , and P have to be taken from the Ephemeris.

## Calculate the rotation period of the Sun

$$S = 360^{\circ} \frac{\Delta t}{\Delta L}$$
$$P = (S * 365,25)/(S + 365.25)$$

Where  $\frac{\Delta t}{\Delta L}$  could be calculated with: the next formula or just dividing the time between two images by the number of degrees that a sunspot has been moving.

$$\frac{\Delta L}{\Delta t} = \frac{\cos^{-1}\left(\frac{x_3 - \varphi}{\varphi}\right) - \cos^{-1}\left(\frac{\varphi - (x_2 - x_1)}{\varphi}\right)}{(t_2 - t_1)}$$

Wolf number

W = k(10g + s)