

Distributed User Support at ESO

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Head of the User Support Department



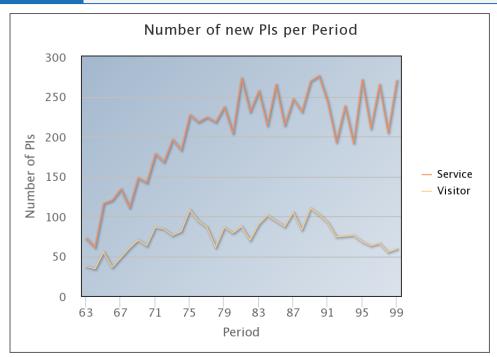
User Support Department

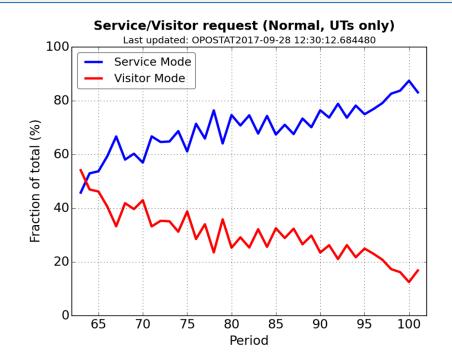
- Main interface to the external users and to the Observatory
 - > Scientific and technical support in proposal and observation preparation
 - Proposal and observation preparation, post-observation support
 - Up-to-date status information services
 - observing queues, observation progress reports and programme management, execution problems follow-up and solution, operations metrics
 - Specification and operation of the front-end infrastructure
 - tools, User Portal, helpdesk, user satisfaction surveys
 - Visiting Astronomers travel
 - ESO Users' Committee
 - New instruments: Reviews, Science Verification, Advanced Commissioning





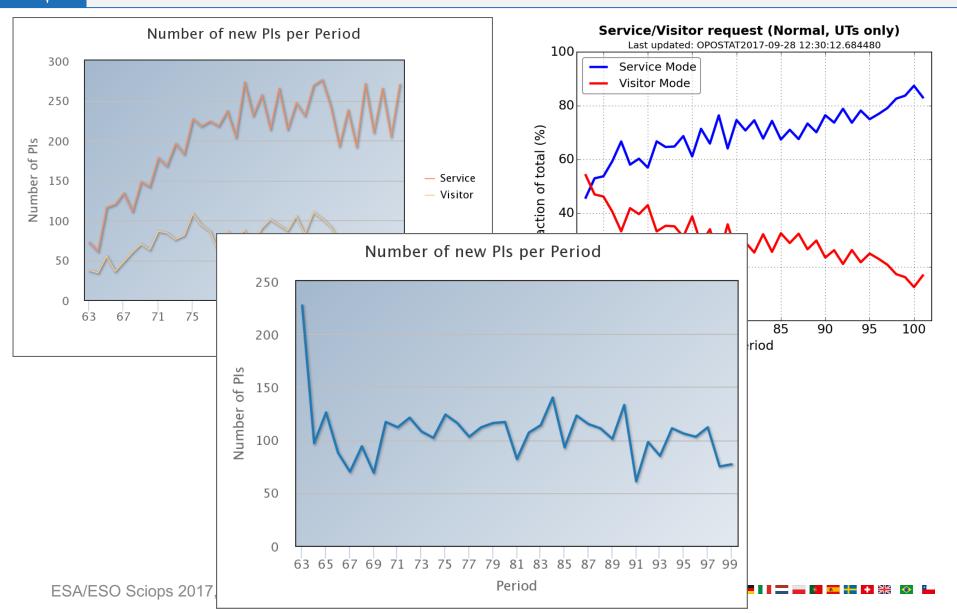
In numbers: users and proposals







In numbers: users and proposals

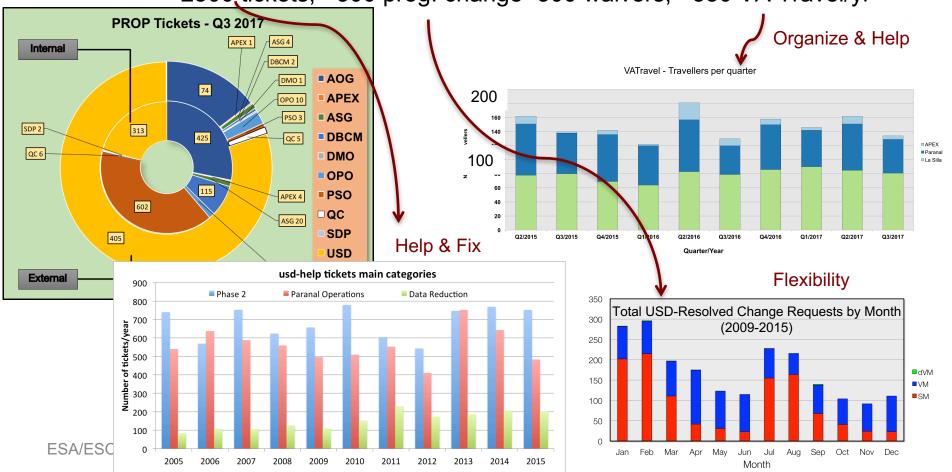




And more operations numbers

Smooth operations and support

- ~900 new Service Mode observing runs per year
- 17 instruments approx 1 new instrument/year coming to operations
- ~2500 tickets; ~300 prog. change+300 waivers; ~550 VA Travel/yr





Visiting Astronomer Travel Office A vital ancillary to ESO's Observing Model

L. Ruiz¹, S. Mysore², S. Marteau² (European Southern Observatory)

1 Visiting Astronomer Travel Office 2 Operations Support Group (User Support Department)

Abstract: The Visiting Astronomer Travel (VAT) office is an important interface between ESO and its user community. The VAT office is a part of the User Support Department and is responsible for organizing all aspects of travel for astronomers on observing missions to ESO observatories. The VAT office is the first point of contact for the Visiting Astronomer (VA) and continues to be the central source of information until all travel arrangements are made. Thus, the VAT office provides a high level of individual service to each traveller while ensuring that ESO financial regulations and Observatory policies are strictly adhered to at all times. There are multiple steps in a chain of events that must take place in order to ensure that the VAs make the best use of their allocated telescope time. Bringing over 550 VAs to La Silla, Paranal, and APEX observatories each year requires excellent coordination among all parties including travel and logistics teams distributed over multiple ESO sites. We present here the central role played by the VAT office in making for a straightforward preparation process and hassle-free travel for the astronomer on an ESO observing mission.



ESO Observing Model



astronomer is physically present at the observatory to conduct observations. Programmes needing real-time decisions to be taken for their scientific success are allocated time in VM. Each approved VM run is allocated specific calendar nights during the observation period of a semester.

Visitor Mode (VM): classical observing mode where the Service mode (SM): gueue observing mode where the lineup of observations to be conducted is continuously updated to optimize the use of available time and the prevailing observing conditions. SM slots in the schedule are allocated per telescope for the semester. During these time slots, observations belonging to SM programmes with highest scientific priority and under the required observing conditions are carried out first by ESO staff astronomers and instrument operators.



VLT/VLTI telescopes on Paranal have VM nights and SM slots interspersed through the duration of the observing period. La Silla is fully operated in VM.

Traveller's task list Before and after an observing mission Approve international flight offer from VAT Office

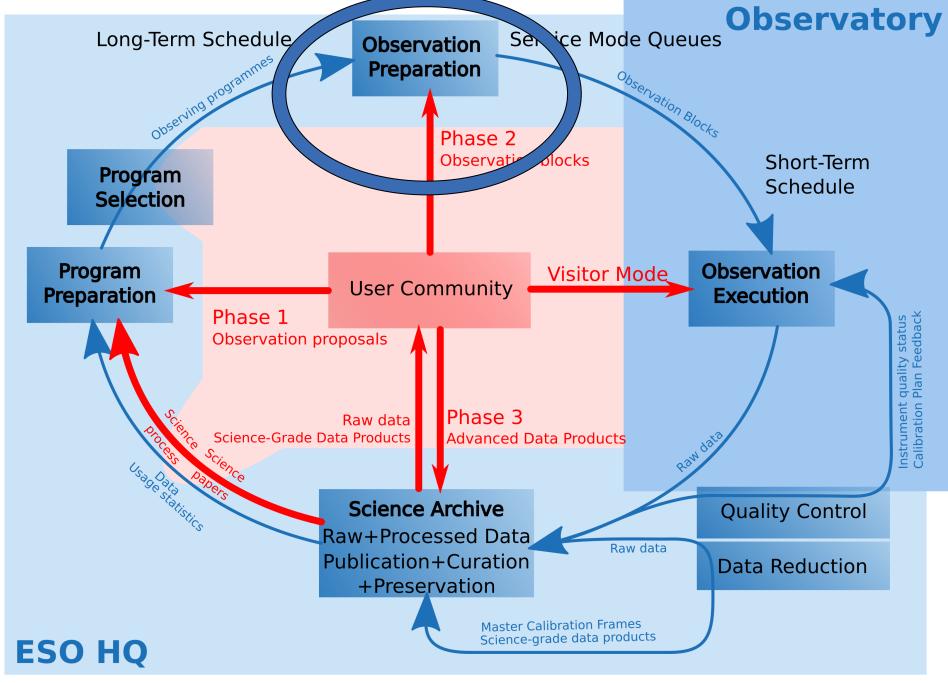
Partners/Agencies interacting with VA Travel Office

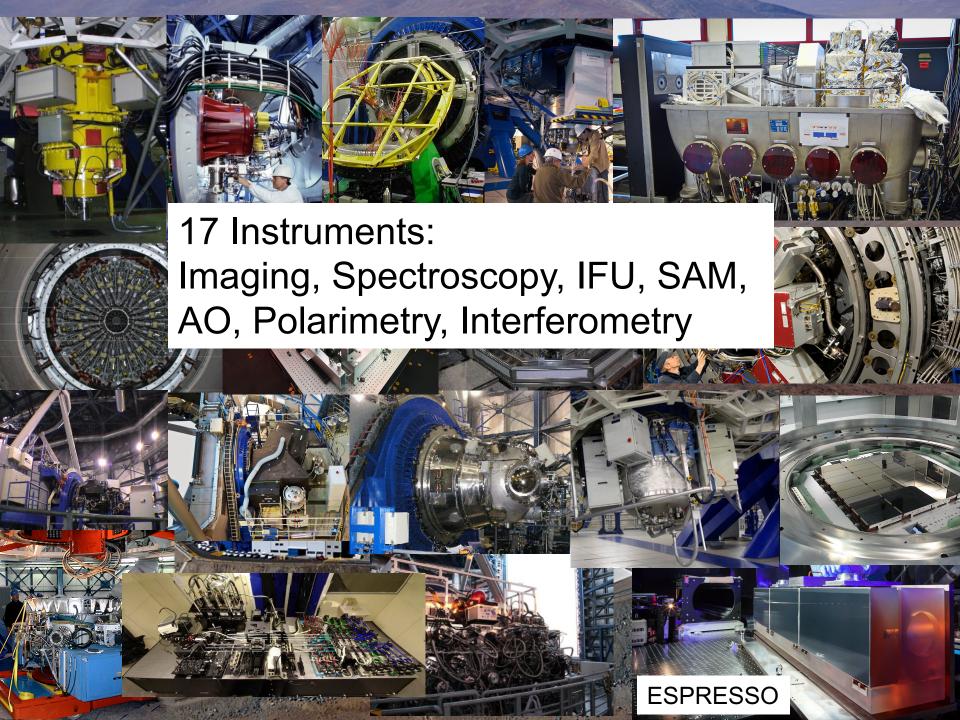
Behind the scenes

Chain of events distributed over multiple sites

La Silla: SciOps and Logistics Paranal: SciOps and Logistics **APEX: SciOps and Logistics**

Verify schedule: Run + Travel dates + 2nd Prepare visa support letter (if needed) Receive Travel Form







Interfaces and distributed support

- Science policies
- Proposals, selection, scheduling
- Software
 - > Instruments preparation
 - ➤ User Portal, Observation preparation, ETCs, P2, ...
- Science Operations, Instrument Operation Teams
- Quality Control: calibrations, instrument health
- Instrument pipelines
 - software development & scientific validation
- Archive
 - Internal & External Data Products



Some of the key ingredients

- Communication & coordination
 - Weekly operations telecons
 - Bi-yearly Science Operations Meetings
 - Face-2-face working visits: Service Mode shifts
 - Helpdesk: Ticketing system(s) internal and external
 - webpages, twiki, mailing lists per group/area
 - Reporting: internal, committees, external
 - Workshops
- Priorities, Projects & Processes (CCBs)
- Documentation & Training
- Standards, Rules & Procedures
- Flexibility and science focus



Diversity and flexibility

Observing Modes

- Visiting Mode (VM)
 - Visiting Astronomers Travel
 - Support at the observatory
 - Designated Visitor Mode, Eavesdropping
- Service Mode (SM)

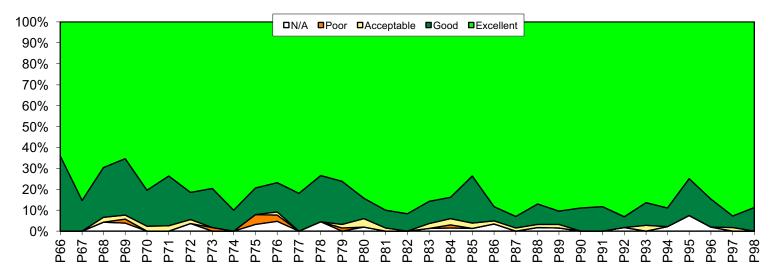
Observing programme types

- Normal high priority vs filler type
- Monitoring Programme relatively small but long time
- Target of Opportunity, Rapid Response Mode
- ➤ Large Programme 10 nights or longer
- Public Surveys imaging, spectroscopic surveys
- Calibration Programme

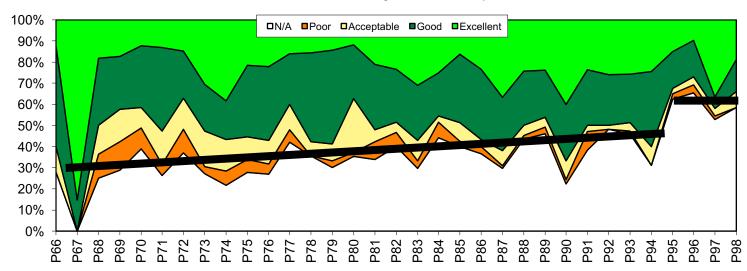


Measuring Success: VM Feedback – 15 years trends

EoM Stats History - Astronomer Support



EoM Stats History - On-line Pipeline

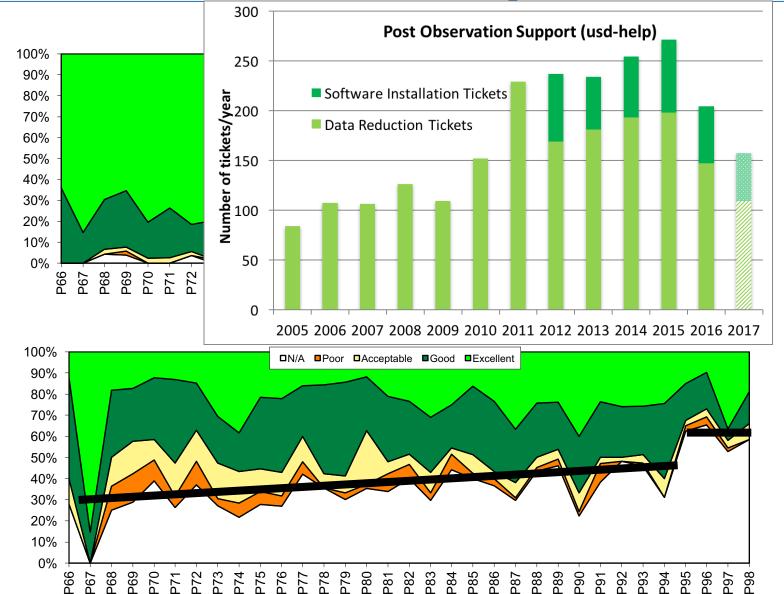






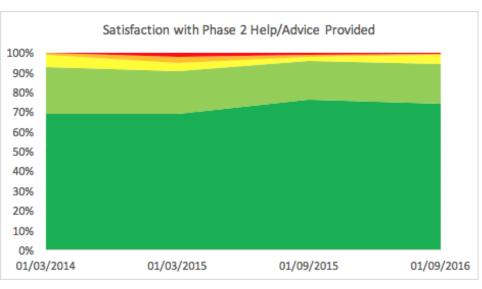
ESA/

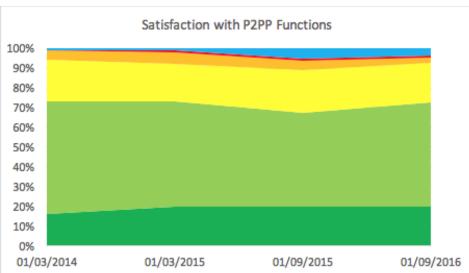
Measuring Success: VM Feedback – 15 years trends

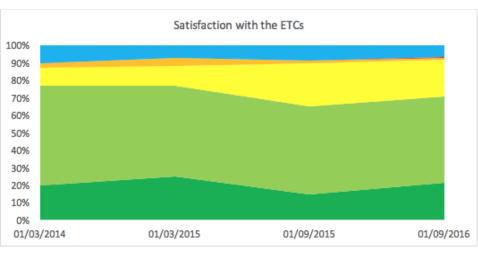


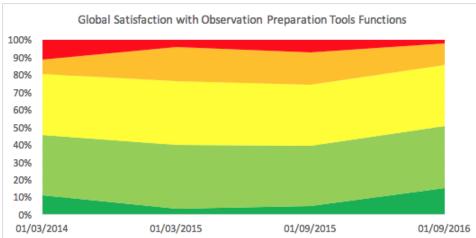


SM User Satisfaction Trends



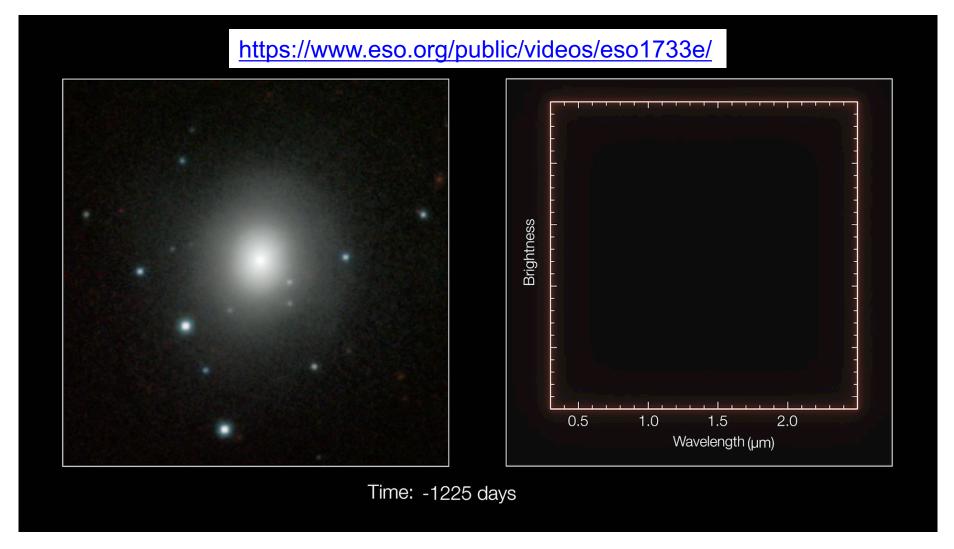








Focus on Science GW170817: VIRCAM & X-SHOOTER





Measuring & Analysing

Telescopes and Instrumentation

Fifteen Years of Service Mode Operations: Closing the Loop with the Community

Francesca Primas¹
Lowell Tacconi-Garman¹
Stéphane Marteau¹
Vincenzo Mainieri¹
Marina Rejkuba¹
Sangeeta Mysore¹
Christophe Dumas¹
Andreas Kaufer¹
Ferdinando Patat¹
Michael Sterzik¹

¹ ESO

The first Service Mode (SM tions with the VLT were ma in April 1999. Since then, n ments have become opera first generation ones replace 12 VLT foci and feeding the ometer and its four Auxilian Efficiently operating such a of instruments, installed an every night of each year, on telescopes offers many cha Although it may appear tha changed since 1999, the ur operational model has evol to accommodate different from the user community a of new instruments. As ESC Member States approach r ations with ALMA, and at the prepare for the next challer struction of the E-ELT, it se to take a closer look at wha brought to the scientific are

Telescopes and Instrumentation

The Scientific Return of VLT Programmes

Michael Sterzik¹
Christophe Dumas²
Uta Grothkopf¹
Andreas Kaufer¹
Bruno Leibundgut¹
Stephane Marteau¹
Silvia Meakins¹
Ferdinando Patat¹
Francesca Primas¹
Marina Rejkuba¹
Martino Romaniello¹
Felix Stoehr¹
Lowell Tacconi-Garman¹
Ignacio Vera¹

- 1 ESO
- ² TMT Observatory Corporation, Pasadena, USA

An in-depth analysis of the public from 8414 distinct scheduled VLT serving programmes between Ap and March 2015 (Periods 63 to 94 presented. The productivity by me

(Visitor or Service Mode) and to

Telescopes and Instrumentation

The Growth of the User Community of the La Silla Paranal Observatory Science Archive

Martino Romaniello¹
Magda Arnaboldi¹
Cristiano Da Rocha²
Carlos De Breuck¹
Nausicaa Delmotte¹
Adam Dobrzycki¹
Nathalie Fourniol¹
Wolfram Freudling¹
Laura Mascetti²
Alberto Micol¹
Jörg Retzlaff¹
Michael Sterzik¹
Ignacio Vera Sequeiros¹

copes on La Silla and the Atacama
Pathfinder Experiment (APEX) antenna on
Chajnantor. Also available through the
SAF are data from selected La Silla
instruments, for example, the GammaRay burst Optical/Near-infrared Detector
(GROND), the Fibre-fed Extended Range
Echelle Spectrograph (FEROS) and the
Wide Field Camera (WFI), together with
the raw data for the UKIDSS WFCAM
survey obtained at the United Kingdom
Infrared Telescope (UKIRT) in Hawaii.

rithms. The Reflex environmet al., 2013) allows the data veniently organised, so tha run through the processing actively assess the quality and, if needed, iterate on the state of the sta

Processed data are also at the Science Archive Facility used directly for scientific r thus alleviating the need fo do any data processing of mentioned above, the SAF



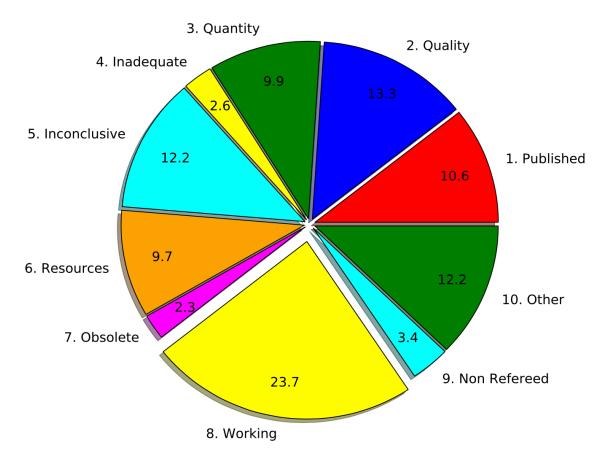
SNPP: Survey of Non-Publishing Programmes

- Why were you not able to publish the results of your observations in a refereed paper?
 - 1. I did publish a refereed paper (provide hyperlink)
 - 2. Insufficient data quality
 - 3. Insufficient data quantity
 - 4. Inadequate ESO data reduction tools
 - 5. Null or inconclusive results
 - 6. Lack of resources on the PI side
 - 7. Science case no longer interesting
 - 8. I am still working on the data (provide time estimate)
 - 9. I published a non-refereed paper (provide hyperlink)
 - 10. Other (provide details)



SNPP results (Patat et al., in prep)

- 1278 targeted programmes (P78-P90)
 - ➤ Replies from 965 (+70 bounced) → 80% return!





Evolving support & role of the community

- Public Surveys
 - > Data reduction & data products from the community

ESPRESSO

- Scheduling multiple telescopes
- Further flexibility: challenging the predictive scheduling
- New or updated tools (EOPS, P2 API)

4MOST

- Changing the paradigm: one observation one programme/one PI
- User support & Data processing
- SOX & NIRPS at La Silla
- E-ELT as "just another telescope"