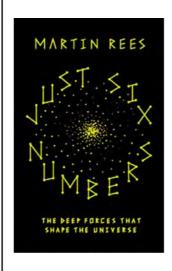
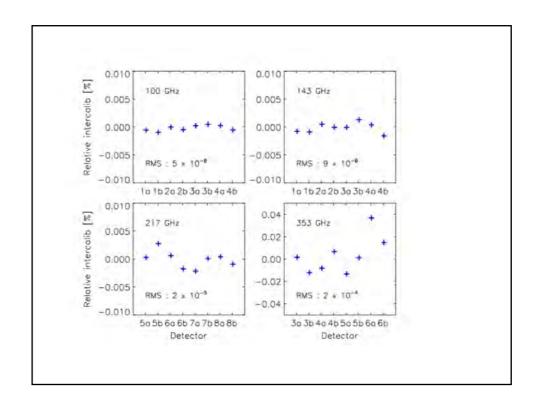
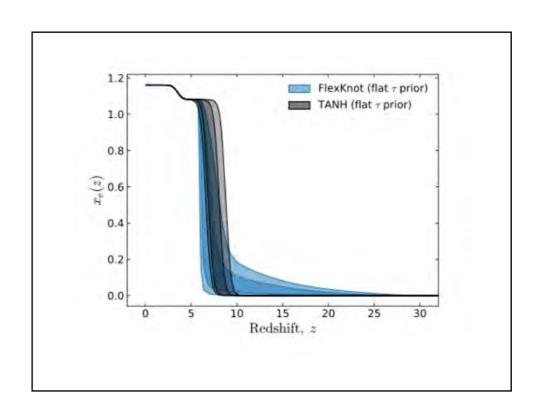


Other science Clusters of galaxies infrared luminous and radio galaxies cosmic infrared background synchrotron, free-free and `anomalous' dust emission Galactic dust emission in temperature and polarization Maps of CO emission, evidence of `dark' HI Magnetic field structure of the Galaxy Cold `clumps' in our Galaxy



- 1. Baryon density $(\Omega_b h^2)$
- 2. Cold dark matter density $(\Omega_c h^2)$
- 3. Angular scale of acoustic peaks (θ)
- 4. Amplitude of fluctuations (A_s)
- 5. Spectral index of fluctuations (n_s)
- 6. Reionization optical depth (τ)





BASE ACDM MODEL		
Parameter	TT	TT,TE,EE+lensing
$\Omega_{b}h^{2}$	0.02212±0.00022	0.02237±0.00015
$\Omega_{\text{c}} h^2$	0.1206±0.0021	0.1200±0.0012
100 θ_*	1.04077±0.00047	1.04092±0.00031
τ	0.0522±0.008	0.0544±0.0073
n_s	0.9626±0.0057	0.9649±0.0042
H _o	66.88±0.92	67.36±0.54
Ω_{m}	0.321±0.013	0.3153±0.00073
σ_8	0.8118±0.0089	0.8111±0.0060
z _{re}	7.50±0.82	7.67±0.73

Planck 2013:

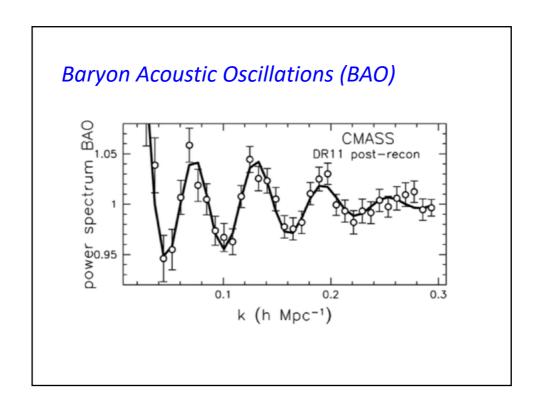
- · good agreement with Planck lensing
- consistent with BAO
- ~2σ tension with Ia SNe
- $\sim 2.5\sigma$ tension with H₀ $\times \times 3.8\sigma$ AGR etal
- tension with measures of σ_8 including:

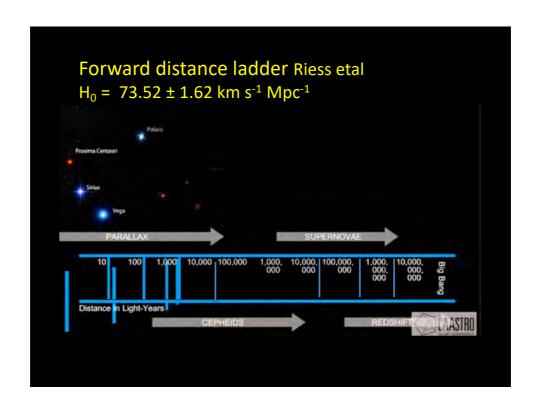
weak lensing

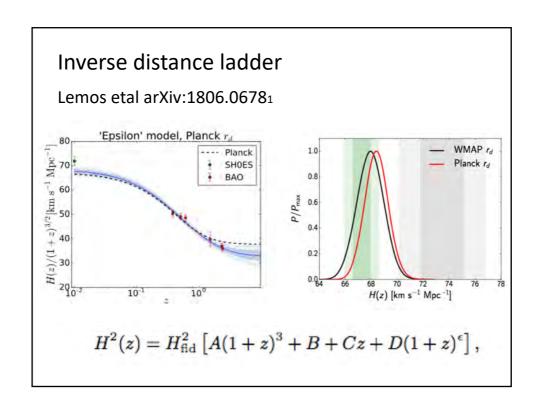
cluster counts ?

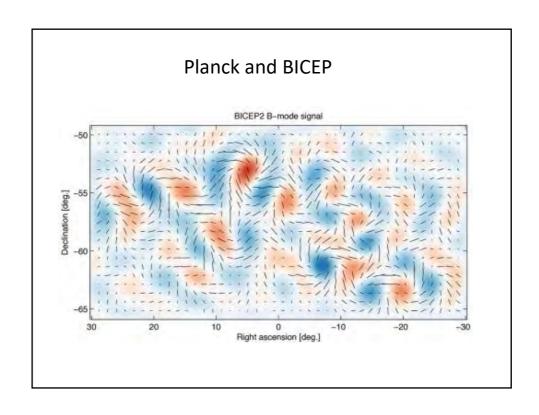
redshift space distortions

Some skeptics have even doubted the fidelity of the Planck data!



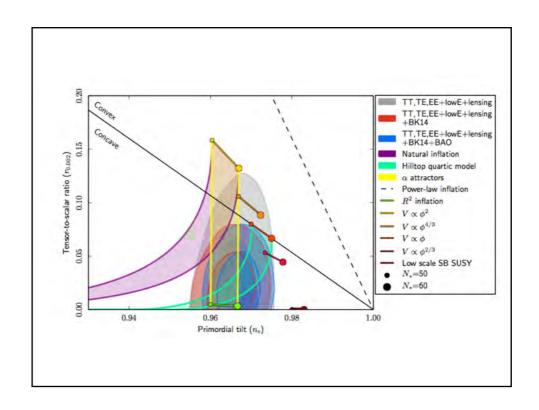












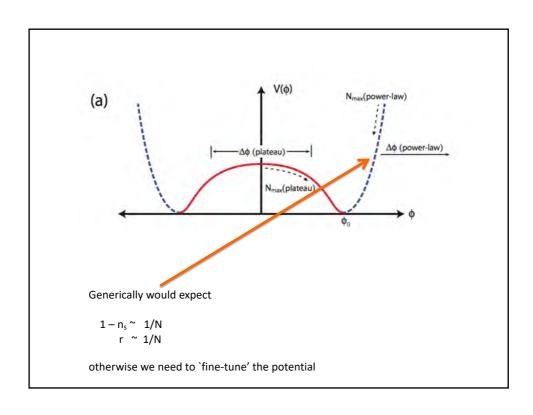


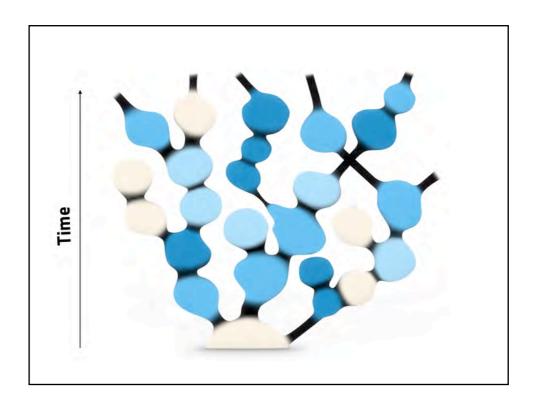
Single field inflation:

$$V(\phi)=\lambda\phi^{\alpha}$$

$$1\text{-}n_s=(\alpha+2)/N, \quad r=4\alpha/N$$
 For $n_s=0.965,\,N{\approx}60,\quad \alpha\approx2.2,\,r\approx0.15$

$$\begin{split} \epsilon &= (m^2{}_{pl}/16\pi) \; (V'/V)^2 \,, \, \eta = (m^2{}_{pl}/8\pi) (V''/V) \\ n_s &= 1 - 6\epsilon + 2\eta \,, \, \, r = 16\epsilon \,, \, \, \, n_t = \, -2\epsilon/N \end{split}$$





Are Ijjas, Steinhardt and Loeb right?

Yes! Inflation is an incomplete theory

Initial conditions, Measure problem,

Infinities

......

No! This does not mean that inflation is `in

trouble'.

It just means that there are many, many

open problems in cosmology,

and we should acknowledge this when we teach

inflation.

- ACDM is not Fake News!
- Are there alternatives to inflation?
- What sets the initial conditions for inflation?
- Multiverse and measures?
- Are there signatures from a 'pre-big b phase?

'This is not the end. It is not even the beginning of the end. But perhaps, the end of the beginning." Winston Churchill

