

# HEROES: Strong lensing of Lyman-alpha emitters

Günther Hasinger

### ESAC Science Seminar, 28. June 2018



# The X-ray survey eROSITA (2019)

Russian-German Collaboration Survey will be 30 times deeper than ROSAT





Former and current PI

Peter Predehl and G. H. with eROSITA in MPE cleanroom.

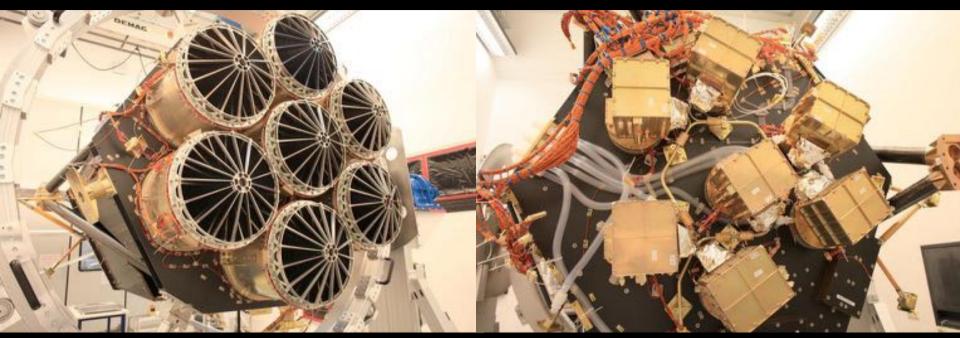
Currently eROSITA is at Lavochkin in Moscow

Launch is foreseen on April 6, 2019

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# X-ray technology from Germany/Italy





X-ray mirrors and CCD X-ray detectors are based on the XMM-Newton technology, further developed at the Max-Planck-Institute for extraterrestrial Physics with MPI Halbleiterlabor and Medialario ESA UNCLASSIFIED - For Official Use

# eROSITA Exposure Map



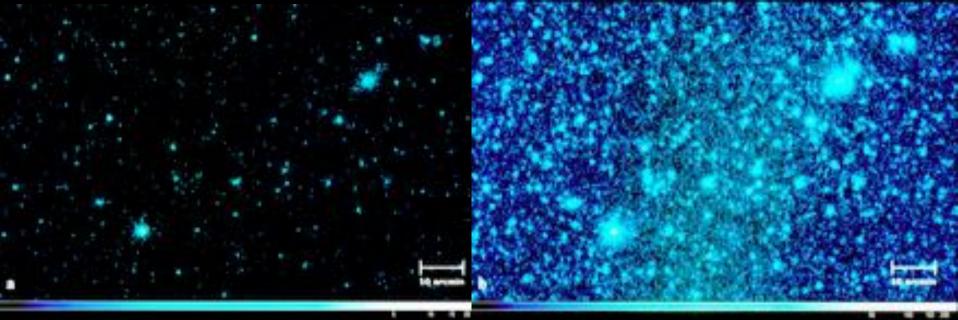
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# **eROSITA Simulations**





All-Sky (2 ksec)

Deep Survey (30 ksec)

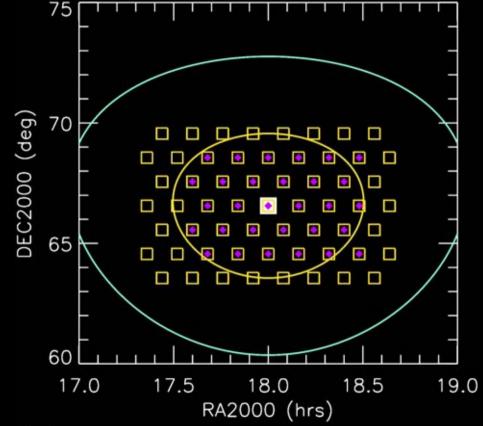
### Clusters of galaxies clearly recognizable as diffuse X-ray sources

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# The Hawaii EROsita Ecliptic-pole Survey





Subaru Hyper SuprimCam, CFHT coverage in U,g,r,i,z,y,NB821, NB916,J

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# HEROES



### Table 1. HEROES specifications

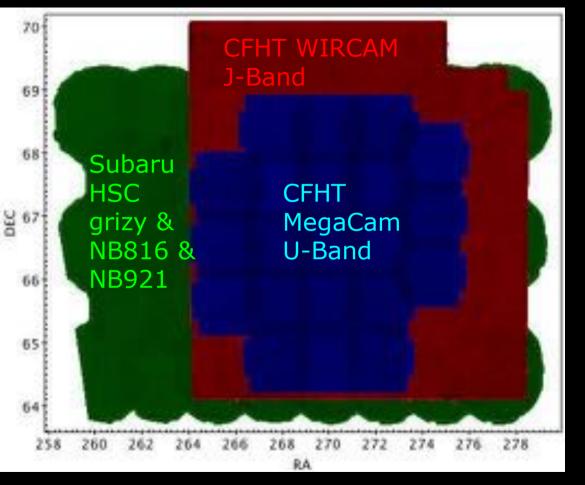
Instrument (1)	Filter (2)	Redshift Range (3)	Exposure/pixel (mins) (4)	$5 \sigma \text{ mag}$ (5)
HSC SUBARU	g		10	26.5
	r		10	26.1
	i		15	25.7
	z		20	25.1
	У		20	24.4
	8140	5.67-5.75	10	24.1
	9210	6.53-6.62	20	24.0
WIRCAM CFHT	J	202	5	22.1
MEGAPRIME CFHT	U		20	25.5

# UH Large Program Status 2016-2018: First observations June/July 2016

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# Current status of HEROES



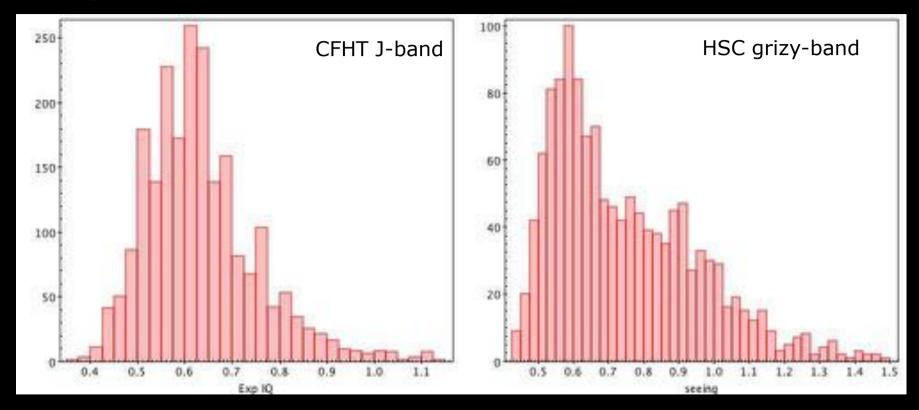


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# Seeing Distributions





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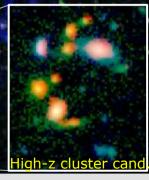
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HEROES – The Hawaii EROsita Ecliptic pole Survey

UH Large Collaborative Program: G. Hasinger (PI), A. Barger, E. Hu, M. Takamiya, R. Sunyaev, A. Cowie, R. Griffiths, L. Cowie & Conor McPartland

NGC6543 PN



see: www.ifa.hawaii.edu/HEROES

NGC6552

# Keck Cosmic Web Imager (KCWI)

## Image Slicers L, M, S

-16.5 x20.4

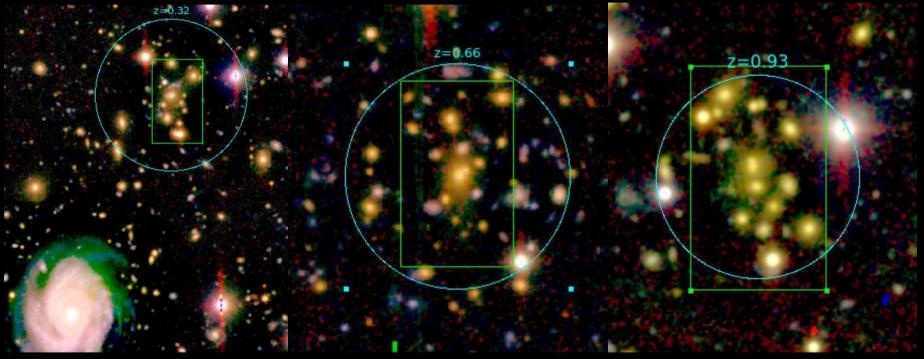
3″x20.4

8.4 "x20.4

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# 2017B KCWI Proposal



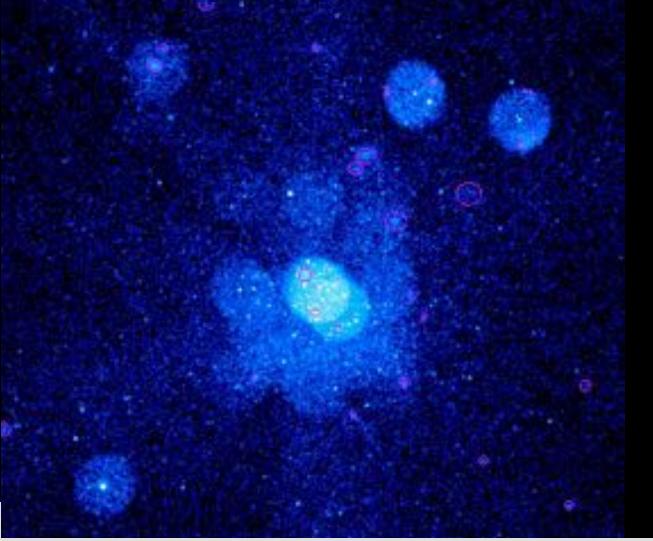


### ASTR699 work from Erica Bufanda

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ROSAT Coverage of the North-Ecliptic Pole

First joint analysis of ROSAT All-Sky survey and pointed raster scan data in the NEP Field

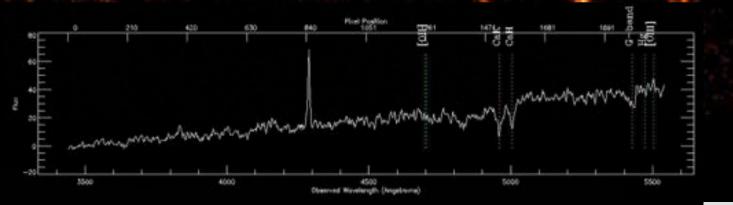
# ROSAT Cluster 15/ Redshift 0.26

HEROES HSC giy data

### Sky-subtracted KCWI raw spectra from the Large slicer

Emission line objects are close, but typically not <u>in</u> the bright continuum sources.

They are not commensurate with the galaxy redshifts. Most likely Ly-alpha at z~2.5-3.



# Flattened 2D Spectrum of KCWI Data Cube (M-slicer)

### OII in cluster A STORAGE WATER STORES

# OII outside

The states

Ly-a

SALAR DESCRIPTION OF THE PARTY OF THE PARTY

Ca K+K

# ROSAT Cluster S1548 z=0.26

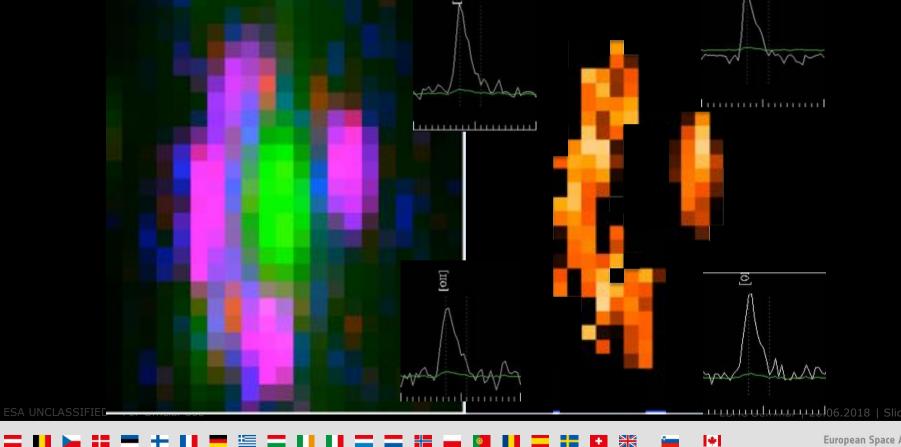


3 strongly lensed Ly-a emitters (1 Einstein cross)

2 OII emitters (one in cluster, one in front)

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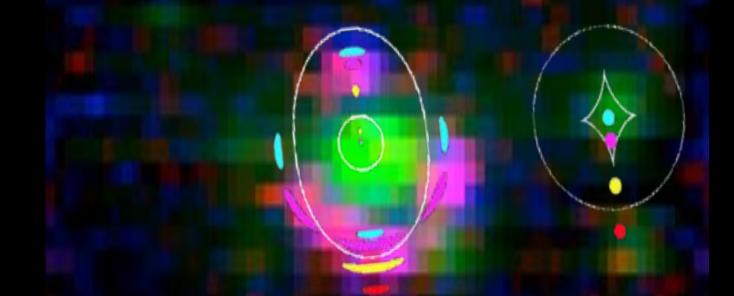
# **KCWI Velocity Map**



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# Possible Lensing Configuration





In an elliptical Dark Matter potential there can be four magnified images plus one demagnified inside the critical lines. Images close to caustics are magnified most. Outside the caustics there is only one magnified image.

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# ROSAT Cluster V6a z=0.34



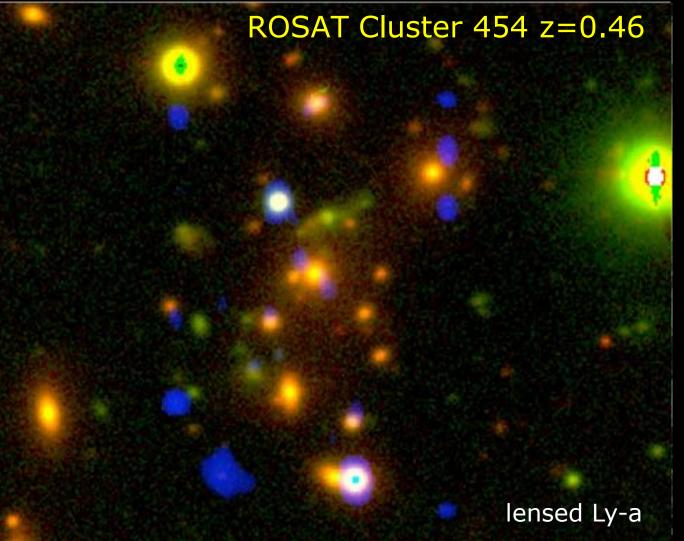
2 strongly lensed Ly-a emitters (1 triple lens)

1 OII emitter in cluster

1 broad [MgII] emitter in background QSO

OII in cluster

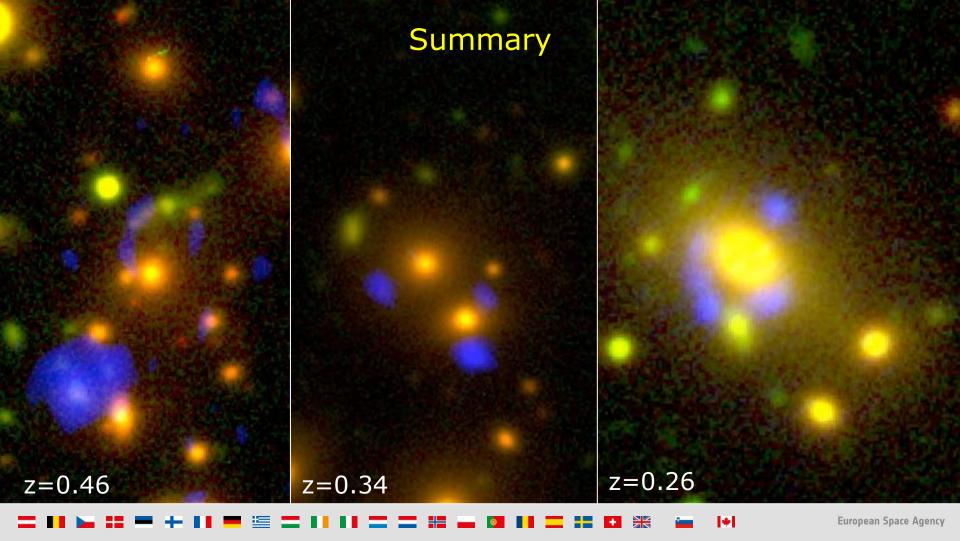
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3 strongly lensed Ly-a emitters

3 OII emitters (two in cluster, including one stripped, one in front)



# Summary and Conclusions



KCWI Integral-field spectroscopy of clusters of galaxies provides an unprecedented sensitive new probe for emission line galaxies.

Morphology of gas inside the cluster accessible through OII line spectroscopy.

Background Ly-a emitters become accessible through lensing amplification.

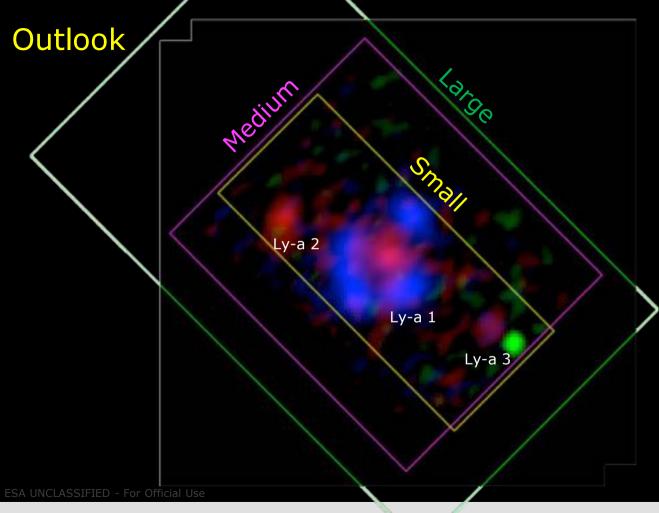
Practically every massive cluster seems to have one multiply lensed system and several single lensed.

Unique new way to identify Einstein crosses.

Statistics of strong lensed sources indicates a high surface density of these background emitters. Consistent with independent measurements with DEIMOS and MUSE.

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I will try to do Large and Medium image slicers for a total of ~10 clusters, and Small for the most interesting ones.

Keck2, 10-12. July

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# Thank you very much!