

# **LIGHT-CONES WITH EMISSION LINES FOR J-PLUS SURVEY**

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**SYNERGY BETWEEN  
N-BODY SIMULATIONS, SEMI-ANALYTICAL  
MODELS & EMISSION LINE MODELING**

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**ESA “*Simulated skies for new-generation spectroscopic surveys*”**  
**24 April 2018**  
**Madrid, Spain**

# MOTIVATION

- **Check pipelines**

- \* Is there any bias in our code or selection?
- \* What are the typical errors?

- **Test new ideas**

- **Physics behind our results**

# MOTIVATION

## MOCKS FOR NARROW BAND SURVEYS



C . Hernandez-Monteagudo

(Wednesday afternoon session)



S . Bonoli

(Wednesday afternoon session)

# Narrow band survey J-PLUS



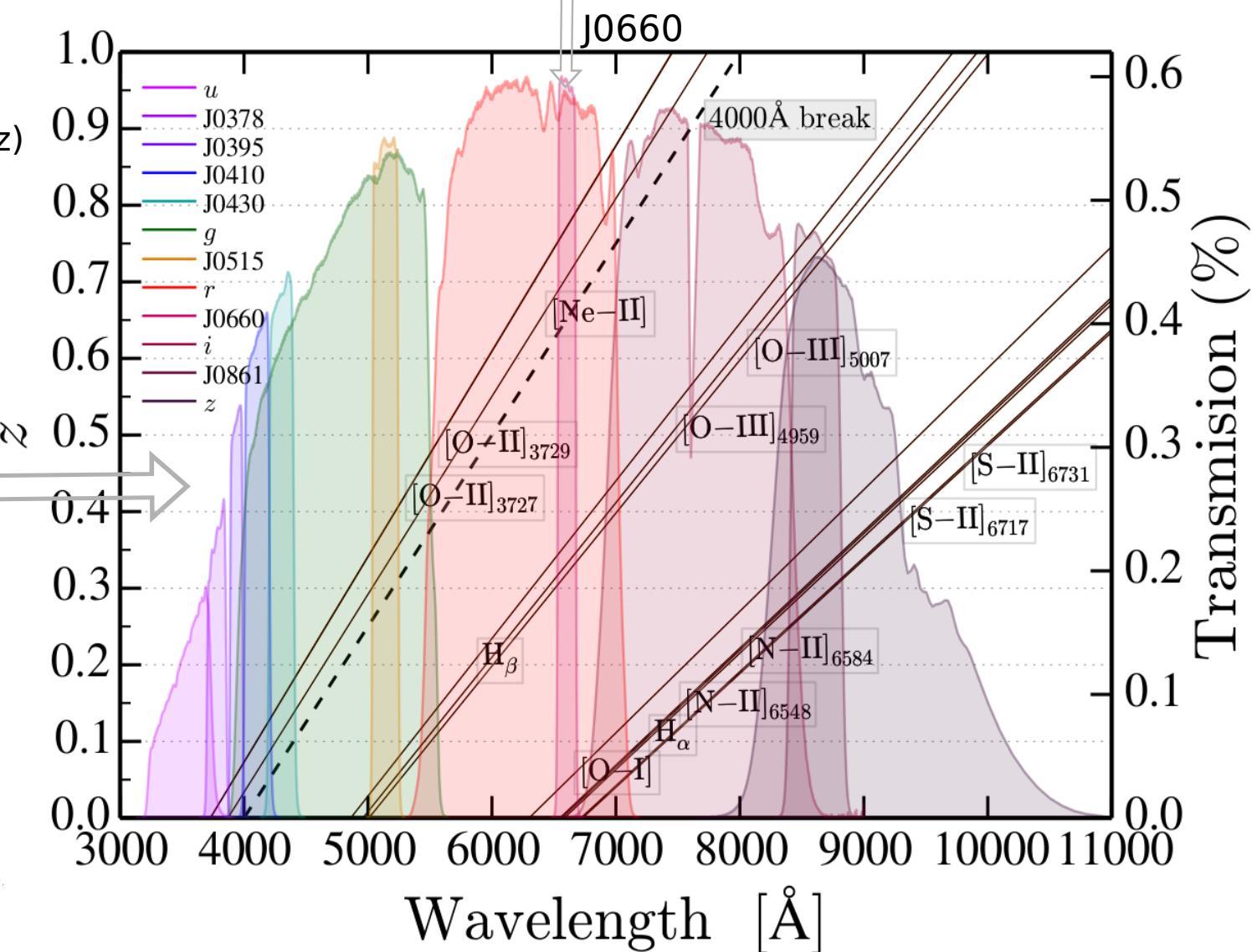
$H_{\alpha}$        $z = 0$   
 $[O-III]$  &  $H_{\beta}$        $z = [0.3 - 0.35]$   
 $[\text{O}-\text{III}]$        $z = [0.77 - 0.8]$

5 BROAD bands (u, g, r i, z)  
 7 NARROW bands

$[O-II]$        $z = 0$   
 $Ly_{\alpha}$        $z = 2.2$

J0395

$$\begin{aligned}
 m_{AB} &= -2.5 \log_{10} (f_{\nu}) - 48.6 \\
 &= -2.5 \log_{10} \frac{\int S(\lambda) \lambda f_{\lambda} d\lambda}{c \int \frac{S(\lambda)}{\lambda} d\lambda} - 48.6
 \end{aligned}$$



# HOW DO WE CREATE A MOCK CATALOGUE?

Mimicking a survey

Predict a reliable  
galaxy population:

**Galaxy formation  
model**

Predict a good  
galaxy spacial  
distribution:

**Dark matter  
N-body simulation**

Predict a emission lines:

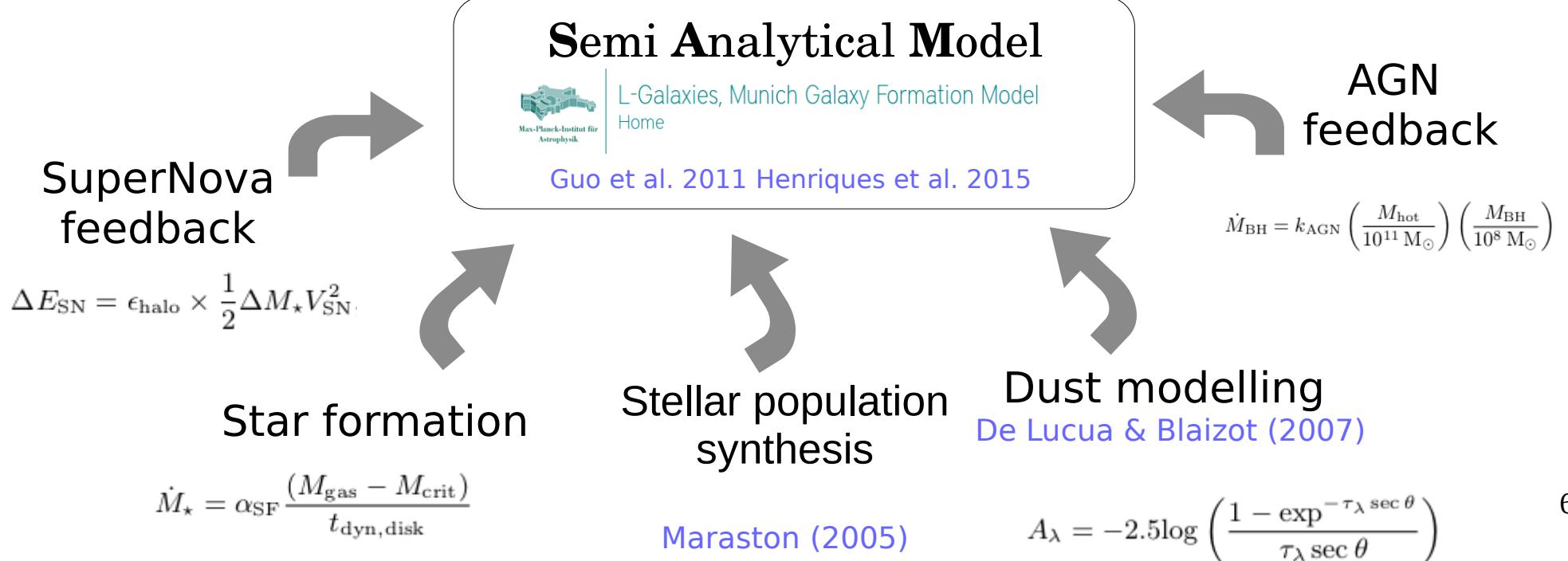
**Emission line models**

# HOW DO WE CREATE A MOCK CATALOGUE?

Predicting a reliable galaxy population.

Allow us to track the cosmological evolution of galaxies

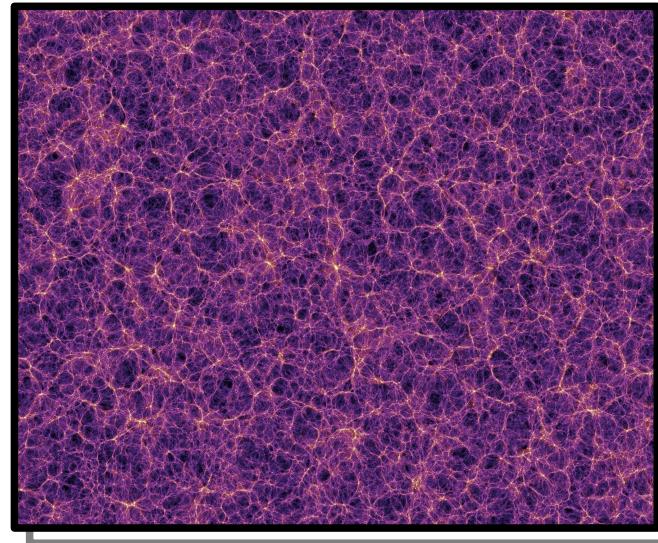
## PHYSICAL MOTIVATED RECIPES



# HOW DO WE CREATE A MOCK CATALOGUE?

Predicting a good galaxy spacial distribution

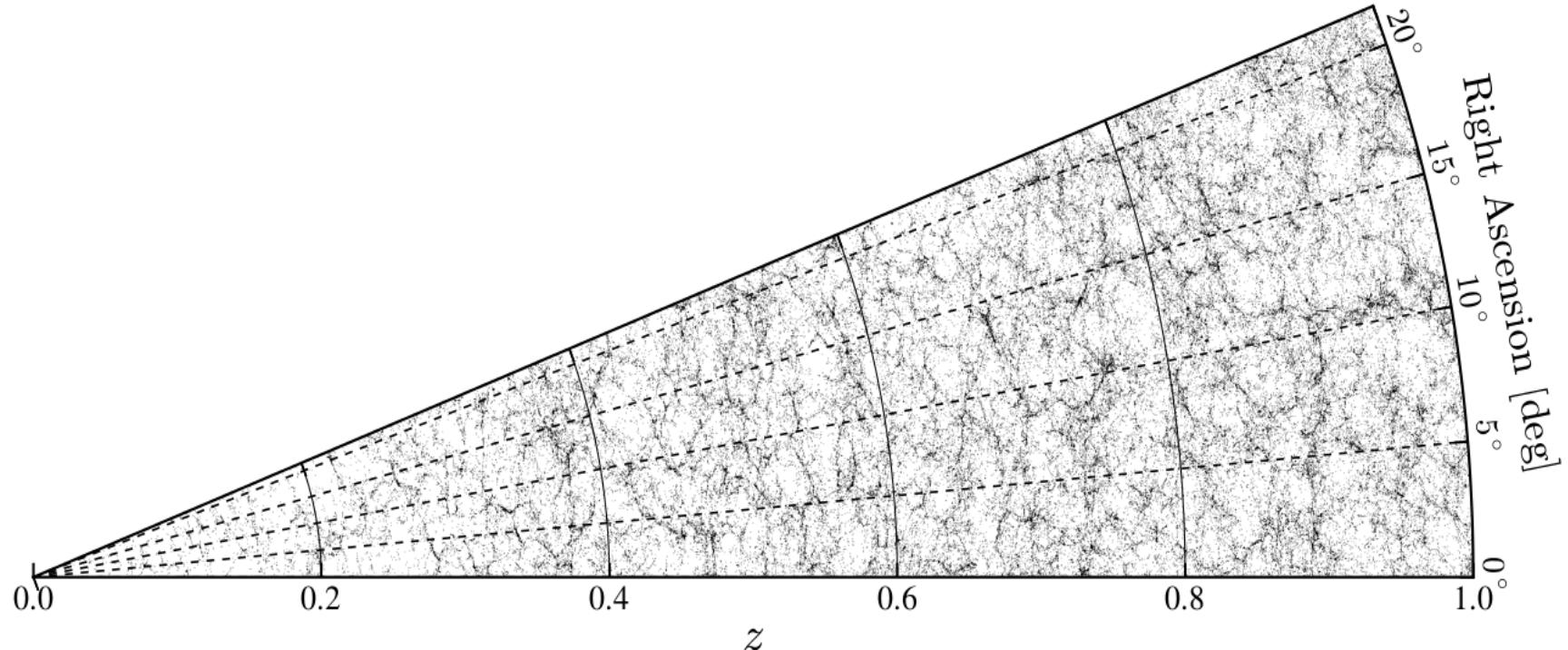
Dark matter simulation



Millennium simulation ([Springel et al. 2005](#))

# HOW DO WE CREATE A MOCK CATALOGUE?

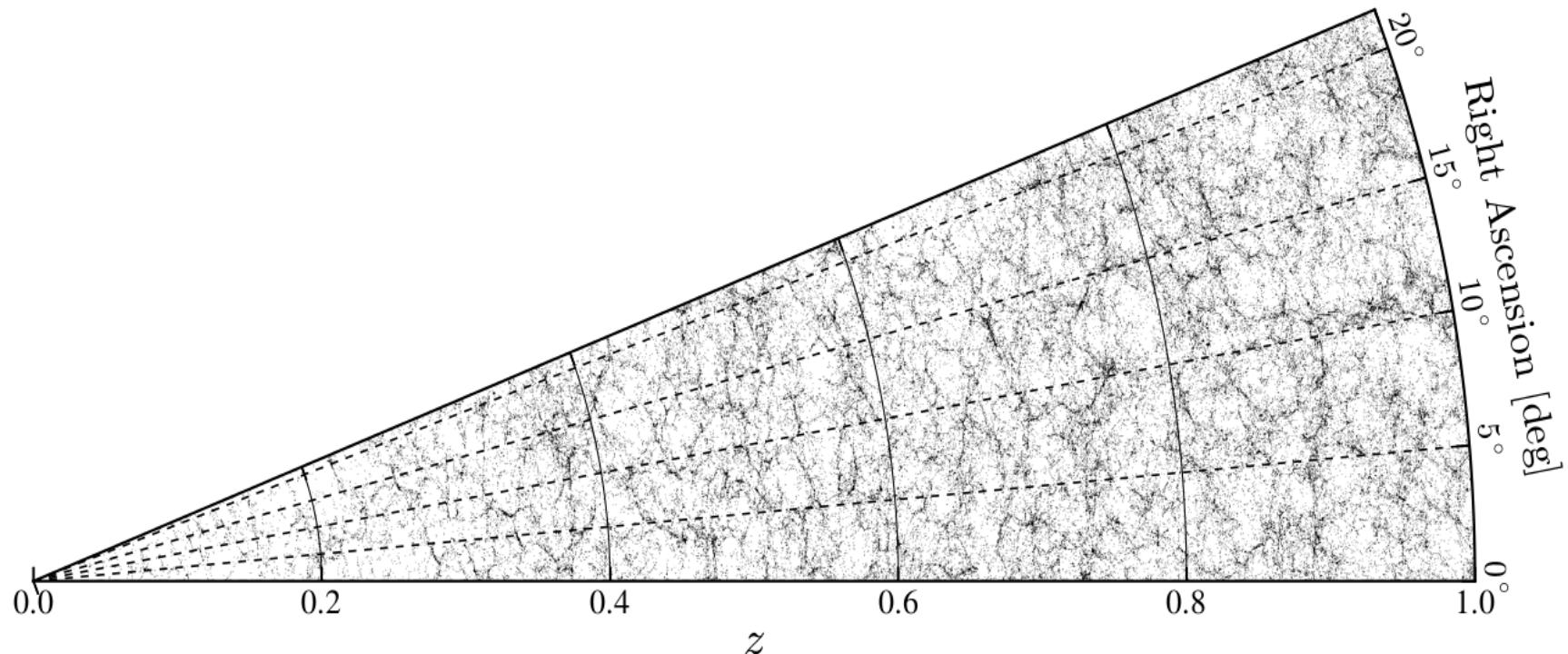
Mock light cone



$\text{Area} \sim 500 \text{ deg}^2$

# HOW DO WE CREATE A MOCK CATALOGUE?

The SAM does not model the lines  
The mock lacks emission lines!



$\text{Area} \sim 500 \text{ deg}^2$

# HOW DO WE CREATE A MOCK CATALOGUE?

Predicting an emission line

## Emission line modelling

Orsi et al. (2014)

Metallicity of galaxies  
( $Z_{\text{cold}}$ )

$$q(Z) = q_0 \left( \frac{Z_{\text{cold}}}{Z_0} \right)^{-\gamma}$$

Ionization parameter  
( $q$ )

$$Z_0 = 0.012 \quad q_0 = 6.5 \times 10^7 \quad \gamma = 0.8$$

**Line Flux**  
Models of

Levesque et al. (2010)

$$F_\lambda(\lambda, q, Z)$$

$$H_\alpha, H_\beta, Ly_\alpha, [O-II], [O-III]$$

$$[OI], [NIII], [SII], [NeIII], [CII], [NII]$$

INSTANTANEOUS

SFR

$$L(\lambda_j)$$

**Luminosity**

10

14 different lines!

# HOW DO WE CREATE A MOCK CATALOGUE?

Predicting an emission line

## Emission line modelling

Orsi et al. (2014)

**Line Flux**  
Models of  
Levesque et al. (2010)

$$F_\lambda(\lambda, q, Z)$$



**Luminosity**

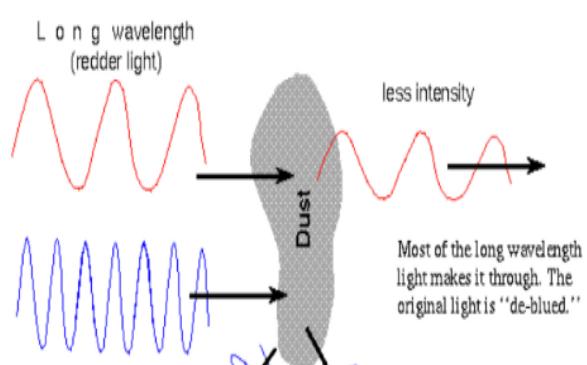
$$L(\lambda_j)$$



**Luminosity DUST  
ATTENUATION**

De Lucua &  
Blaizot (2007)

$$L_{\text{corrected}}(\lambda_j) = L(\lambda_j) * \frac{1 - e^{\tau_\lambda \sec \theta}}{\tau_\lambda \sec \theta}$$

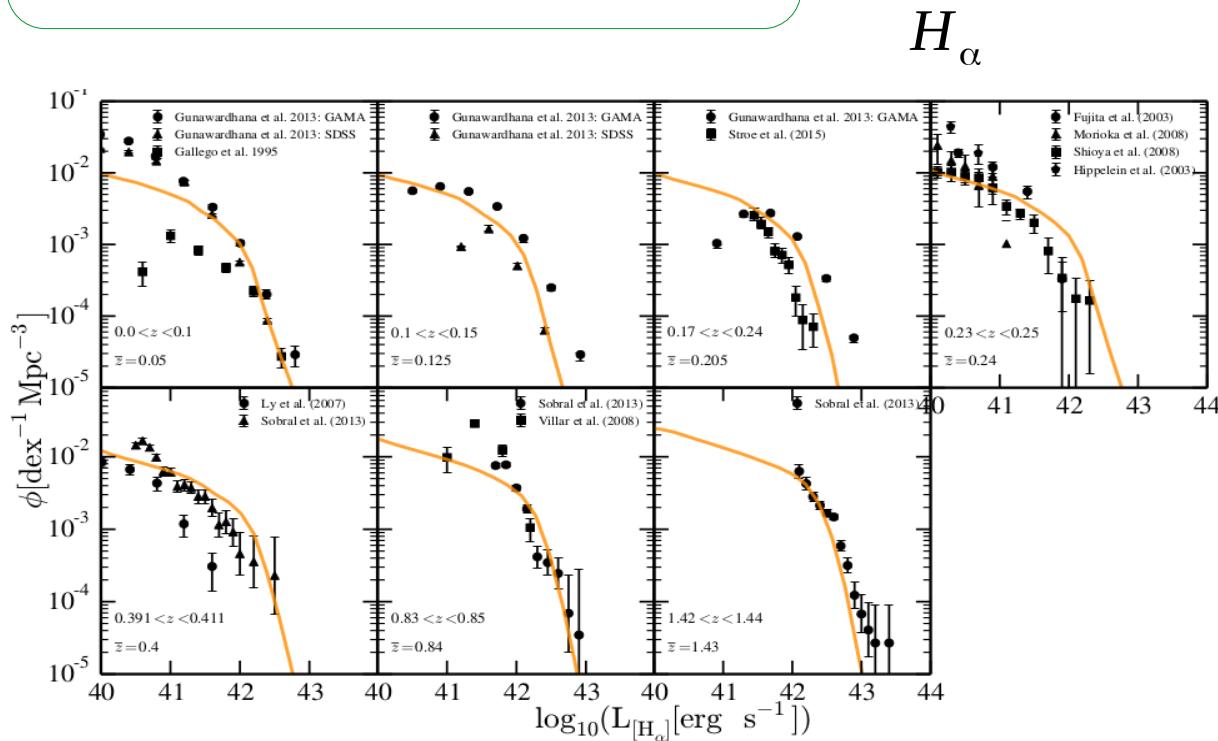


$$\tau_\lambda = C(z) Z_{\text{cold}} \frac{A_V}{A_B} \frac{A(\lambda)}{A_V}$$

$H_\alpha, H_\beta, [O-II], [O-III]$

# HOW DO WE CREATE A MOCK CATALOGUE?

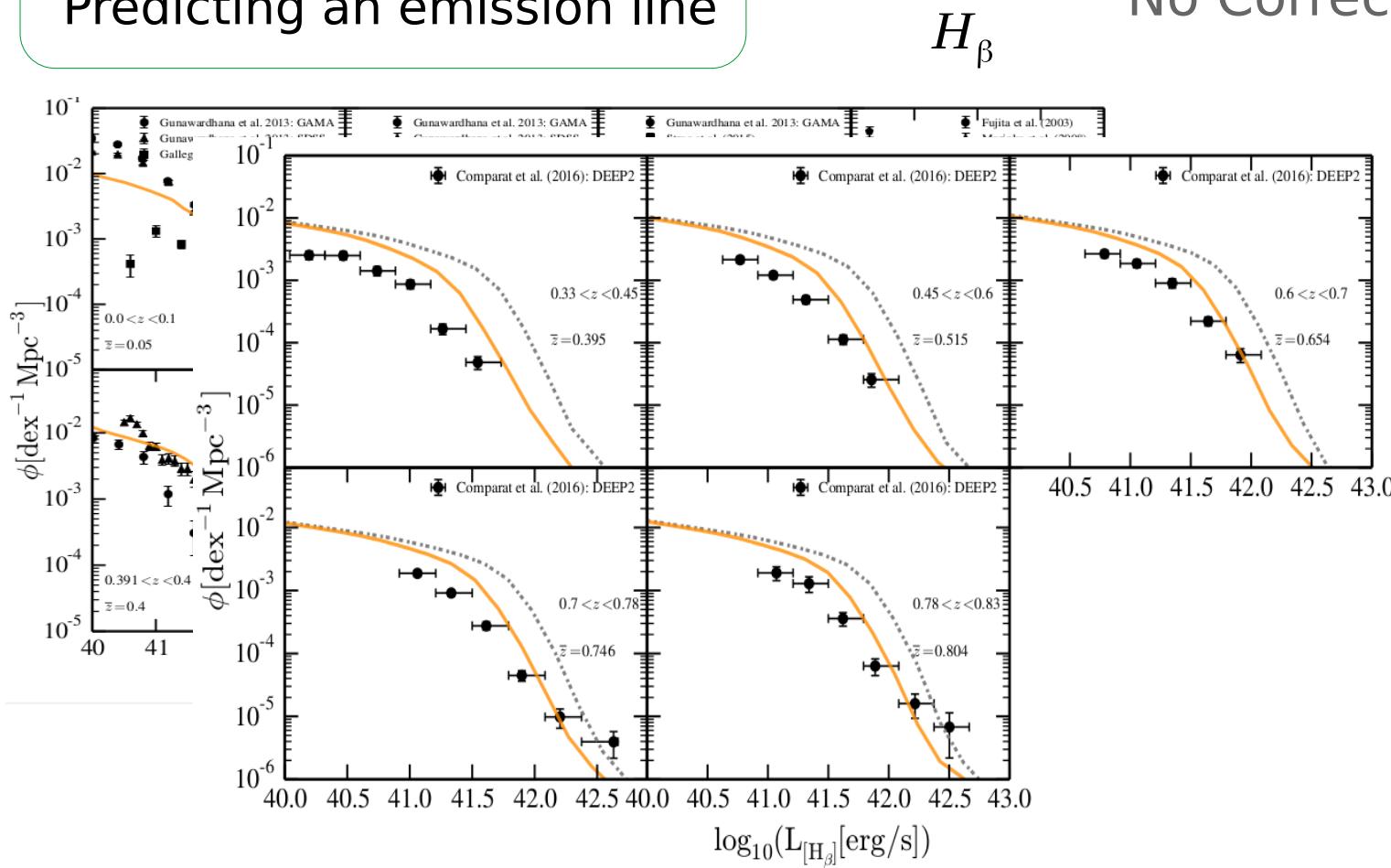
Predicting an emission line



# HOW DO WE CREATE A MOCK CATALOGUE?

Predicting an emission line

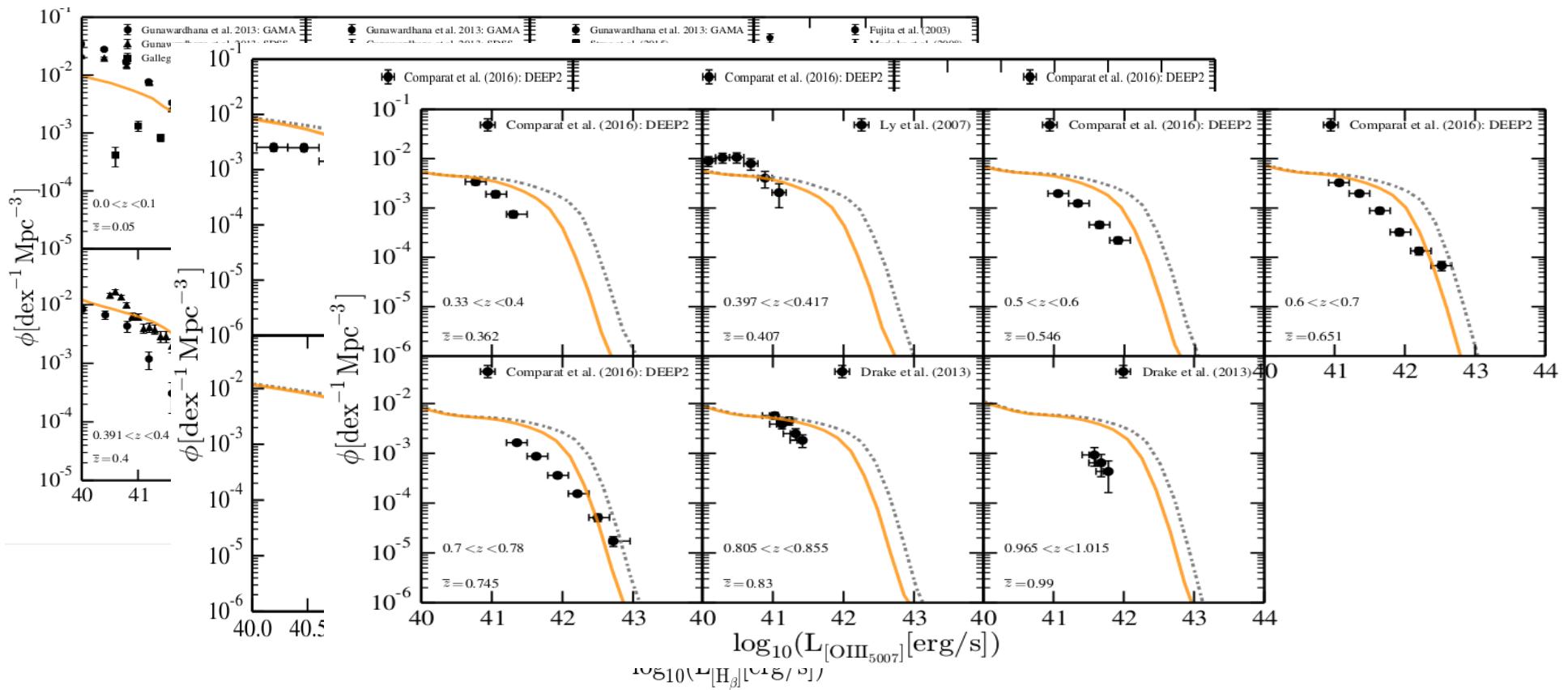
Dust corrected  
No Corrected



# HOW DO WE CREATE A MOCK CATALOGUE?

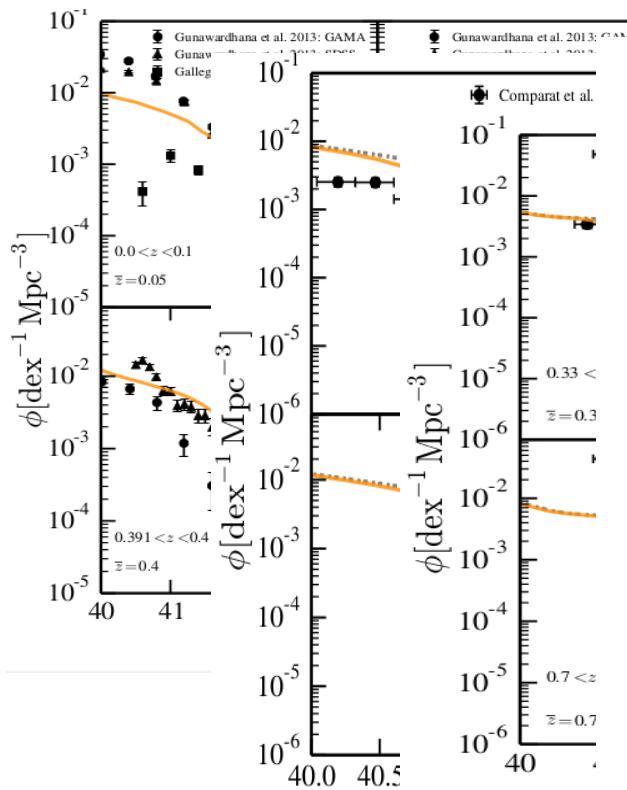
Predicting an emission line

Dust corrected  
No Corrected  
 $[O-III]_{5007}$



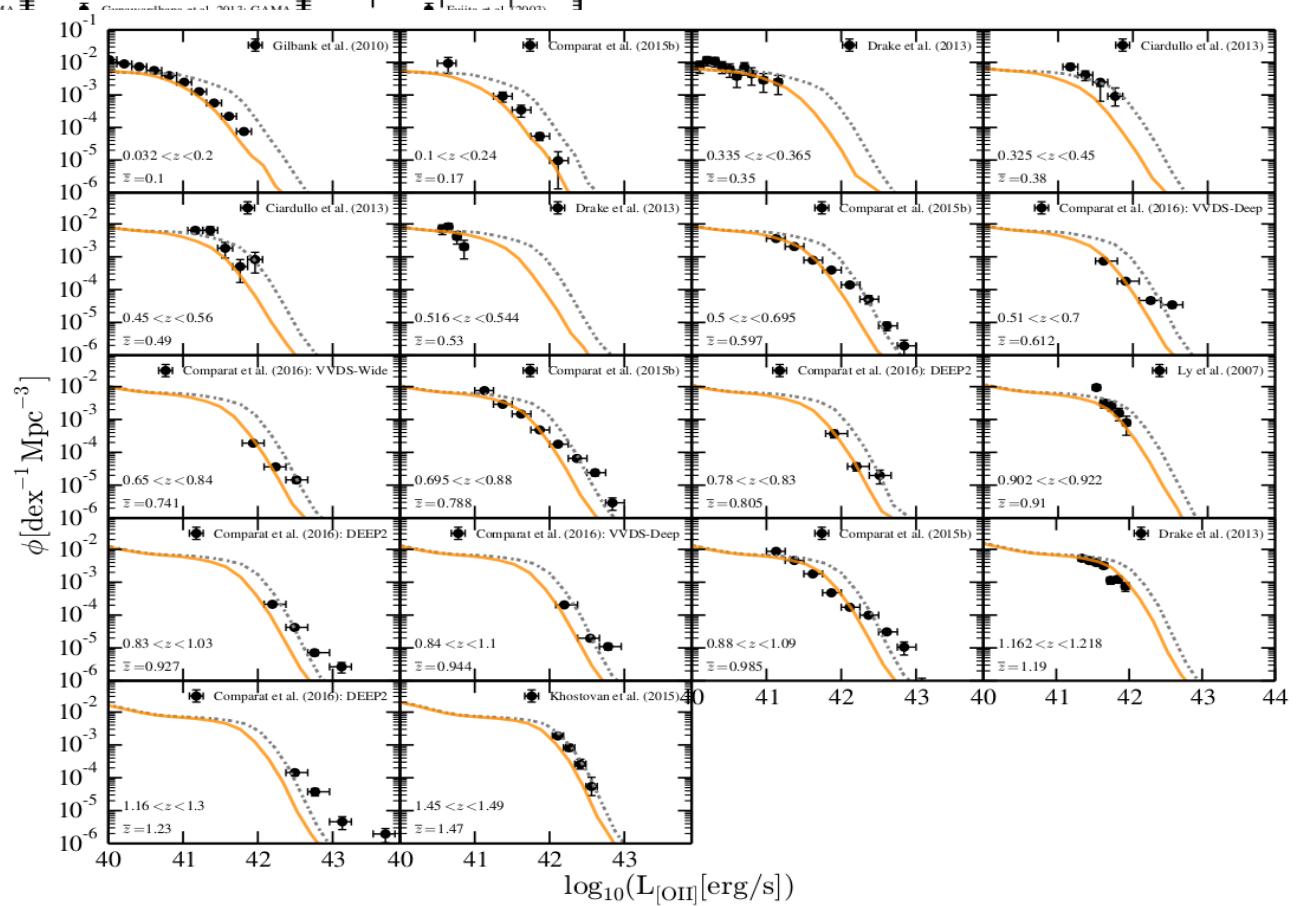
# HOW DO WE CREATE A MOCK CATALOGUE?

Predicting an emission line

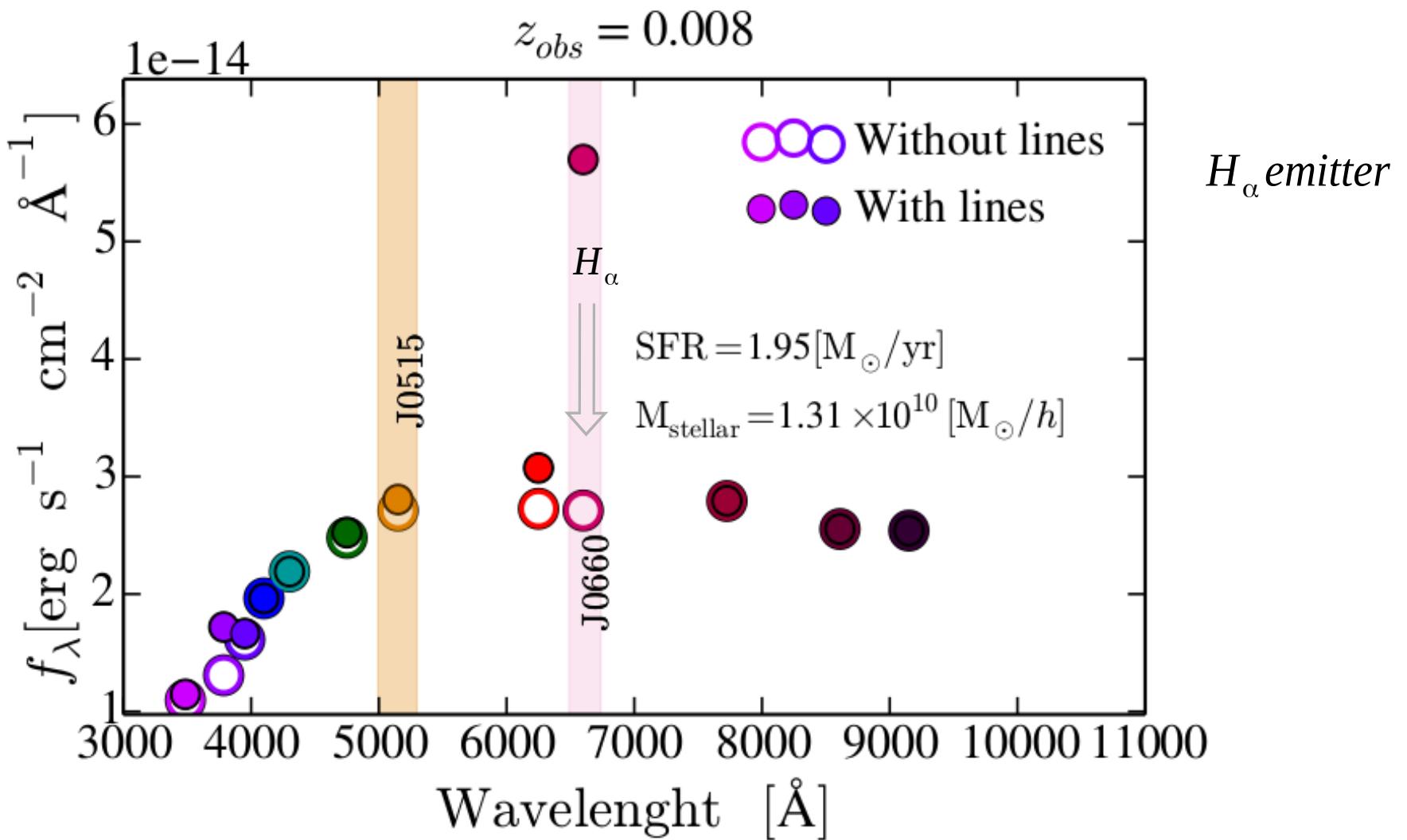


[O-II]

Dust corrected  
No Corrected

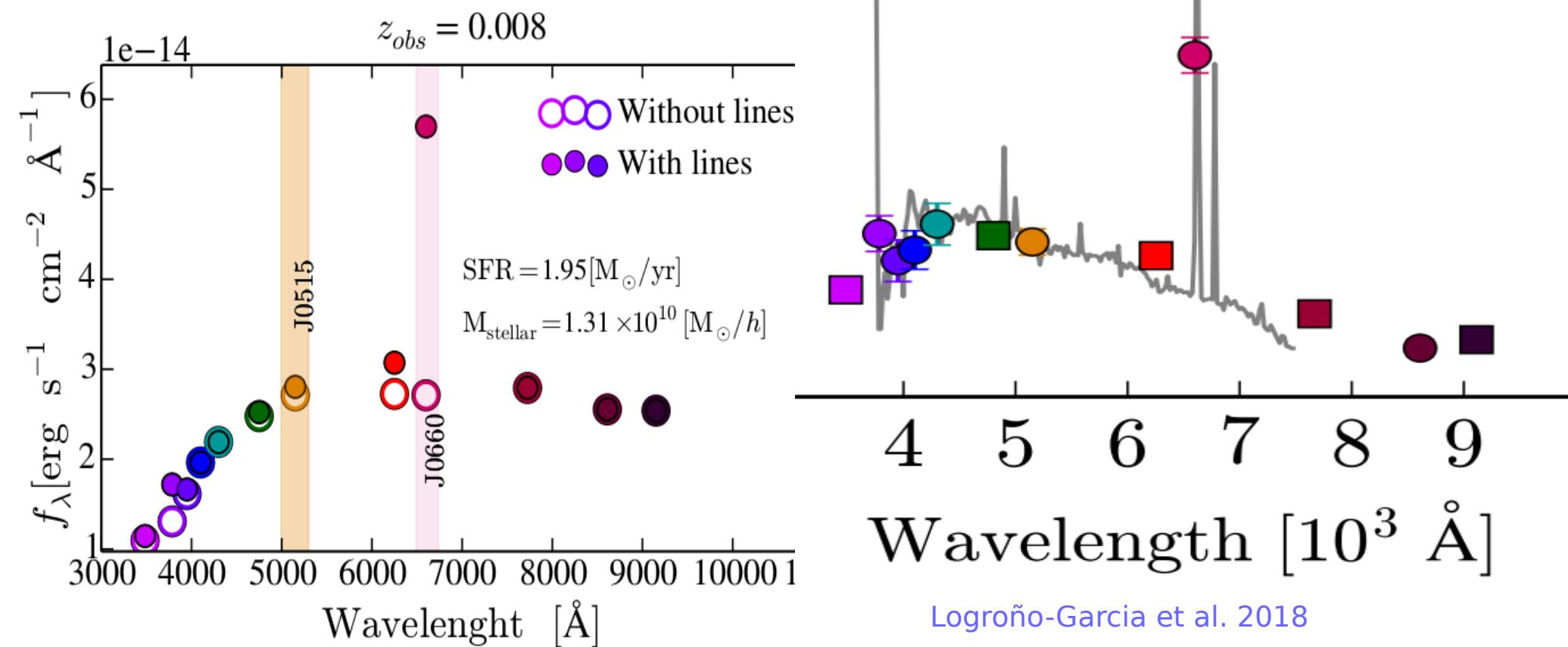


# EXAMPLES

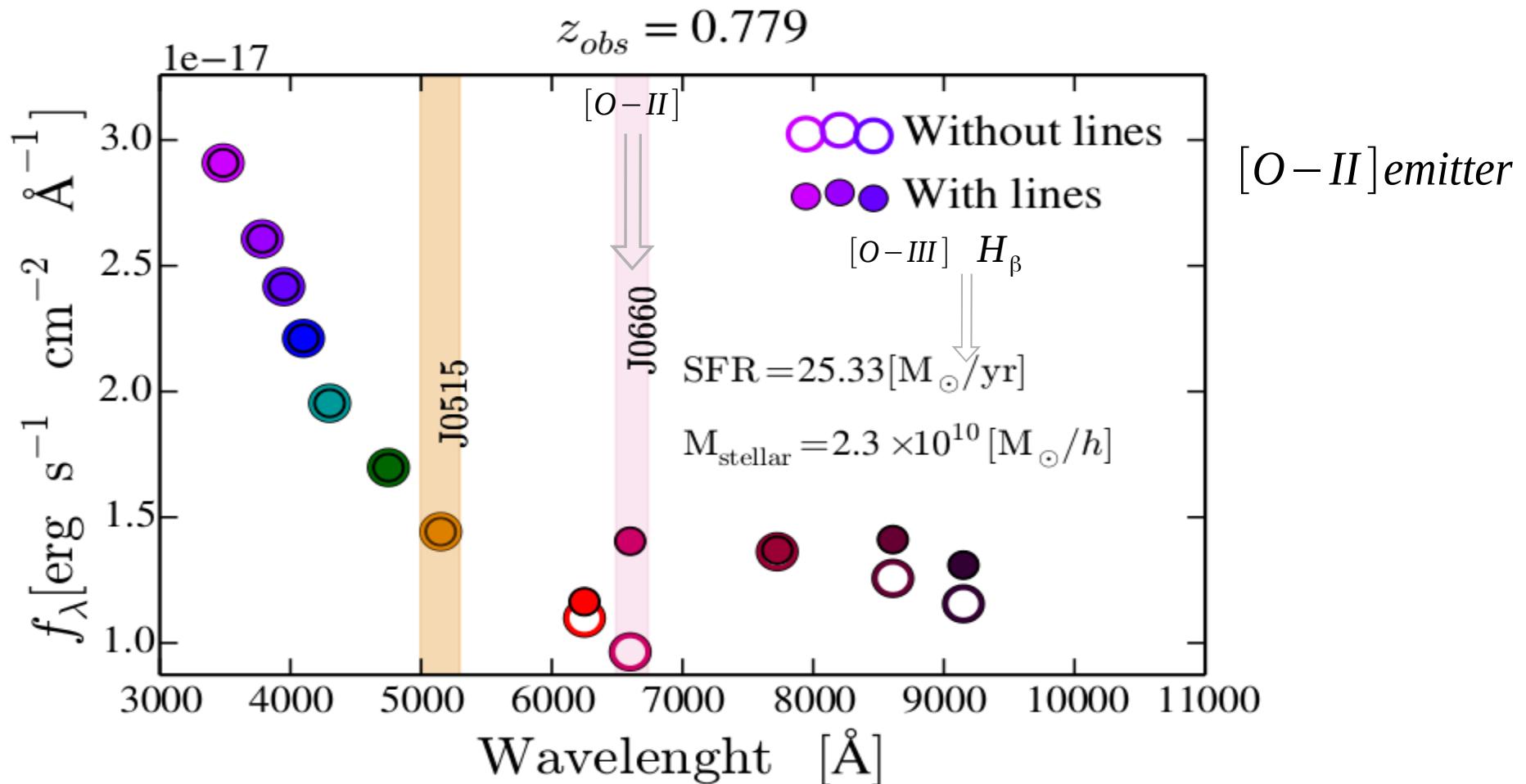


# EXAMPLES

8



# EXAMPLES



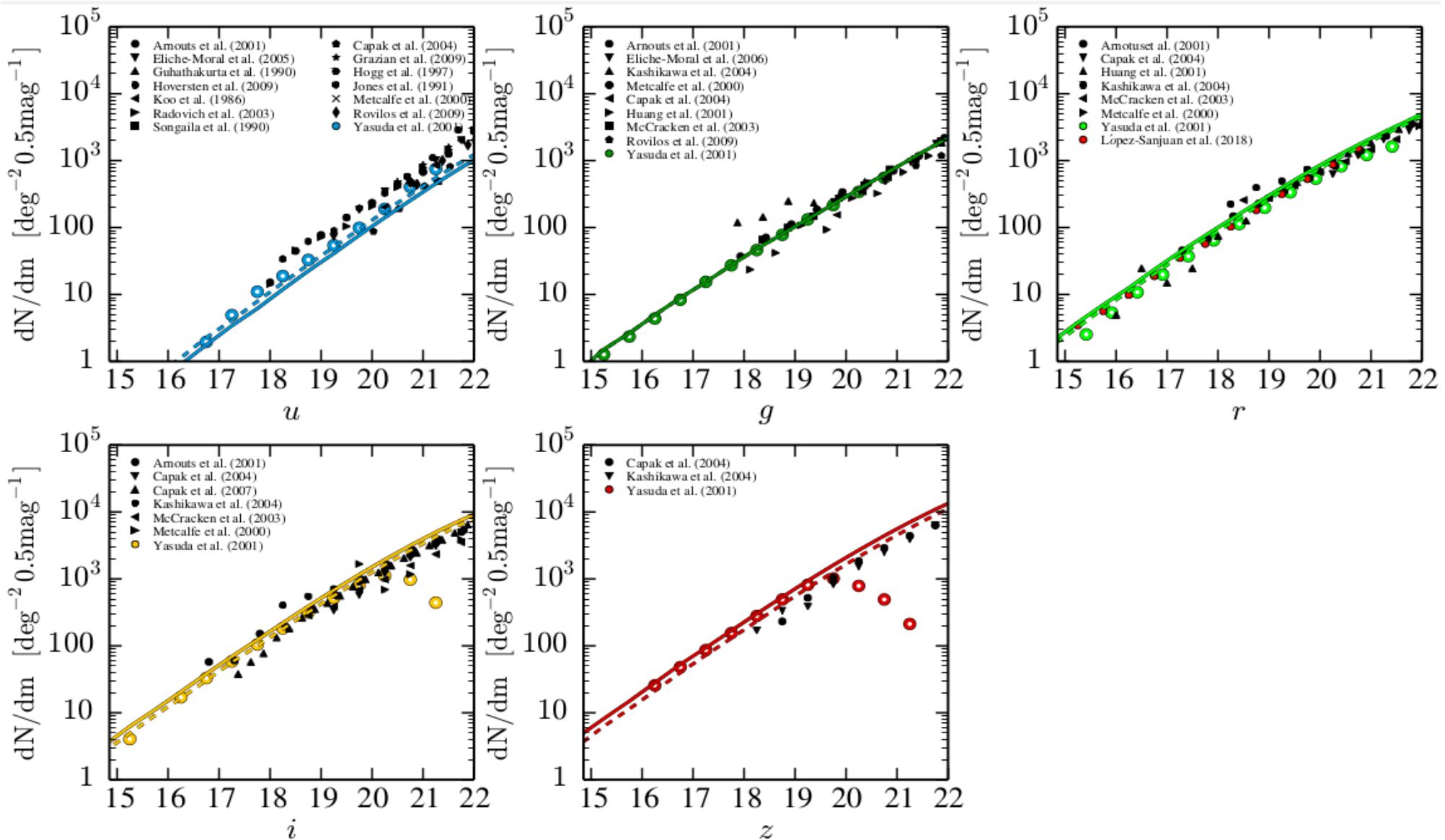
## Use of JPLUS mock

We have a virtual survey

We can do tests!

# Use of JPLUS mock

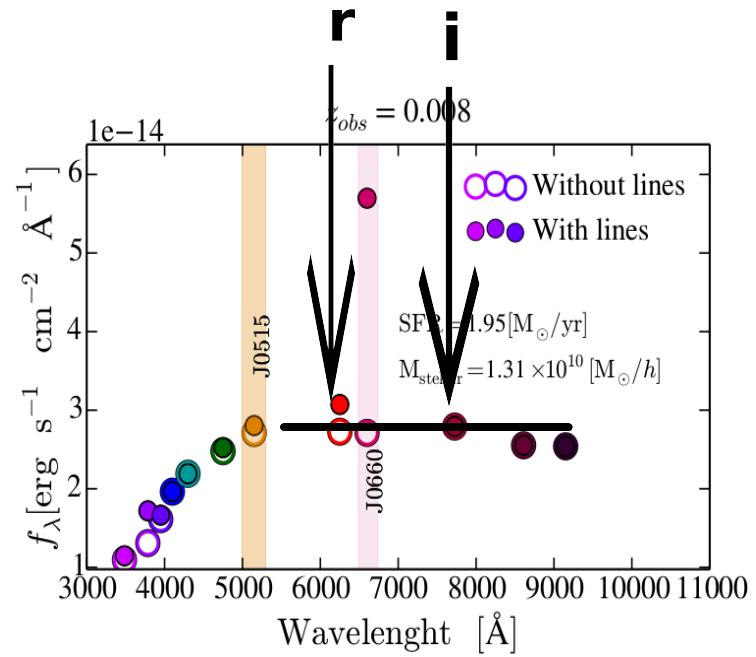
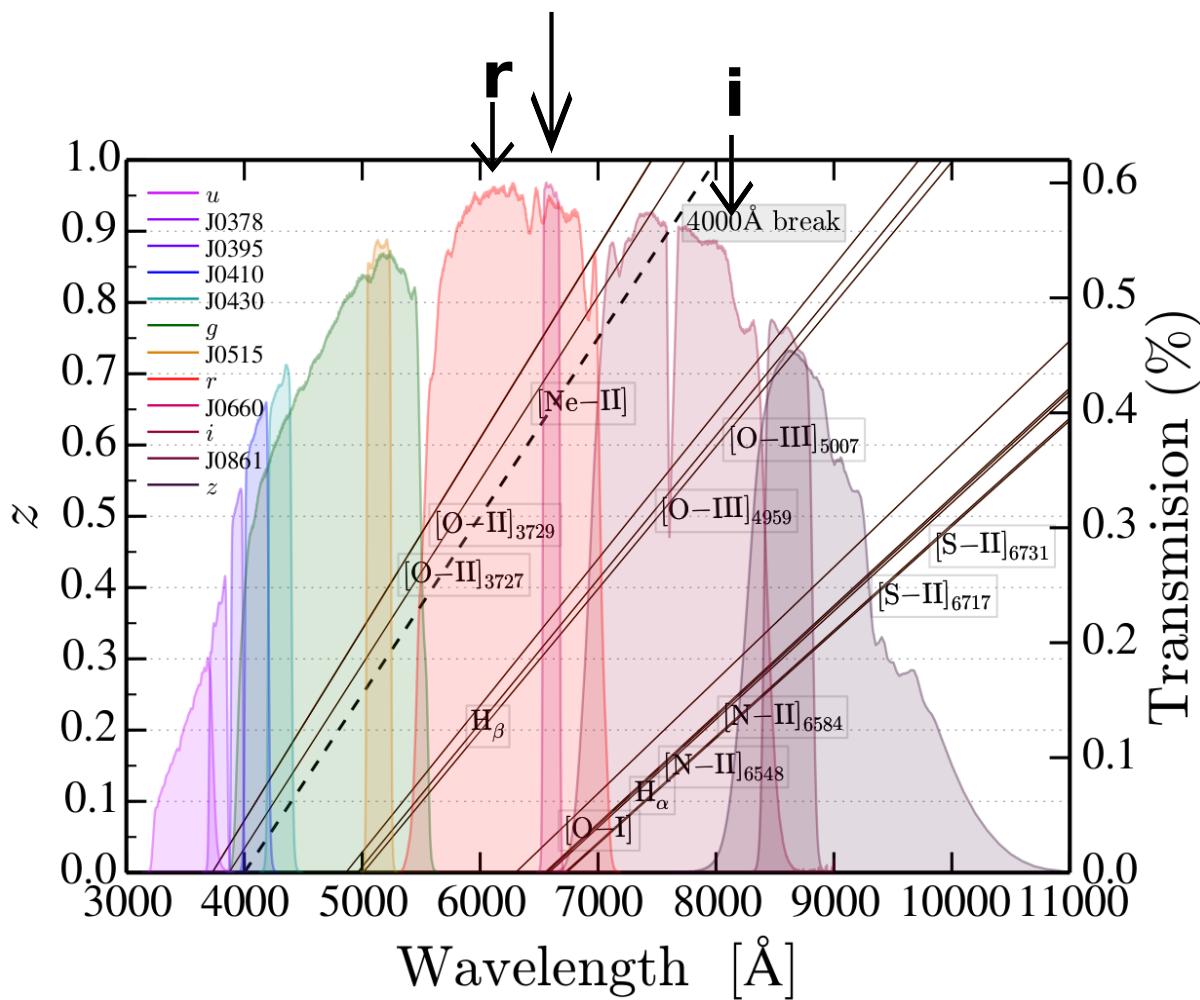
## Check: NUMBER COUNTS



# Use of JPLUS mock

**Three filter method to extract emission lines** ( Vilella-Rojo et al. 2015 )

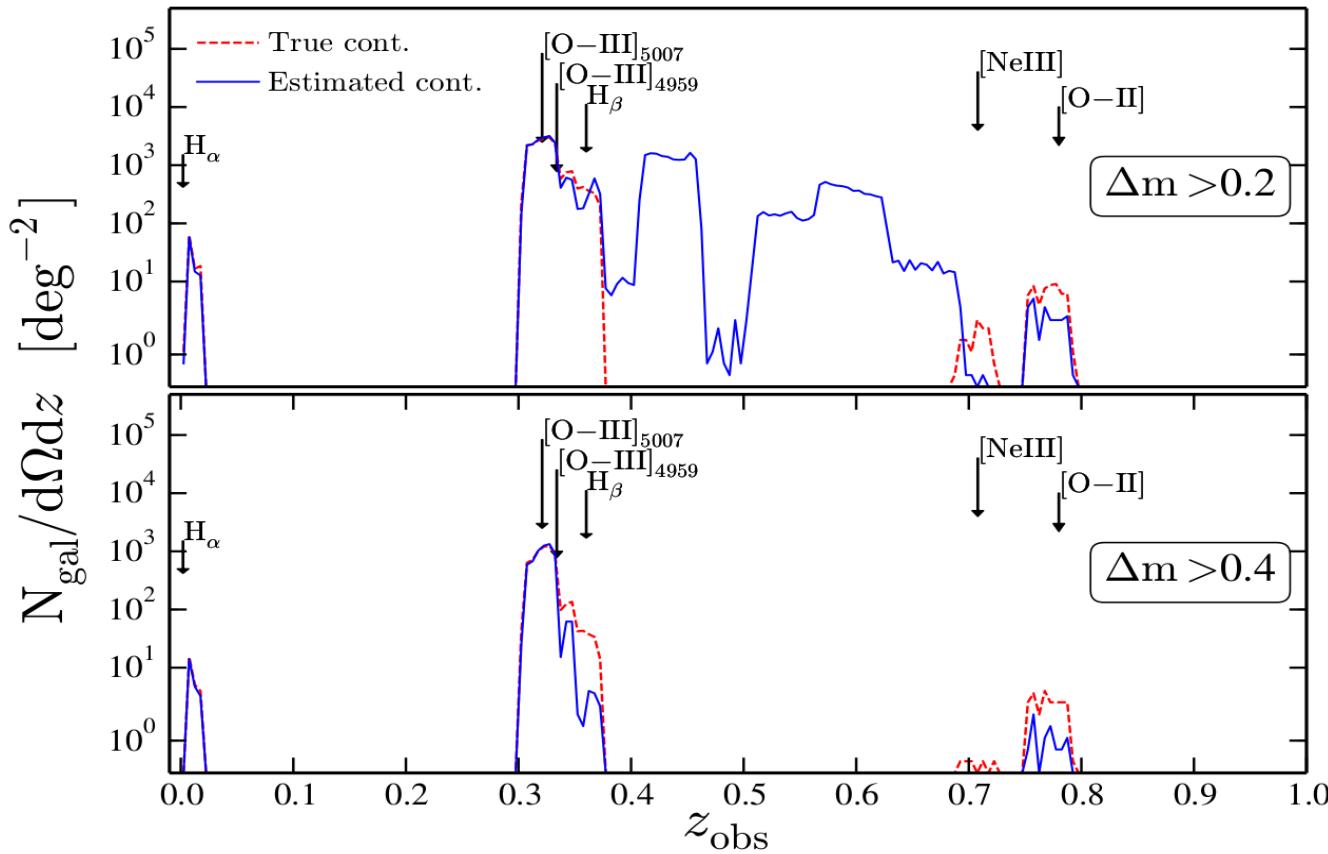
**F6660** → Sensible to emission lines!



# Use of JPLUS mock

## EMISSION LINE DETECTION

### Magnitude excess selection



$$\Delta m = m^{J660} - m_C^{J660}$$

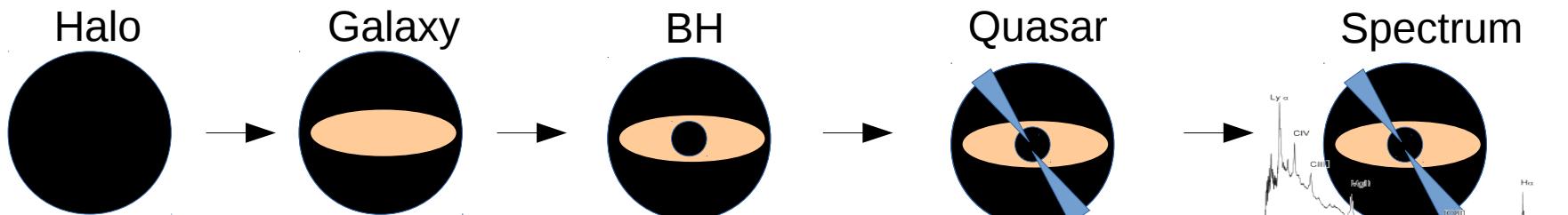
$$m_C^{J660, 3FM}$$
  

$$m_C^{J660, \text{TRUE}}$$

Interlopers 4000 Å break

# Use of JPLUS mock

## QUASAR MOCKS

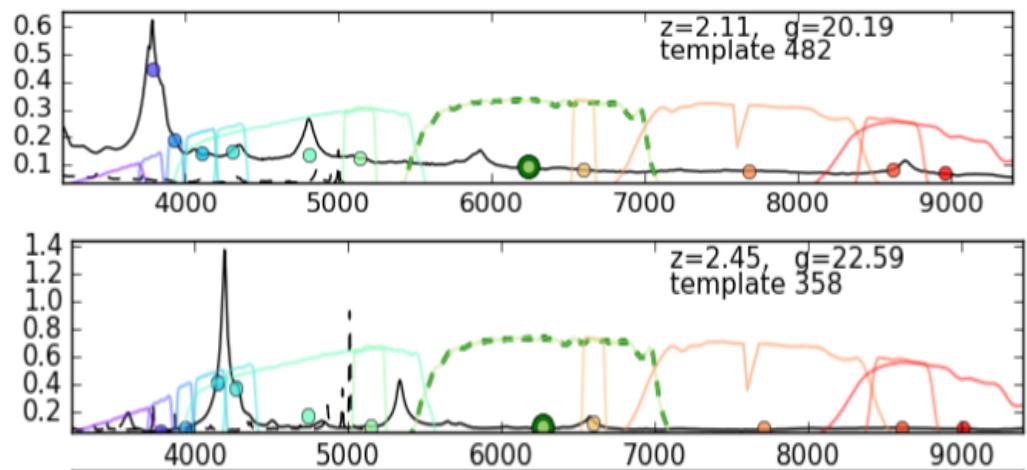
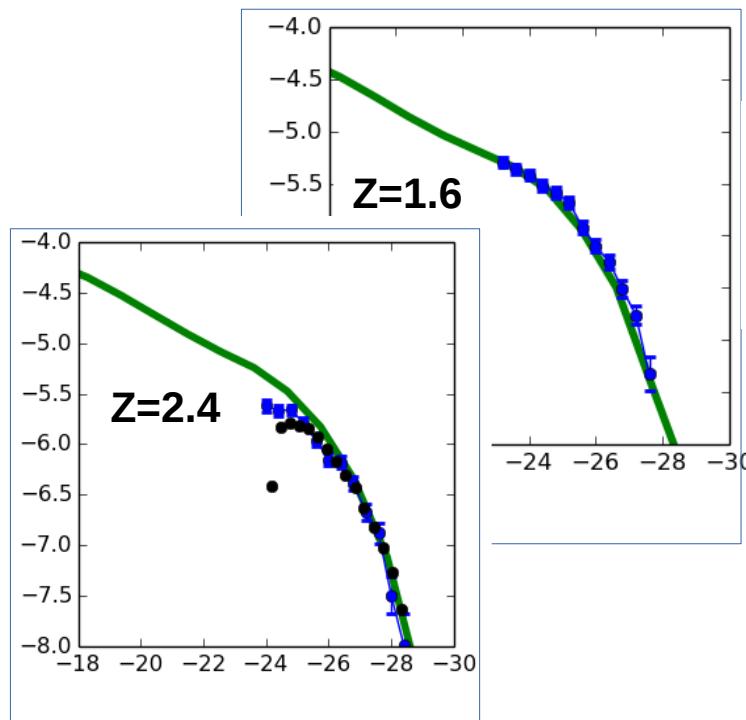


$M_{\text{halo}}-M_{\star}$   
relation

$M_{\star}-M_{\text{BH}}$   
relation

- Eddington distribution
- duty cycle

Quasar  
templates



# Conclusion

- 1 - Semi analitical model
- 2 - Dark matter simulation
- 3 - Emission line modelling
- 4 - Dust correction



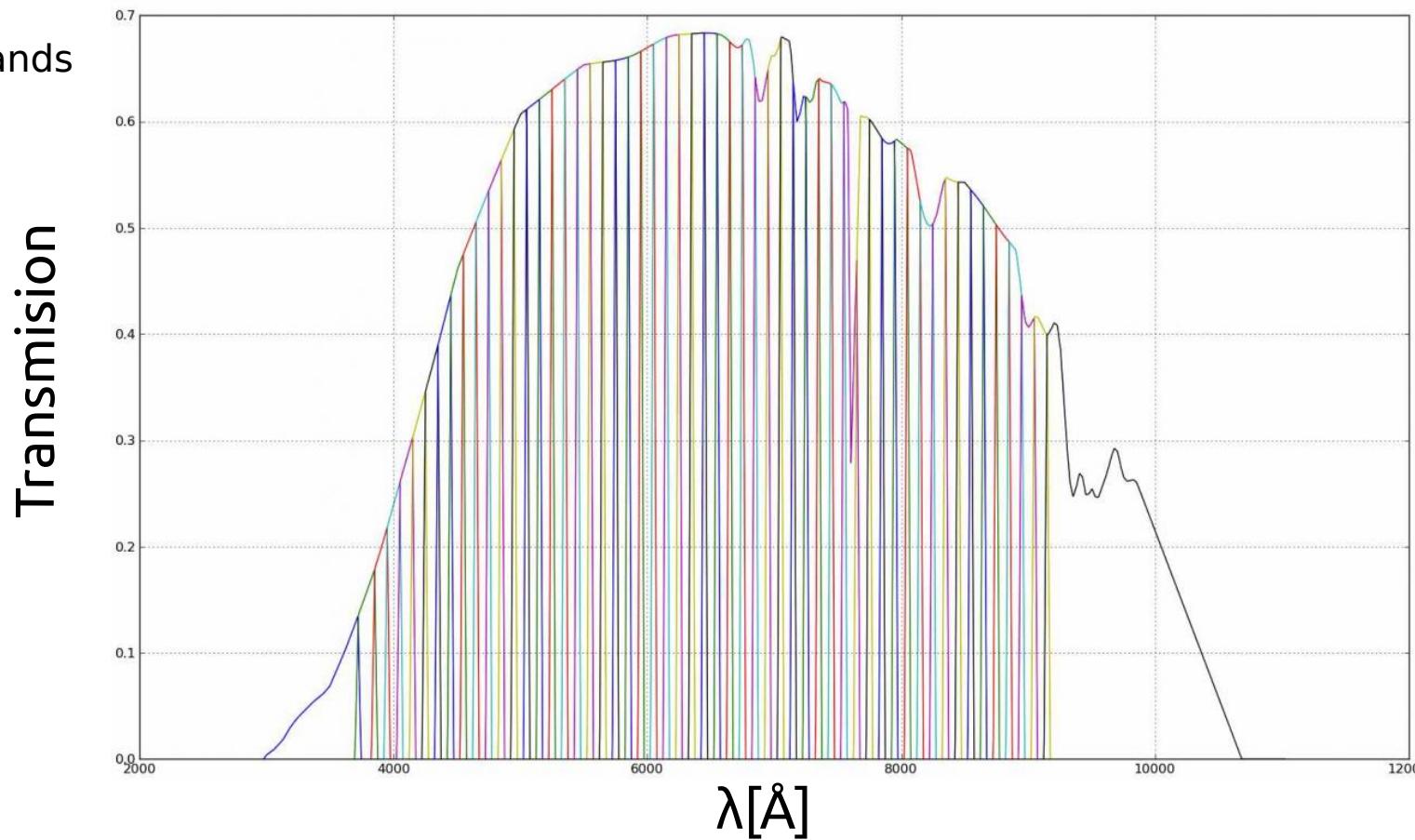
Mock catalogues for  
**PHOTOMETRIC SURVEYS**

# Future work

Apply the machinery to J-PAS survey

5 BROAD/MEDIUM bands

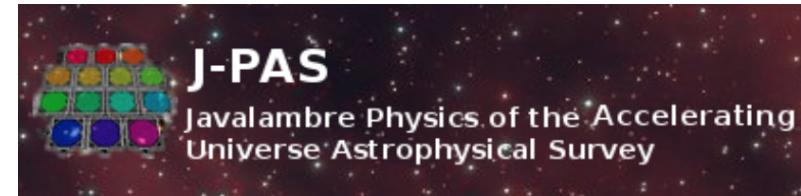
54 NARROW bands





C . Hernandez-Monteagudo

(Wednesday afternoon session)



S . Bonoli

(Wednesday afternoon session)



# THANKS!



**UNIÓN EUROPEA**  
Fondo Europeo de Desarrollo Regional  
*"Una manera de hacer Europa"*



GOBIERNO  
DE ESPAÑA

MINISTERIO  
DE ECONOMÍA, INDUSTRIA  
Y COMPETITIVIDAD

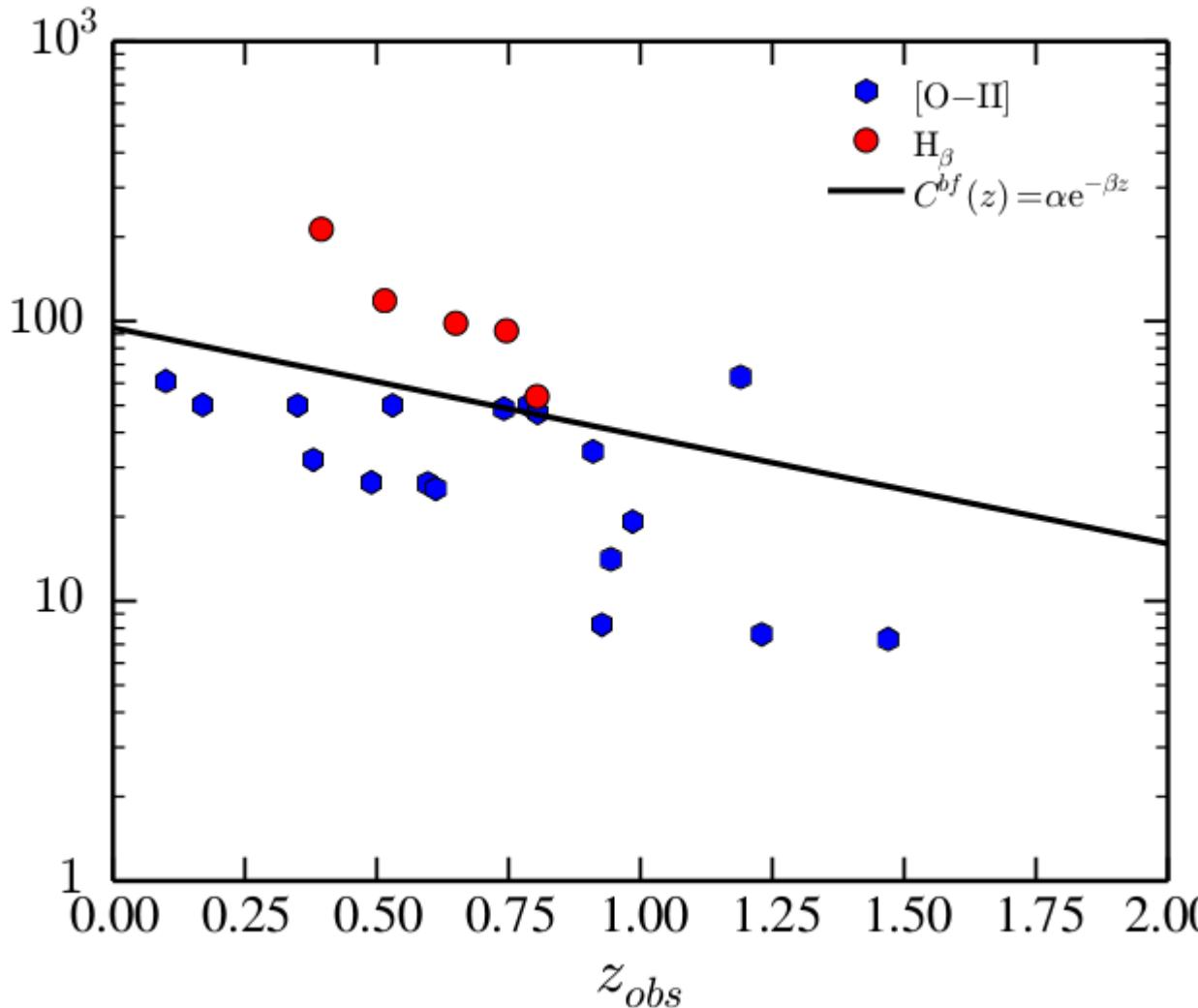
**GOBIERNO  
DE ARAGÓN**  
Actividades de investigación subvencionadas por Gobierno de Aragón

# HOW DO WE CREATE A MOCK CATALOGUE?

Predict a  
**Line Flux**  
 Models of  
 Levesque et al. (2014)

$$F_\lambda(\lambda, q, Z)$$

Dust  
 B



si et al. (2014)

sity  
 Γ  
 TION

$$(\lambda_j) * \frac{1 - e^{-\tau_\lambda \sec \theta}}{\tau_\lambda \sec \theta}$$

$$\frac{A_V}{A_B} \frac{A(\lambda)}{A_V}$$

II],[O-III]