



UPWARDS  
UNDERSTANDING PLANET MARS



Scientific Workshop "From Mars Express to Exomars"  
27 - 28 February 2018  
ESAC (ESA), Madrid, Spain

# Synergistic retrieval of H<sub>2</sub>O vapor in Mars' atmosphere: the path to a systematic 3D exploration of H<sub>2</sub>O vapor

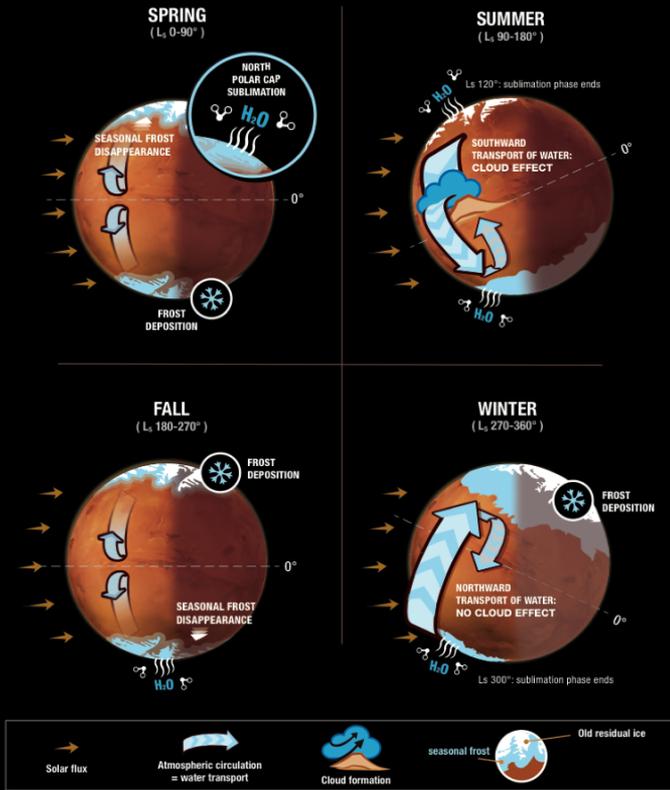
Franck Montmessin (LATMOS)



Stéphane Ferron (ACRI-ST)



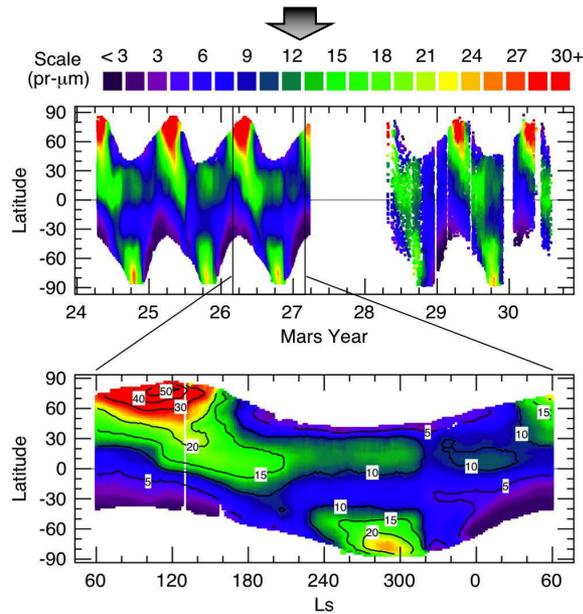
## MARS WATER CYCLE



Montmessin et al. (2017)

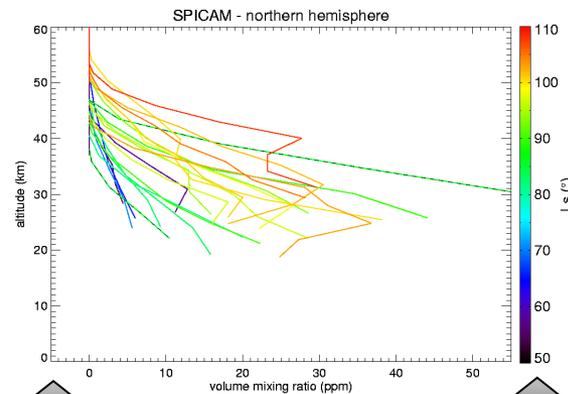
Exploring H<sub>2</sub>O vapor  
in every dimension

Very dense & wide coverage  
of nadir observations



**NADIR**  
retrieves the column  
abundances of water  
and essentially explores  
(*x, y*) - variability.

Figure: Multi-annual  
compilation of nadir TES (TIR)  
and CRISM (NIR) results  
covering > 5 Martian Years.



Vertical profiles only  
at sunrise/sunset  
and only in selected places

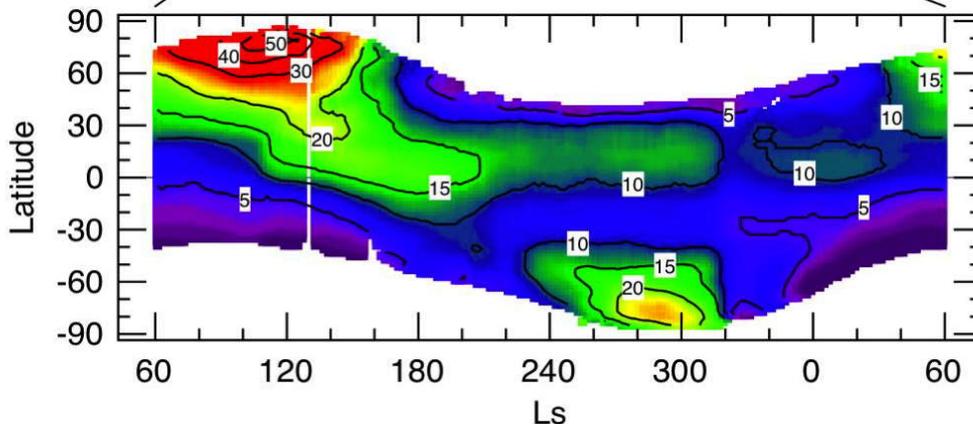
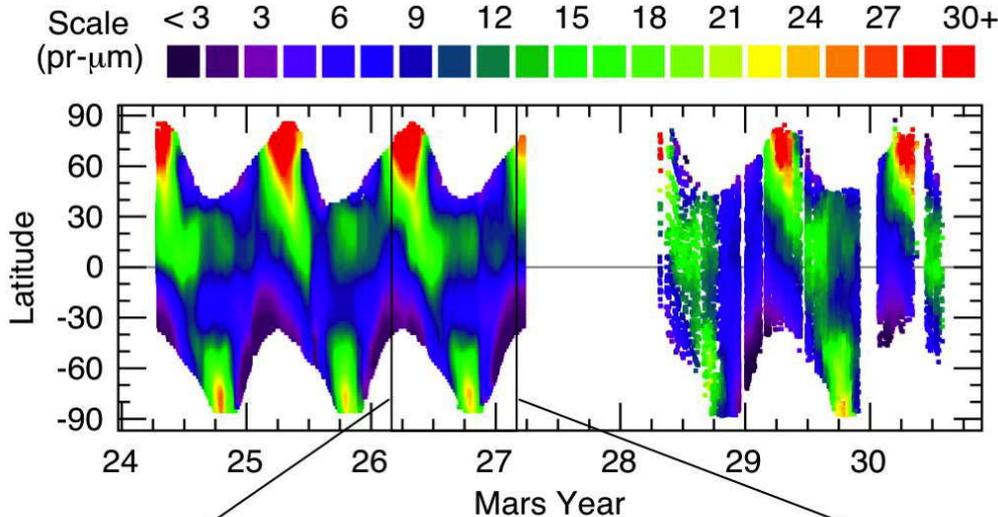
**SOLAR OCCULTATION**  
retrieves the local  
concentrations of water  
and essentially explores  
(*z*) - variability.

Figure: Altitude profiles  
collected by SPICAM-IR.  
(Maltagliati et al., 2011).

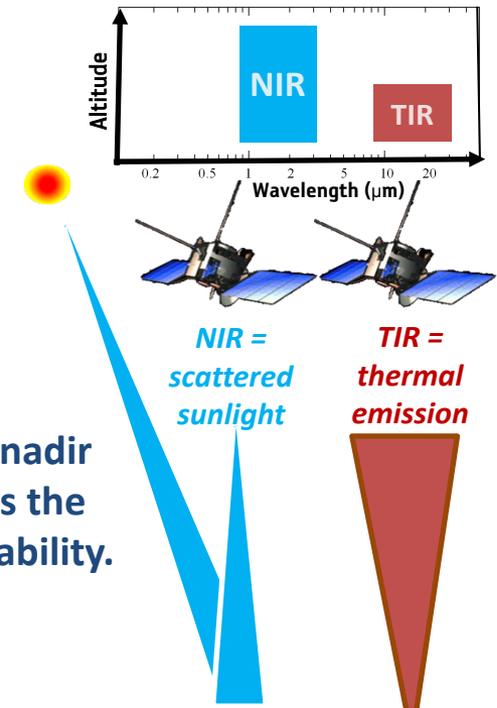
## Synergistic retrieval

=

Simultaneous inversion of H<sub>2</sub>O in 2 spectral domains: **NIR** & **TIR**



**NIR + TIR** in nadir might access the (x, y, z) - variability.



Only Mars Express enables synergy with SPICAM and PFS.

Study funded by the UPWARDS project led by M.L. Valverde  
(Horizon 2020 – EU program)

## Study Goals:

1. Assess theoretical performances of synergy
2. Apply it on actual MEX datasets



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## Methodology:

- M1. Build a **forward model** to solve the radiative transfer problems in the TIR and NIR bands
- M2. Create **synthetic datasets** representative of existing ones (SPICAM and PFS on Mars Express)
- M3. Insert the forward model into a **bayesian optimization scheme** to retrieve **optimal parameters** for H<sub>2</sub>O:
  - (1) column-abundance
  - (2) parameter related to vertical distribution (partial column in the first 5 km)

## Forward model:

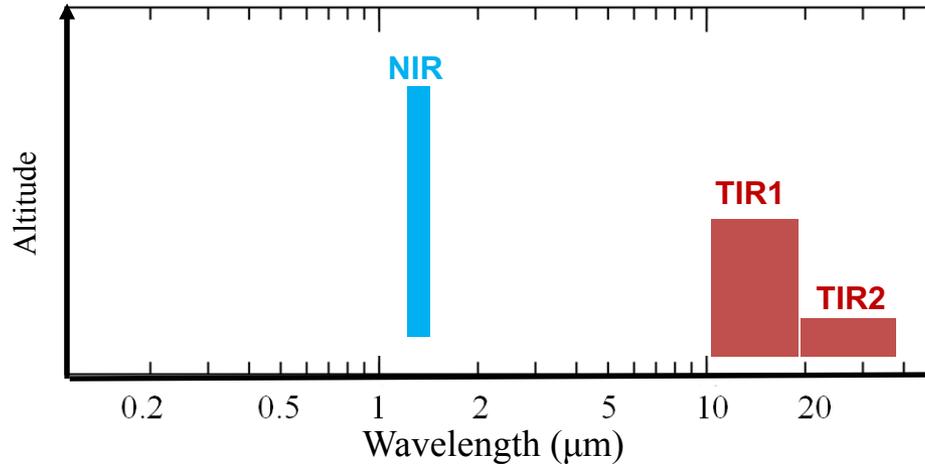
- LIDORT solver
- Dust model from Wolff & al. (2009), no ice cloud included
- Gaseous absorption (Correlated-K D approach, HITRAN 2012 + Brown 2007)
- Lambertian surface assumed
- MCD v5.2 used to constrain *prior* state vector for T<sub>s</sub>, T<sub>atm</sub>, etc.

## Synthetic Data set (Goal 1):

- Reproduce SPICAM-type data @ 1.38 μm (NIR)
- Reproduce PFS-type data @ **20 μm for H<sub>2</sub>O (TIR2)** and **15 μm for Temperature (TIR1)**

## Real Data set (Goal 2):

- Made of **233 different joint SPICAM and PFS observations** performed during Martian Year 27.
- Each observation corresponds to the compilation of 9 spectra.



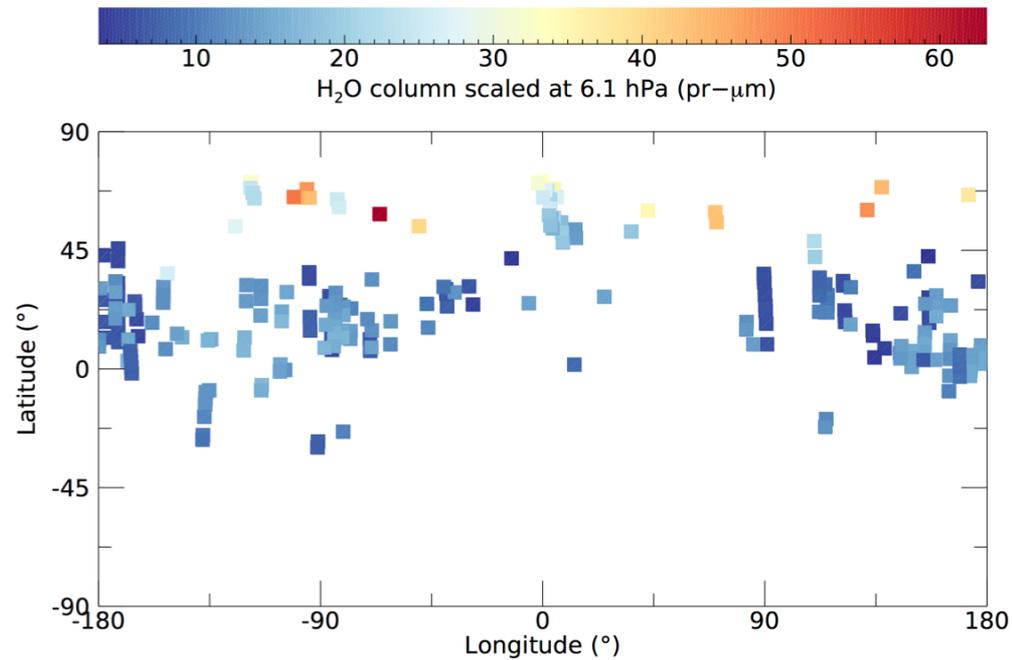
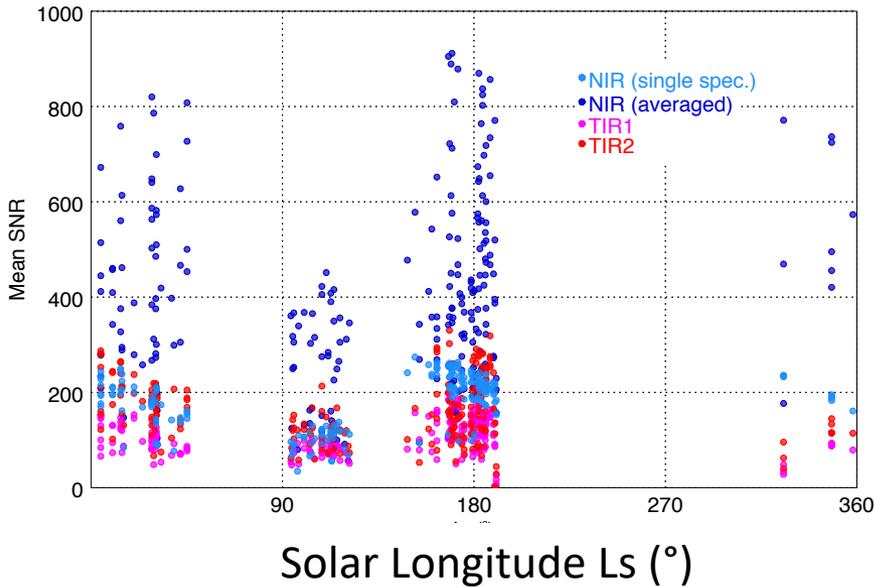
Domain	Wavelength Range	Spectral features	Retrieved Parameters	Instrument
NIR	1.35-1.43 μm	H <sub>2</sub> O	H <sub>2</sub> O ( z ) Albedo Dust	SPICAM-IR
TIR1	12.2-19.2 μm	CO <sub>2</sub>	T(z) Surface Temperature Dust Opacity	PFS-LW
TIR2	20-34 μm	H <sub>2</sub> O	H <sub>2</sub> O ( z )	PFS-LW

The number of independent parameters (Degree Of Freedom-DOF) is given by **the trace of the covariance matrix** (computed through optimization procedure).

	NIR -SPICAM-	TIR -PFS-	NIR+TIR -SPICAM+PFS-
<i>Wet area (moderate dust) – north pole in summer</i>			
<b>DOF</b> (Degree Of Freedom)	1.13	0.87	<b>1.68</b>
<i>Wet area (no dust) – north pole in summer</i>			
<b>DOF</b>	1.14	1.16	<b>1.87</b>
<i>Dry area (low dust) – southern mid latitude in summer</i>			
<b>DOF</b>	1.22	0.36	<b>1.56</b>

- With synergy, nearly **2 independent parameters** for H<sub>2</sub>O can be retrieved compared to 1 (whole column) with NIR or TIR nadir observations alone.
- **Theoretically, H<sub>2</sub>O vertical distribution can be constrained by NIR+TIR synergy**

Testing made on 233 samples of MY27.  
Each sample consists of 9 spectra.



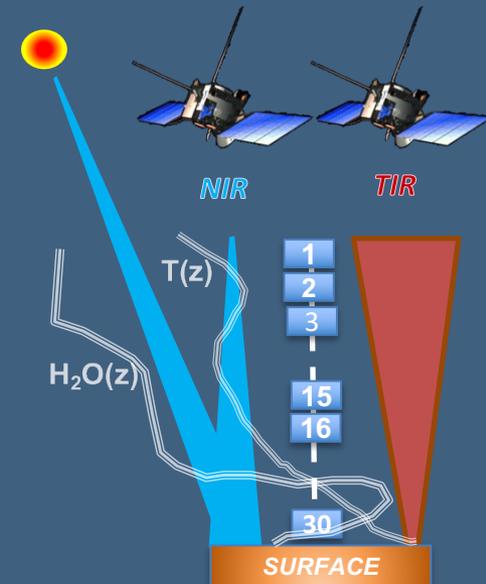
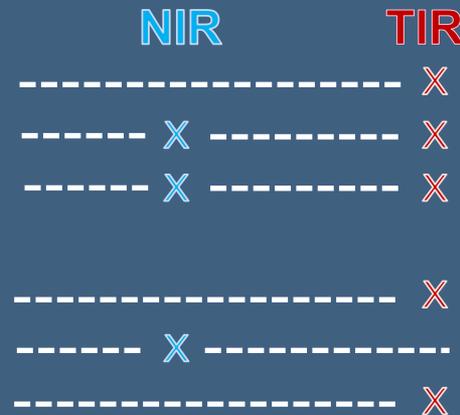
Retrieval is based on a bayesian estimation with simultaneous adjustment of 64 parameters with a prior assumption on averaging kernels and on parameters (constrained by MCD).

Atmosphere

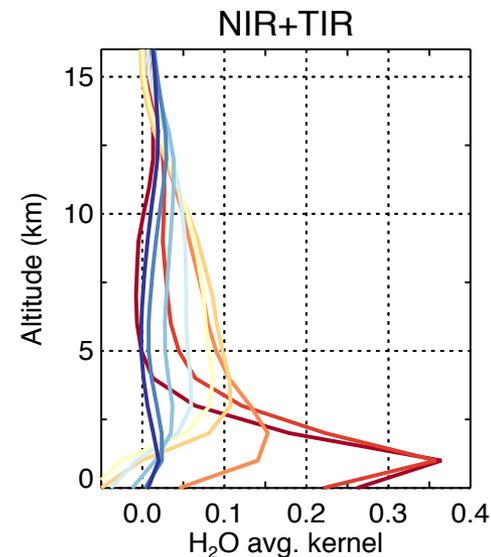
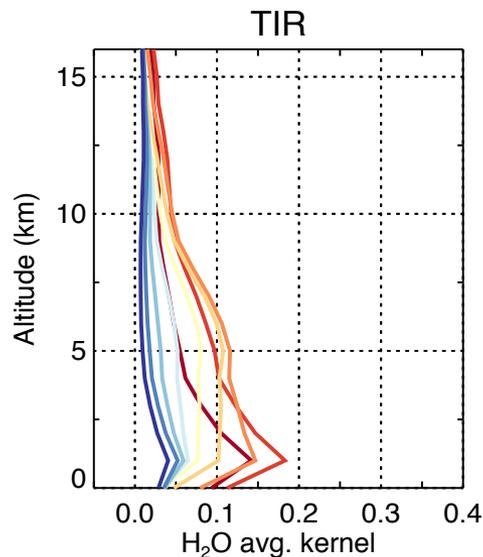
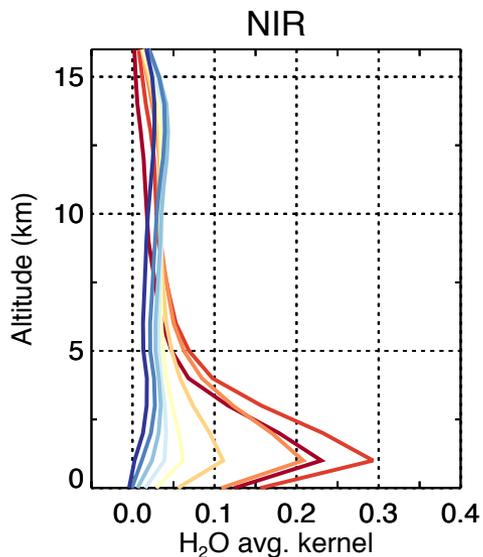
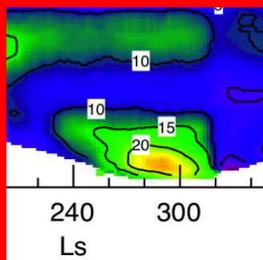
- Temperature: 30 levels
- H<sub>2</sub>O mixing ratio: 30 levels
- Dust column-opacity

Surface

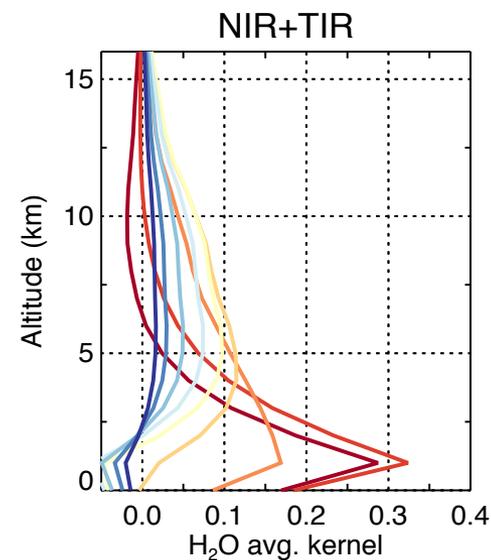
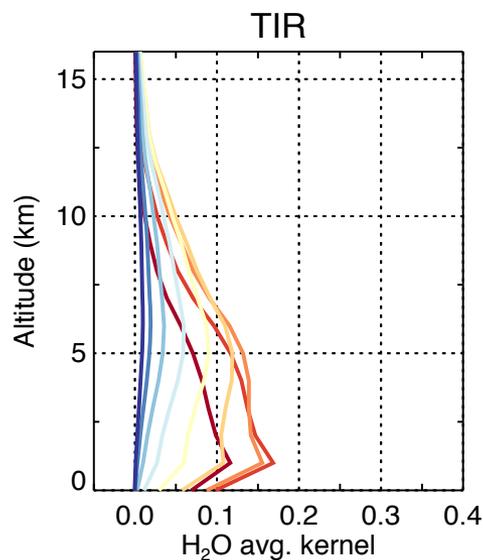
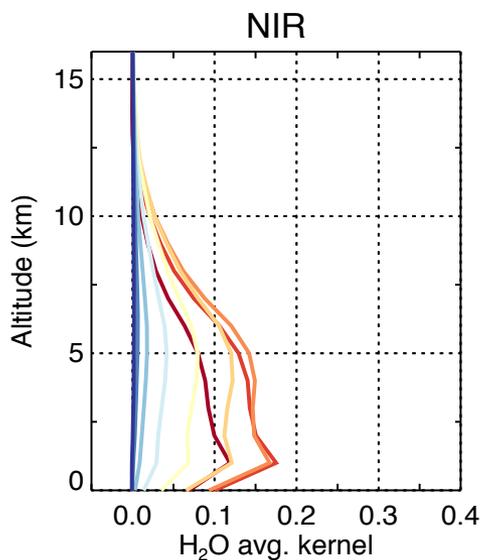
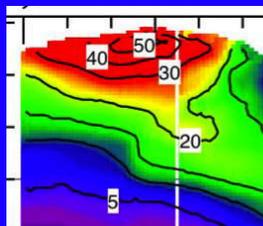
- Temperature
- Albedo
- Emissivity



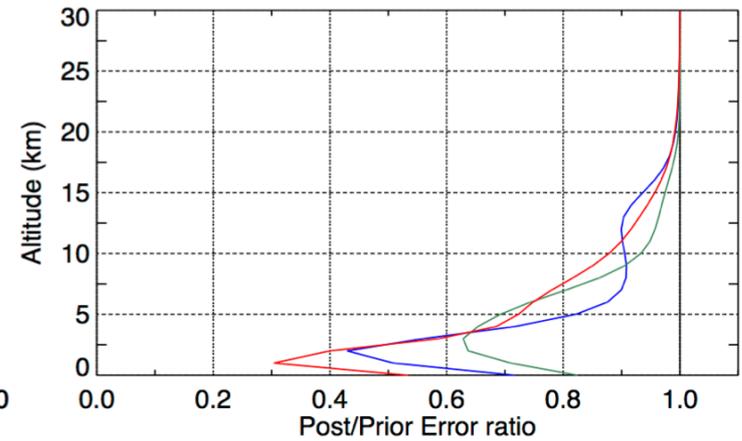
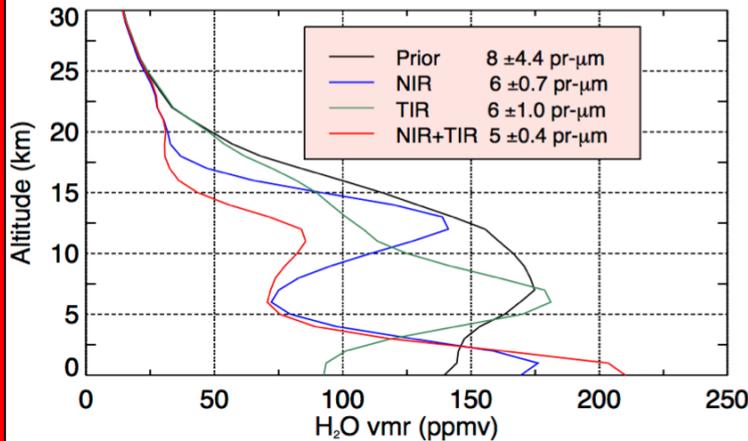
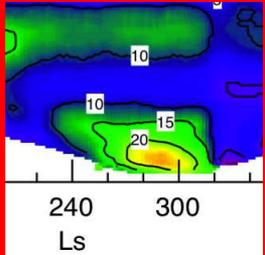
**Dry Case**  
(southern summer  
mid-latitude)



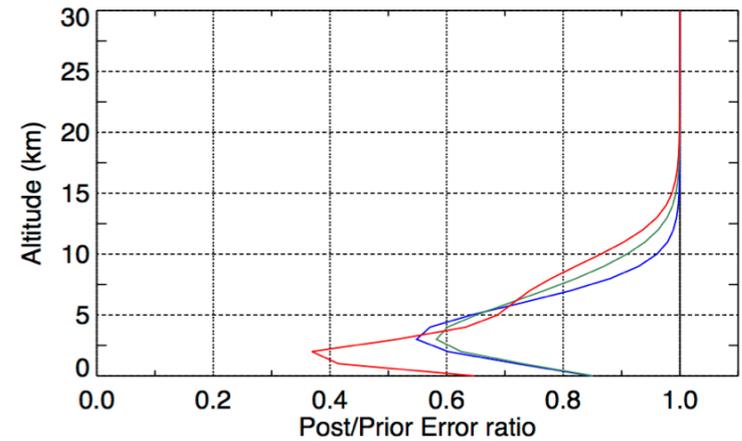
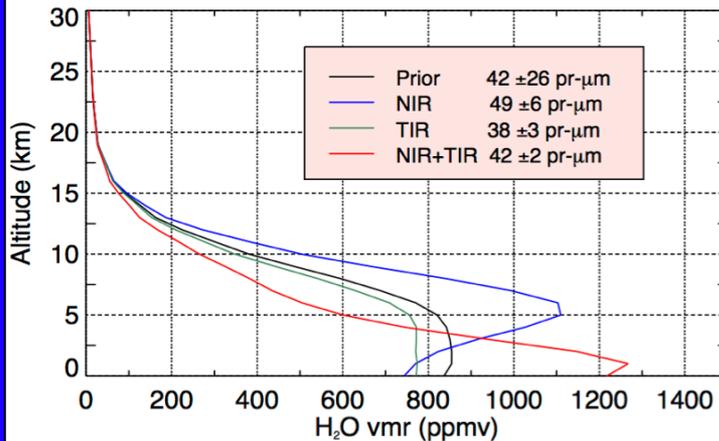
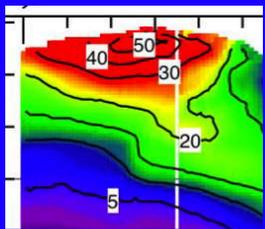
**Wet Case**  
(northern summer  
pole)



**Dry Case**  
(southern summer  
mid-latitude)

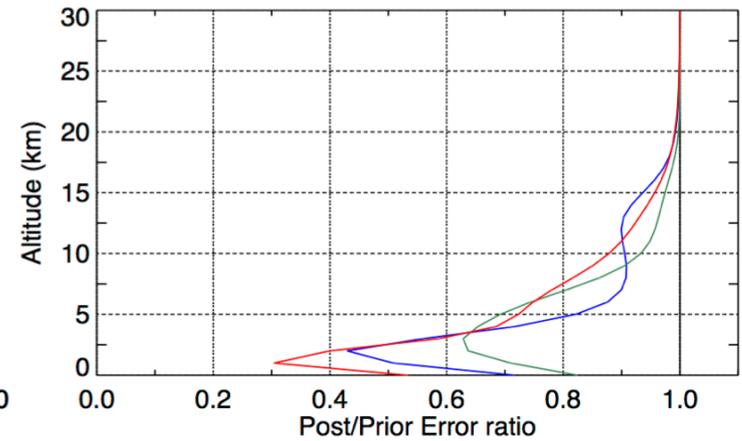
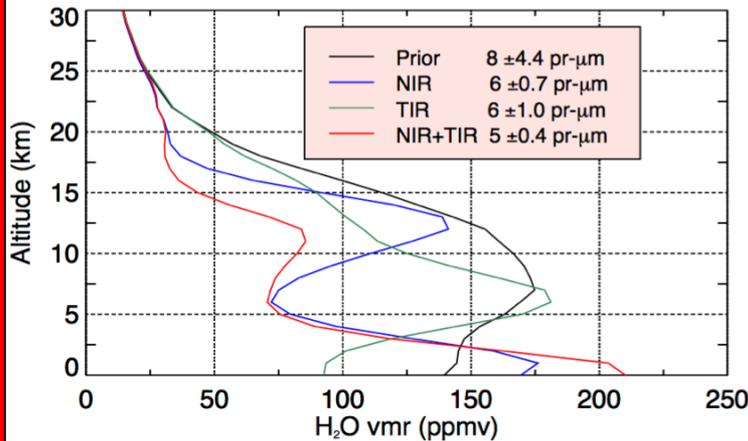
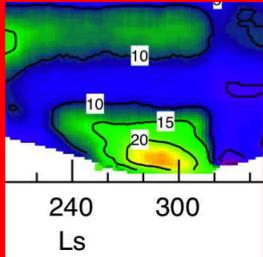


**Wet Case**  
(northern summer  
pole)

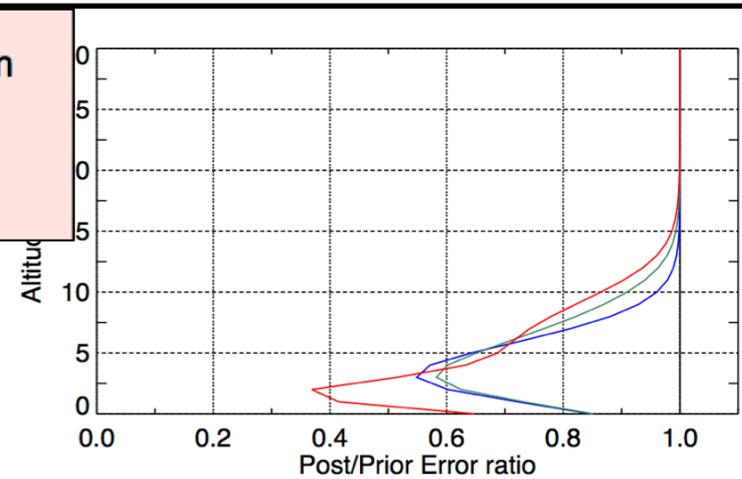
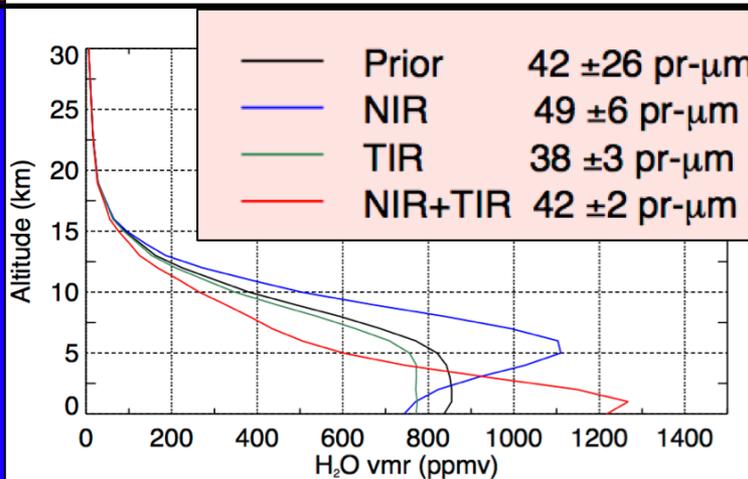
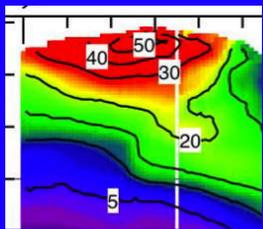


**The NIR+TIR synergistic inversion gives stronger constraints on the shape of the profile and it also refines the estimate of the full column.**

**Dry Case**  
(southern summer  
mid-latitude)

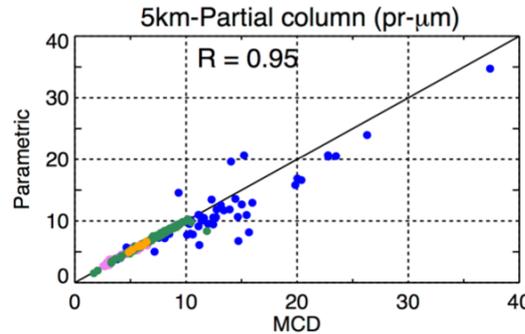
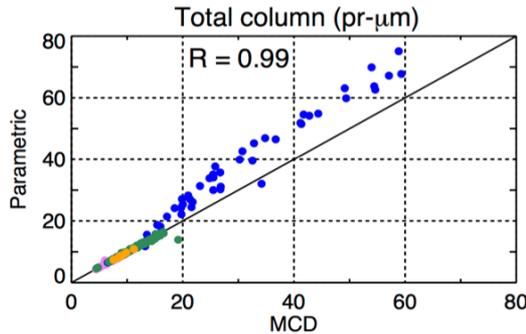


**Wet Case**  
(northern summer  
pole)

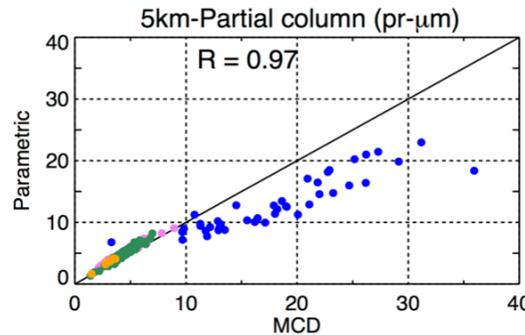
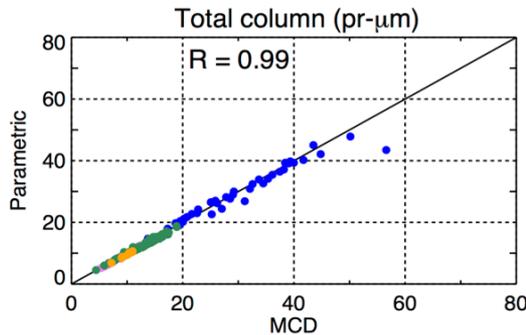


**The NIR+TIR synergistic inversion gives stronger constraints on the shape of the profile and it also refines the estimate of the full column.**

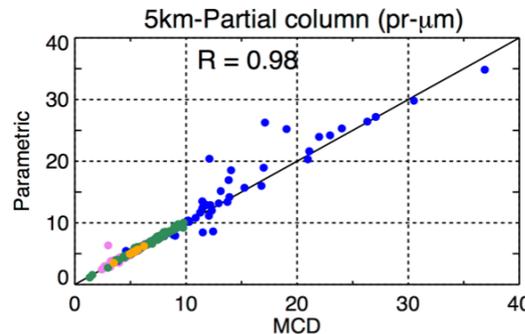
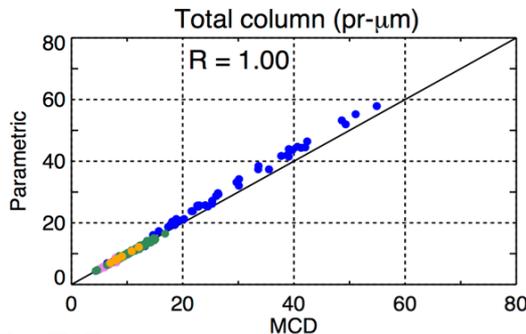
NIR



TIR



NIR+TIR

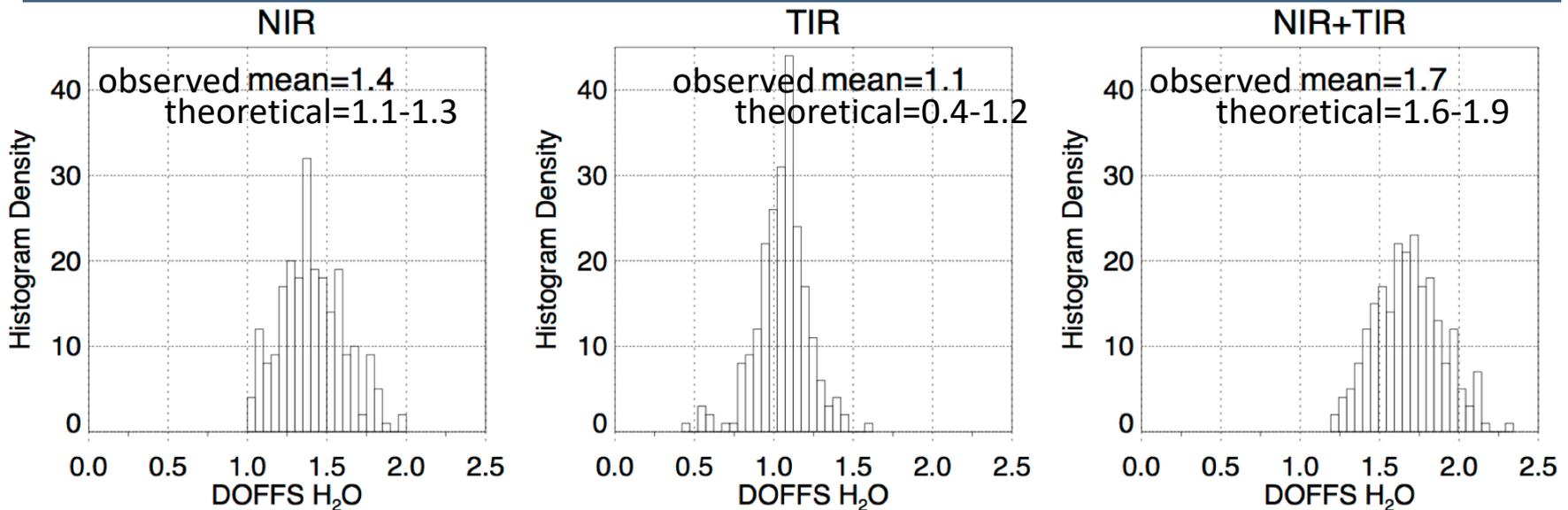


Full columns (*left*)  
and Partial columns (*right*)  
of H<sub>2</sub>O  
derived from observations  
(NIR, TIR, NIR+TIR)  
compared (correlation level)  
to predictions (MCD)

Correlation (R) are given between  
model & observations.

Color code refers to  
the various  $L_s$  shown previously.

- A model for the synergistic retrieval of H<sub>2</sub>O from SPICAM and PFS datasets has been developed and validated based on a terrestrial methodology (Wörden et al., 2010)
- Our study shows that **synergy increases the content of retrieved information on H<sub>2</sub>O by 20 to 80%** compared to NIR or TIR alone.



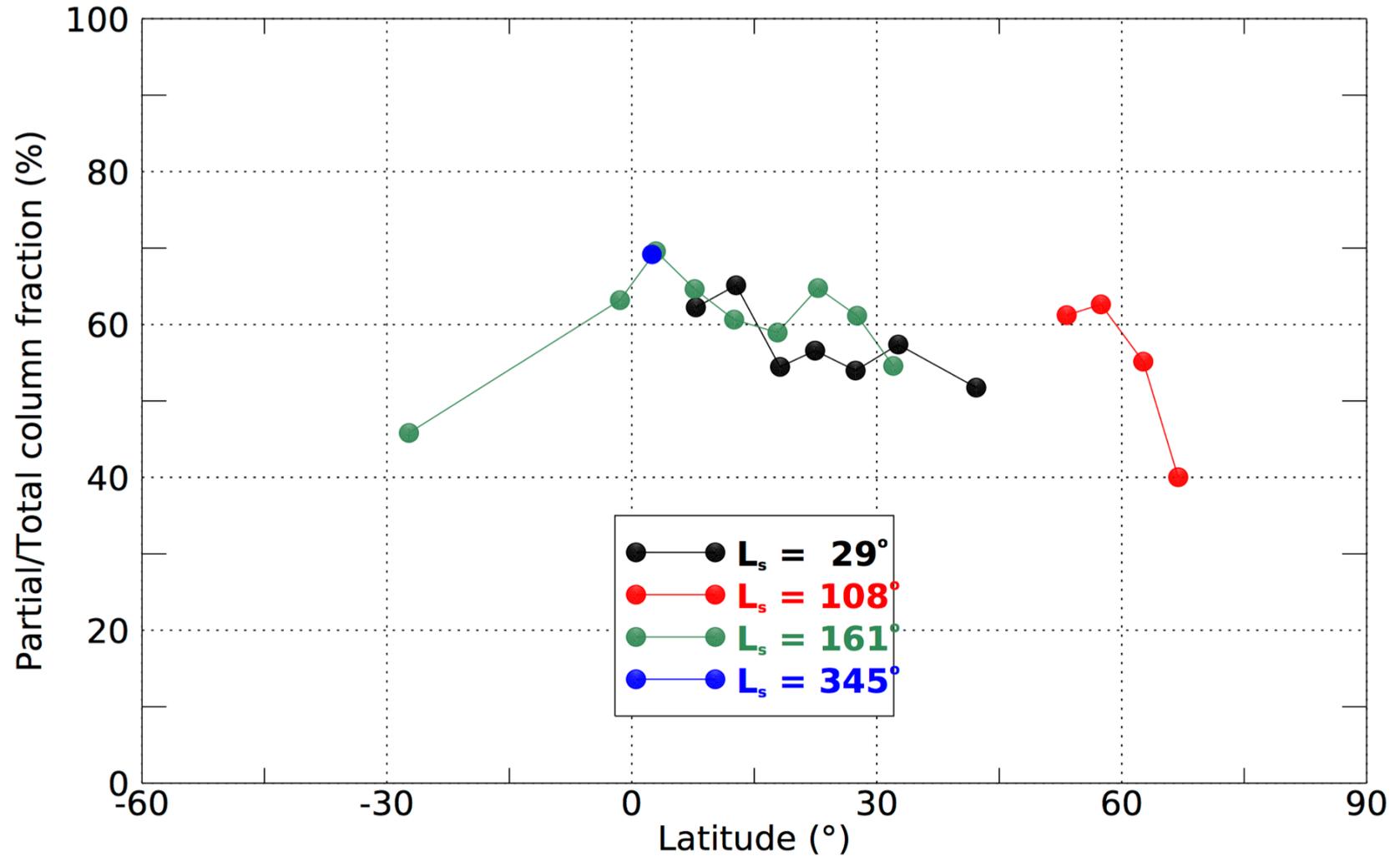
- A model for the synergistic retrieval of H<sub>2</sub>O from SPICAM and PFS datasets has been developed and validated based on a terrestrial methodology (Wörden et al., 2010).
- Our study shows that **synergy increases the content of retrieved information on H<sub>2</sub>O by 20 to 80%** compared to NIR or TIR alone.
- So far NIR or TIR have only provided estimates of the full H<sub>2</sub>O column.
- Now with synergy:
  1. **a new parameter for H<sub>2</sub>O can be extracted** from nadir observations and it is now possible to estimate with no bias the partial column in the **first 5 km** above the surface
  2. **a more (>40%) reliable estimate of the full H<sub>2</sub>O column** can be obtained when combining instrument data together

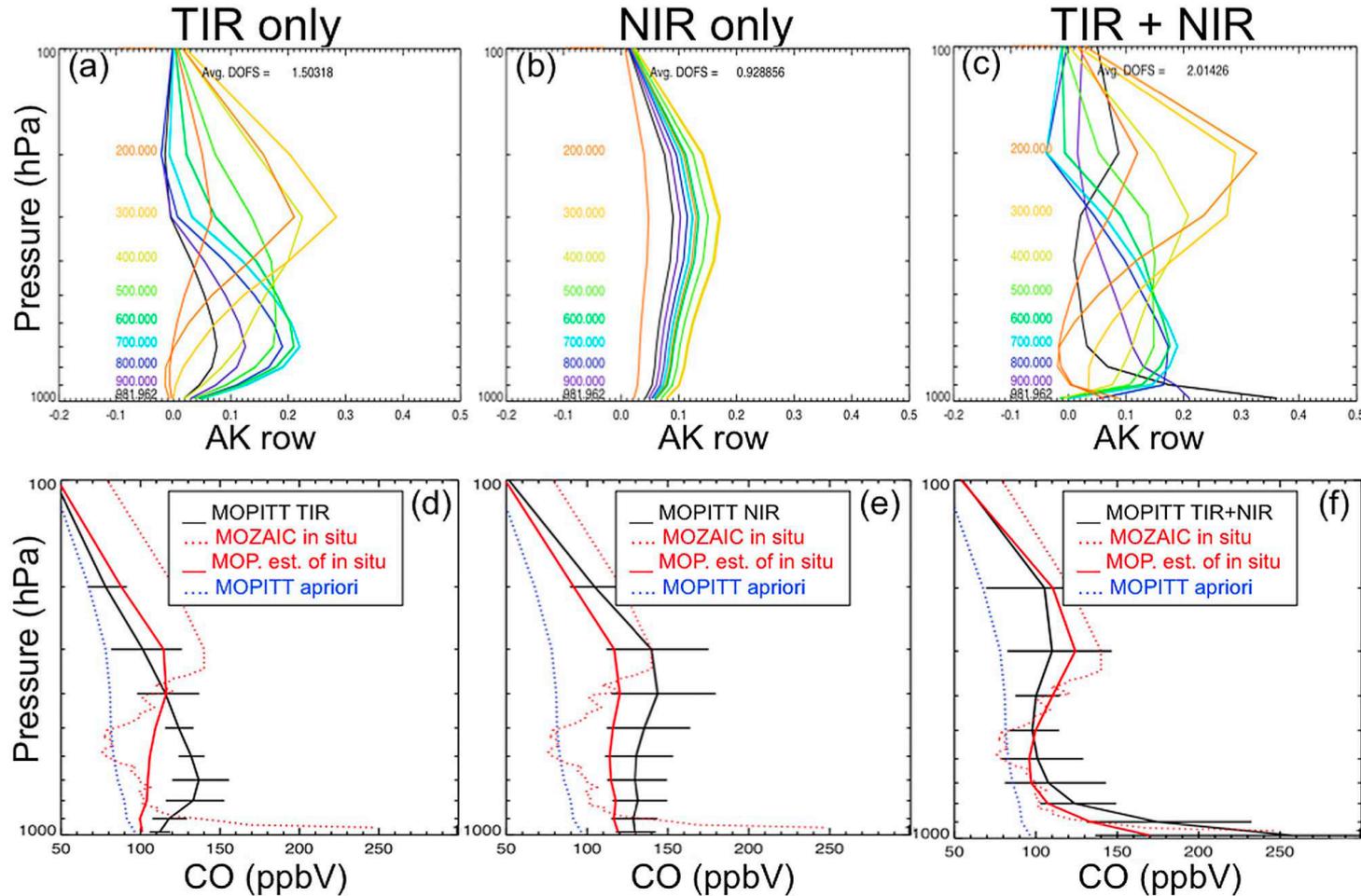
## These results open the door to a range of extensions:

1. Only a small (<10%) subset of MEX data has been used here, 90% remain to be done.....while both PFS and SPICAM still operate
  2. Dust opacity could be similarly tentatively analyzed using SWIR scattering and TIR emission (synergistic retrieval of CO attempted by S. Robert et al., 2017)
  3. Future concepts of mission and/or instrumentation should explore this type of NIR+TIR combination
- Submitted article: Montmessin & Ferron, “*A spectral synergy method to retrieve Martian water vapor column-abundance and vertical distribution applied to Mars Express SPICAM and PFS nadir measurements*”

# Back-Ups

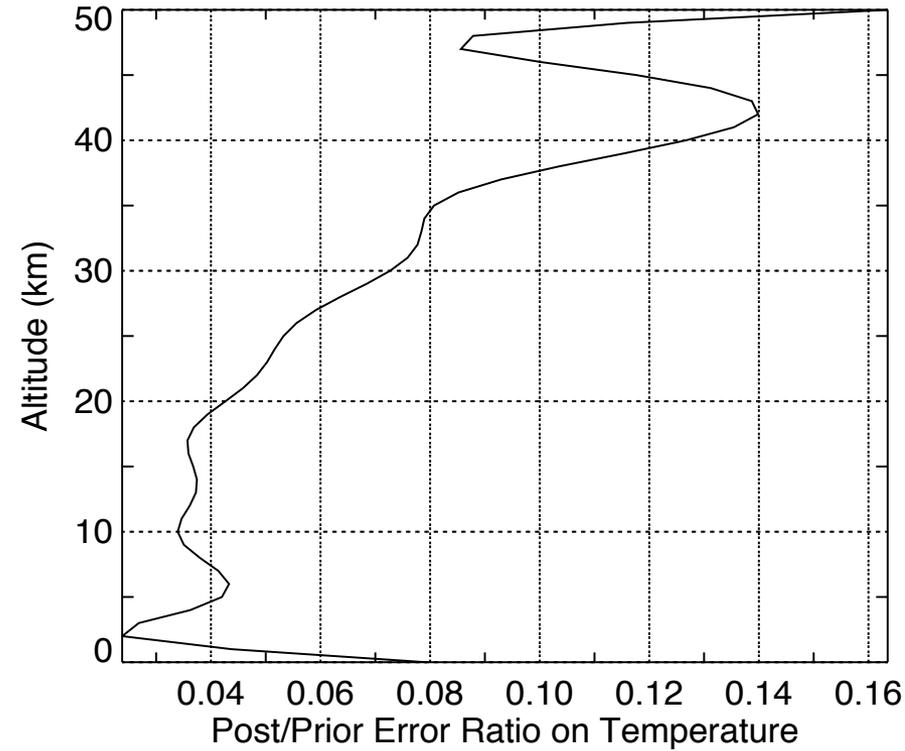
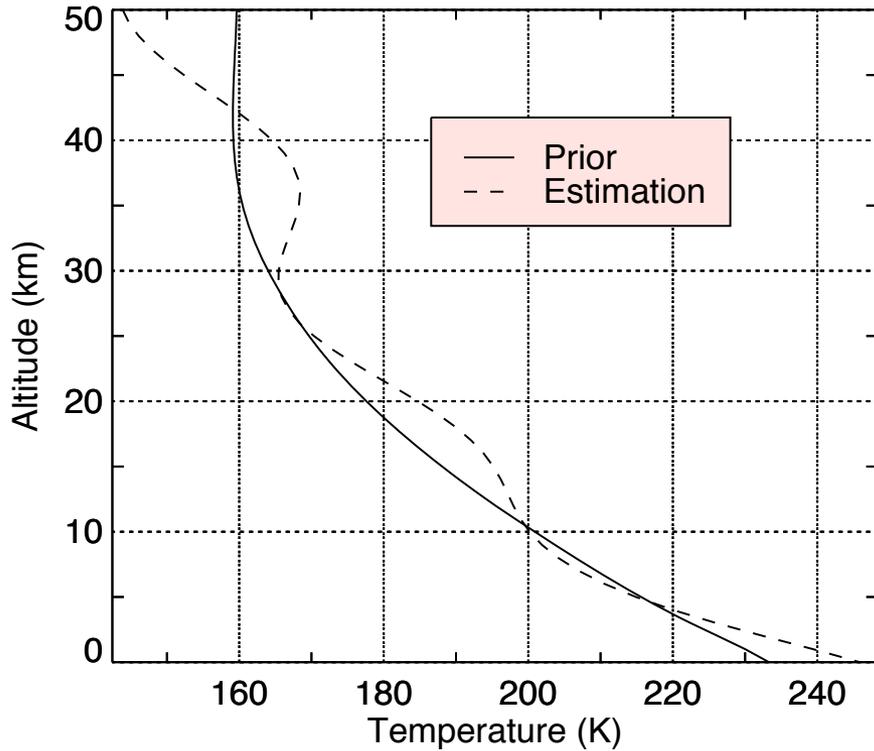
### H<sub>2</sub>O near-surface confinement (<5km)

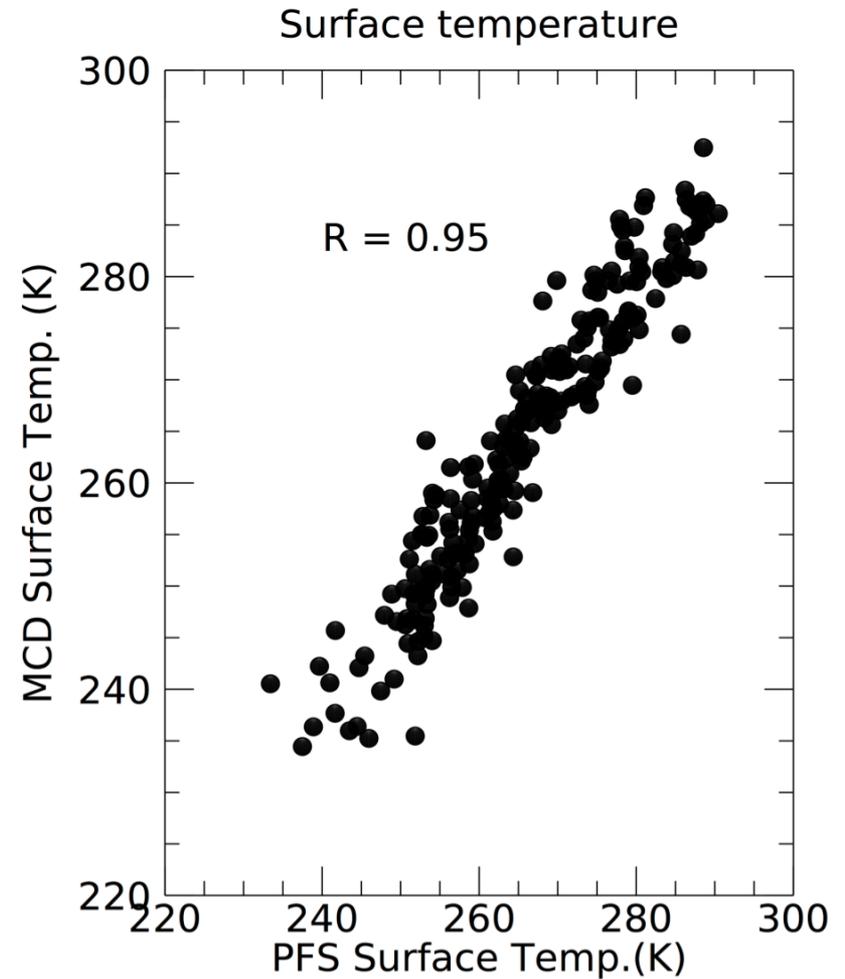
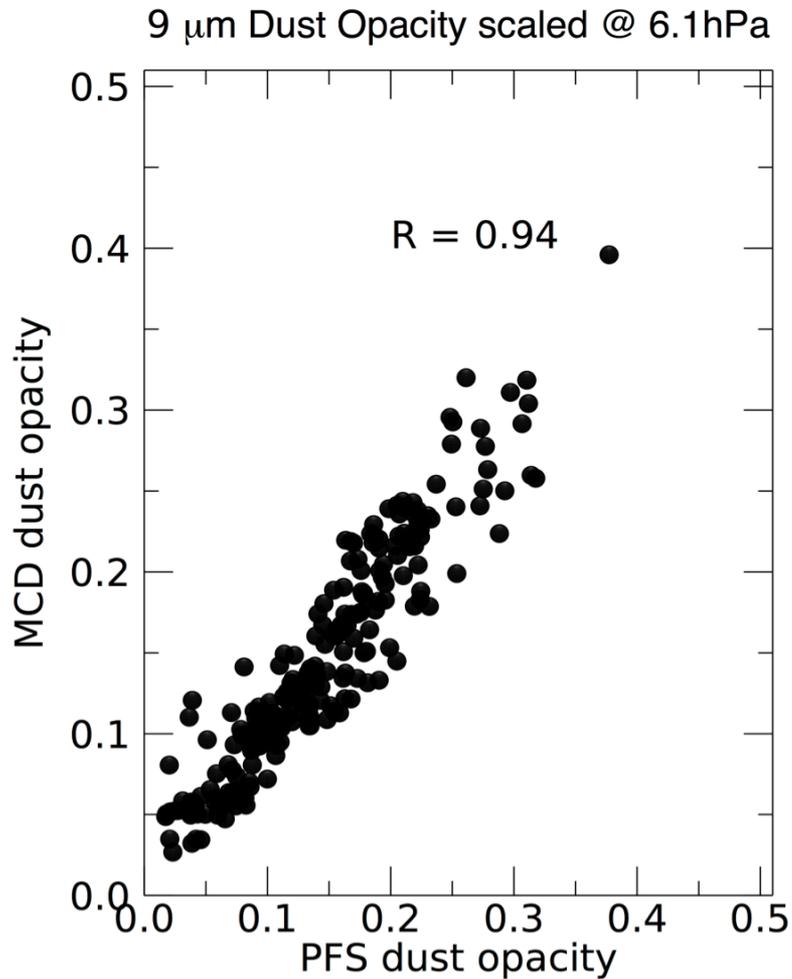




Wörden et al. (2010)







As a test of consistency, the dust opacity and the surface temperature derived from the PFS data (as by-products of the synergy) are compared to MCD predictions.

The high degree of correlation between the two is a good indicator that the retrieval method works well.