



# Extraction and validation of water ice indicators from Mars Express / OMEGA spectral imagery – determination of the diurnal CLOUD LIFE cycle

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# 1. Introduction

- **Objective of this study**

- Construct a 4D gridded (spatio-temporal) water ice cloud database extracted from Mars Express/OMEGA spectro-imager data
- Use the derived products to determine the diurnal cloud life cycle.

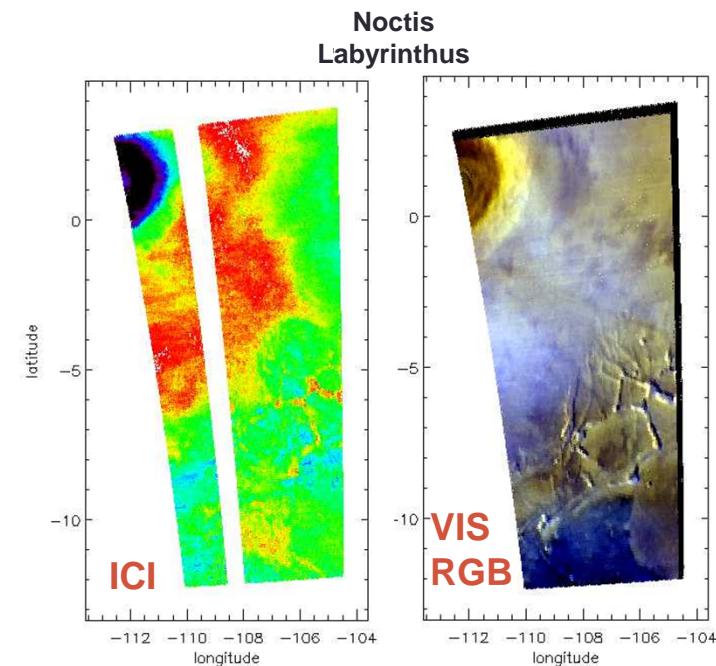
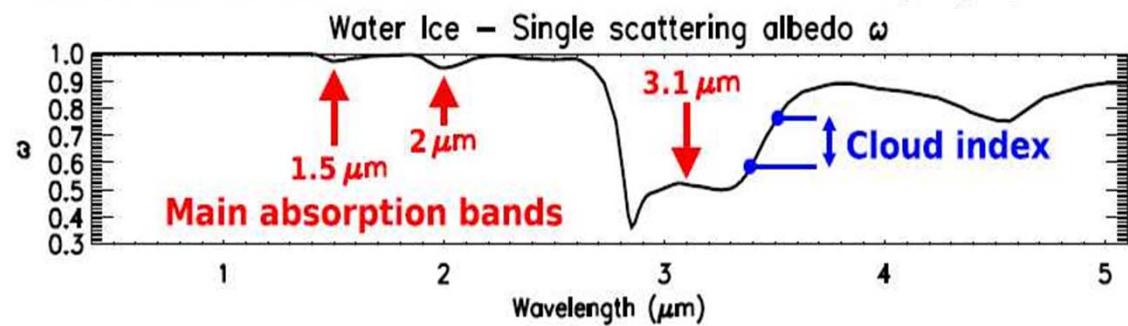
- **OMEGA instrument :**

- VNIR + SWIR (0.36 – 5.2  $\mu\text{m}$  ; 352 spectels)
- Long period of observation : MY 26-32 (1/2004 - 4/2015)
- Non-heliosynchronous orbit ==> Better temporal coverage.

## 2. Ice Cloud Index and Percentage of Cloudy Pixels

### Calculation of the Ice Cloud Index (ICI)

- Pixel-based
- Derived from Langevin et al., 2007



- Slope → original ICI :  $\text{IClo} = I_{3.38\mu\text{m}} / I_{3.52\mu\text{m}}$
- Normalized IceCloudIndex :  $\text{ICI} = 1 - \text{IClo}$

# Construction of a 4D cloud database

N summer solstice  
(Ls=45-135° )

## 1) Definition of a regular grid

1° longitude X 1° latitude X 5° Ls X 1 h  
LT

## 2) Binning of individual (pixel-based) ICIs onto grid

## 3) Average of ICI on each gridpoint

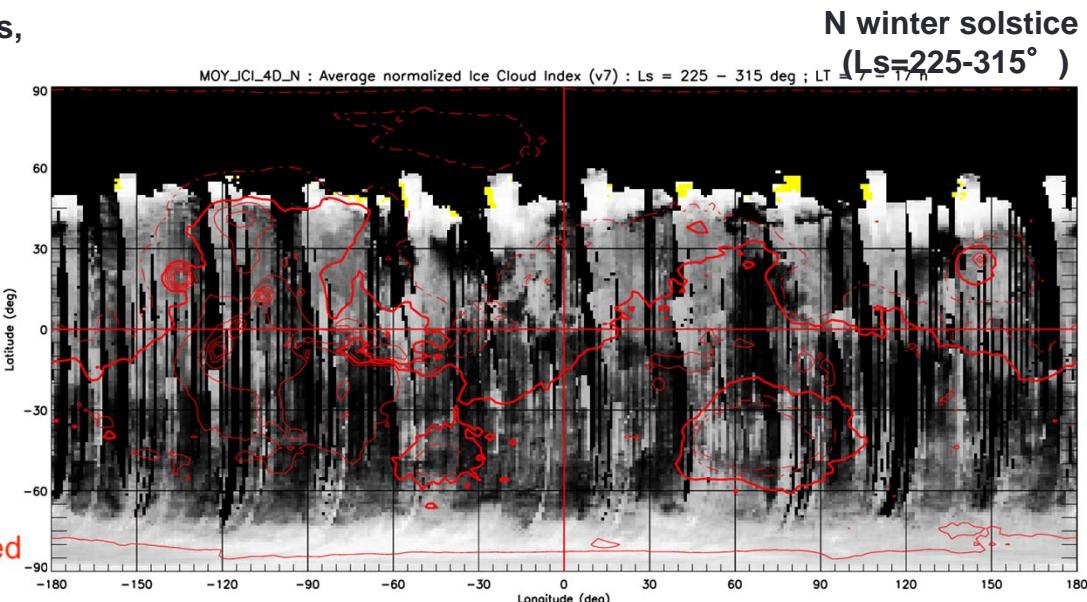
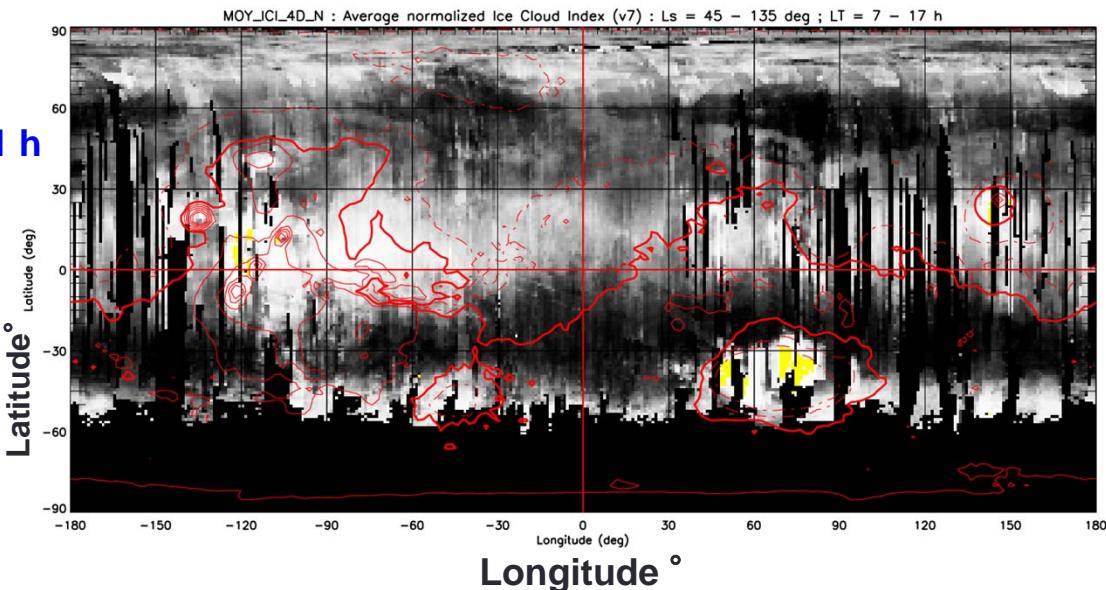
## 4) Illustration : 2D ICI maps at solstices

-Main cloud features identified : aphelion belt, Hellas, cloud edges of polar hoods.

-Average ICI higher in Northern hemisphere than in Southern hemisphere (in tropics and non-polar hood and non-Hellas midlatitudes).

**But :**

Only ~2% of daytime gridpoints have ICI values



# Percentage of cloudy pixels (PCP)

## 1) Selection of a threshold to extract cloudy pixels :

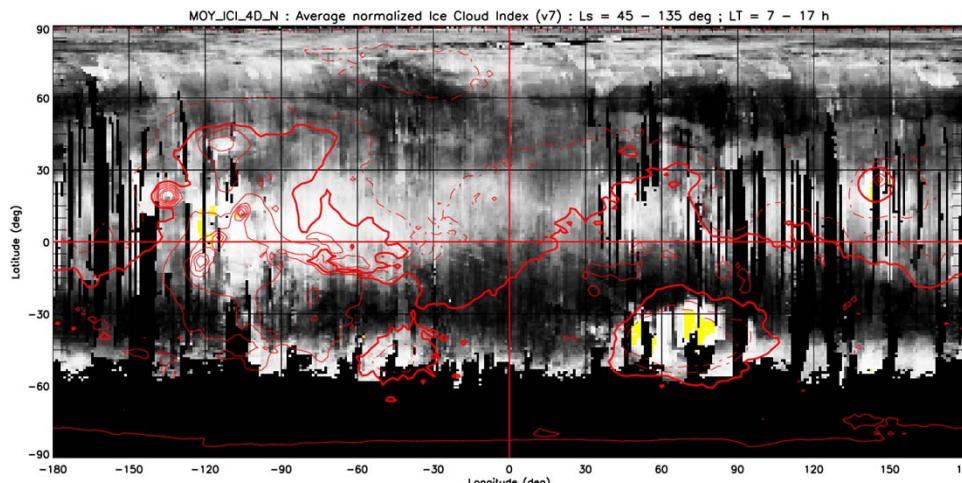
- $ICI > ICI_{thr}$  → the pixel is cloudy
- $ICI_{thr} = 0.28$  (based on Madeleine et al., 2012;  
Audouard et al., 2014)

## 1) Percentage of cloudy pixels (PCP) ⇔ cloud coverage

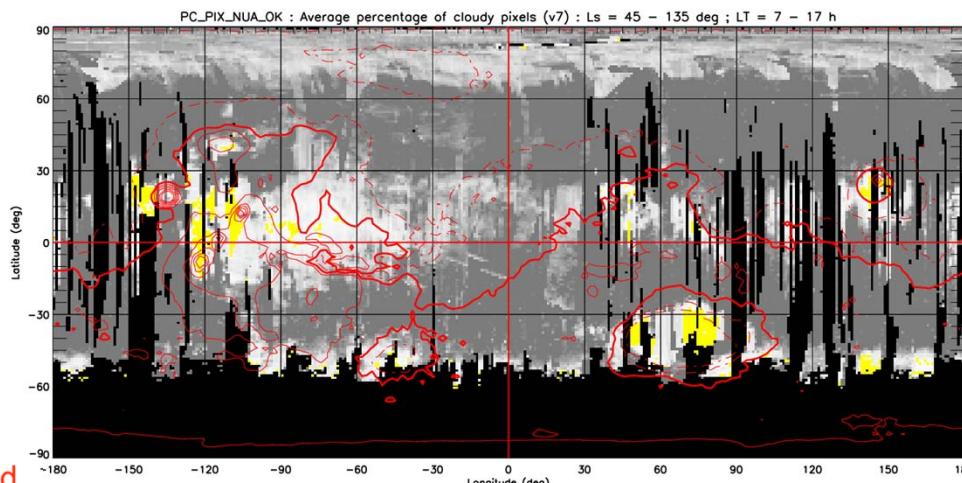
$$PCP = 100 \times N_{\text{cloudy\_pixels}} / N_{\text{all\_pixels}} (\%)$$

## 1) PCP filters out areas with limited average cloud coverage and thin clouds (thin cirrus = lower ICI)

Main elevation contours : red  
 Highest value : yellow

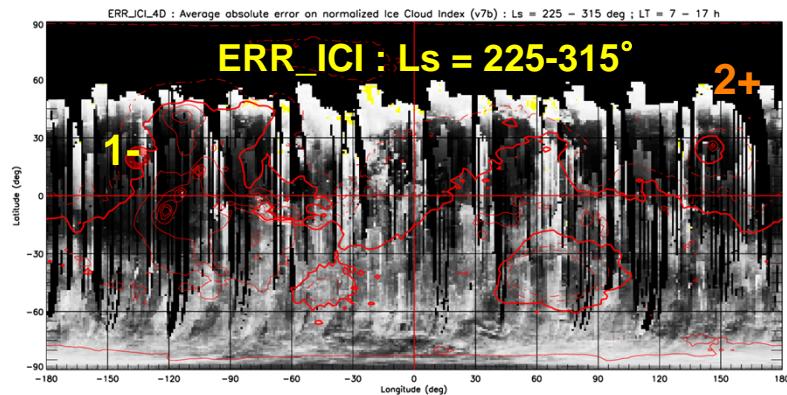
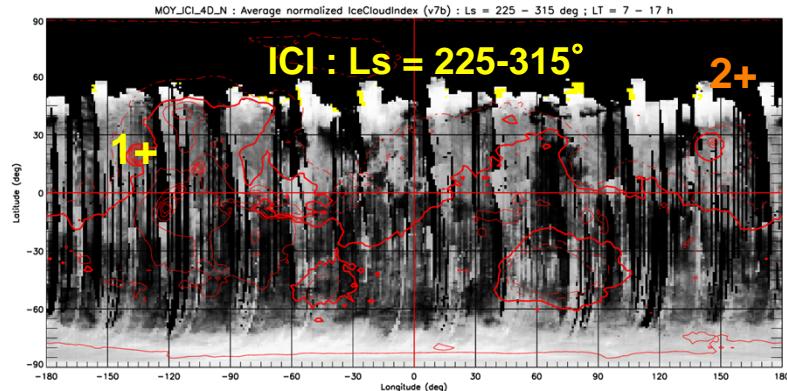
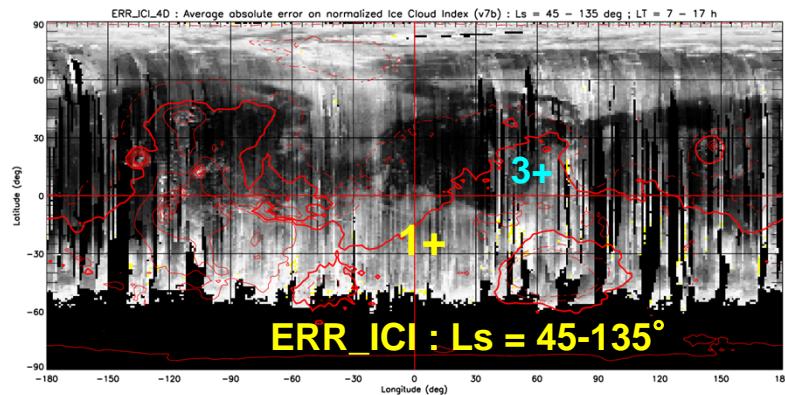
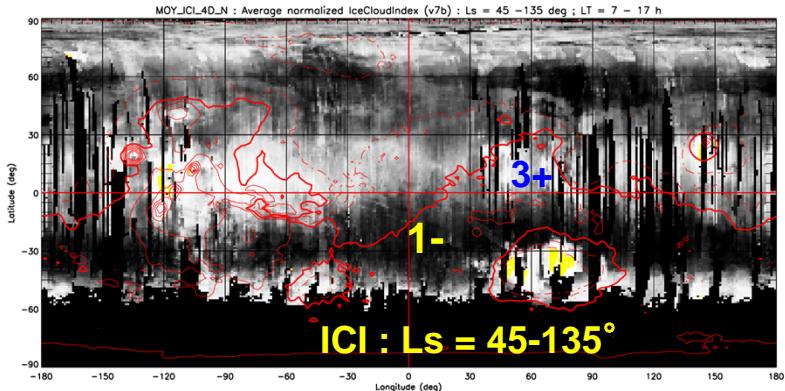


Average Ice Cloud Index ( $Ls = 45 - 135^\circ$ ,  $LT = 7 - 17$  h)



Average Percentage of Cloudy Pixels

# Error Estimation



- Main observation-based rules on ICI error bar (ERR\_ICI)

- 1) Higher (resp. lower) ICI  $\iff$  lower (resp. higher) ERR\_ICI
- 2) Exception 1 : high ERR\_ICI for high ICI in high-latitude areas at edge of the polar hood.
  - Cause : reduced solar illumination  $\implies$  reduced reflectance  $\implies$  higher relative instrumental error
- 3) Exception 2 : specific areas of relatively high ICI  $\implies$  Relatively high ERR\_ICI
  - Dark areas  $\iff$  Reduced reflectance  $\implies$  higher relative instrumental error. (Syrtis Major)
- 4) Sunlit polar regions : intermediate and variable ERR\_ICI (around summer solstice)

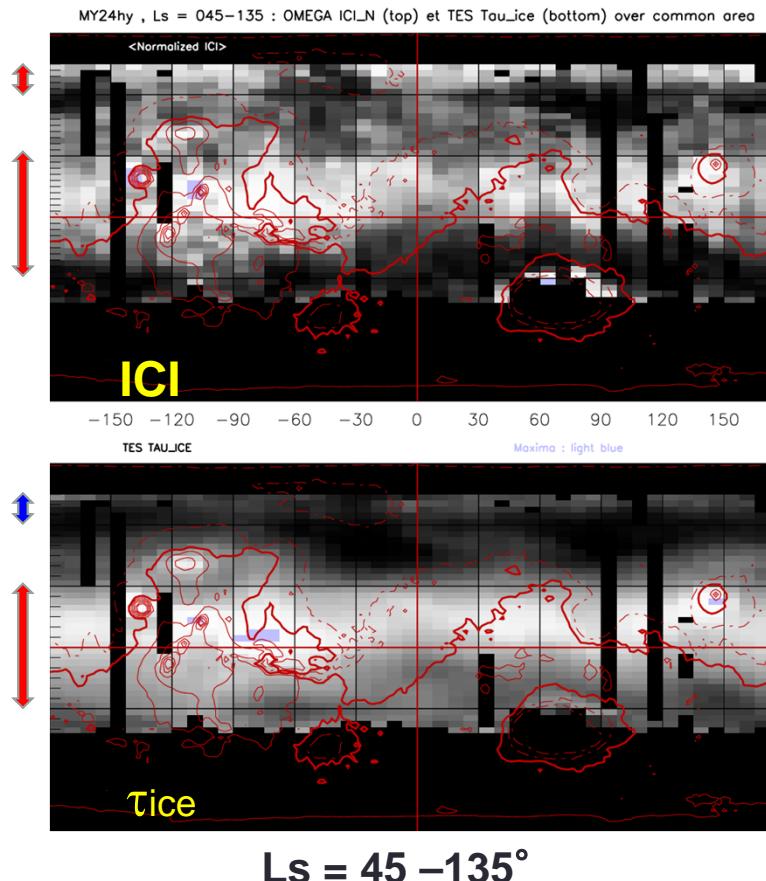
### 3. Comparison with other datasets

- **ICI vs. TES water ice optical thickness ( $\tau_{\text{ice}}$ )**

- ICI gridded onto same resolution as TES

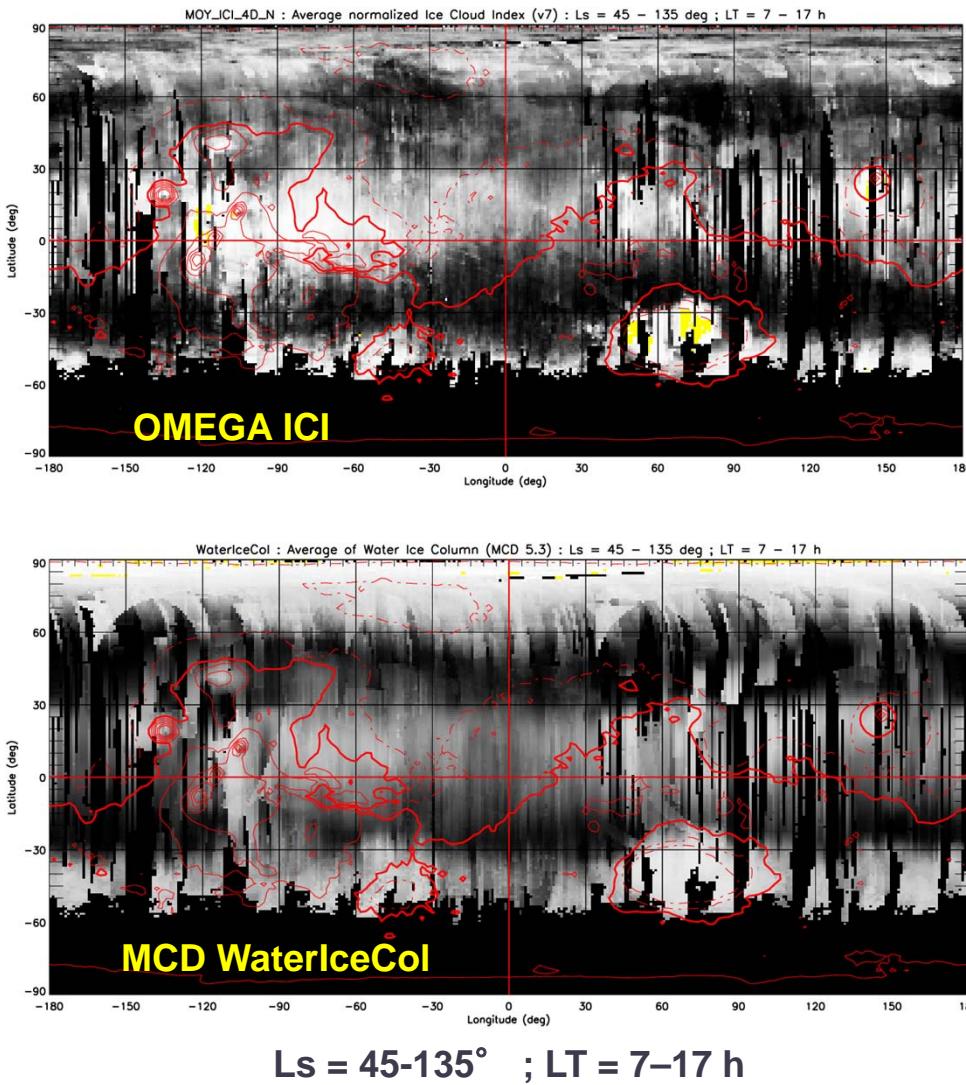
climatological  $\tau_{\text{ice}}$  (Smith, 2004) :

- $7.5^\circ \text{ lon} \times 3^\circ \text{ lat} \times 5^\circ \text{ Ls} \times \text{LT}=14 \text{ h}$
- $\tau_{\text{ice}}$  : Optical thickness at 12 mm.
- TES  $\tau_{\text{ice}}$  of MY 24 hybrid year :
- Similar cloudy features and cloudless areas
- Differences :
  - Reduced  $\tau_{\text{ice}}$  values in comparison with ICI between high latitudes and tropics
  - Possible explanations : biased estimation of  $\tau_{\text{ice}}$  derived from very low temperatures, poor detection of low-level clouds for TES.



# Comparison : ICI vs. MCD Water Ice Column

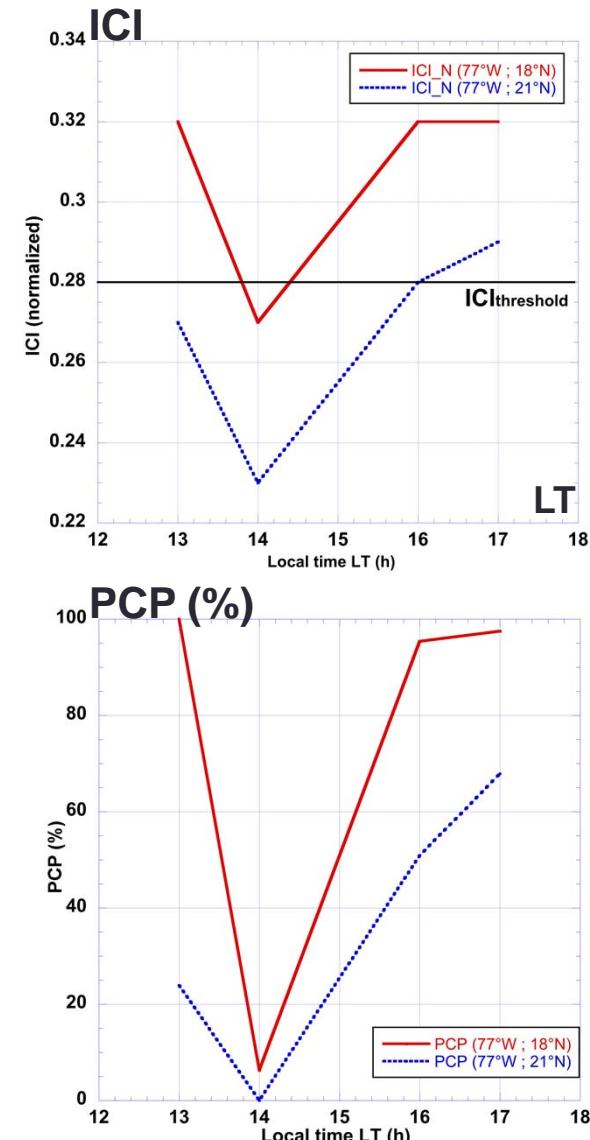
- **ICI 4D dataset**
- **Mars Climate Database (MCD 5.3)**
  - Derived from Martian General Circulation Model from the LMD.
  - MCD value to compare : water ice column (`WaterIceCol`), extracted on the same 4D grid as ICI
- **Comparison :**
  - Overall similar cloudy and cloudless areas on both 2D images
  - Smoother aspect of `WaterIceCol` image



## 4. Diurnal cloud life cycle

- **Principle**

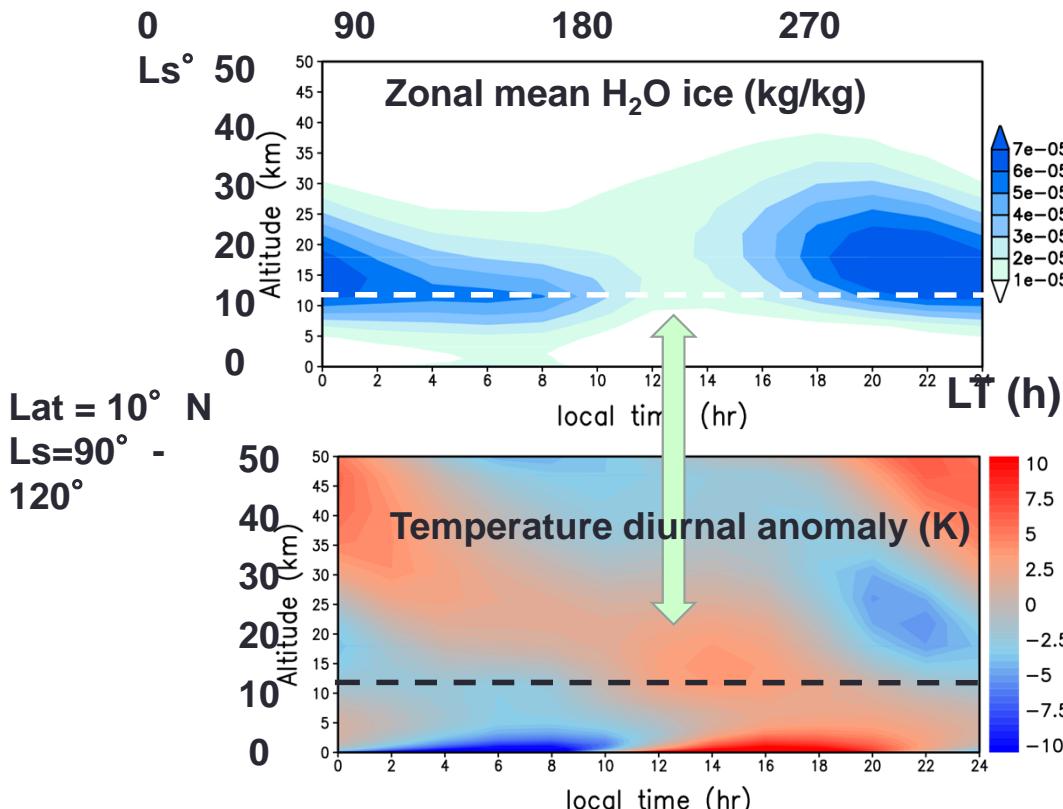
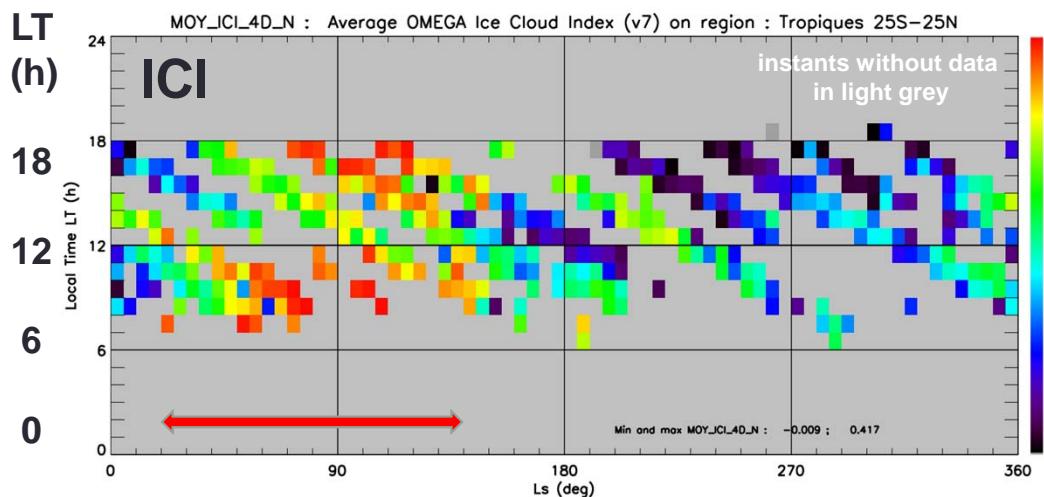
- 1 individual gridpoint (lon, lat, Ls) ==> maximum 4 h  
LT available in the tropics
- Average of ICI, PCP, WaterIceCol values over larger spatial areas (lon, lat) ==> longer daily sequences.
- **21 areas defined :**
  - Small areas : single volcano + surroundings (Arsia, Olympus, Elysium)
  - Intermediate size areas : Hellas, Argyre, Chryse Planitia...
  - Large climatic areas : tropical (lat < 25° ) : midlatitudes (25° N-55° N),... (over all or a large band of longitudes).



Diurnal cycle : 77° W ; 18 and 21° N ; Ls=80

# Tropical region ( $25^{\circ}$ S – $25^{\circ}$ N ; all longitudes)

- **N summer solstice :**
  - Most cloudy period
  - Dominant cloud structure : **aphelion belt**.
  - Diurnal cycle : important cloudiness in the morning and later in the afternoon, reduced around noon.
- **GCM-based interpretation :**
  - Clouds tend to form above hygropause (10-20 km) at minimal temperatures.
  - Temperatures are controlled by thermal tides.
  - Max. temperature at cloud altitude around midday => cloud minimum around noon (not due to radiative heating of surface by the sun).





# 5. Conclusion and prospects

- **ICI and PCP : complementary products**

- Ice Cloud Index : general indicator of presence of water ice clouds.
- Percentage of Cloudy Pixels = cloud coverage : better adapted to discriminate thick clouds and partial cloud coverage.
- ICI and PCP are robust products for water ice clouds global and regional studies.
- ICI (+error) and PCP in PDS-4 format soon available in the frame of UPWARDS project WP 4.
- ICI : basic element for derivation of cloud optical thickness and effective radius of ice particles (presentation of Kevin Olsen et al.)

- **Diurnal cloud life cycle**

- 4D products : unprecedented dataset of water ice clouds for Martian climatology.
- Regions: trade-off between spatial and temporal coverage.

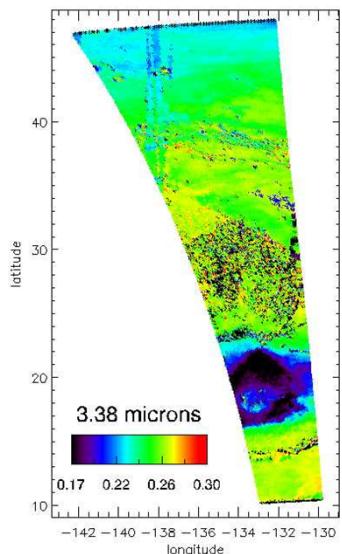
- **In the future :**

- Use of OMEGA ICI and PCP products for the validation of high-resolution Martian GCMs (Pottier et al., 2017) and comparison with other datasets.

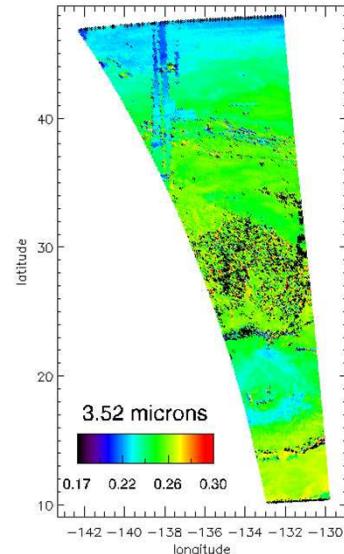
- **Acknowledgement :**

- This study was partially funded in the frame of the UPWARDS project.

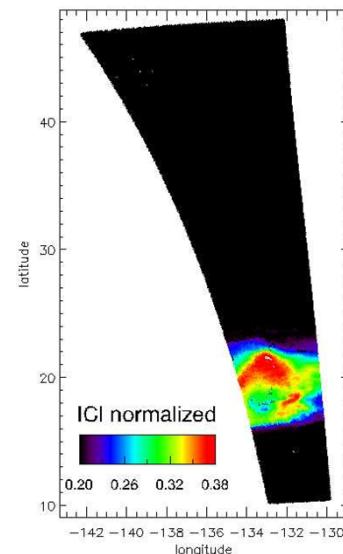
# Thank you



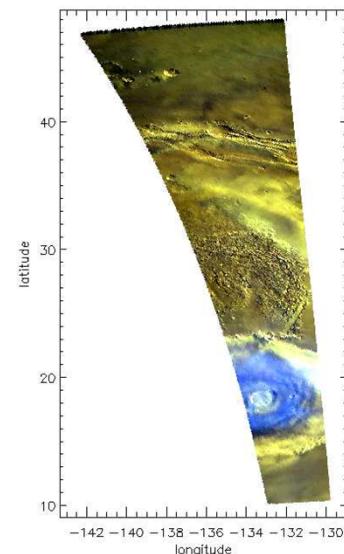
Albedo (3.38  $\mu\text{m}$ )



Albedo (3.52  $\mu\text{m}$ )



ICI



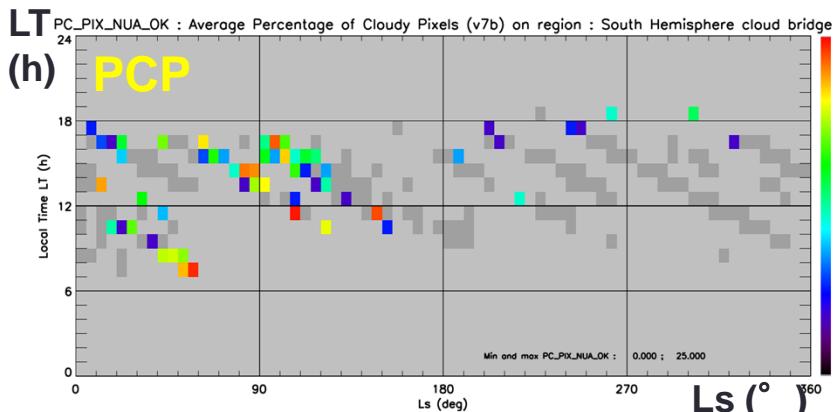
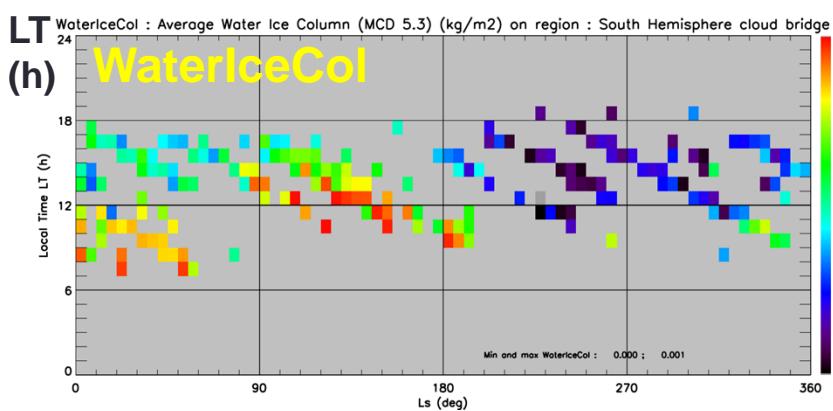
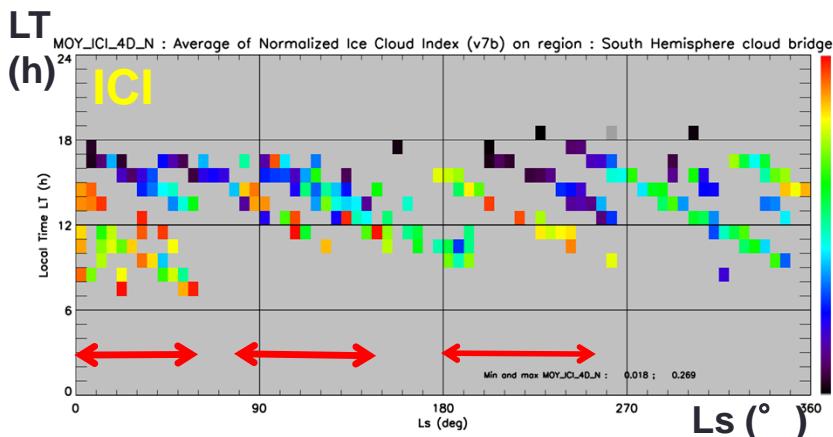
VISIBLE



# Example : South Hemisphere cloud bridge

- SH cloud bridge :  $35^{\circ}$  S –  $20^{\circ}$  S ;  $150^{\circ}$  W –  $60^{\circ}$  W :

- Reduced cloudiness :  $ICl_{max} = 0.27$  (vs. 0.41 in tropical belt)
- Partial cloud coverage frequent ( $PCP_{max} = 25\%$ )
- Cloud presence during spring ( $Ls=0-60^{\circ}$ ), summer ( $80-150^{\circ}$ ) and autumn ( $180-250^{\circ}$ ).
- Diurnal cycle :
  - Spring : presence in the morning, dissipation around noon (LT).
  - Spring and summer : clouds dissipate earlier and earlier during the day.
- Reflects the limited water transport between the aphelion belt and the south polar hood, decreasing in summer and autumn, and during the day.



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