

# **Limb data from OMEGA/MEX**

Characterization and dust retrievals

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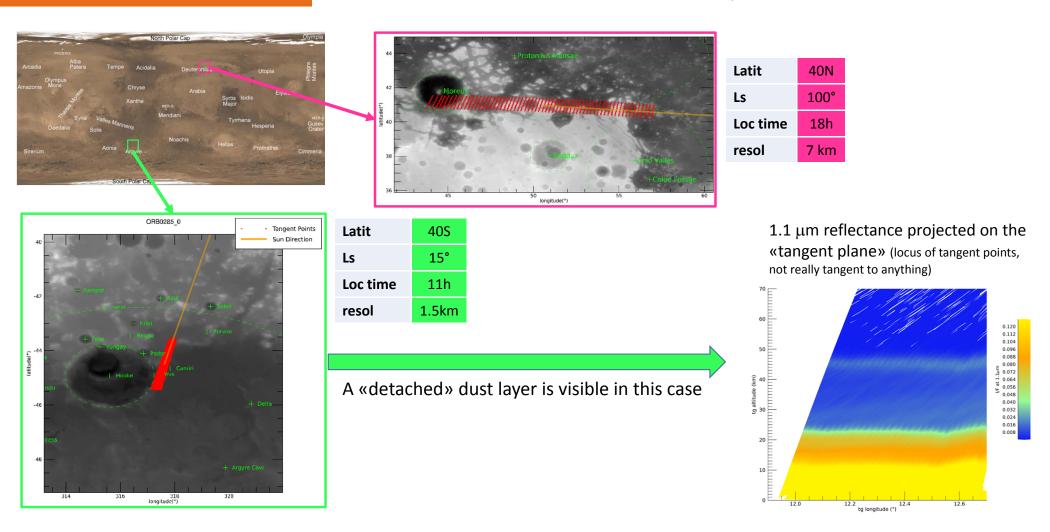
*IAPS-INAF, Rome, Italy* 

## Outline

- Assessing information content in OMEGA limb data
- Addressing issues for quantitative retrievals
- Testing retrieval of dust vertical profiles
- Planning application to the whole OMEGA limb dataset

## Two working cases

### Two OMEGA cubes have been selected in very different conditions



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### Instrumental issues

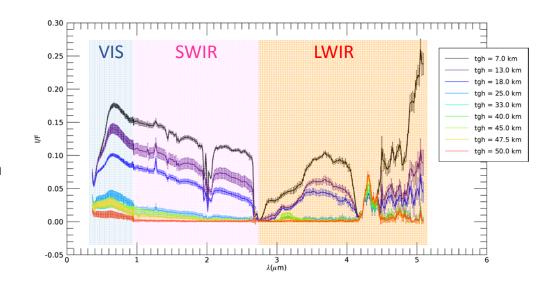
Spectra from the 3 OMEGA channels can be averaged in bins of tangent altitude to achieve homogenous spectral scans in the range  $0.4-5~\mu m$ .

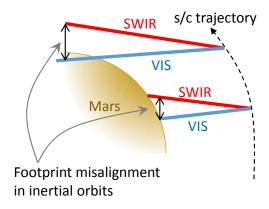
The pixel footprints differ among channels, due to:

- Pointing misalignment
- IFOV slight different
- Push-broom vs. whiskbroom acquisition scheme

In the **nadir** case the channels are usually co-registered with a **rigid shift** by matching *surface features*.

In the **limb** case the **surface curvature** makes the footprint misalignment larger and not uniform in the same orbit.





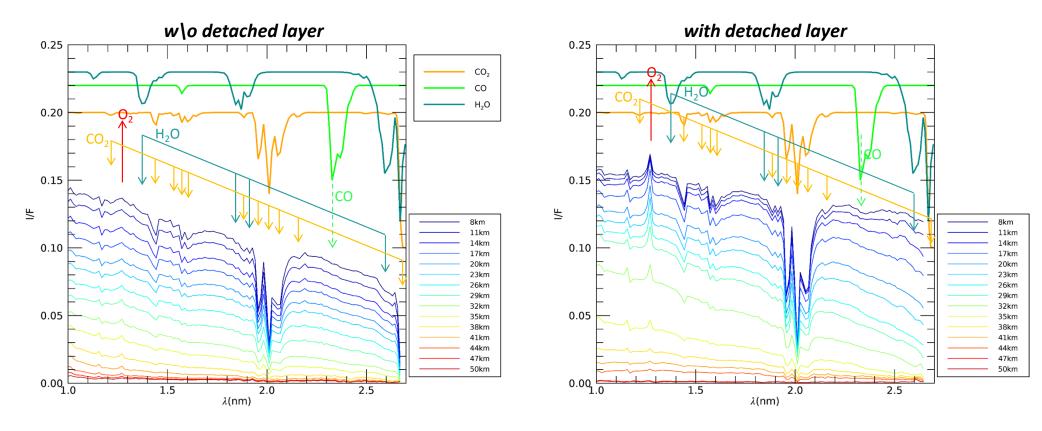
Moreover spectral junctions are often affected by enhanced <u>straylight</u>. Deep space spectra can be partially useful for correcting this effect, but straylight is an additive signal depending on the distance from the target.

## Spectral features in SWIR channel

Limb spectral scan after vertical binning

#### Main features are

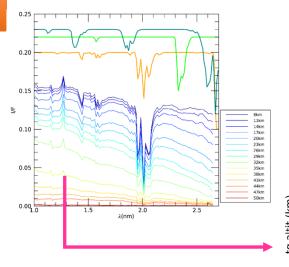
- O<sub>2</sub> emission at 1.27 μm
- $H_2O$  gas absorption at 2.6  $\mu m$
- CO 2.35 μm band undetected

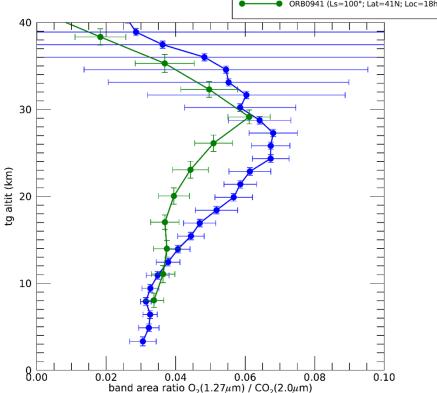


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1.27 $\mu$ m emission can be mapped to produce  $O_2$  vertical profiles





 $O_2$  band area at 1.27 µm is shown here ratioed by the  $CO_2$  band area at 2 µm, to grossly normalize the path length.

The peak altitude is related to chemical equilibrium among H<sub>2</sub>O dissociation products and is known to be variable with season, latitude and local time.

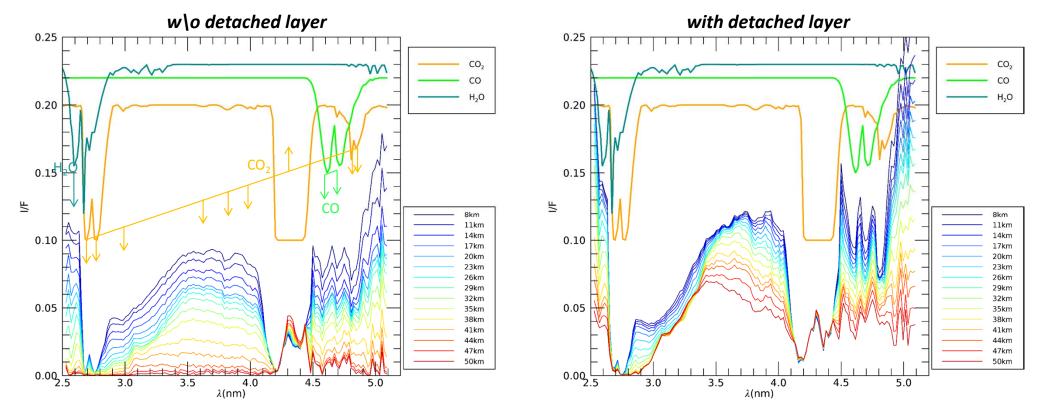
O<sub>2</sub> abundances can be retrieved with <u>non-LTE modeling</u> and compared with H<sub>2</sub>O profiles retrievable from the same observations (not shown here, see e.g. Mahieux et al. in this workshop)

## Spectral features in LWIR channel

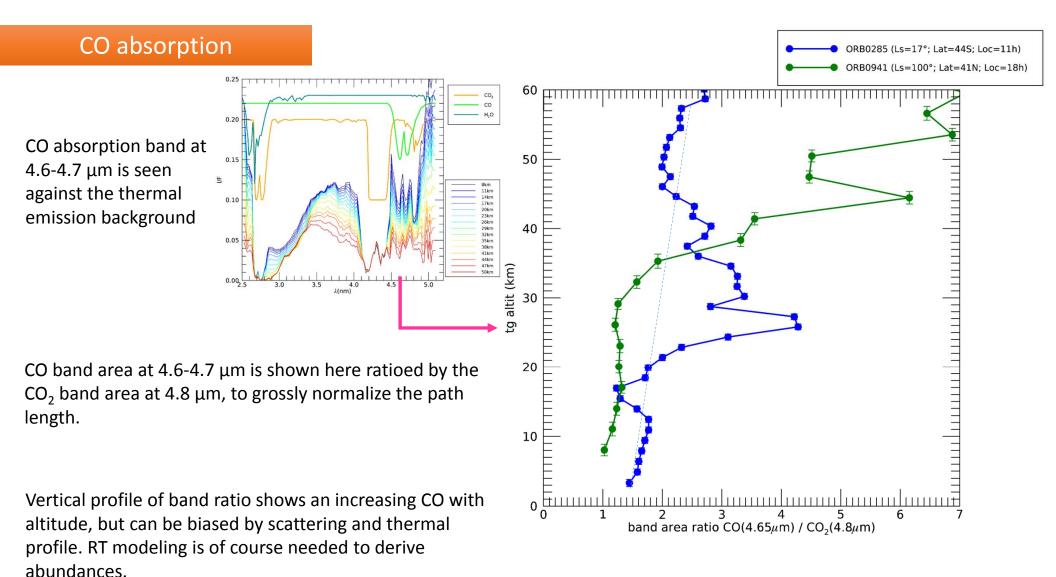
Limb spectral scan after vertical binning

#### Main features are

- $CO_2$  emission at 4.3 µm
- CO absorption at 4.6-4.7 μm
- H<sub>2</sub>O gas absorption at 2.6 μm



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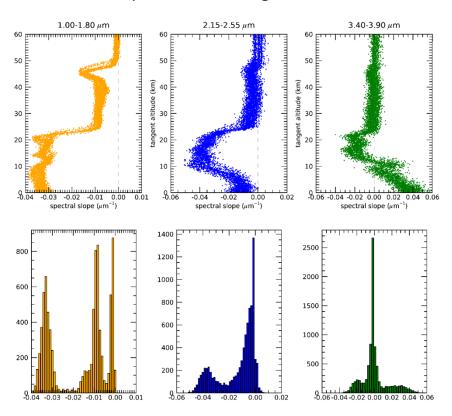


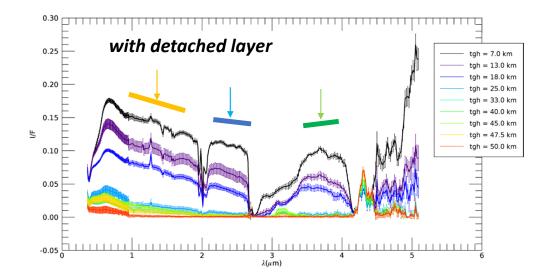
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### Dust

The overall spectral shape at limb is dominated by scattering by dust (and molecules in the blue range).

Spectral slopes in between the main of gaseous absorptions can be related the composition and/or grains size of the scatterers.





Slopes identify changes in scattering regimes with altitude, but quantitative retrievals require full modelling of radiative transfer with multiple scattering in spherical geometry...

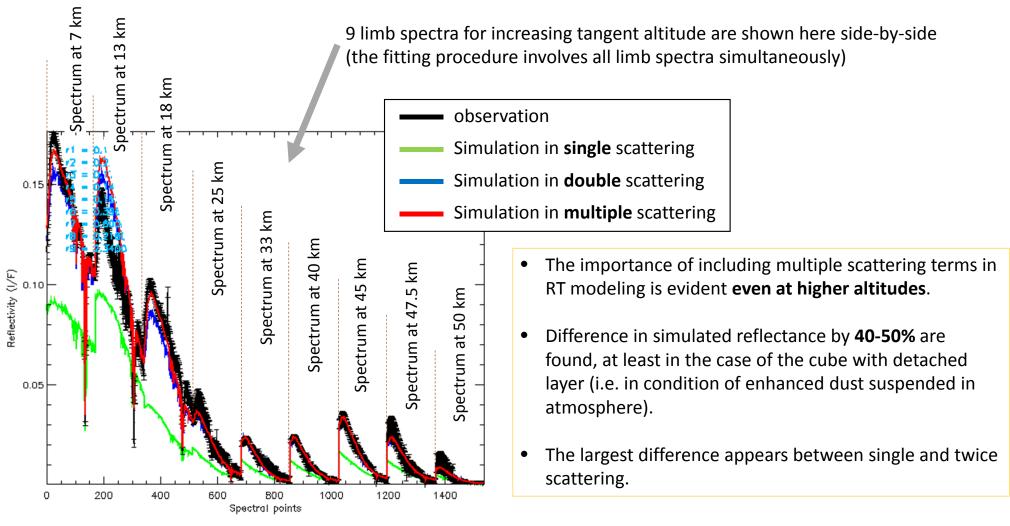
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## Radiative Transfer at limb – modeling scheme

Pros and cons of some Radiative Transfer schemes Plane-parallel **Onion-peeling Pseudo-spherical Monte Carlo** to the observe. multiple scattering couplings Multiple scattering Multiple scattering Fast convergence Multiple scattering Many solvers available Single scattering PPA approach Accurate physics Small incidence and Complex angles management Time consuming emission angles MITRA tool Backward Monte Carlo as forward RT model

Bayesian inversion for retrieval of dust properties (and gases)

### Radiative Transfer at limb – multiple scattering

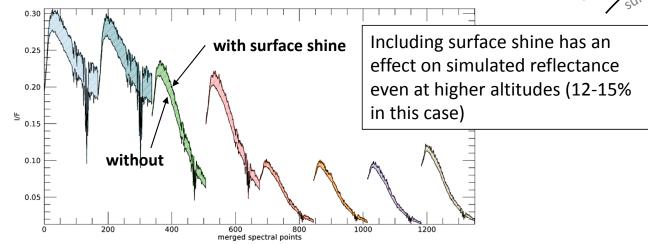


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### Radiative Transfer at limb – surface shine

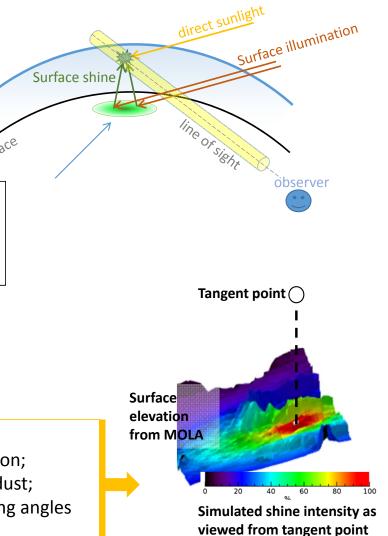
The whole limb atmospheric column is illuminated from the surface shine. *Full treatment* would require a 3D RT model.

<u>Simplified treatment</u>: average amount of light scattered from the surface toward a given tangent altitude of the limb observation.



We obtained spectra for surface shine by:

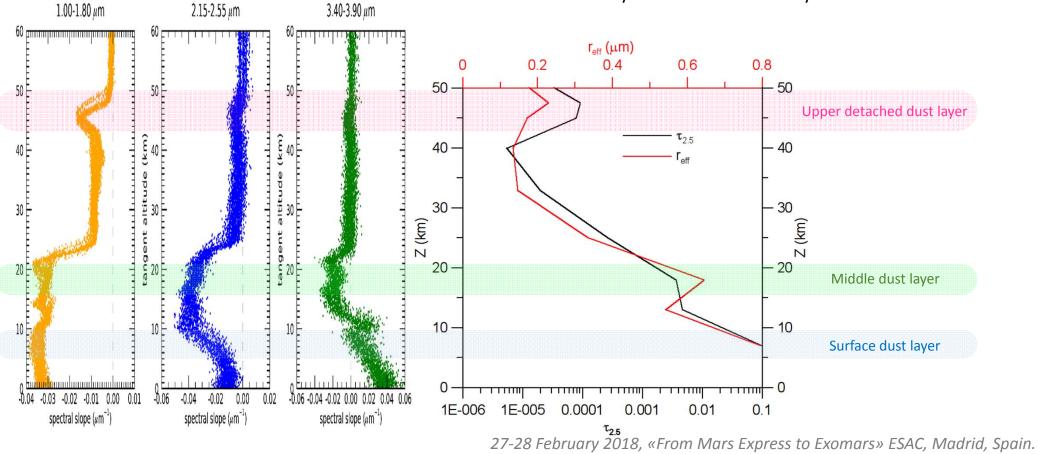
- a) Selecting OMEGA nadir cubes covering the region below the limb observation;
- b) Correcting them with SAS method to remove contribution from gases and dust;
- c) Averaging this dust-free surface albedo weighted by illumination and viewing angles relative to the given tangent point.



## Dust profiles – first results for detached layer cube

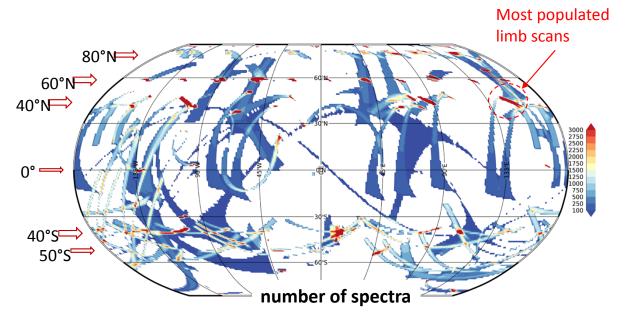
Retrievals obtained from 0.5 to 2.5 micron, by assuming:

- single population of dust grains (lognormal constant Veff)
- constant composition (Wolff et al.)
- Two enhanced dust layers at 45 km (detached) and 18 km (not detached).
- Increase of both density and size within the layers.



### Conclusions

- $\triangleright$  OMEGA limb data are promising for retrieval of **vertical profiles** of both **gases** (H<sub>2</sub>O, CO, O<sub>2</sub>), **dust** properties (abundance and grain sizes) and **water ice** when present.
  - > Several issues must be addressed for quantitative retrievals, both instrumental (misalignment, straylight) and for Radiative Transfer modelling (multiple scattering in spherical geometry, surface shine).
    - The first attempt to study a limb spectral scan in presence of a <u>detached layer</u> was successful in retrieving two dust layers overlapped to an exponential background dust distribution.



- ➤ OMEGA data include <u>several limb vertical</u> <u>spectral scans</u> that will be studied in the future. They appear mostly concentrated at mid latitudes (*red zones*), but are characterized by variable vertical resolution (about 1.5 to 10 km/pixel) and observing conditions (incidence, phase, season).
- Final results may be valuable in providing insight on atmospheric vertical structure in a climatological framework.