A multi-spacecraft reanalysis of the atmosphere of Mars S. R. Lewis, J. A. Holmes and M. R. Patel, The Open University

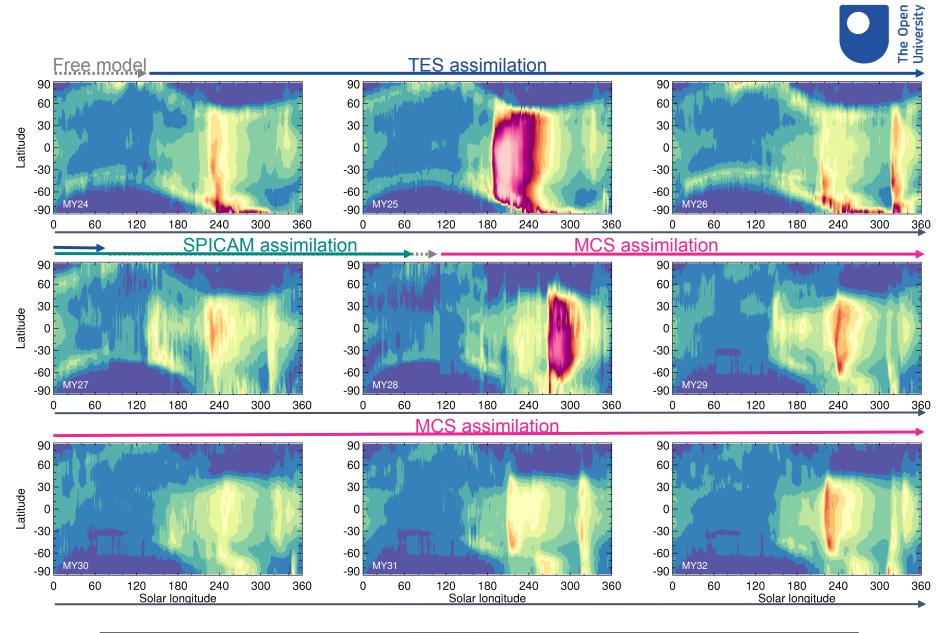


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A multi-spacecraft reanalysis of the atmosphere of Mars

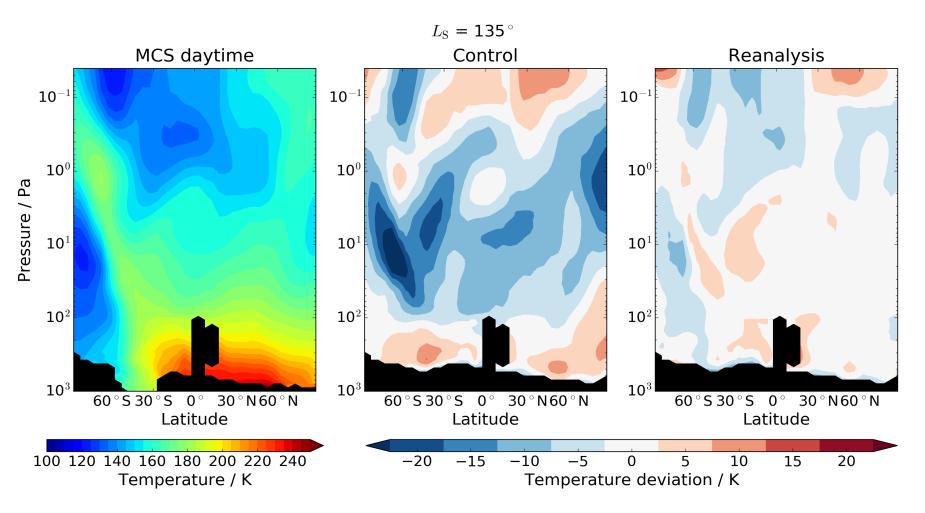
- MGS, MEx, and MRO data, covering over ~8 MY, have been assimilated in a continuous 9 MY integration with the latest UK version of the LMD Mars Global Climate Model
- Includes water cycle and improved dust representation
- Includes ozone photochemistry
- Thermal profiles and total dust opacity (TES) or dust profiles (MCS) are assimilated on 10 km (TES) or 5 km (MCS) vertical scales
- Water vapour is assimilated by matching total column abundance with MGCM prediction
- Ozone is assimilated by matching total column abundance with MGCM prediction
- Baseline assimilations at T31 resolution (roughly 3.75° or < 225 km horizontal grid) with 35 vertical levels
- Focused studies at higher horizontal and vertical resolution (100 levels)







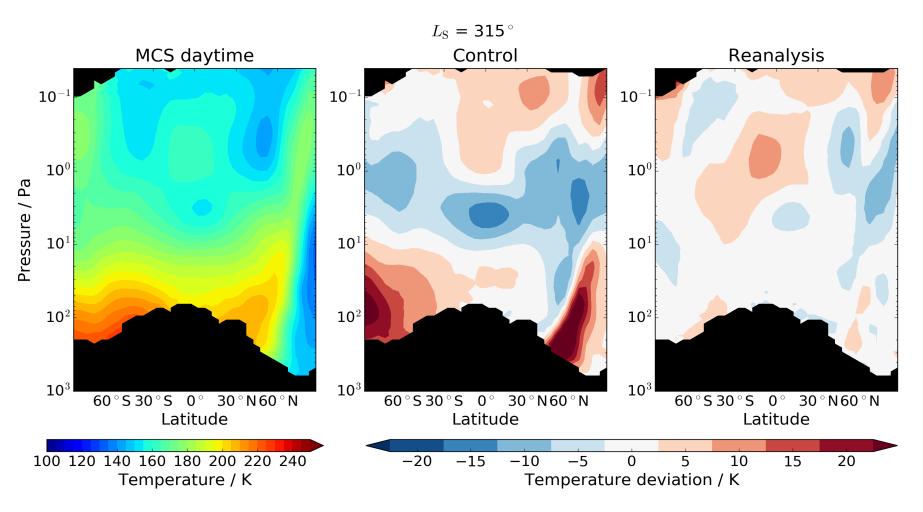
Validation of the 9-year reanalysis



(Left) MCS daytime retrievals, (middle) temperature deviations in the free-running GCM and (right) temperature deviations in the new reanalysis averaged over $L_s = 120-150^{\circ}$



Validation of the 9-year reanalysis

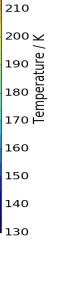


(Left) MCS daytime retrievals, (middle) temperature deviations in the free-running GCM and (right) temperature deviations in the new reanalysis averaged over $L_s = 300-330^\circ$



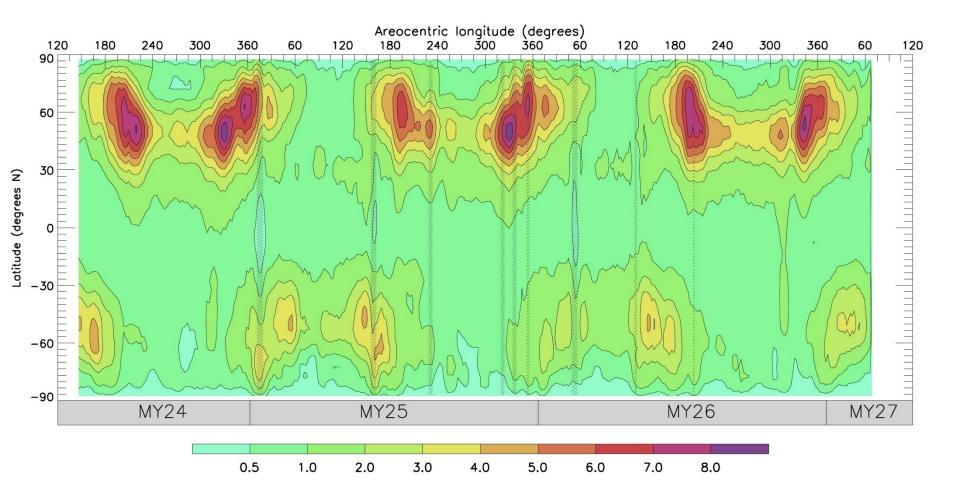
MY 24-26 60 ° N 30 ° N Latitude 30°S 60°5 o MY 27-29 60 ° N 30 ° N Latitude 30°5 60°S MY 30-32 60 ° N 30 ° N Latitude 30°5 60°S ٥- $L_{\rm S}$

Sample output of atmospheric temperature at 3 pm and 60 Pa altitude





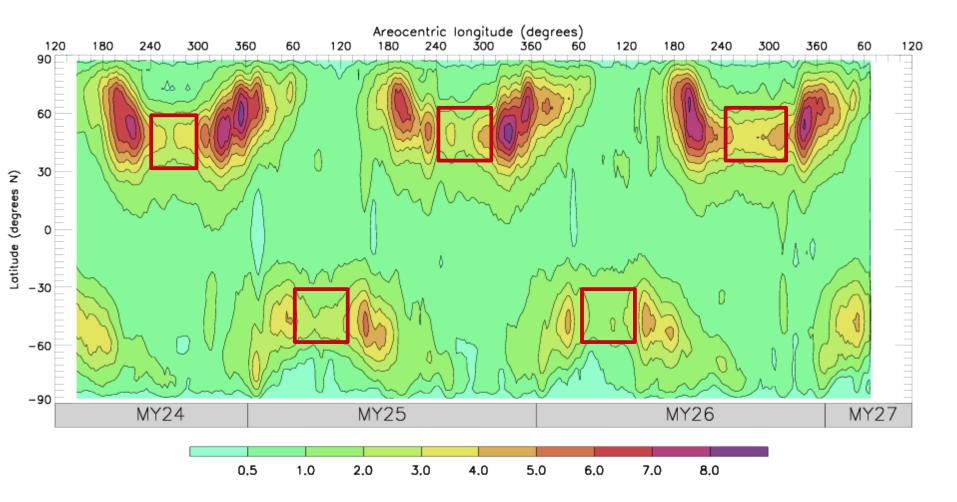
Lewis et al. (2016), Icarus 264, 456-464.



RMS temperature variability at 2.5 km altitude from the TES-period assimilation, band-pass filtered to periods of 1.5–30 sols and with a 20-sol running mean applied. Hatching indicates no TES data. (Lewis et al. 2016, http://dx.doi.org/10.1016/j.icarus.2015.08.039)



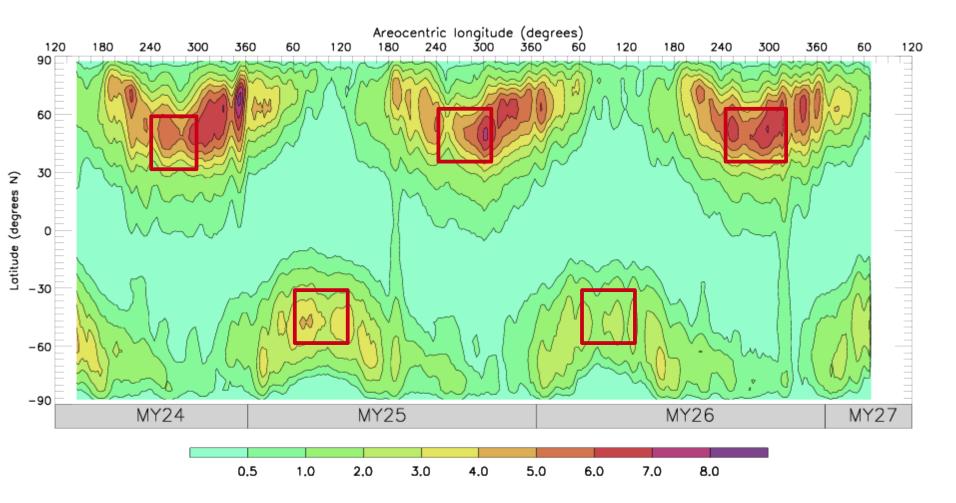
Same time period (TES) but from new reanalysis using updated GCM



RMS temperature variability at 2.5 km altitude from the TES-period assimilation, band-pass filtered to periods of 1.5–30 sols and with a 20-sol running mean applied.



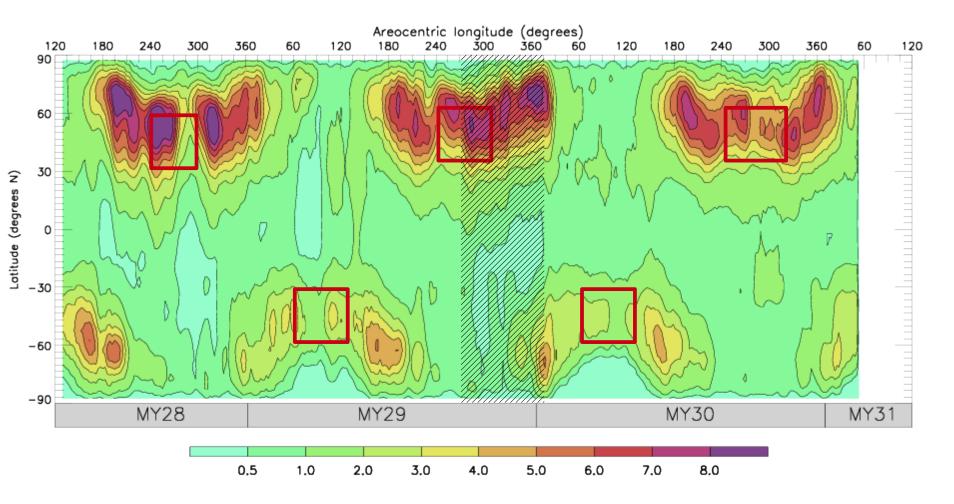
Same time period (TES) but free-running model using updated GCM



RMS temperature variability at 2.5 km altitude from the TES-period assimilation, band-pass filtered to periods of 1.5–30 sols and with a 20-sol running mean applied.



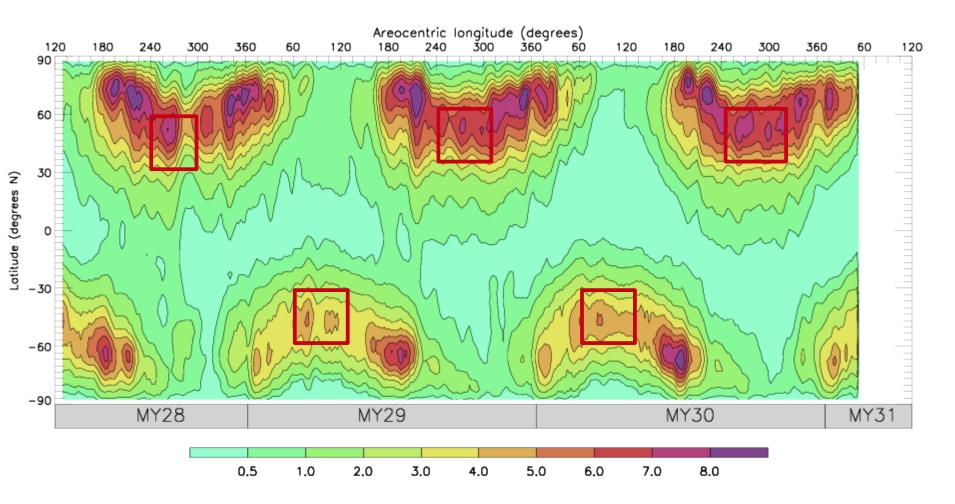
New time period (MCS) from new reanalysis using updated GCM



RMS temperature variability at 2.5 km altitude from the MCS-period assimilation, band-pass filtered to periods of 1.5–30 sols and with a 20-sol running mean applied.



New time period (MCS) but free-running model using updated GCM



RMS temperature variability at 2.5 km altitude from the MCS-period assimilation, band-pass filtered to periods of 1.5–30 sols and with a 20-sol running mean applied.

- MGS, MEx, and MRO data, covering over ~8 MY, have been assimilated in a continuous 9 MY integration with the latest UK version of the LMD Mars Global Climate Model
- Inter-annual variability in the solsticial pause investigation extended from TES period to now also include MCS period
- Preliminary results show enhanced wave activity during MY28 global dust storm, with brief solsticial pause present. General increase in wave activity during MCS time period
- Free-running GCM unable to adequately reproduce solsticial pause