

Discussion

Mars Express -Trace Gas Orbiter collaboration



Logistics



09:15-11:00 Meeting part1

11:00-11:30 Coffee

11:30-13:15 Meeting part2

13:30-14:30 Lunch at ESAC

13:30 Departure of Bus1 to Airport

14:30 Departure of Bus2 to Airport



Agenda

1. D. Titov, H. Svedhem. Introduction (10 min)
2. A. Cardesin, B. Geiger. Analysis of joint observation opportunities (20 min)
3. Discussion (5-10 min/ presentation)
 - *Surface properties and geology*
 - *Lower atmosphere and meteorology*
 - *Upper atmosphere, aeronomy, plasma environment and escape*
 - *Instruments cross-calibration*
 - *Modelling and retrieval tools with multi-mission data*
 - *Comparison of radiative transfer codes*
 - *Collaboration with other missions (MAVEN, MSL, Isight, Exomars 2020, ...).*

Goals of the discussion

- Outline science areas and topics for the most efficient collaboration
- New ideas for collaborative science
- Identify operational requirements and opportunities for joint MEX and TGO observations
- Outline synergies in data analysis
- Identify needs for cross-calibrations of the instruments, algorithms and codes



Joint MEX-TGO collaboration plan

Mars

Distance: 18,678.5 km
Radii: [3,397 3,397 3,375] km

2018-Apr-11 22:10:41 UTC
1,000x time

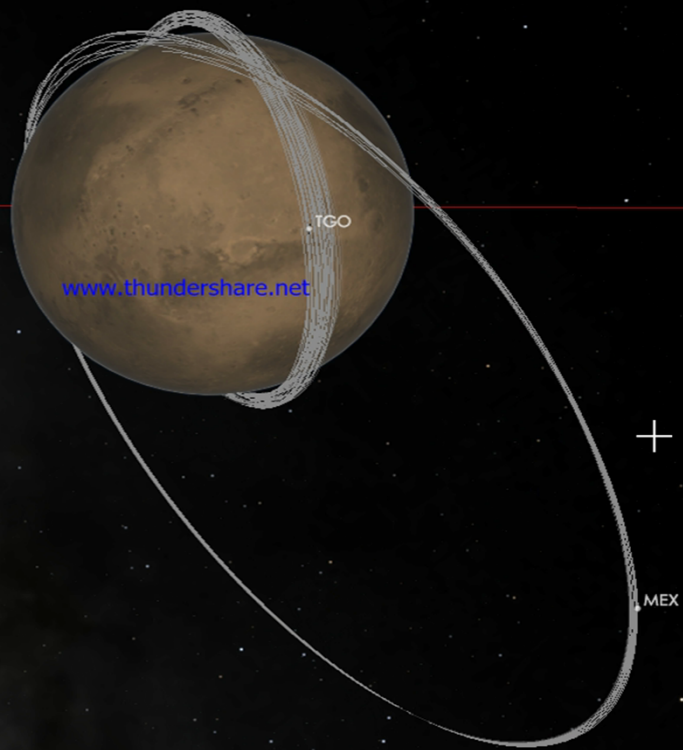


Analysis of observation opportunities /A. Cardesin & B. Geiger/

Vesta

Deimos

Jupiter



www.thundershare.net

Mars Express E



SPICE - Cosmographia

ESA | 26/04/2017 | Slide 5

European Space Agency

Surface properties and geology

- *HRSC and CASSIS*
- *MARSIS, MaRS (BSR) and FREND*
- *Targets for joint observations*

Lower atmosphere and meteorology



MEX-TGO climate synergies

A. Määttänen

Mars Express IDS

MEX Legacy

- Climatologies of:
 - Temperature
 - Dust
 - Water ice clouds
 - Water vapour
 - Ozone
 - Carbon monoxide
 - CO2 ice clouds
 -
- **ADVANTAGE OF MEX:**
access to multiple local times (in contrast to MGS/TES, MRO/MCS...)

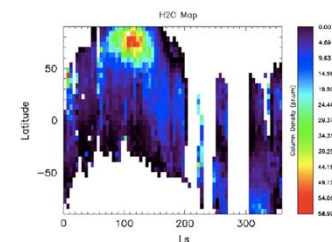
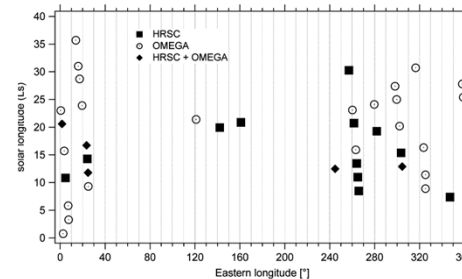
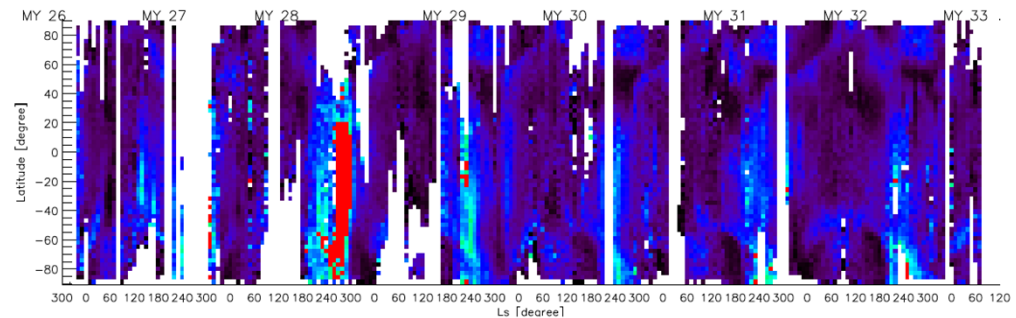
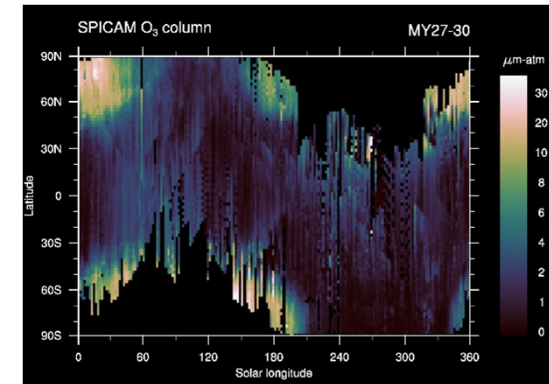
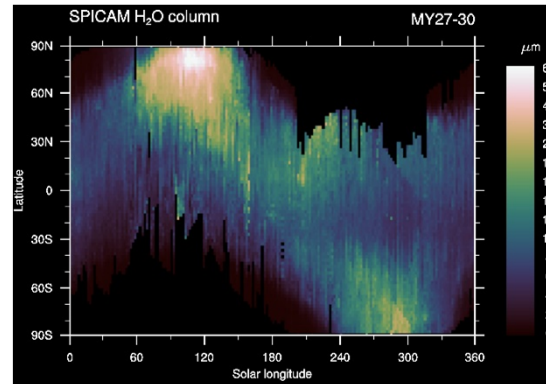


Fig. 8. Map of retrieved abundance of water vapour as a function of Solar Longitude (Ls) and Latitude as observed by PFS/SWC, from orbit 634 to orbit 6537, obtained using spectra averaged in 5° Ls × 5° Latitude bins.

MEX + TGO

- For climate studies long, continuous datasets are important:
 - Statistics
 - Interannual variability
- For climate studies good local time coverage is important:
 - Diurnal variations (examples: thermal tide, cloud formation, dust storm evolution)
- 🗄️ TGO will complement the MEX climatologies
 - continuing the datasets of MEX
 - overlapping observations (intercomparison/cross-calibration possible?)
 - with a different LT coverage (complement the LT coverage)
- if TGO mission extends beyond MEX lifetime, it will also prolong the climatologies

MEX + TGO + ground-based observations

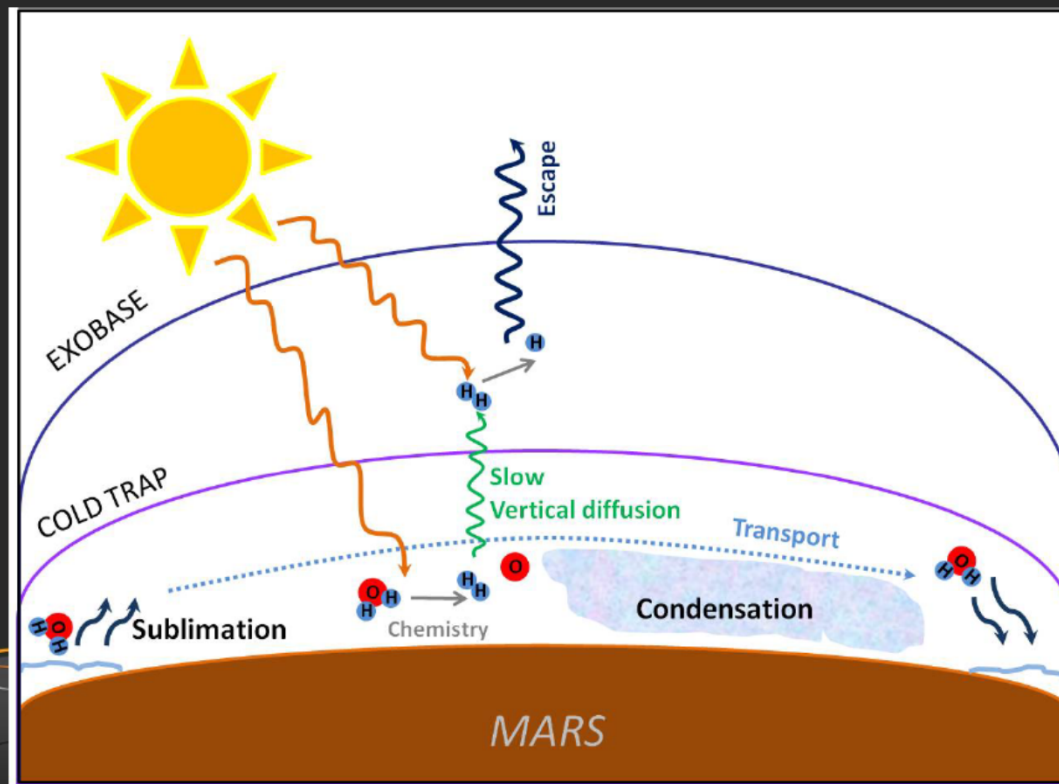
- F. Forget pointed out at the last MEX SWT that MEX could be used in a novel way: full-disk viewing at apoapse
 - VMC is already doing this!
 - using the other instruments would allow for obtaining full-disk snapshots of climatological variables
 - sometimes better lighting conditions (when periapse in the night): increases the observation opportunities during the mission
 - confrontation with higher resolution observations (MEX+TGO)
 - confrontation with (possibly near-simultaneous) ground-based full-disk observations
 - would this be possible?

Upper atmosphere and escape



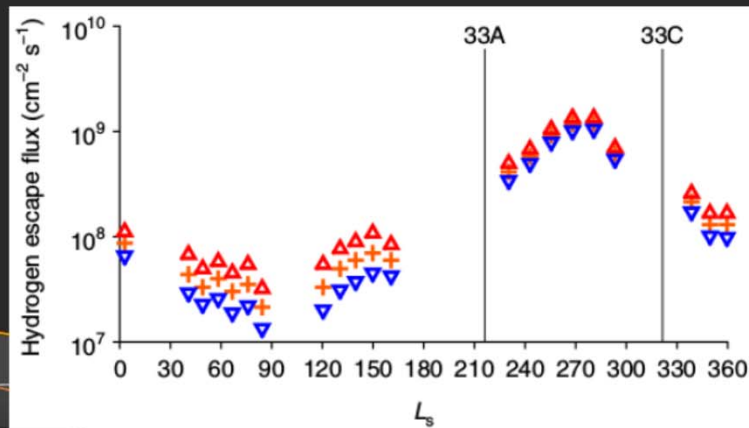
MEX+TGO+MAVEN: Water from surface to escape

- Classical picture of water escape



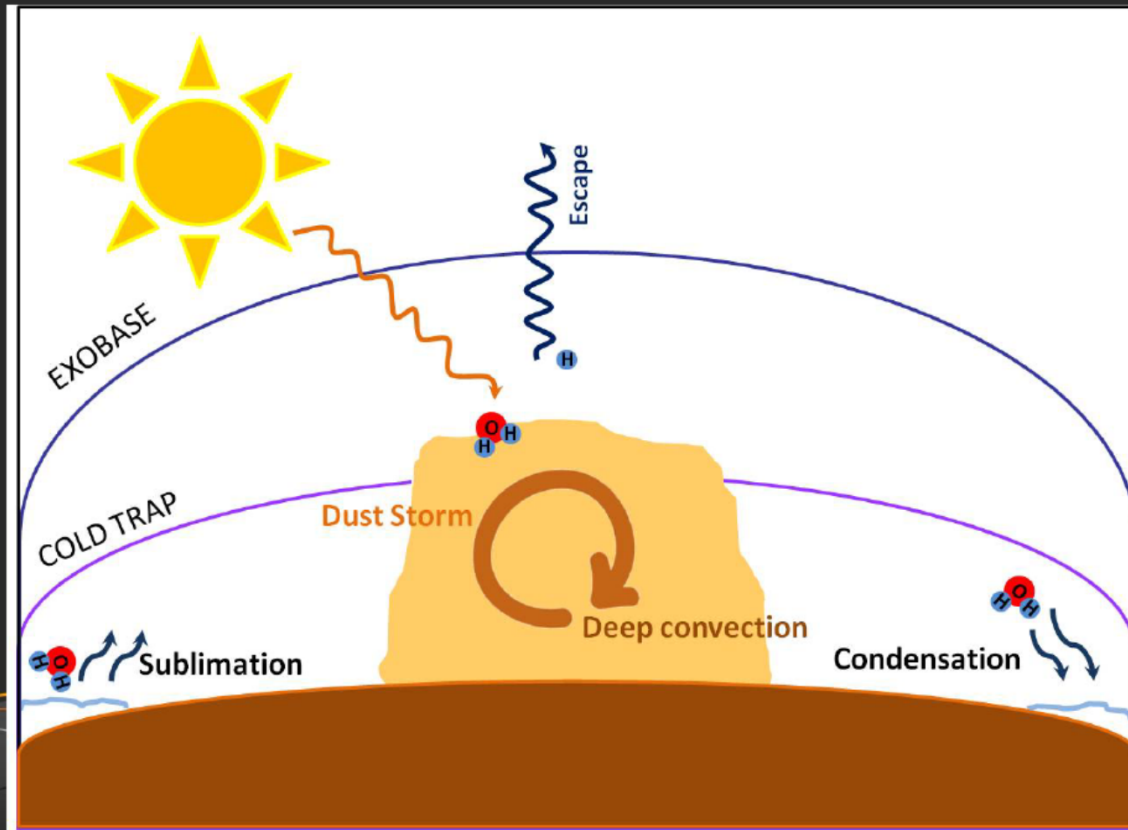
ME_x+TGO+MAVEN: Water from surface to escape

- Classical picture of water escape
- Recent observations show things are more complicated:
 - Strong fast variability in H thermal escape, seasonal and/or dust storms
 - Supersaturated water vapour in the middle atmosphere
 - Water-derived ions



Heavens et al., 2018

ME_x+TGO+MAVEN: Water from surface to escape



MEx+TGO+MAVEN: Water from surface to escape

- Classical picture of water escape
- Recent observations show things are more complicated:
 - Strong fast variability in H thermal escape, seasonal and/or dust storms
 - Supersaturated water vapour in the middle atmosphere
 - Water-derived ions
- Better characterization of links between lower atmosphere and H escape by combining:
 - MEx measurements of water and dust cycles (inc. profiles by SPICAM SO)
 - Regular TGO measurements of H₂O (and HDO) profiles
 - MAVEN measurements of H escape and water ions

Radio Science on TGO vs MEX

- **Not yet an official topic on TGO**
 - Earth Radio occultation measurements can be achieved at a fairly low cost and with a high value
- **MEX → TGO mutual radio occultations**
 - Better spatial and local time coverage
 - Independence on radio-occultation seasons
 - Higher S/N ratio
 - No distortions in the Earth atmosphere/ionosphere and interplanetary medium

Instruments cross-calibration

- *CASSIS - HRSC*
- *ACS & NOMAD - SPICAM*
- *ACS (TRVIM) - PFS*

Comparison of radiative transfer codes and retrieval tools



Collaboration with other missions



Additional slides

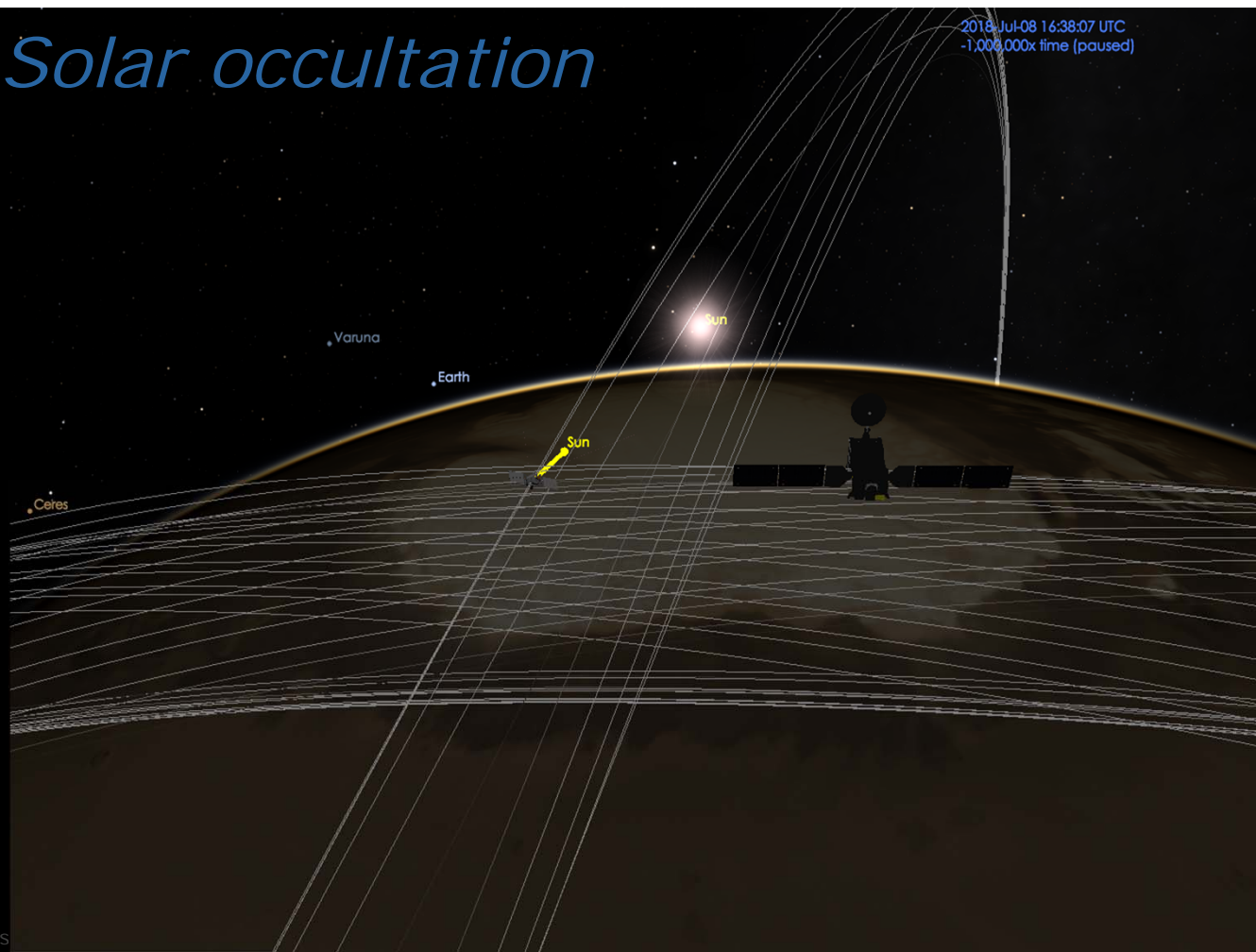


MEX – TGO collaboration



Solar occultation

2018 Jul-08 16:38:07 UTC
-1,000,000x time (paused)



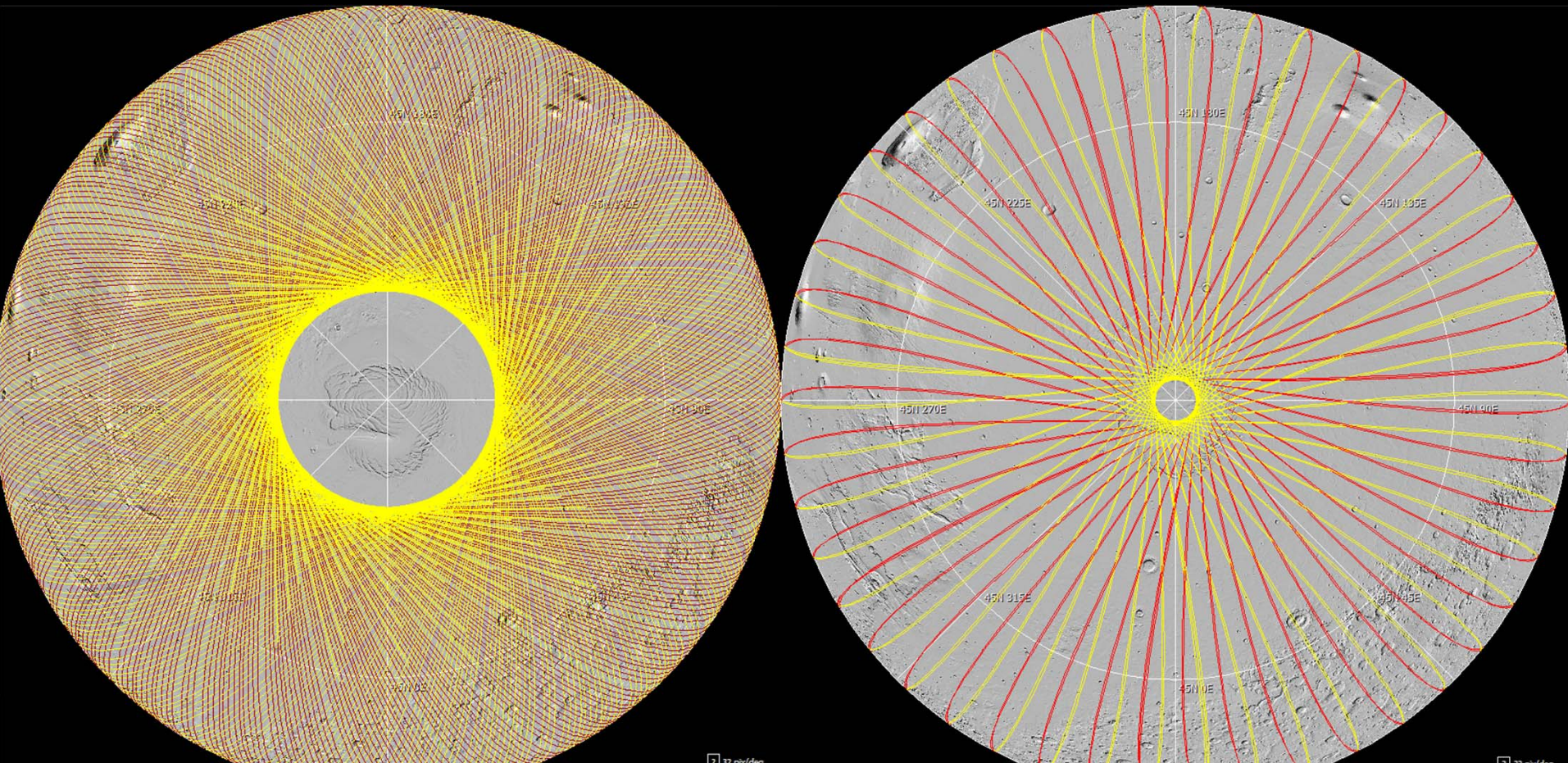
Mars Express Extens

SA | 26/04/2017 | Slide 23



European Space Agency

Nadir observation



Orbits alignment



Topics for the discussion



- climatological parameters (temperature, minor species, dust and ices)
- assimilation of MEX measurements in GCMs
- couplings between atmospheric layers:
from the lower atmosphere to the exosphere
- dependence of escape on the state of the lower atmosphere

- surface: photometric properties and temporal variations
- sub-surface ice and its correlation with surface geology and mineralogy
- evolution of the polar caps and polar atmosphere



Implementation



➤ Collaboration in data analysis

- *Early sharing and discussion of results*
- *Cross comparison of retrieved climatological parameters*
- *Joint use of TGO and MEX data in GCMs*

➤ Collaboration at operations level

- *MEX support by providing data products*
- *Analysis of joint observation opportunities*
- *Joint Long-term activity plan (targets, campaigns)*
- *Simultaneous and co-located observations*
- *Filling gaps in the surface coverage*



Measures to foster MEX-TGO collaboration



- Exchange of Co-Is between TGO and MEX
- Selection of IDS and participating scientists on ExoMars
- “Gap analysis” in both topical and competence domains
- Implementation of operational synergies and collaboration at operational level
- Release of MEX high-level data products
- Conferences and data workshops

