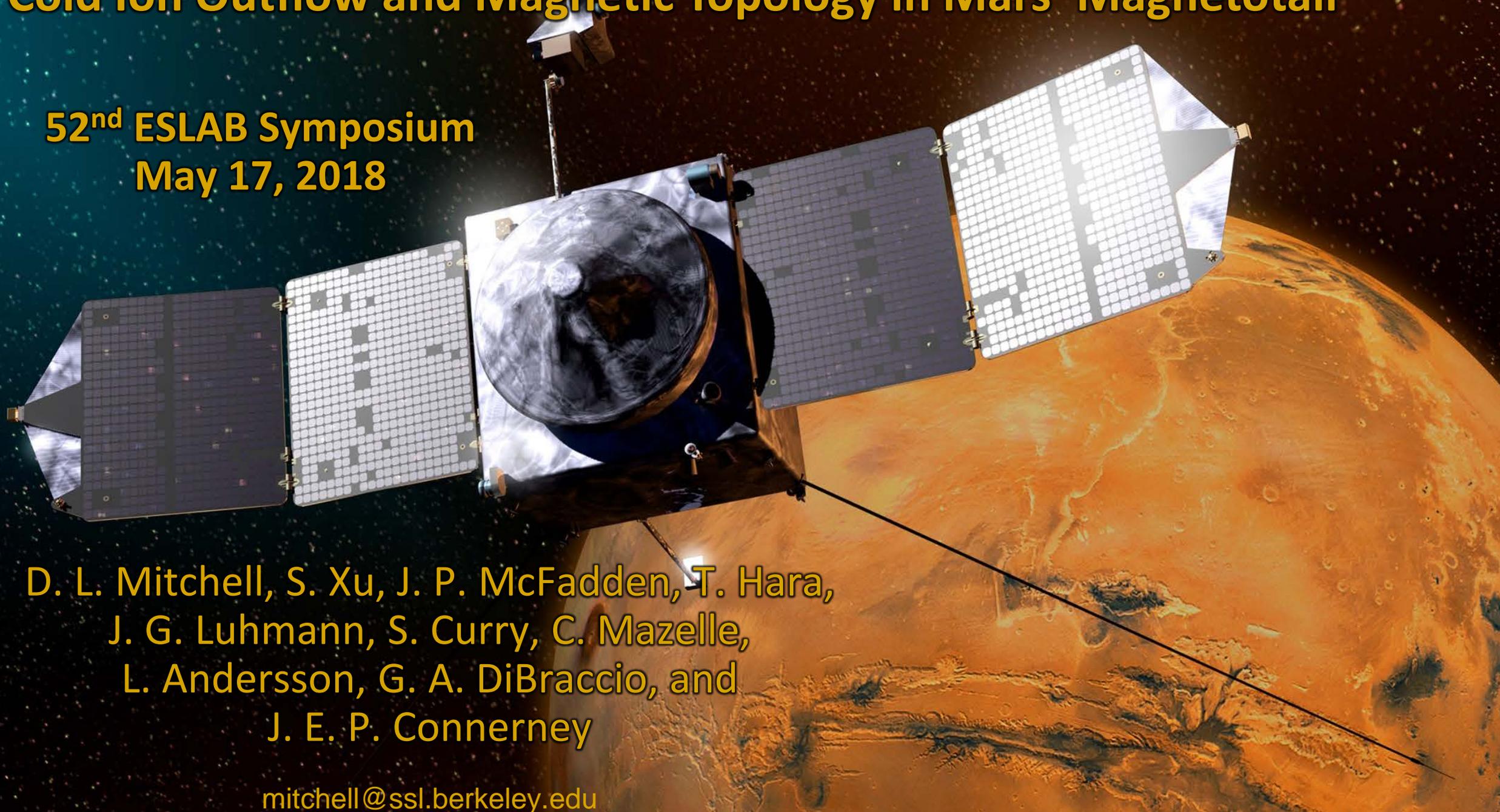


# Cold Ion Outflow and Magnetic Topology in Mars' Magnetotail

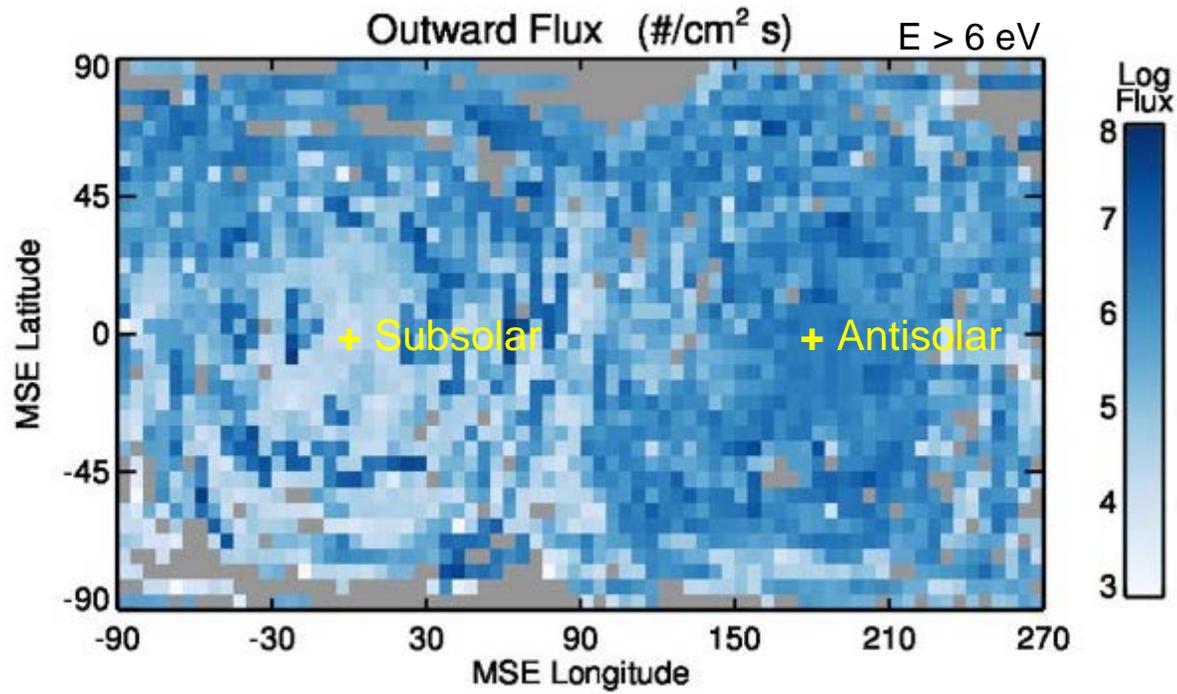
52<sup>nd</sup> ESLAB Symposium  
May 17, 2018



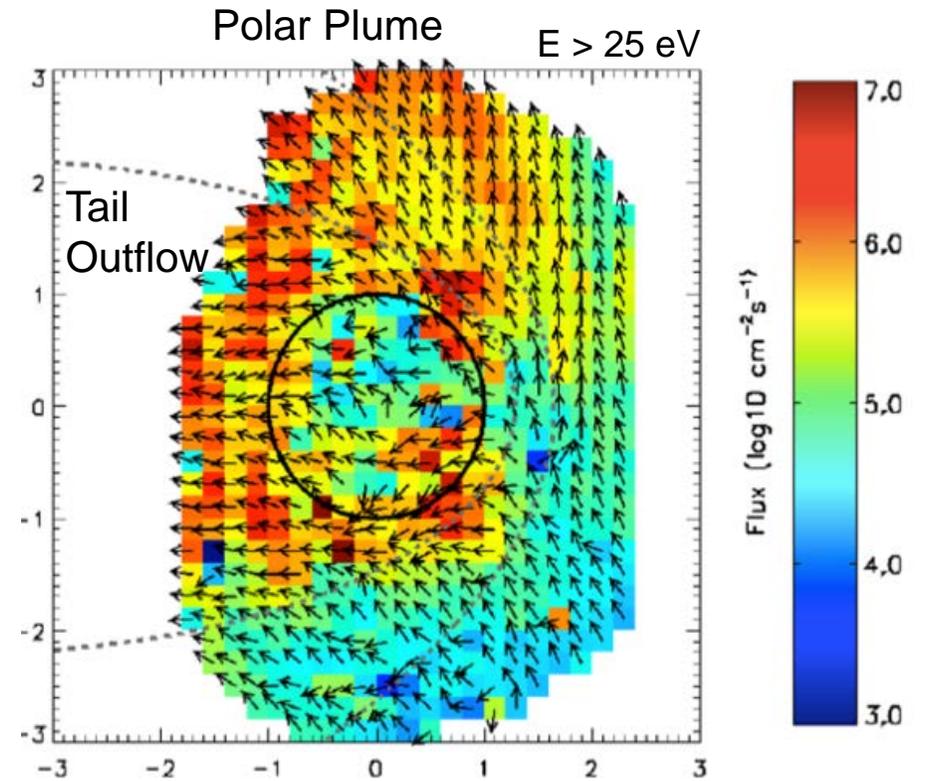
D. L. Mitchell, S. Xu, J. P. McFadden, T. Hara,  
J. G. Luhmann, S. Curry, C. Mazelle,  
L. Andersson, G. A. DiBraccio, and  
J. E. P. Connerney

[mitchell@ssl.berkeley.edu](mailto:mitchell@ssl.berkeley.edu)

# Oxygen Ion Loss from MAVEN

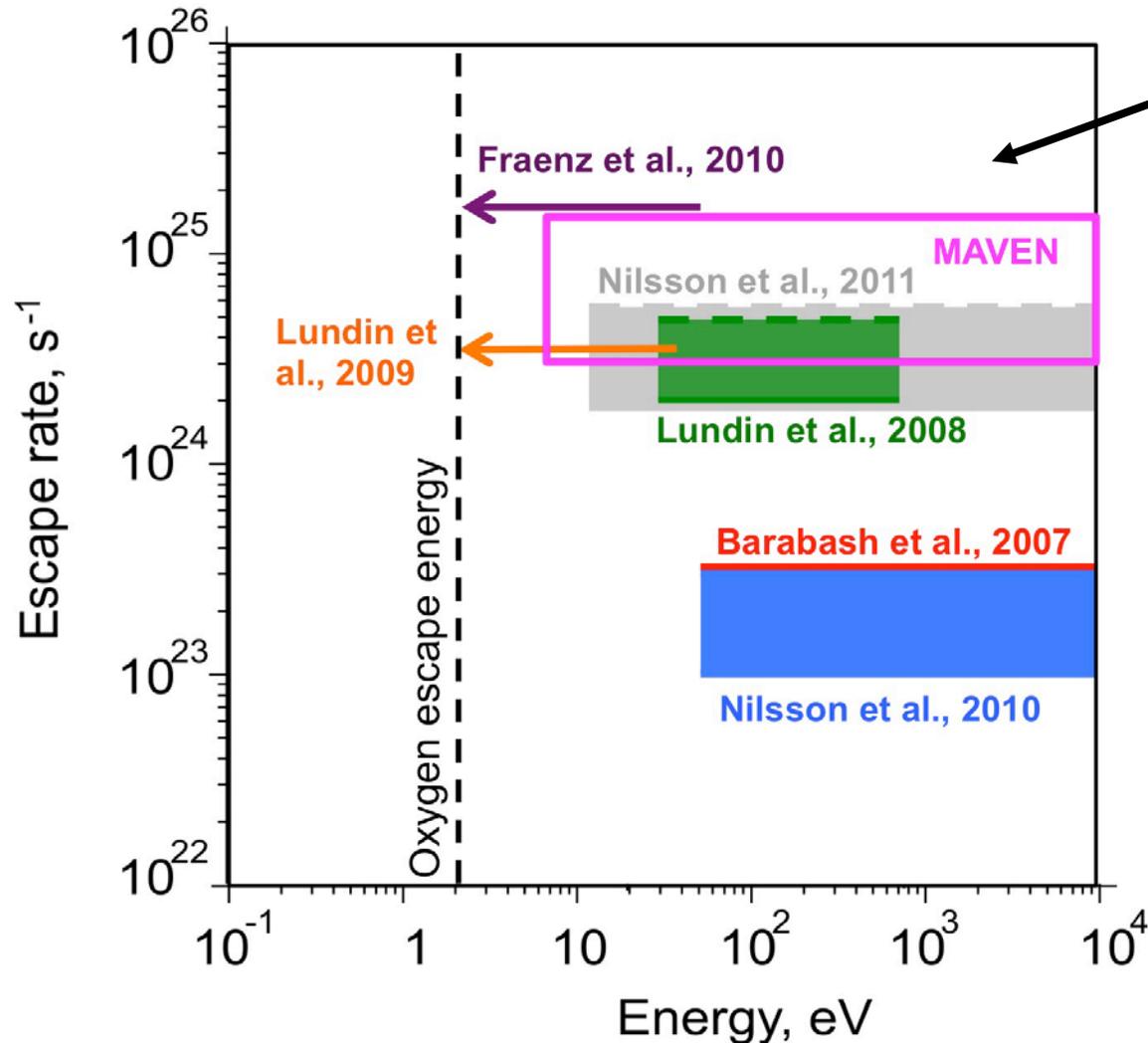


Brain et al. 2017



Dong et al. 2015

# Oxygen Ion Loss from Mars Express



Escape rate increases as lower energy threshold decreases.

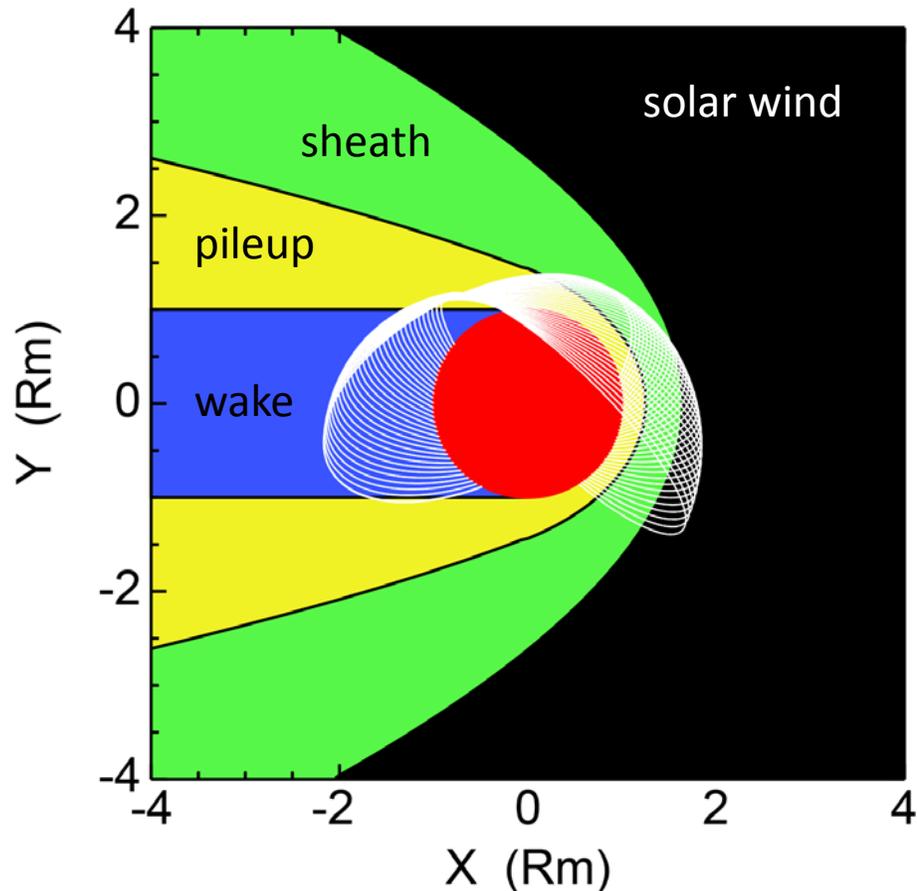
Trans-terminator flow (Fränz et al. 2010)

- Assume flow is strictly tailward with an isotropic temperature.
- Assume loss rate is cylindrically symmetric around terminator.

“Statistical” 3D distributions (Nilsson et al 2011)

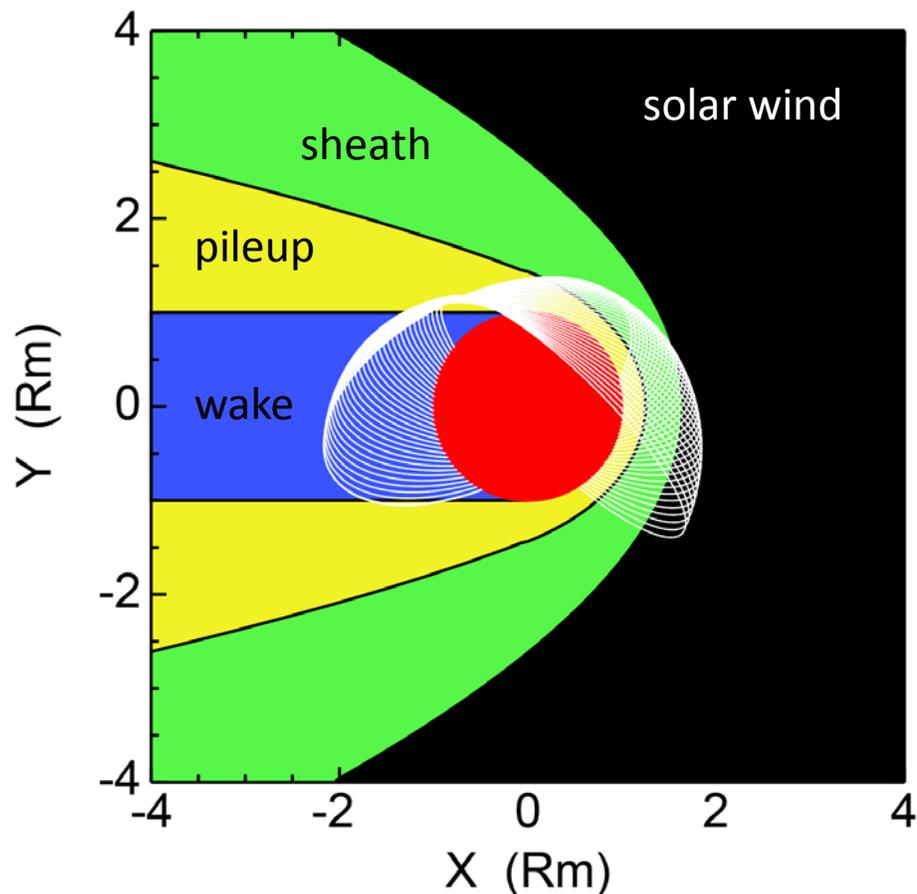
- Combine many 2D observations in different orientations.

# MAVEN Cold Ion Campaigns



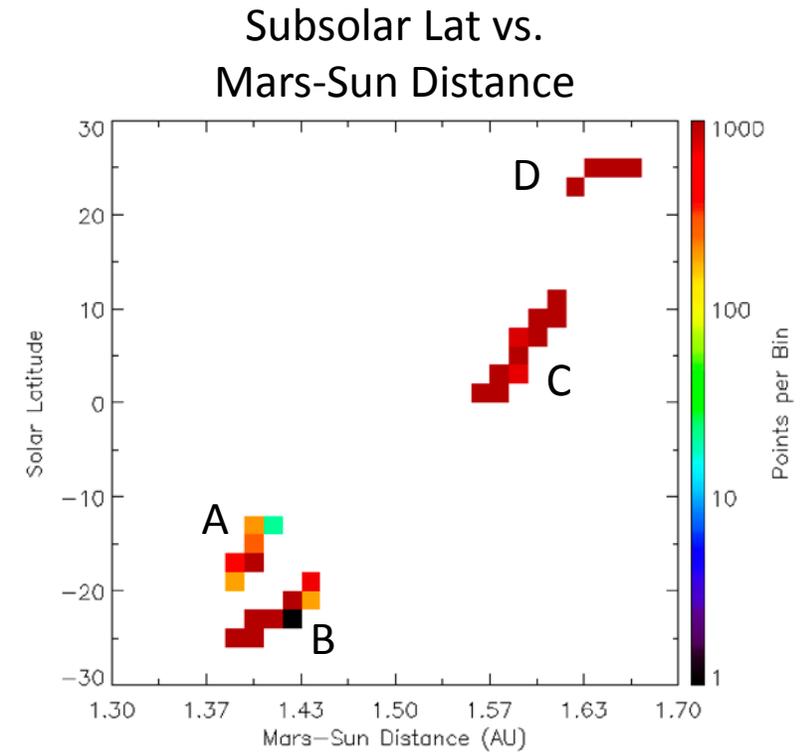
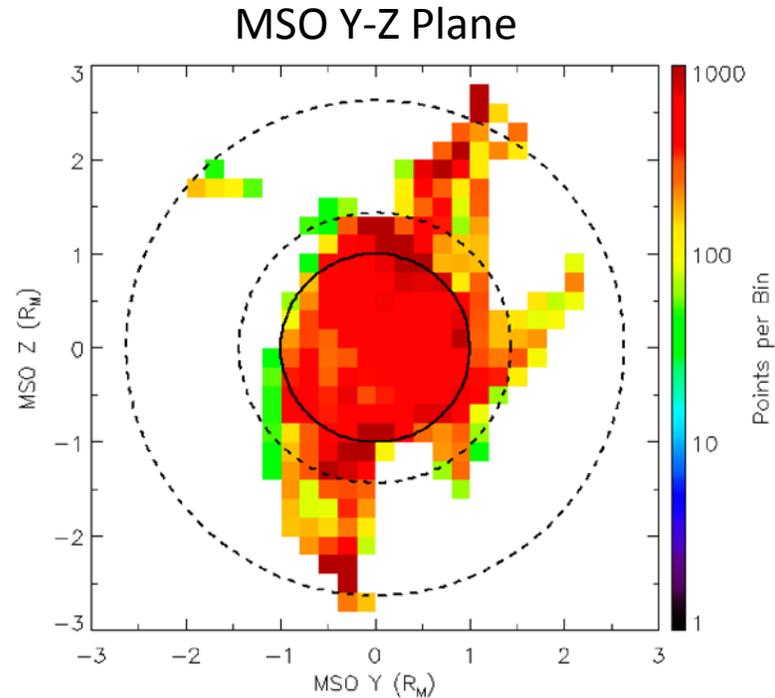
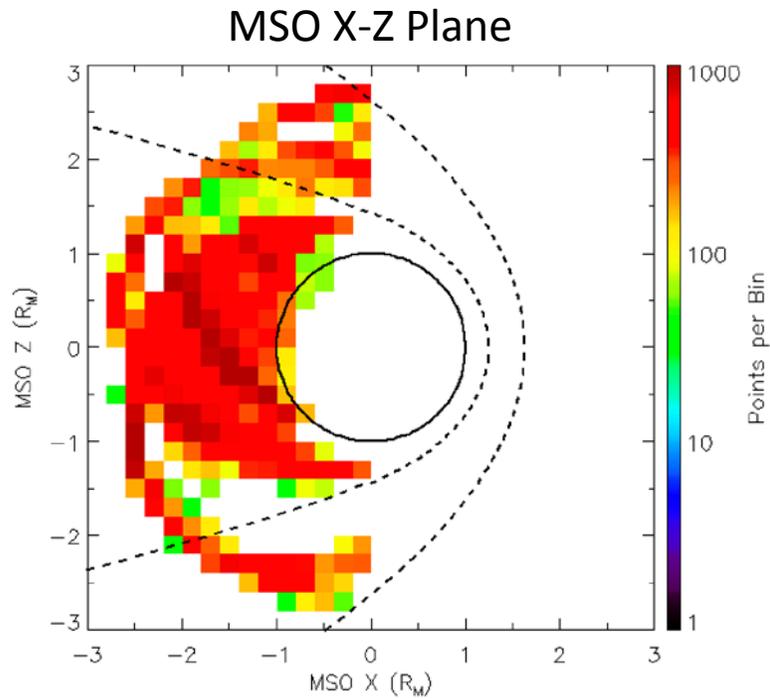
- 5 sweeps through the tail
- 1200 – 5000 km altitude (1.3 – 2.5 R<sub>M</sub>)
- STATIC and SWEA fields of view optimized
- ❖ Magnetic topology from SWEA and MAG
- ❖ Spacecraft potential from SWEA, STATIC, LPW
- ❖ Ion distribution functions from STATIC, corrected for spacecraft potential and motion
- Determine ion outflow down to and below escape velocity

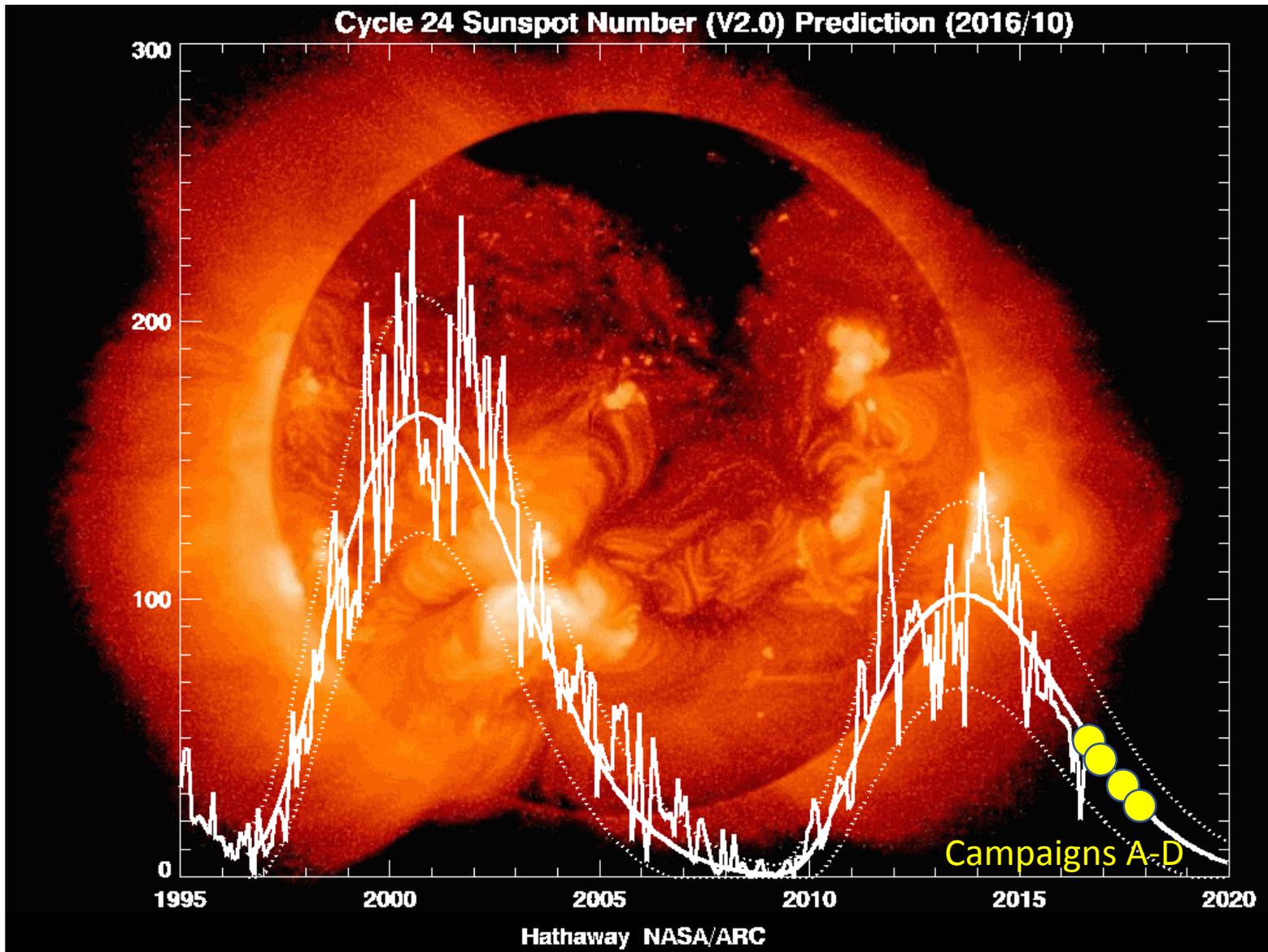
# MAVEN Cold Ion Campaigns



- What are properties of the outflow?
  - composition, temperature, velocity
  - morphology and variability
- How does outflow depend on magnetic topology? What accelerates the ions?
- What are the loss rates of different species and the total loss rate, and how does this compare with previous estimates?
- How does outflow depend on drivers?

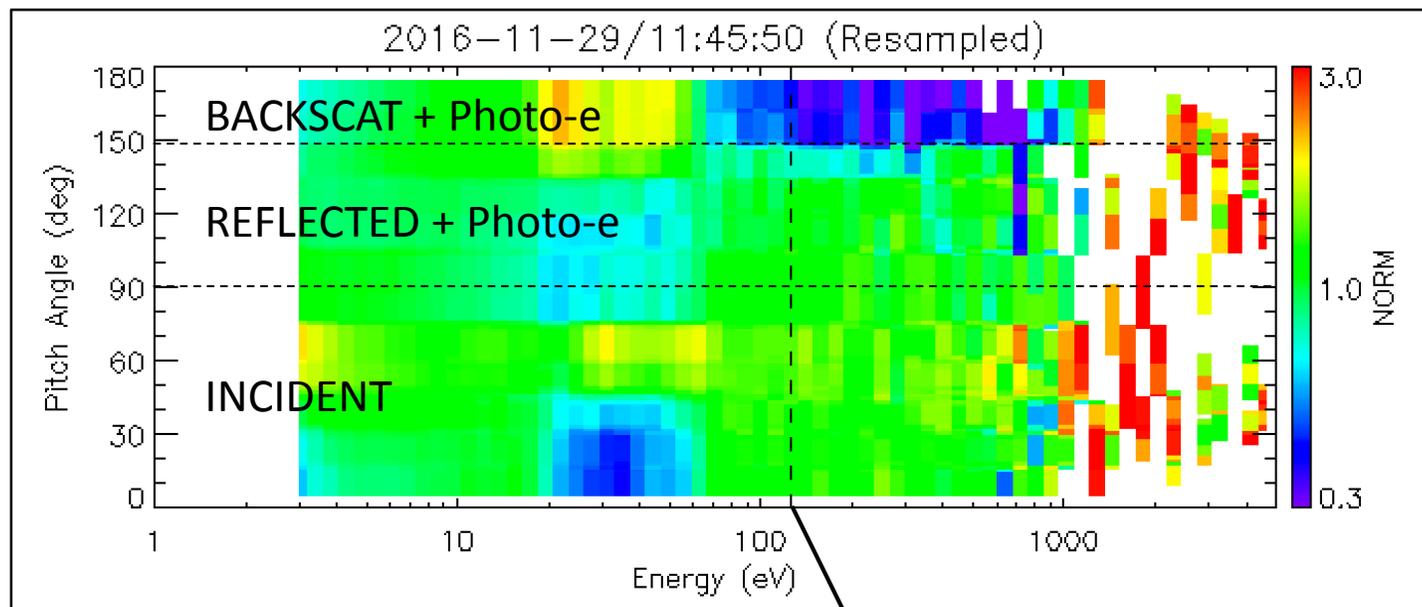
# Sampling for Campaigns A-D



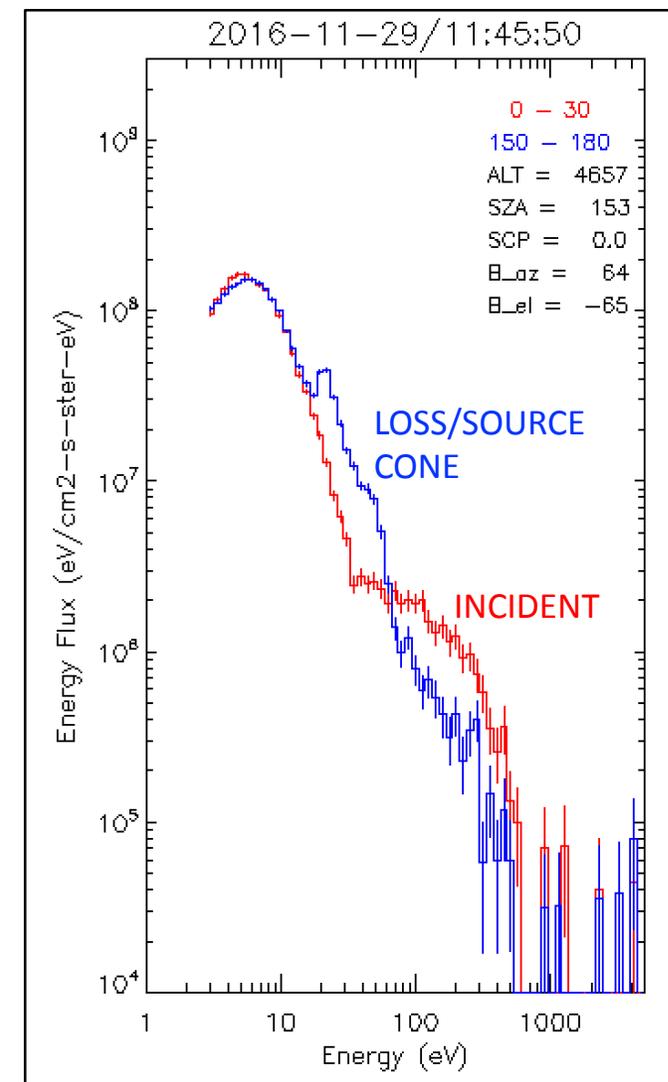
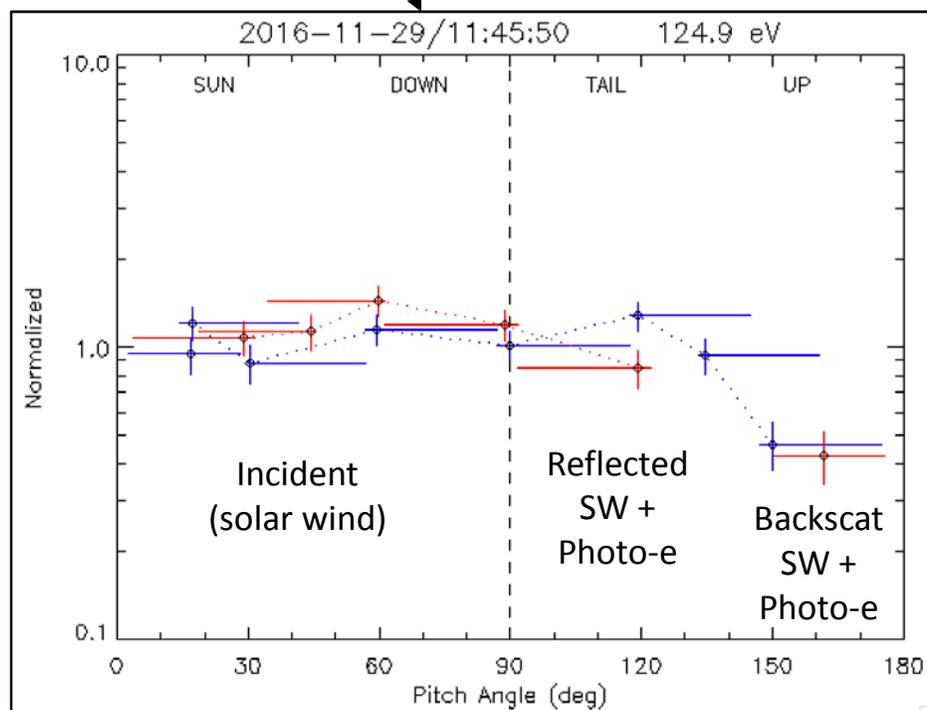


All measurements  
obtained during low  
solar activity.

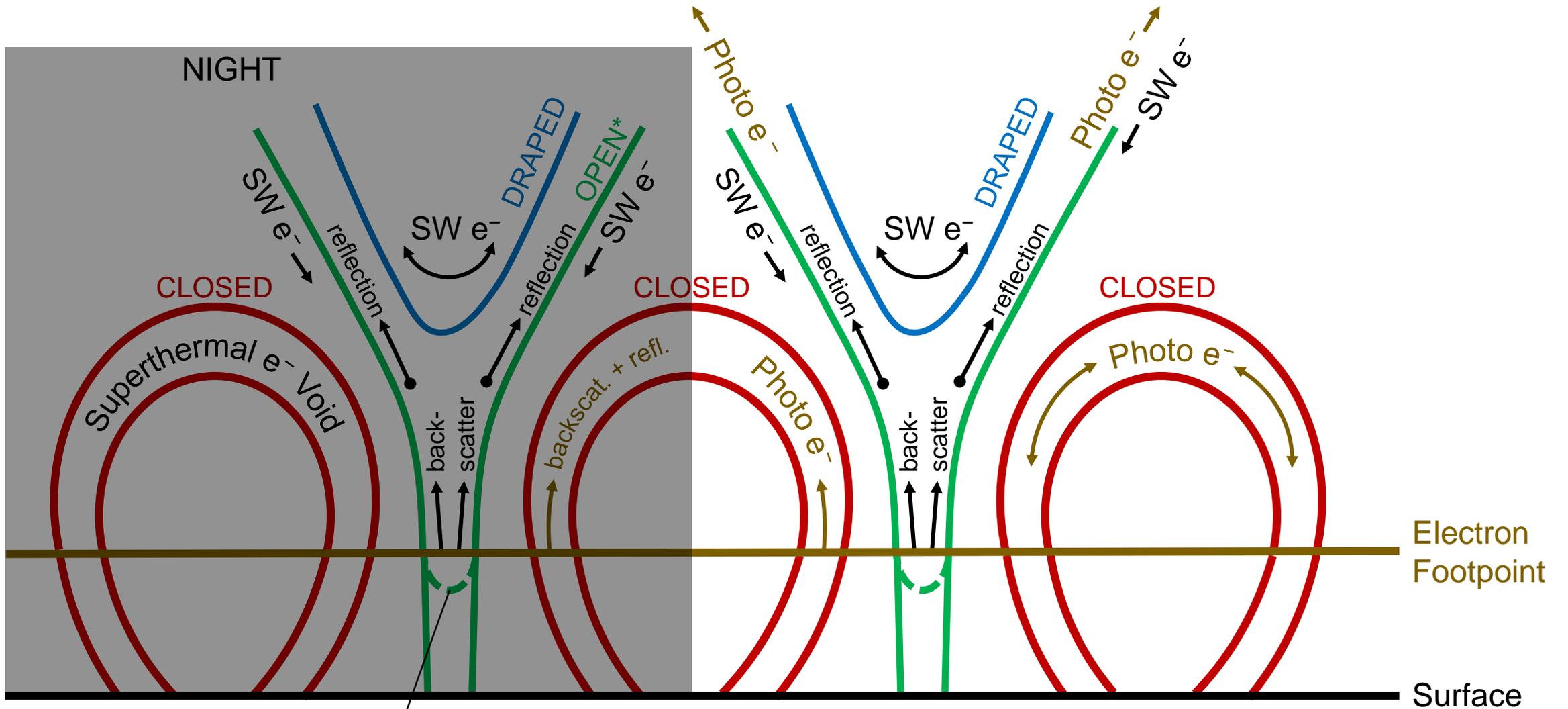
(September 2017 event  
occurred between  
Campaigns C and D.)



PAD Shape  
(Weber and Brain)



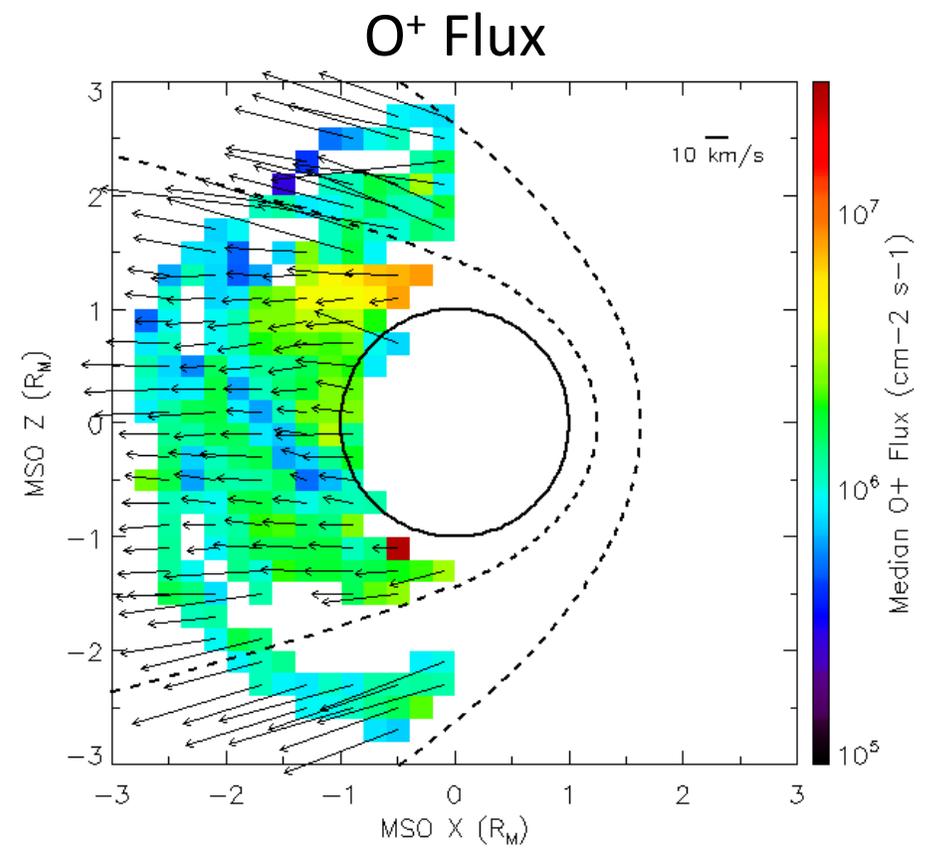
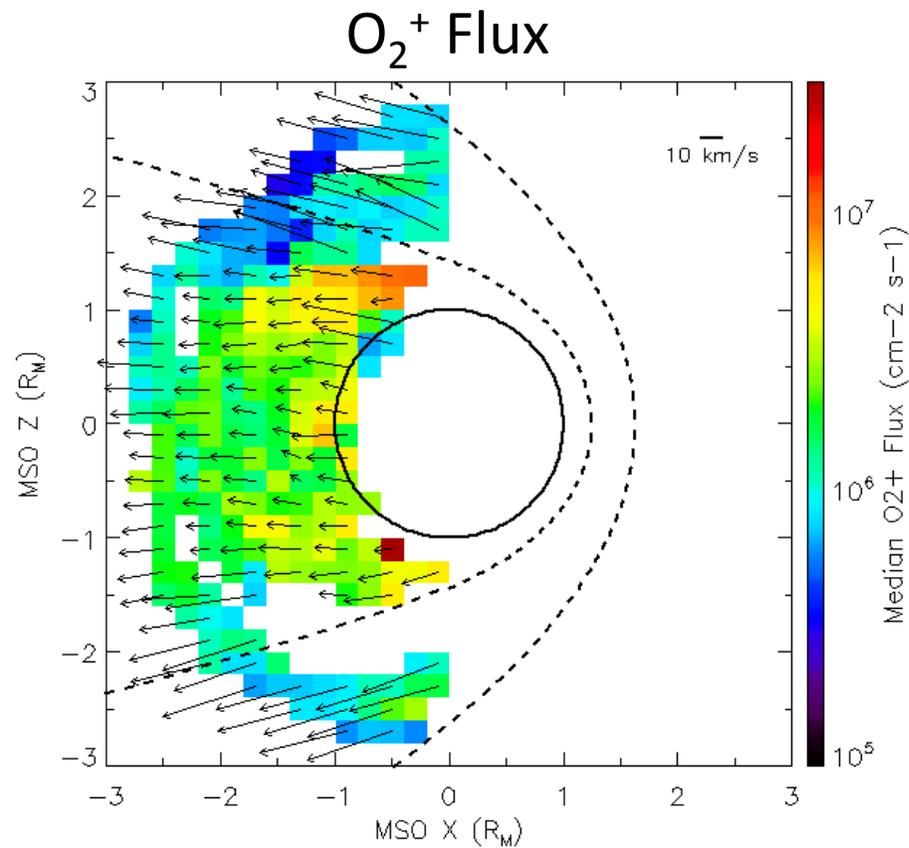
Two-Stream Shape Parameter  
(Xu and Mitchell)



\*or DEEPLY DRAPED

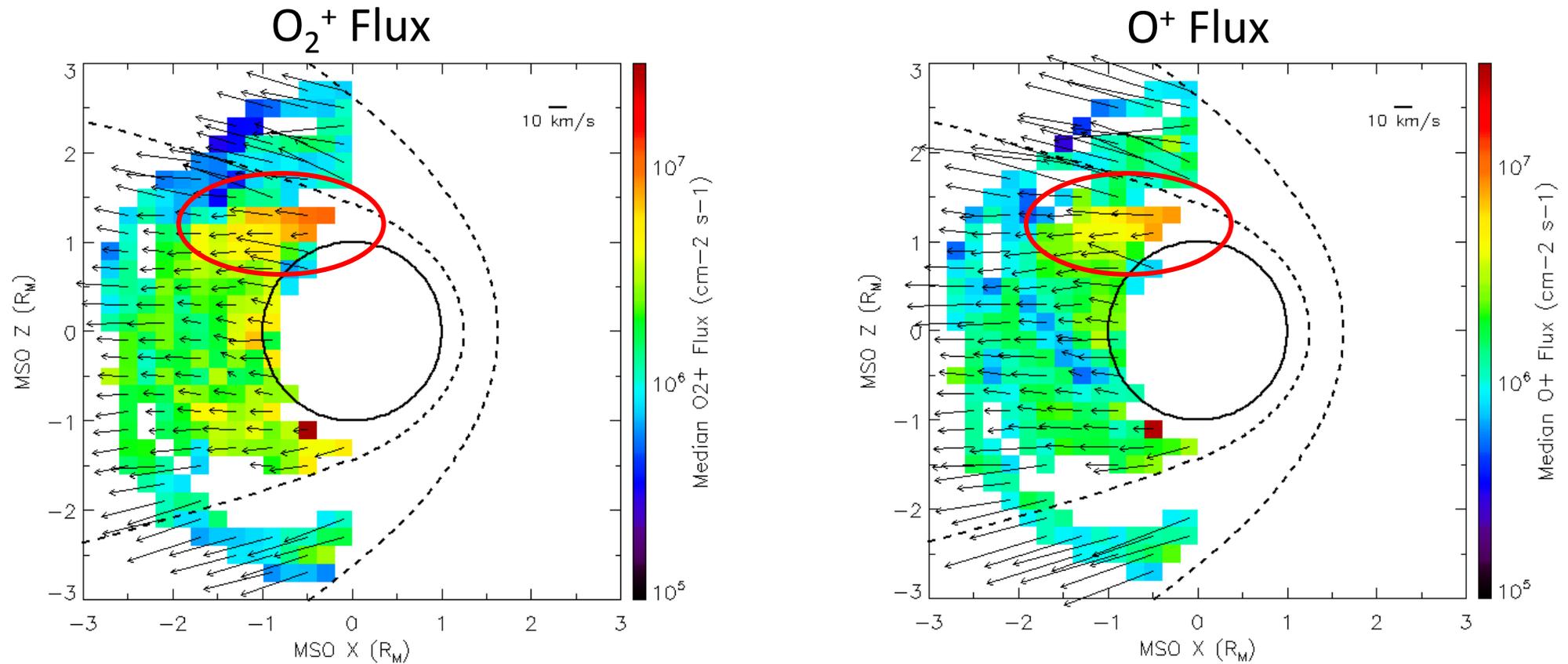
# Ion Outflow in the Tail - Overall

All topologies, all crustal field orientations, all upstream conditions



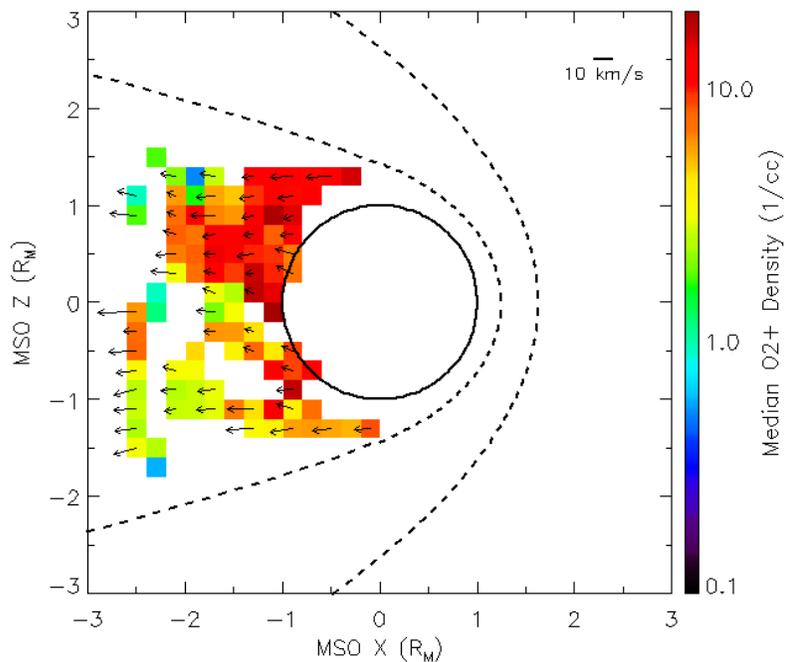
# Ion Outflow in the Tail - Overall

All topologies, all crustal field orientations, all upstream conditions

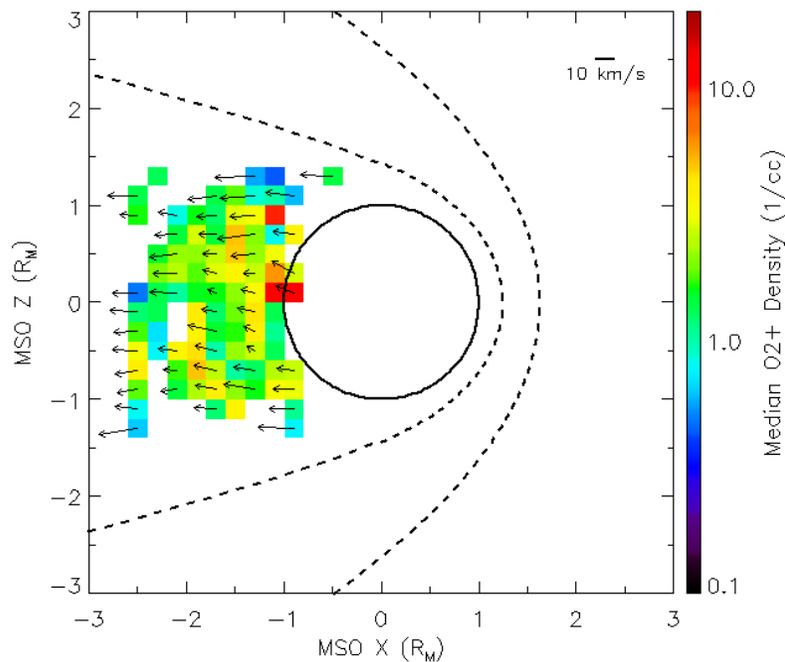


# Magnetic Topology Organizes the Data

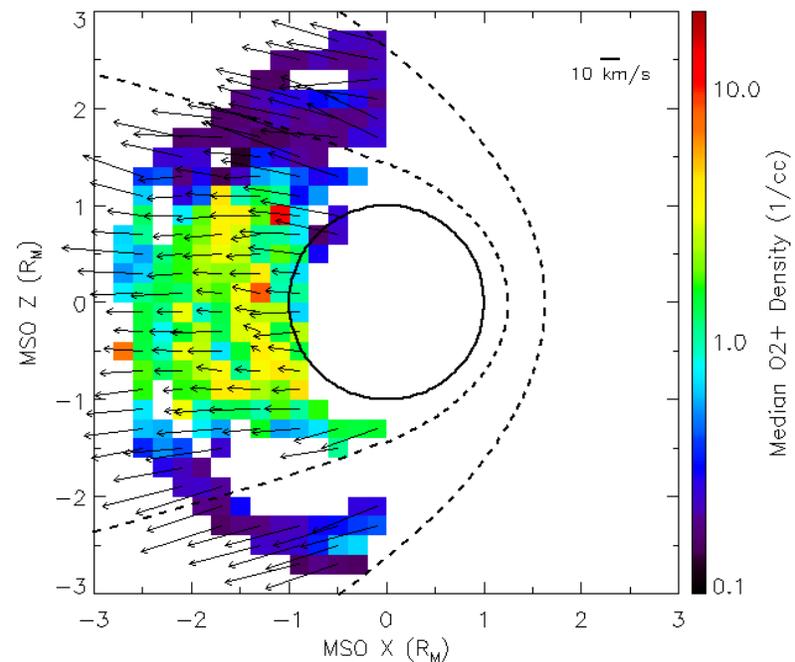
$O_2^+$  Density  
Open to Day



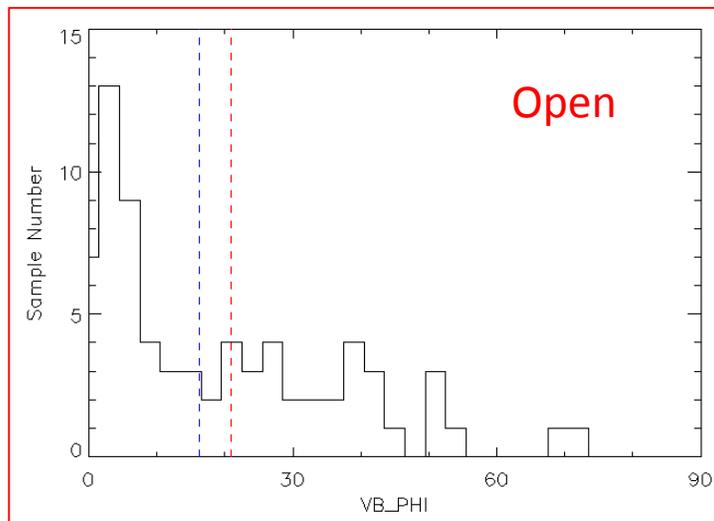
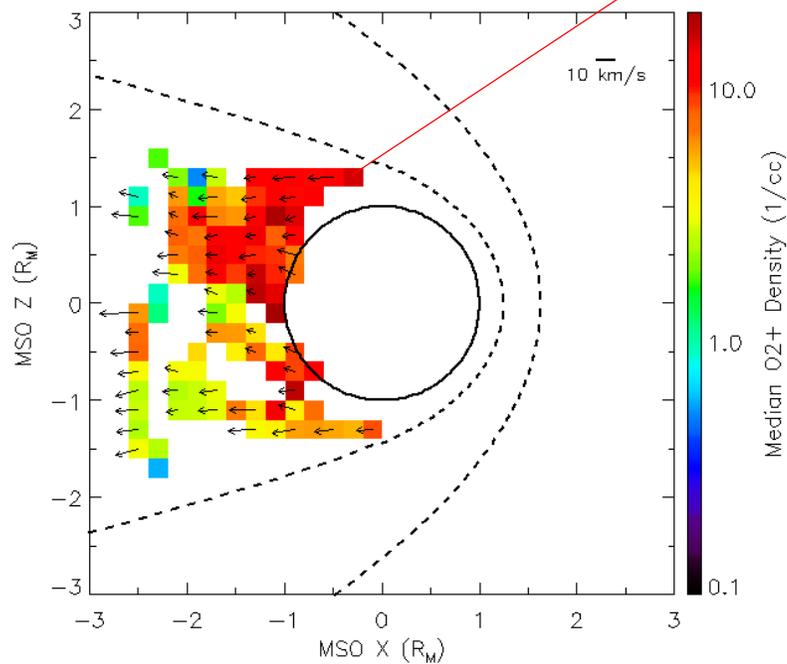
$O_2^+$  Density  
Open to Night



$O_2^+$  Density  
Draped

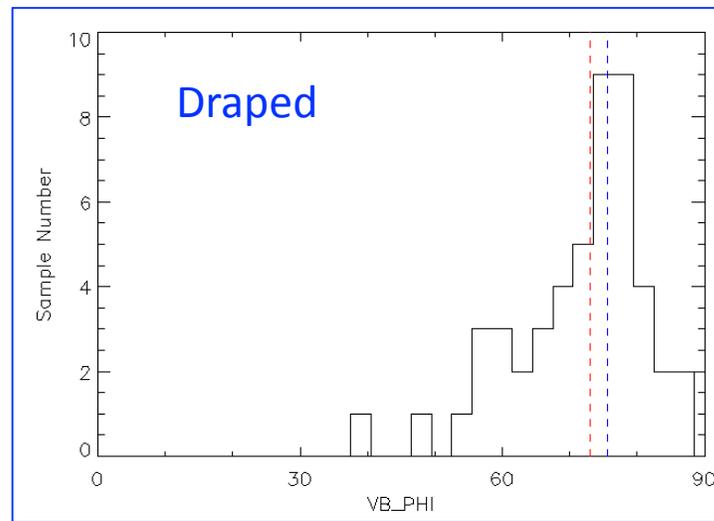
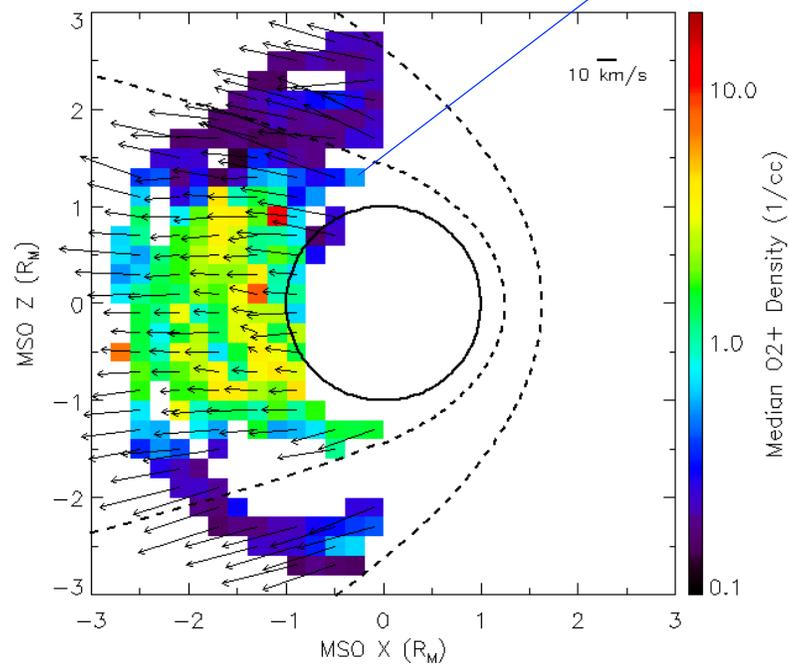


$O_2^+$  Density  
Open to Day



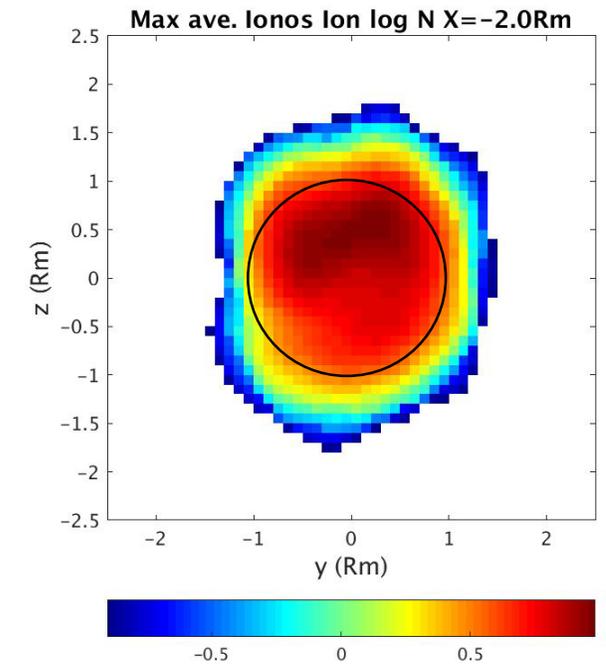
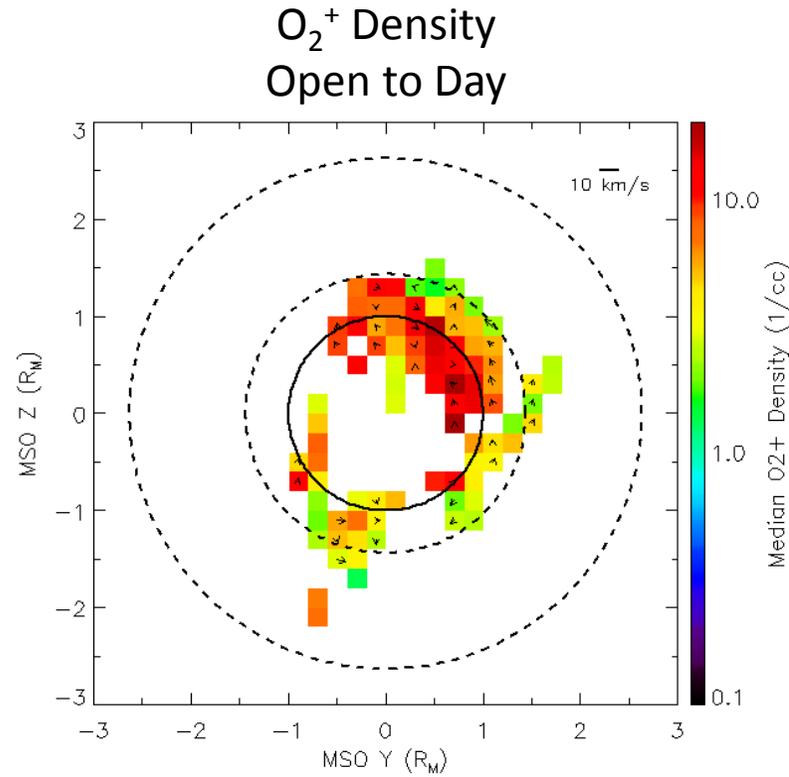
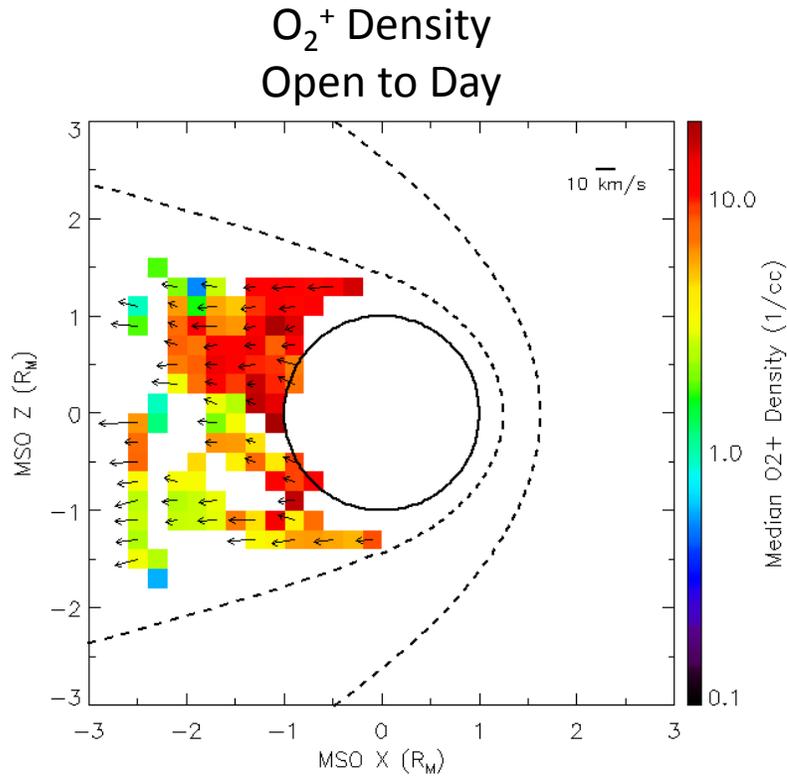
Open

$O_2^+$  Density  
Draped



Draped

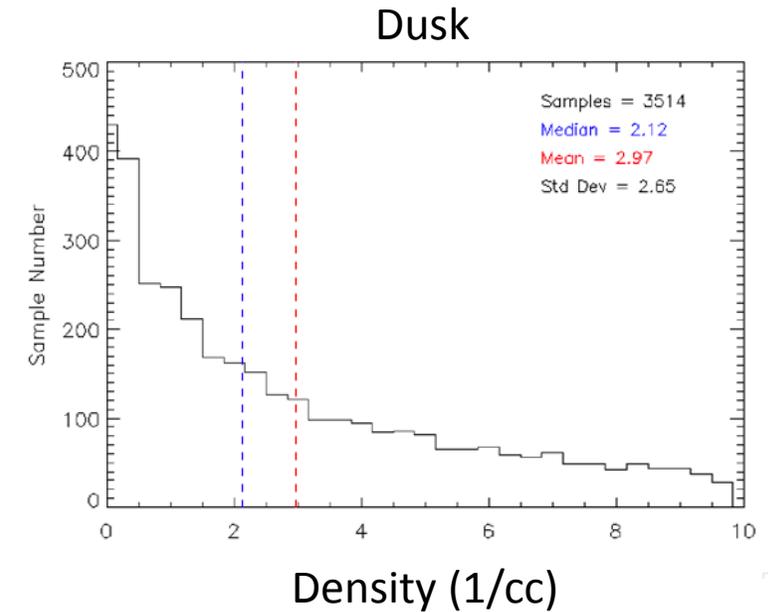
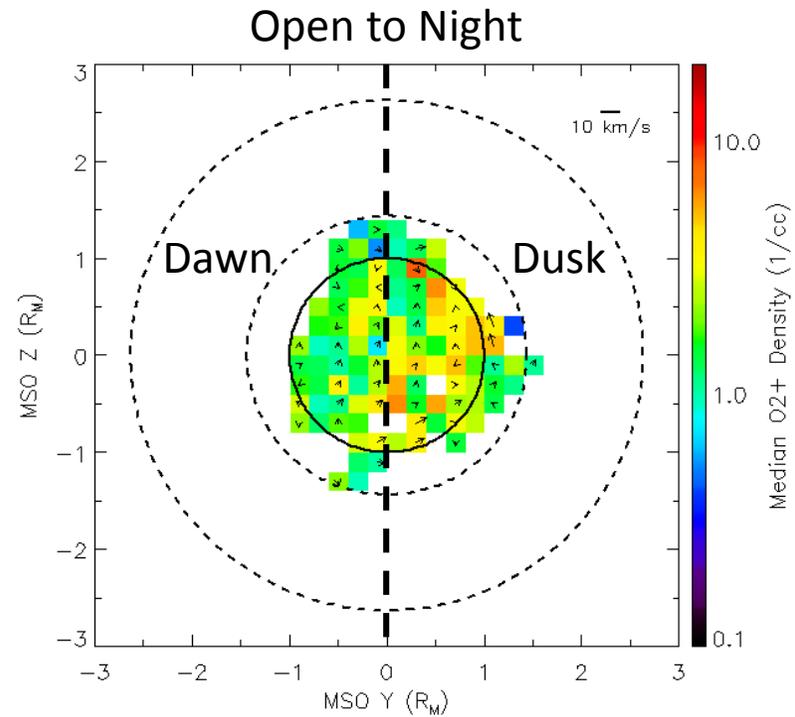
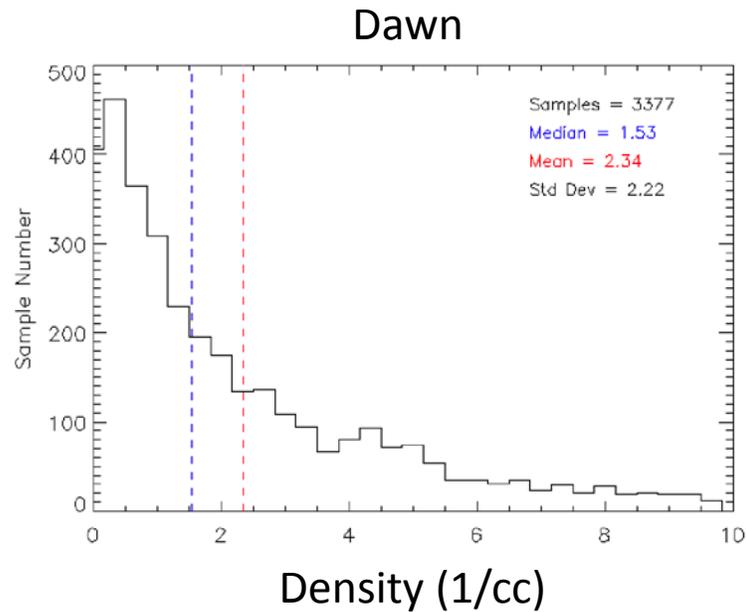
# Cold Ion Loss on Open Field Lines Exhibits a North Bias



**Data:** all upstream conditions and planet rotations

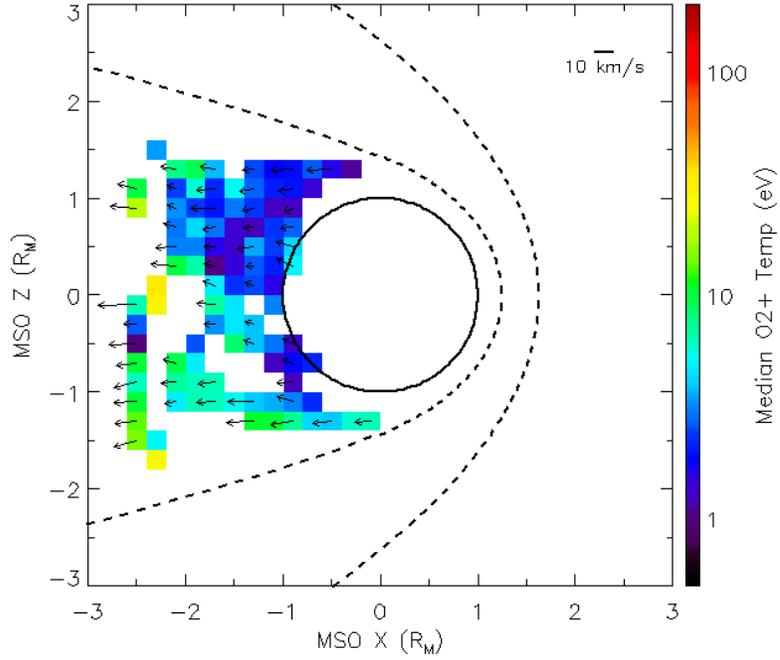
**Simulation:** ensemble of 8 multi-species, single-fluid models (C. Dong, Y. Ma, J. Luhmann)

# $O_2^+$ Night-Side Density: Dawn-Dusk Asymmetry?

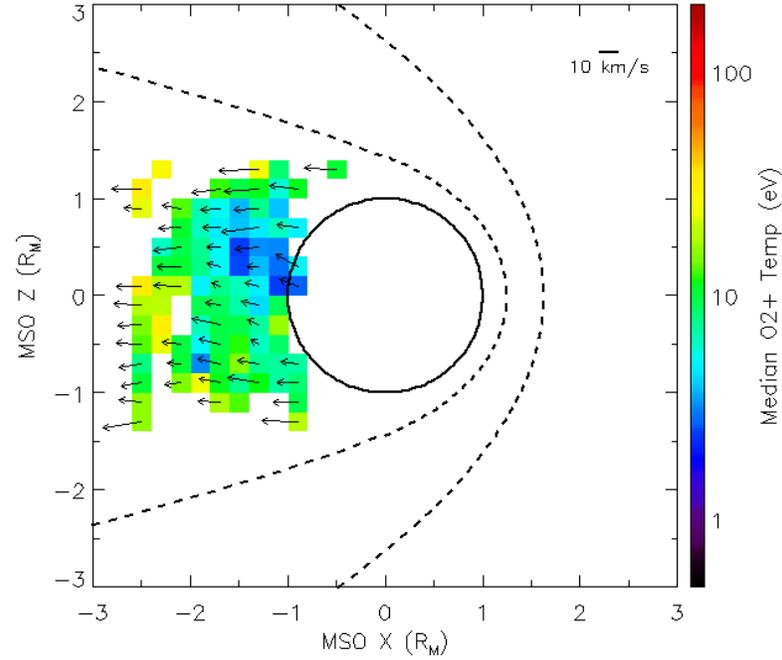


# $O_2^+$ Temperature: How Cold is Cold?

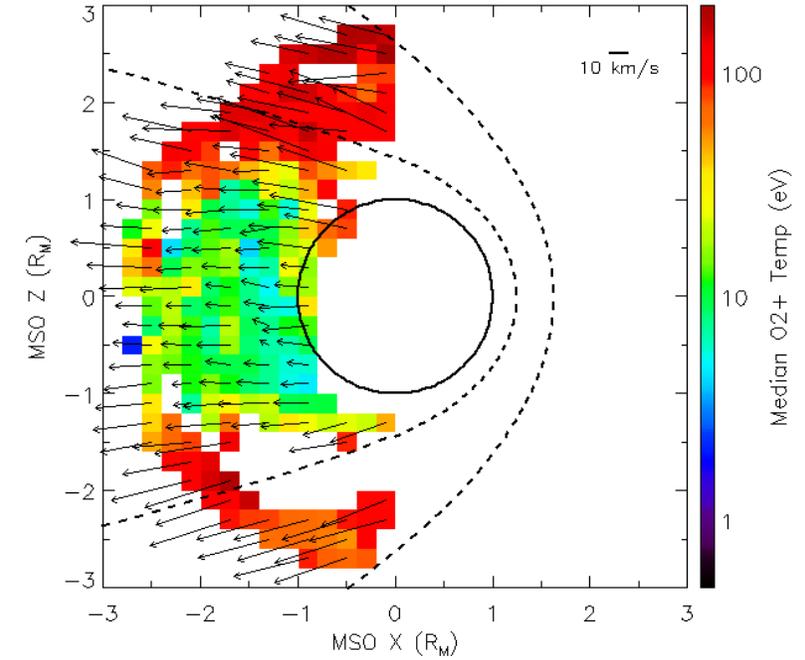
Open to Day



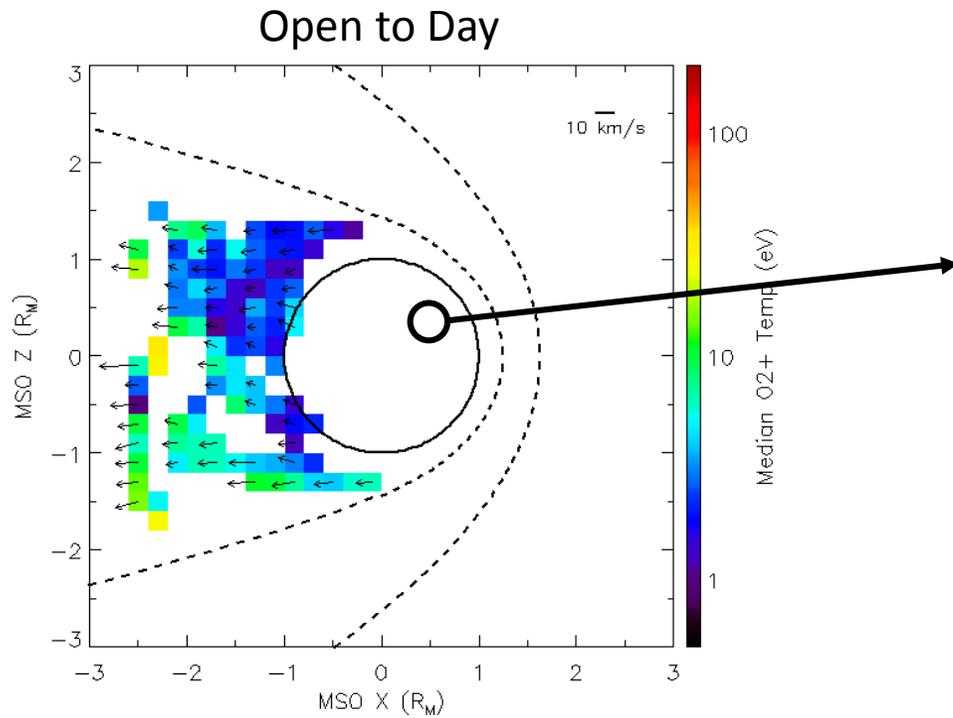
Open to Night



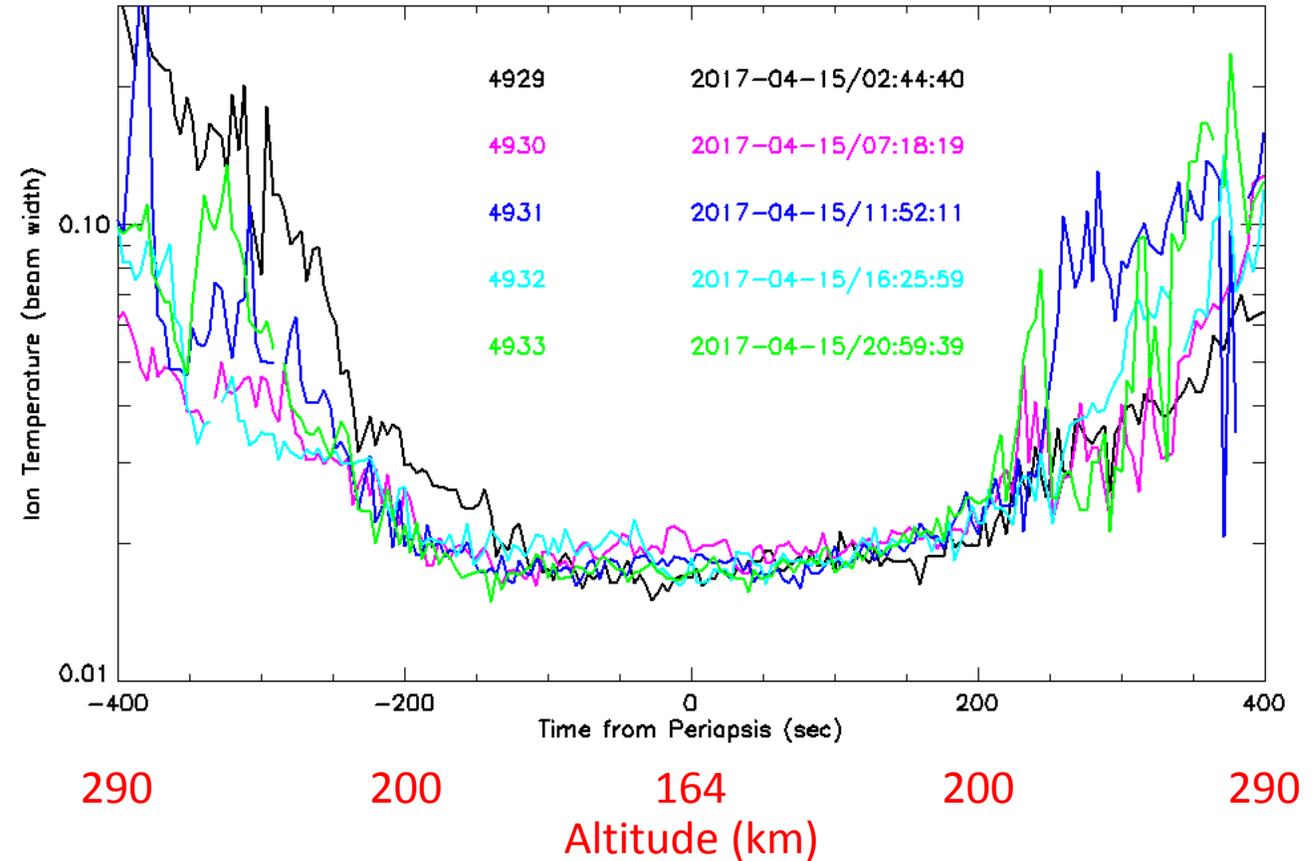
Draped



# O<sub>2</sub><sup>+</sup> Temperature: How Cold is Cold?

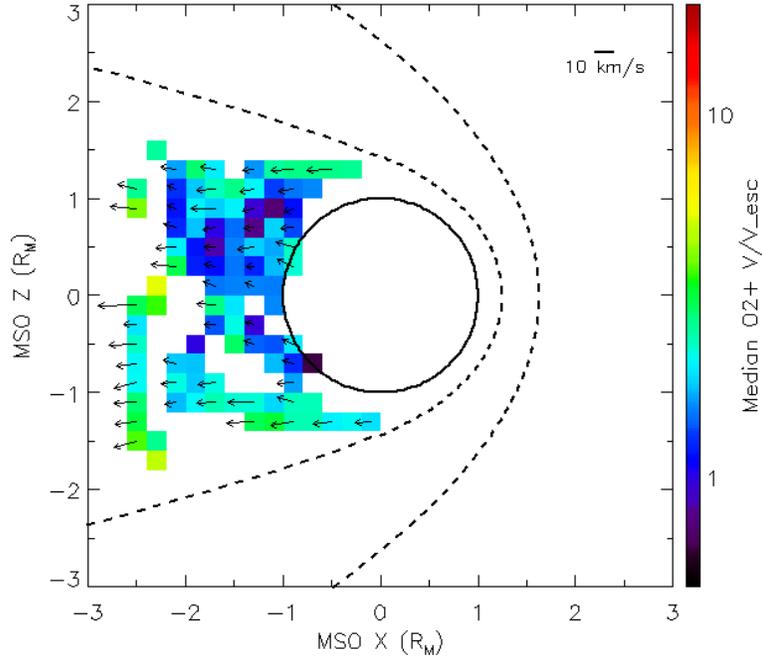


Ion Temperature around Periapsis

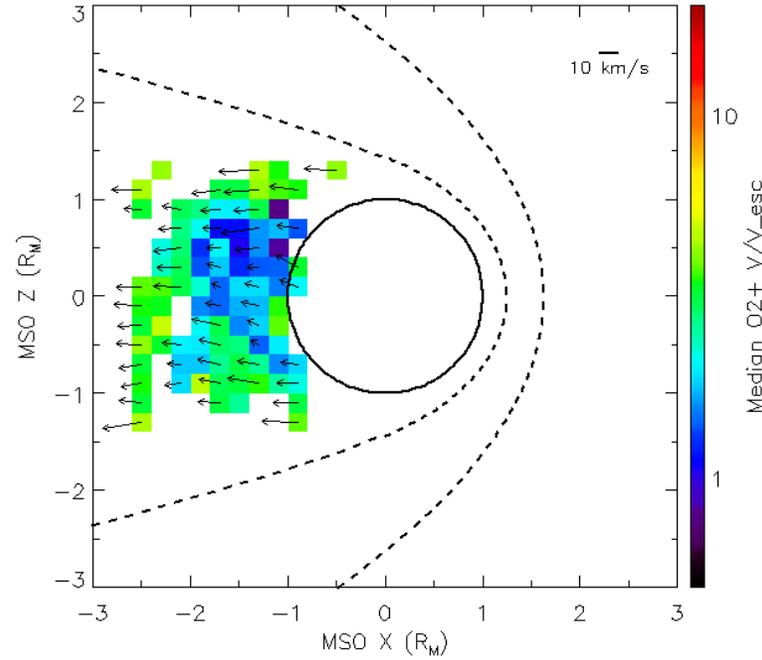


# $O_2^+ V_{\text{BULK}}/V_{\text{ESC}}$ vs. Topology

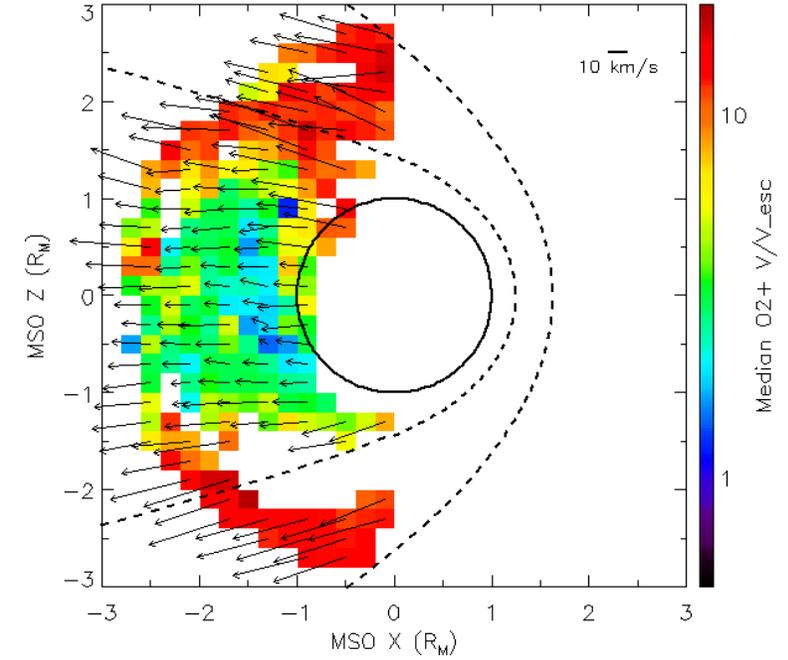
Open to Day



Open to Night



Draped



Statistics for  $R < 1.5 R_M$  and  $-3 < X < 0$  (all upstream conditions) : Topology Distribution

Topology	Number of Measurements	O <sub>2</sub> <sup>+</sup> loss (s <sup>-1</sup> )
Open	18484	$1.9 \times 10^{24}$
Draped	31732	$2.4 \times 10^{24}$
<b>All</b>	<b>50216</b>	<b><math>4.3 \times 10^{24}</math></b>

Topology	Number of Measurements	O <sup>+</sup> loss (s <sup>-1</sup> )
Open	17587	$1.0 \times 10^{24}$
Draped	27951	$1.4 \times 10^{24}$
<b>All</b>	<b>45538</b>	<b><math>2.4 \times 10^{24}</math></b>

O<sub>2</sub><sup>+</sup> / O<sup>+</sup> loss ratio

- Open : 1.9
- Draped : 1.7

Draped / Open loss ratio

- O<sub>2</sub><sup>+</sup> : 1.3
- O<sup>+</sup> : 1.4

Total Oxygen Atom Loss in Tail as Ions:

$$\text{Loss (O)} = 2 \times \text{Loss(O}_2^+) + \text{Loss(O}^+) = 1.1 \times 10^{25} \text{ s}^{-1}$$

Statistics for  $R < 1.5 R_M$  and  $-3 < X < 0$  (all upstream conditions) : Energy Distribution

Energy	Number of Measurements	Fraction of Total O loss ( $s^{-1}$ )
< 5 eV	24468	0.19
< 6 eV	28247	0.27
< 7 eV	32116	0.31
< 8 eV	35627	0.41
< 10 eV	41414	0.46
< 25 eV	64548	0.73
< 50 eV	74829	0.80
< 100 eV	81238	0.84
<b>All energies</b>	<b>96367</b>	<b>1.00</b>

<-- 50% of loss

## Key Points

- Cold ion loss occurs mainly in a cylinder about the MSO X axis with a radius of  $1.5 R_M$
- Loss is organized by magnetic topology
- Loss exhibits a north bias on open field lines
- $\sim 50\%$  of oxygen loss occurs below 10 eV
- $O_2^+$  and  $O^+$  are lost in close to a 2:1 ratio
  - occurs on both open and draped field lines
  - consistent with ionospheric source region near the exobase
- Total oxygen loss in tail:  $\sim 10^{25}$  O atoms/s