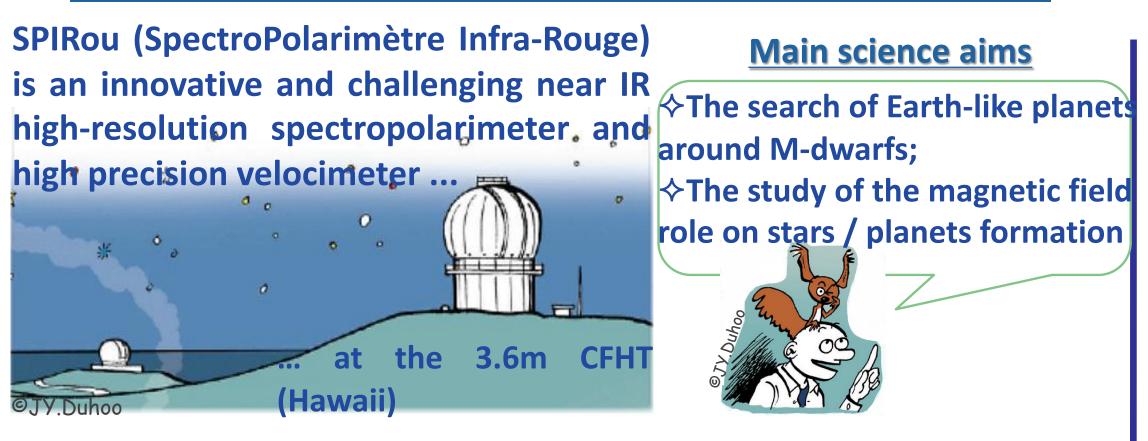


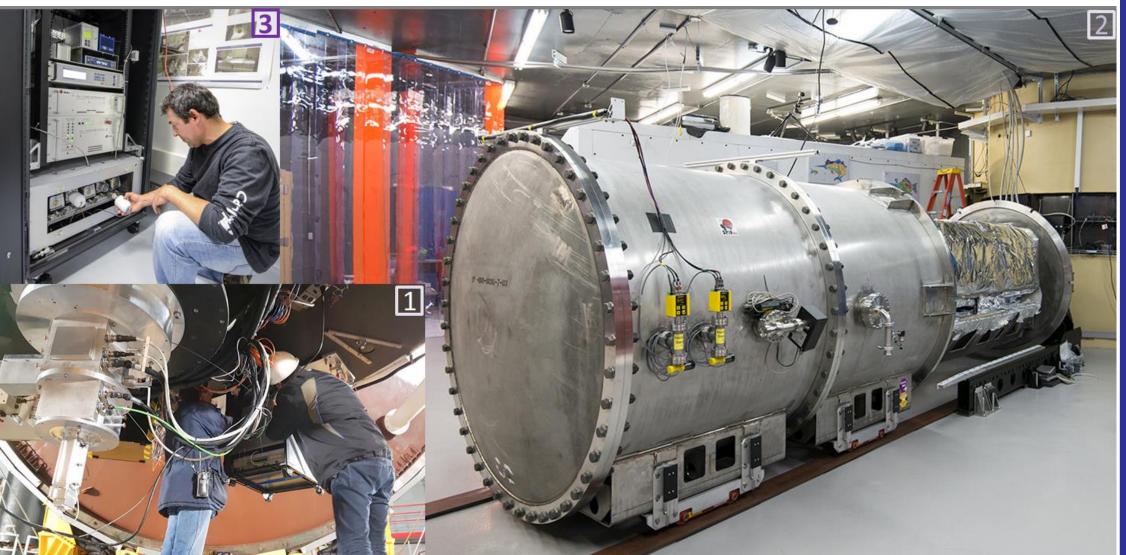
# **SPIRou meets ARIEL**

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# **HOW TO COMBINE SPIRou & ARIEL?**

# **1. HIGH RESOLUTION ATMOSPHERIC CHARACTERIZATION**

♦ SPIRou is optimal for transmission spectroscopy: large nIR domain, high spectral resolution, dry Earth atmosphere, service observing

Atmosphere composition and winds: the combination of SPIRou and ARIEL would resolve model degeneracies (Brogi & Line 2019) ♦ Simultaneous Rossiter-McLaughlin anomaly can be measured to get planet obliquities

## **2. STELLAR ACTIVITY FILTERING**

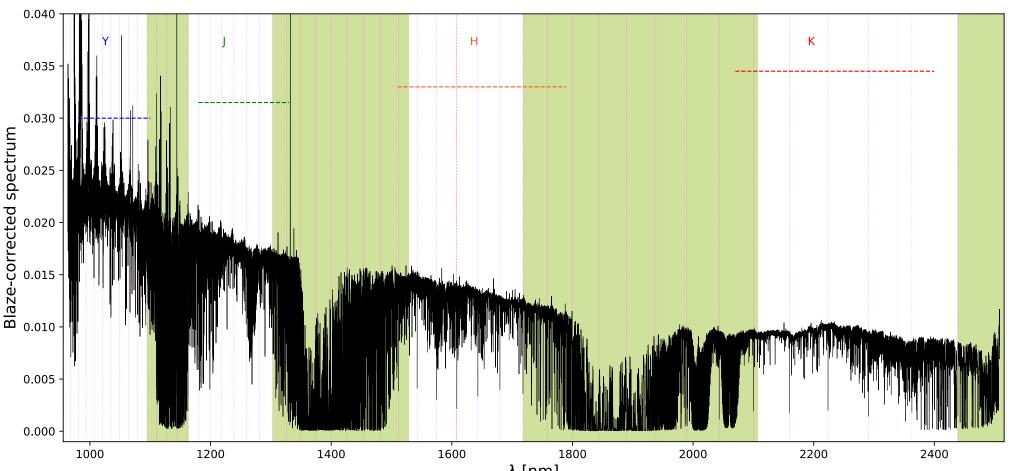
♦ SPIRou measures the circular polarization in stellar lines, producing series of Stokes V profiles (Donati et al. 97) => reconstruction of the magnetic large-scale topology of the stellar surface and distribution of brightness features

 $\diamond$  Observing in the nIR domain offers the possibility to estimate the



of the main SPIRou sub-systems: (1) the Cassegrain unit (2) the spectrograph enclosed in its cryostat; (3) the Calibration module.

Spectral range: 0.96 – 2.48  $\mu$ m in a single exposure, no gaps, **YJH and K bands** 49 orders

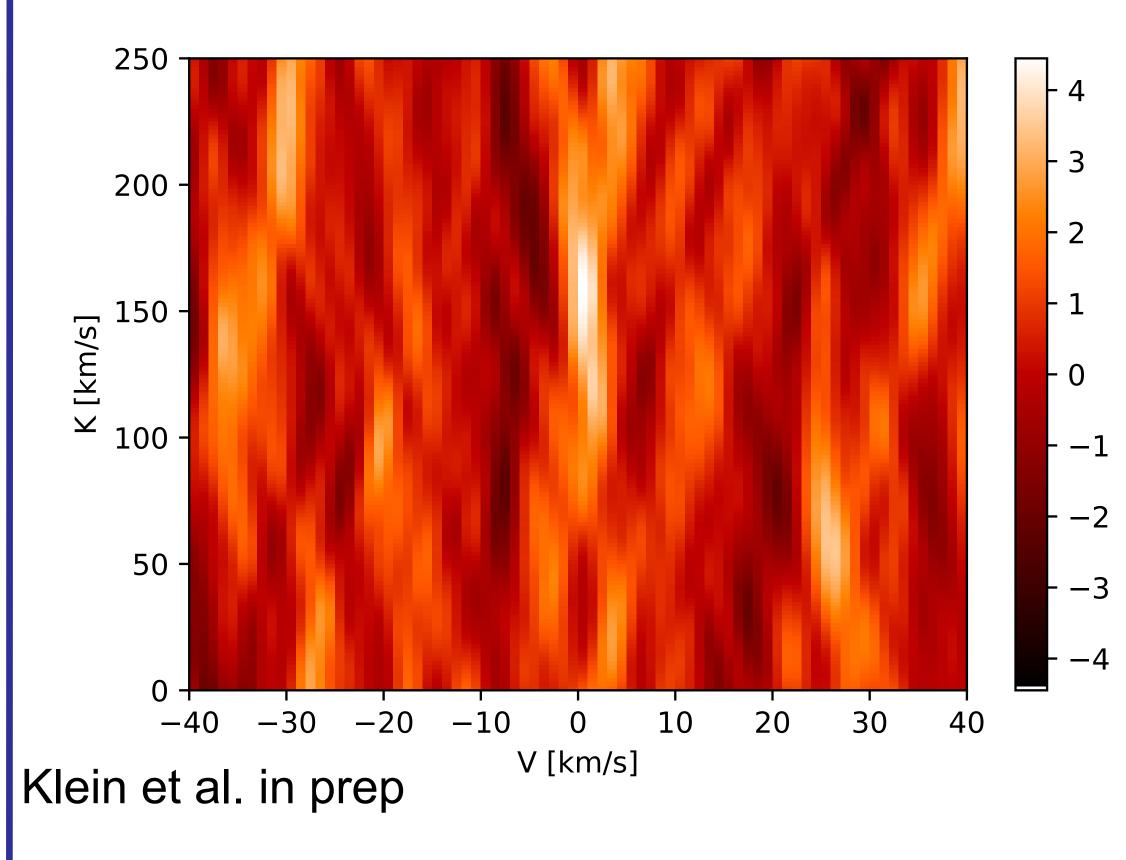


 $\Rightarrow$  Planet's rotation period can be derived (Snellen et al 2014)

♦ Phase curves and eclipse observations: temperature, albedo, atmospheric circulation

**IN PREPARATION AND IN COMPLEMENT** TO ARIEL, SPIROU IS ABLE TO ENLARGE **SIGNIFICANTLY OUR KNOWLEDGE OF** PLANETARY ATMOSPHERES

**Example: cross correlation of high resolution** simulated spectrum of HD 189733b with water lines **Expected 4 sigma detection of water** 



small-scale magnetic field concentrated in spots as well

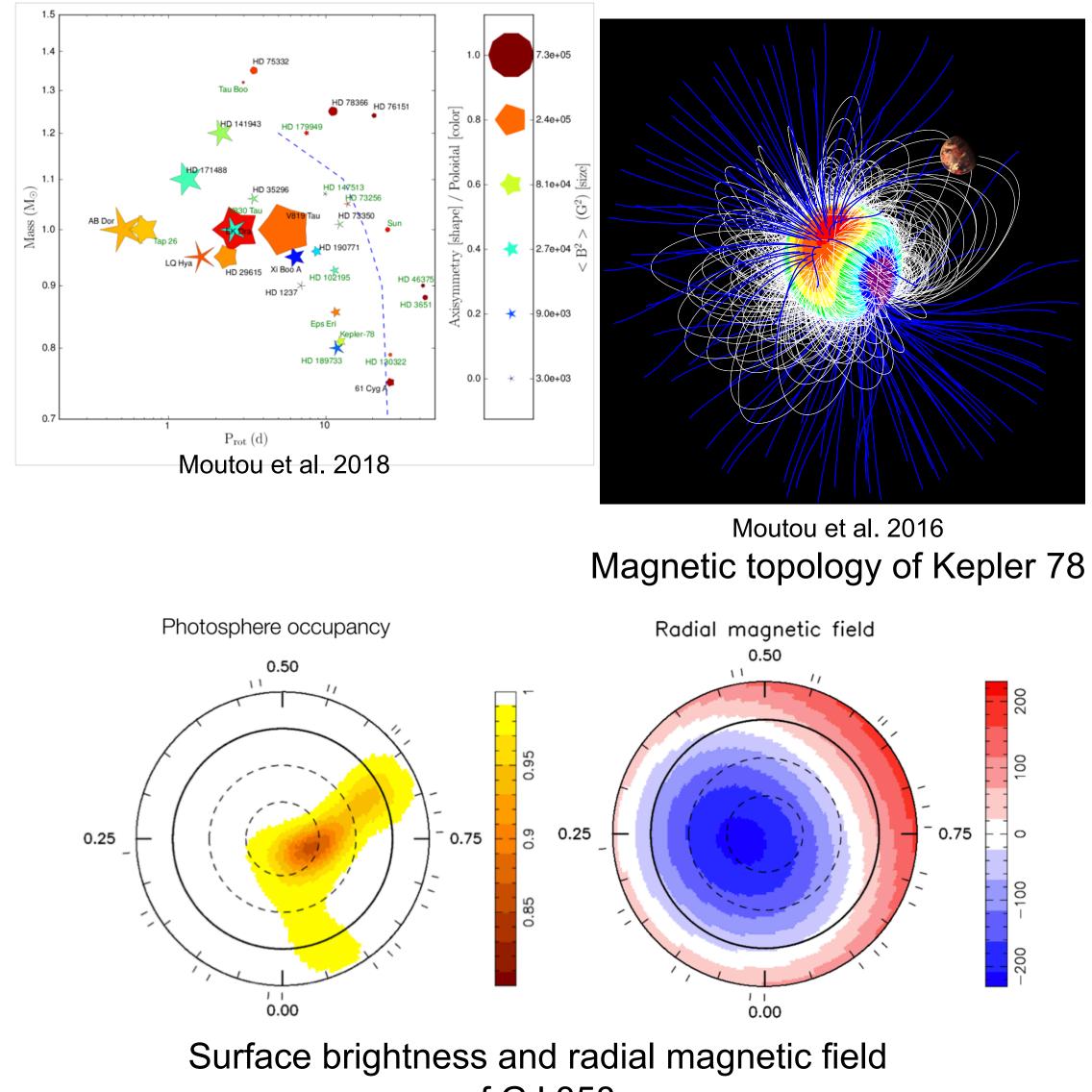
♦ Magnetic field characterization is critical to evaluate stellar « noise »: select best targets for ARIEL, observe simultaneously to correct for stellar variability (Zhang et al 2018)

 $\diamond$  Observations in spectropolarimetry should be simultaneous to planetary characterization measurement for a proper modelling (Fares et al. 2018)

♦ Star-planet interactions can be estimated, and planetary magnetic field (Cauley et al. 2019) - which can be critical for habitability (Vidotto et al. 2013)

**POLARIMETRIC OBSERVATIONS WOULD GREATLY IMPROVE THE ATMOSPHERIC CHARACTERIZATION FROM ARIEL AND HELP SELECT BEST TARGETS** 

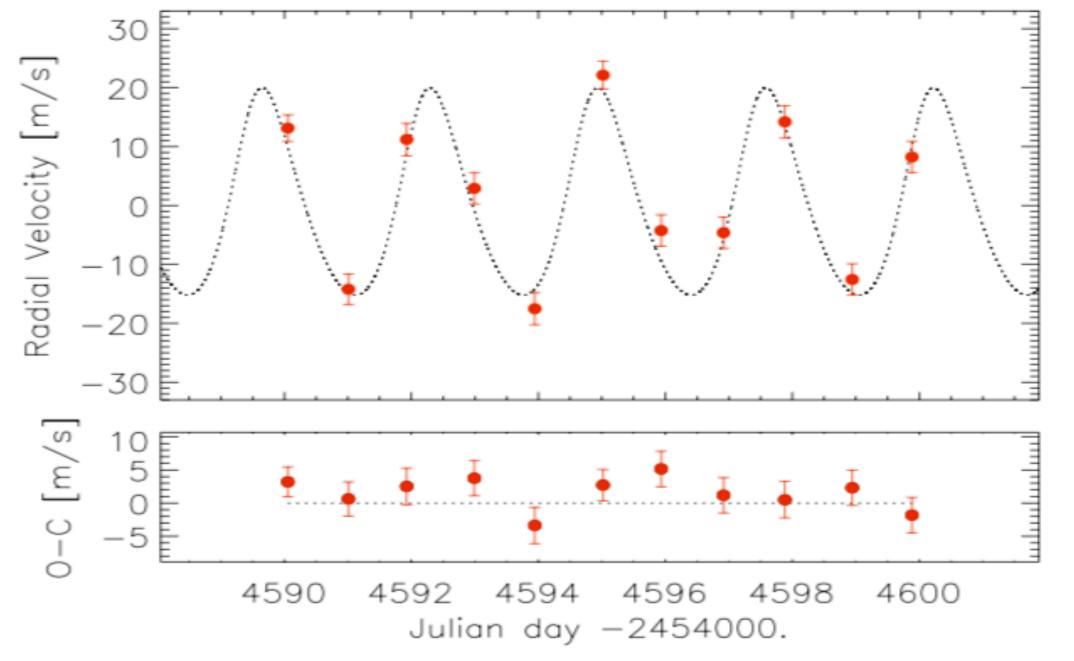
> Example: magnetic topologies of planet hosting stars



### Example of a reduced blaze-corrected Stokes I spectrum obtained for HD189733 in

September 2018. The beginning of each diffraction order is indicated through a magenta vertical dotted line and the Y, J, H and K bands are shown as the horizontal dashed lines on the top. The green vertical bands indicate the spectral ranges dominated by tellurics

> Spectral resolution: 70±5K Radial velocity precision: 2 m.s<sup>-1</sup>



Radial velocity curve of GI436 obtained with SPIRou, confirming the presence of the planet GI436b. Residuals of 2.5 m/s

## **CONCLUSIONS**

 $\diamond$  High resolution spectropolarimeter installed at CFHT  $\diamond$  Current radial velocity precision : 2m/s

## **Combination with ARIEL :**

 $\diamond$  Higher resolution spectrum in the near infrared ♦ Possibility to filter out stellar activity

## **Complement :**

of GJ 358 Hebrard et al. 2016

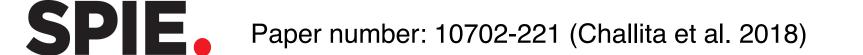
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#### Pictures Copyright: ©S. Chastanet – CNRS/OMP <sup>†</sup>SPIRou is an international project led by France (IRAP/OMP) and involving the CFHT,

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