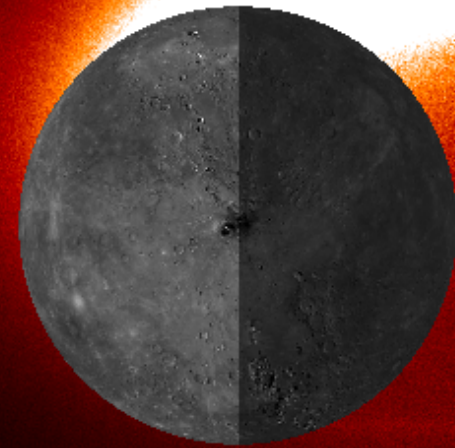
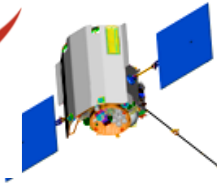


Mercury's Calcium Exosphere in the MESSENGER Era



Matthew Burger, Rosemary Killen, Bill McClintock, Aimee Merkel, Ron Vervack, Jr., Tim Cassidy, Menelaos Sarantos



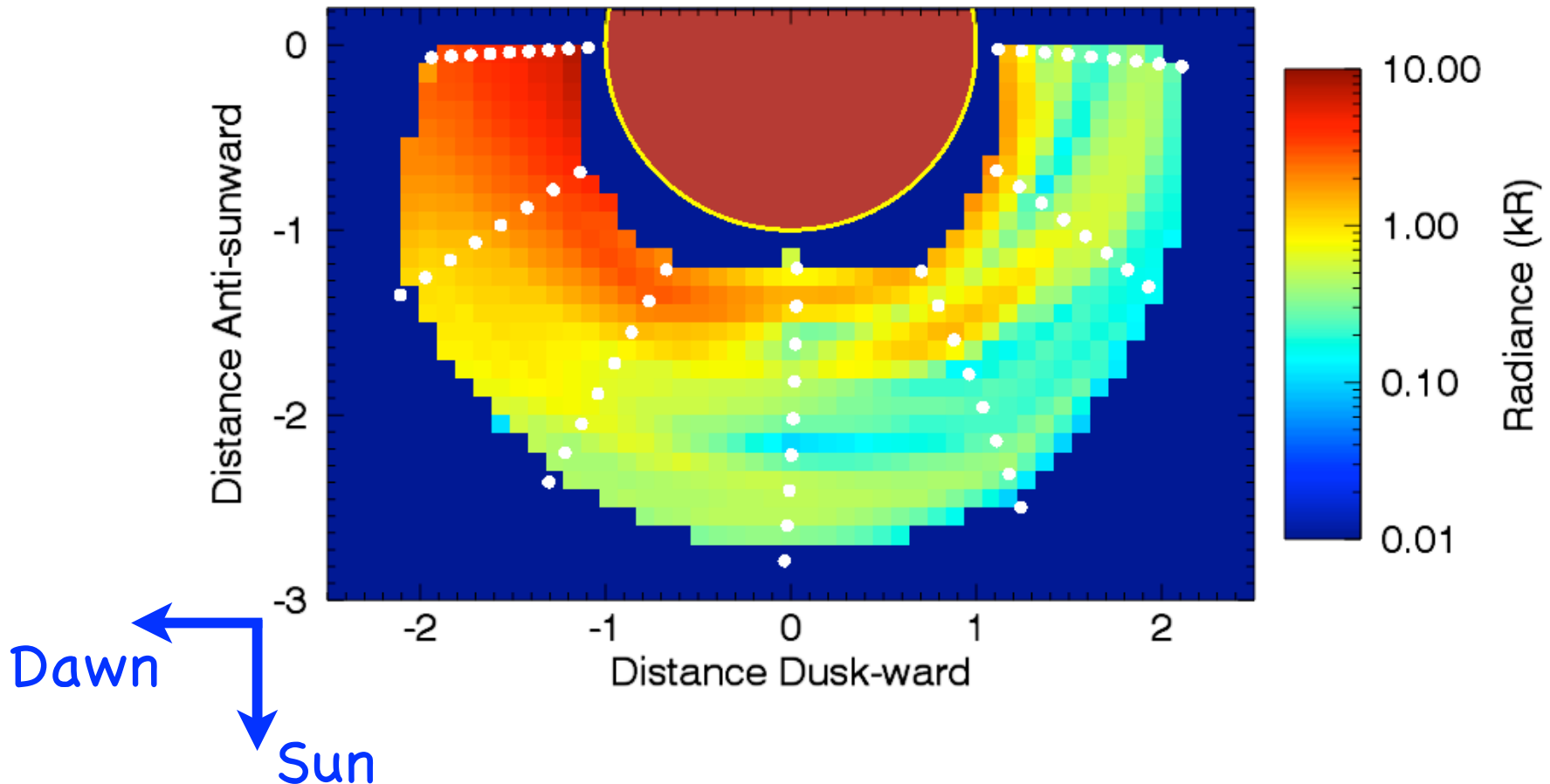
MESSENGER

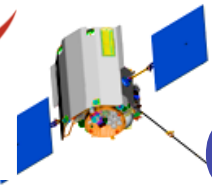
Dawn Ca Enhancement



- Dayside limb scans show persistent dawn enhancement (Burger et al. 2012, 2014)

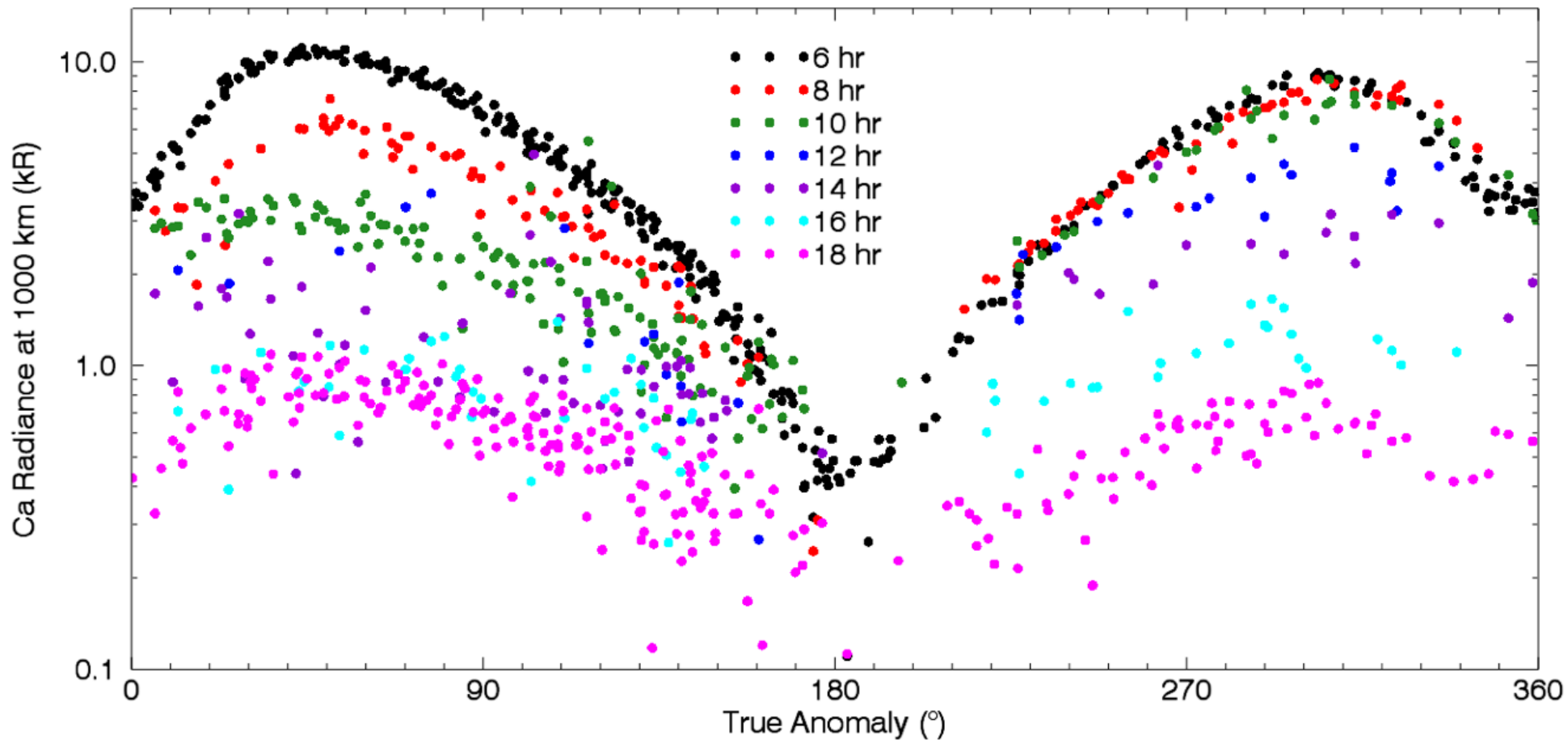
Orbit 36, 2011 APR 05, TAA = 105°

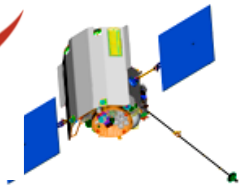




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Ca Radiance at 1000 km





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Dawn Calcium Source

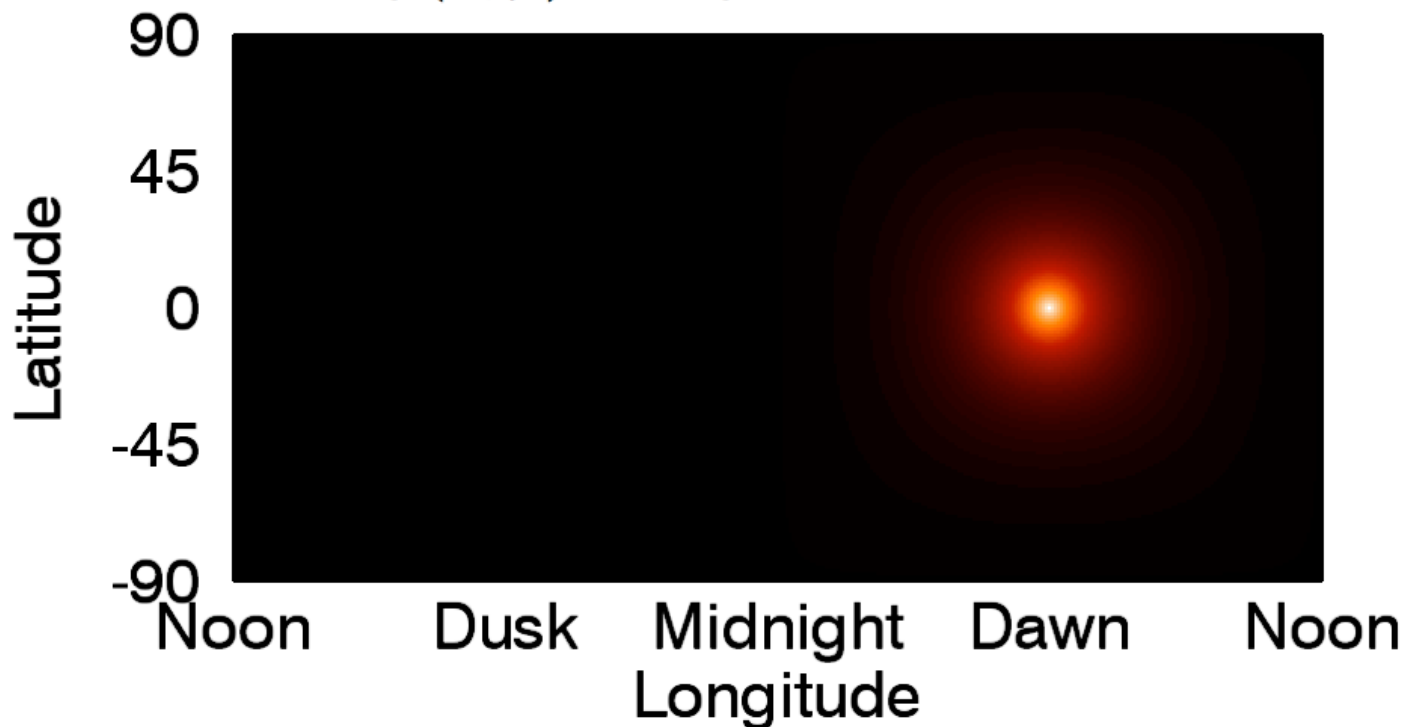


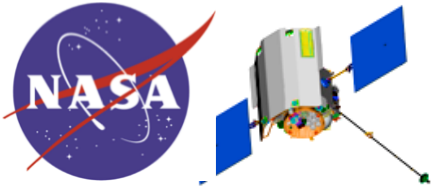
Source is centered at dawn equatorial point and drops off exponentially with width = σ

$$\mathbf{r} = (x, y, z) = (\cos \lambda \cos \mu, \sin \lambda \cos \mu, \sin \mu)$$

$$\cos \phi = \mathbf{r} \cdot \mathbf{r}_0$$

$$f(\lambda, \mu) = f_0 e^{-\phi(\lambda, \mu)/\sigma}$$



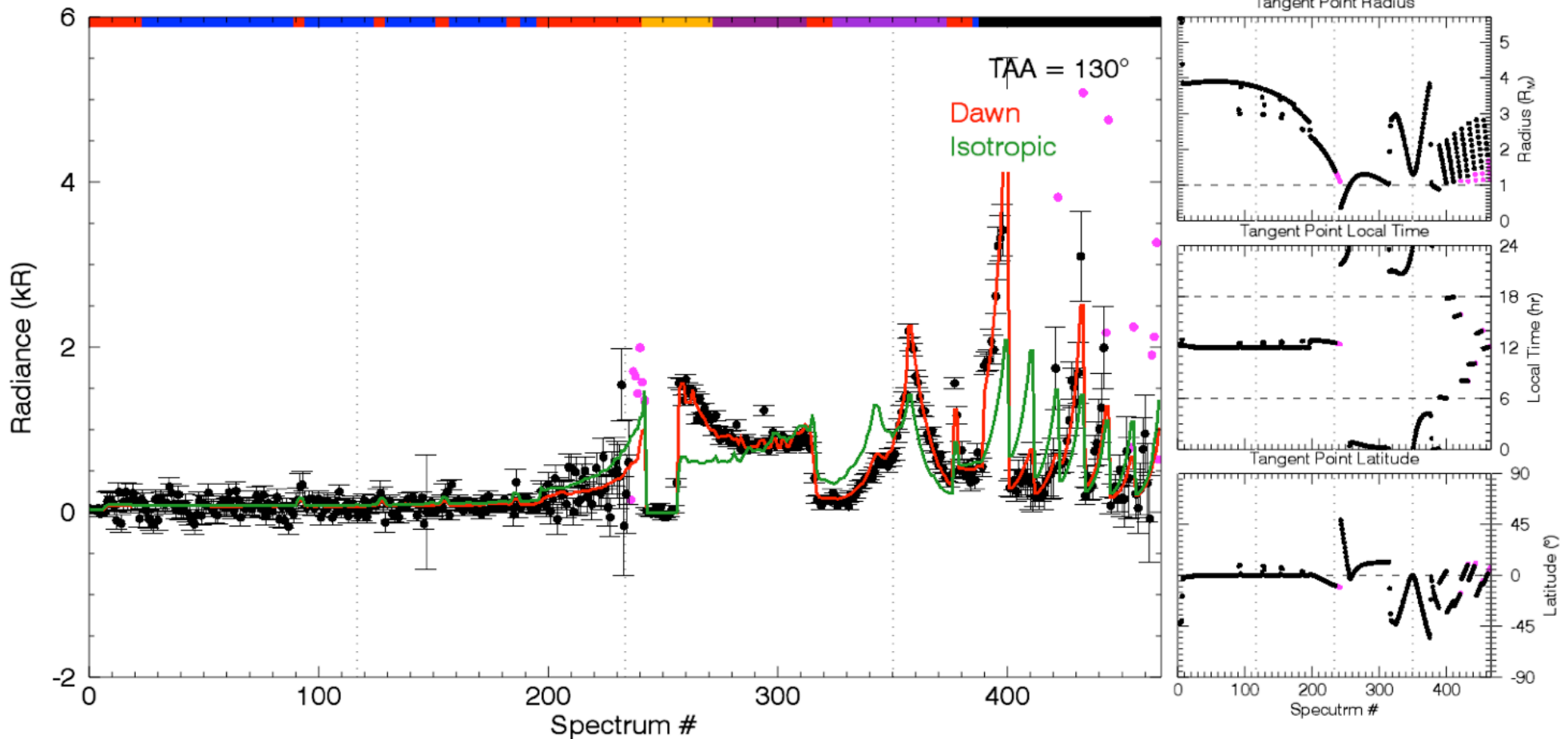


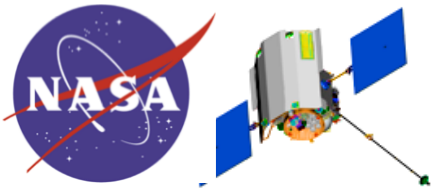
MESSENGER Nominal Model



- Our nominal Ca model uses a Maxwellian source distribution with $T=70,000$ K, $\sigma=50^\circ$, centered on the

Orbit0050, Temp=70000.0 K

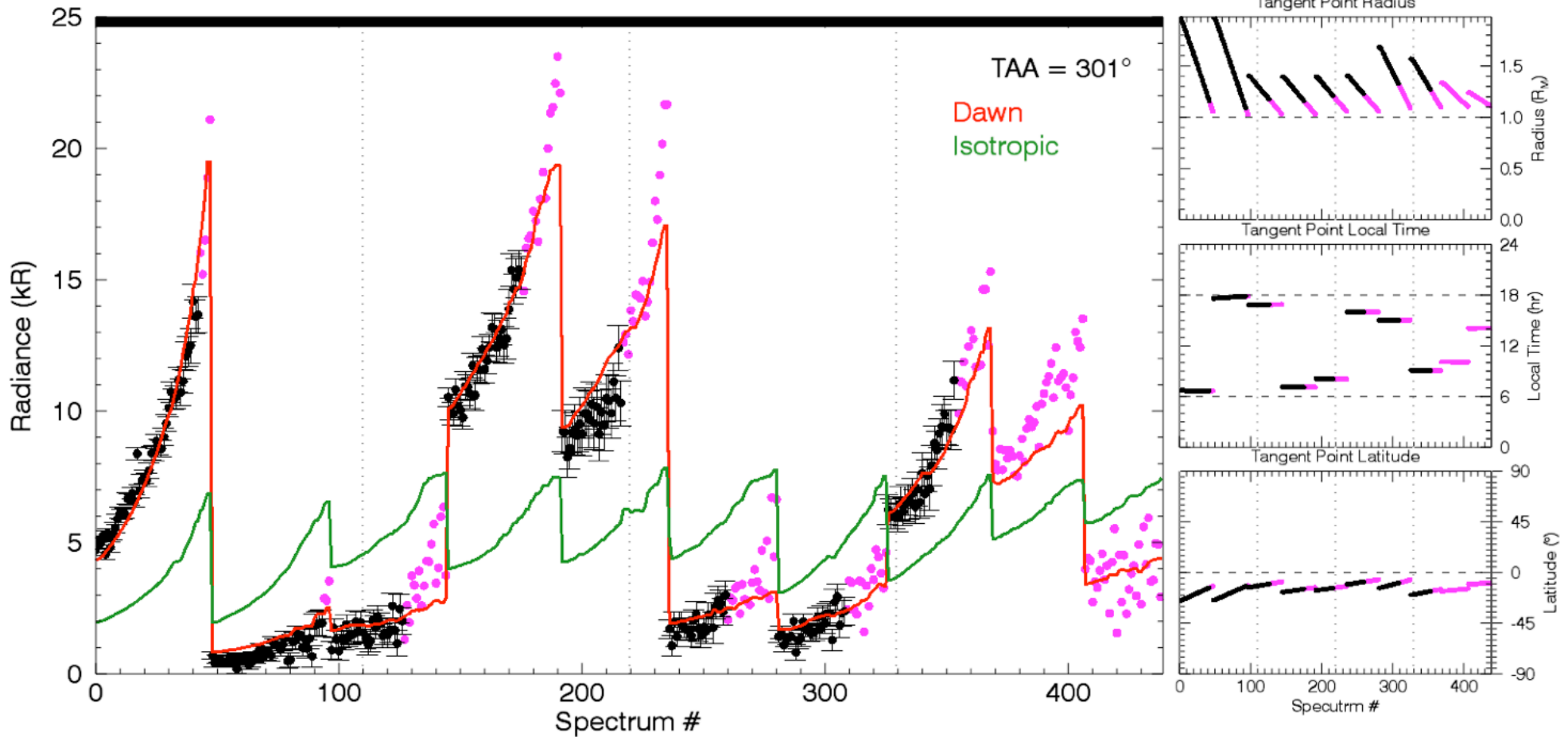


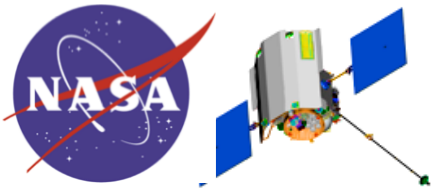


MESSENGER Nominal Model



Orbit0896, Temp=70000.0 K

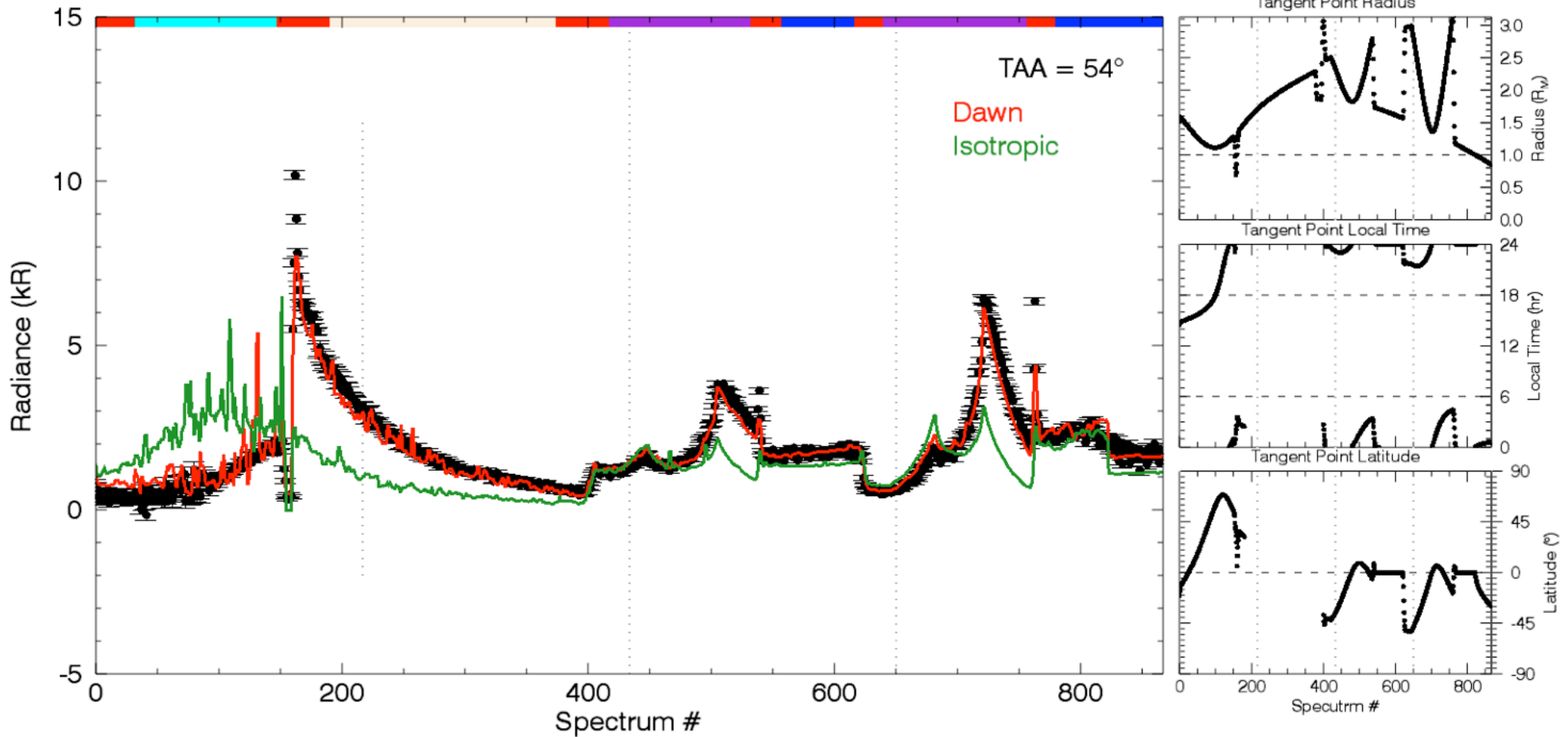


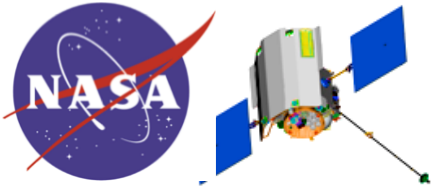


MESSENGER Nominal Model



Orbit2008, Temp=70000.0 K

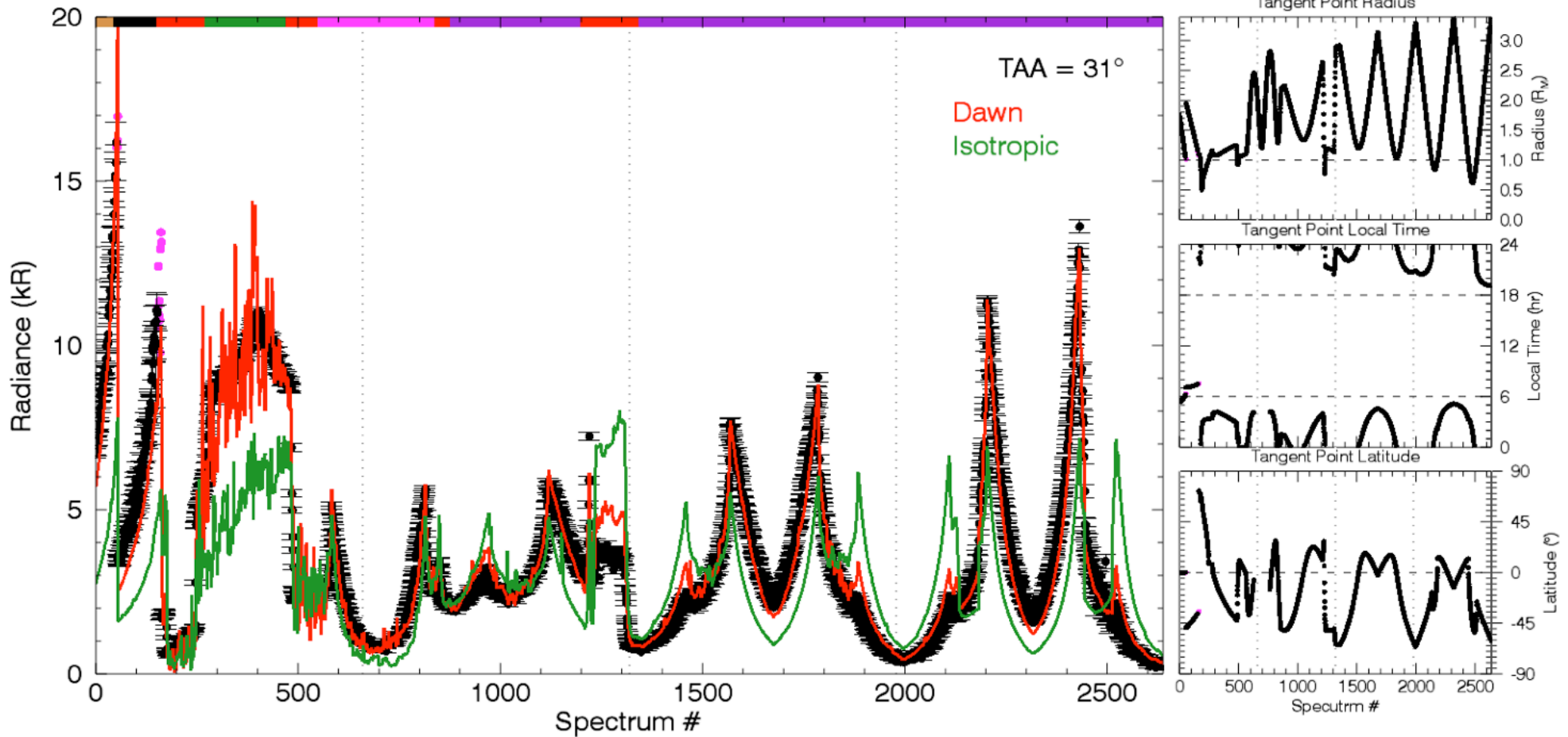


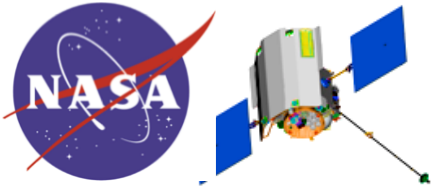


MESSENGER Nominal Model



Orbit3051, Temp=70000.0 K





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Ca Source Rate

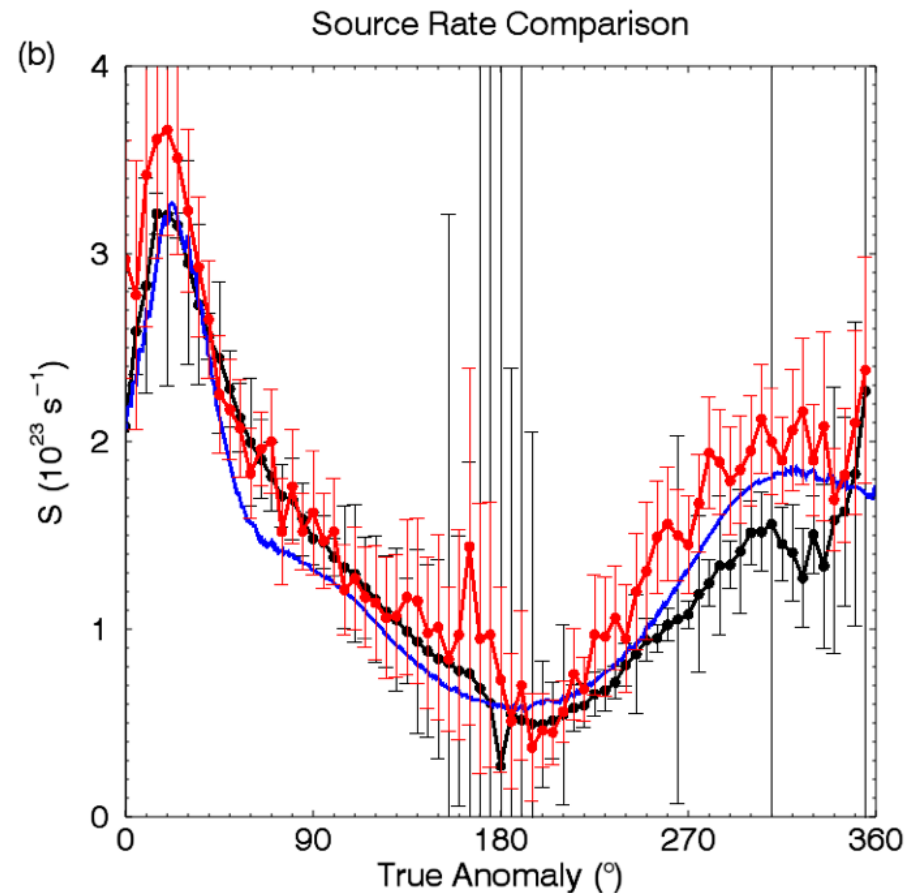
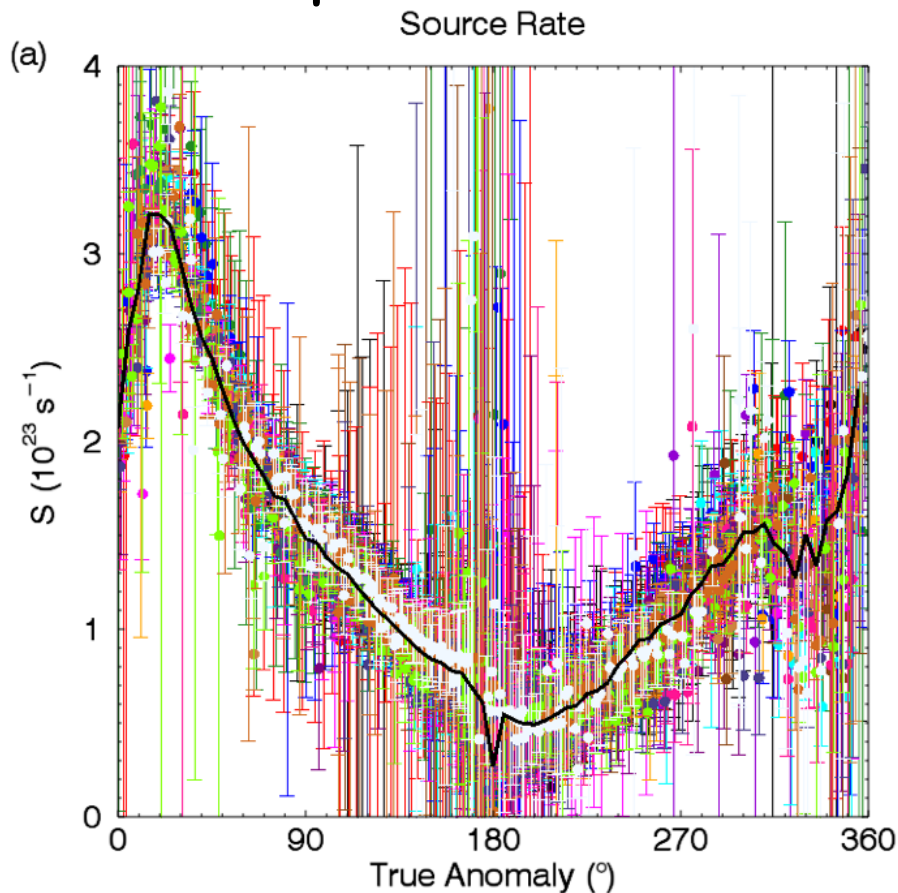


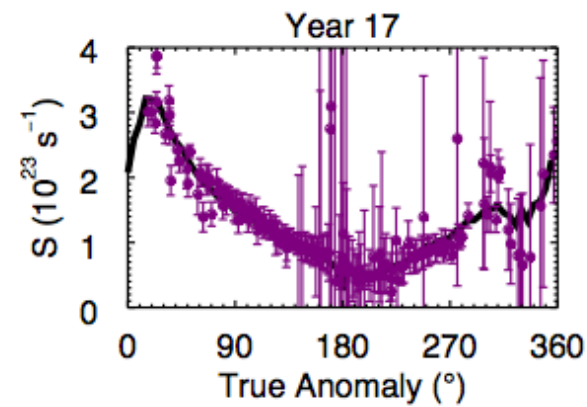
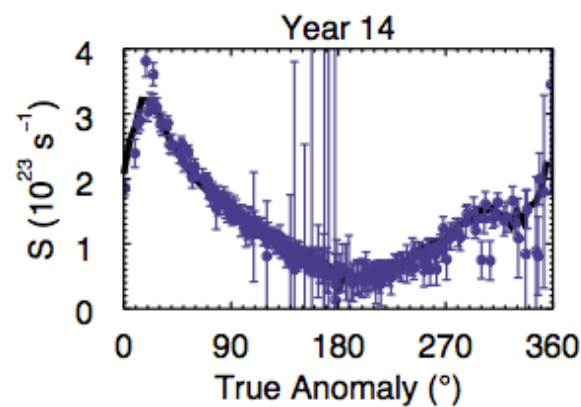
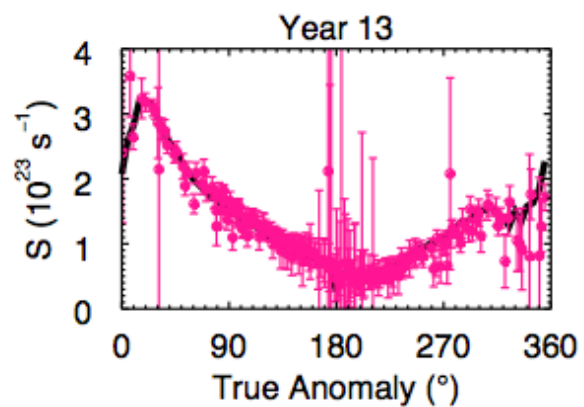
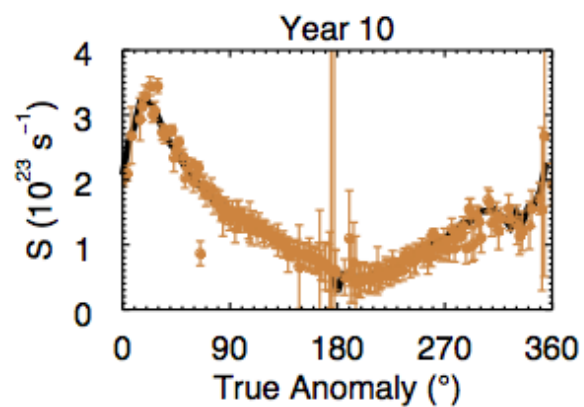
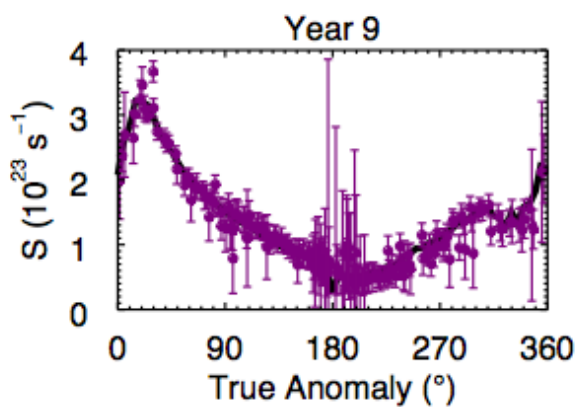
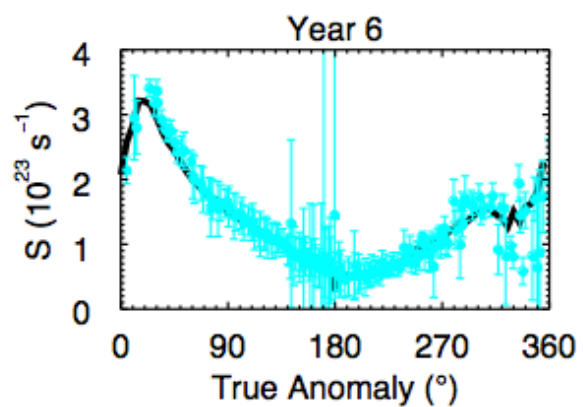
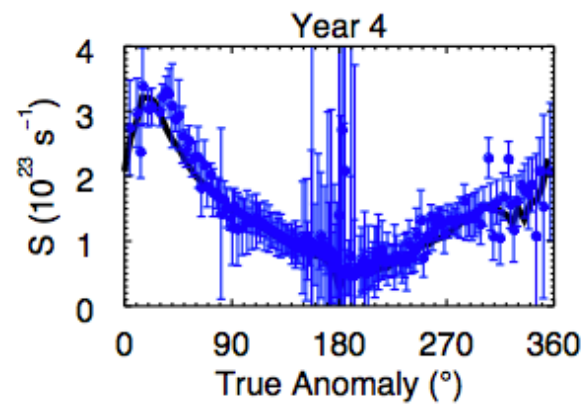
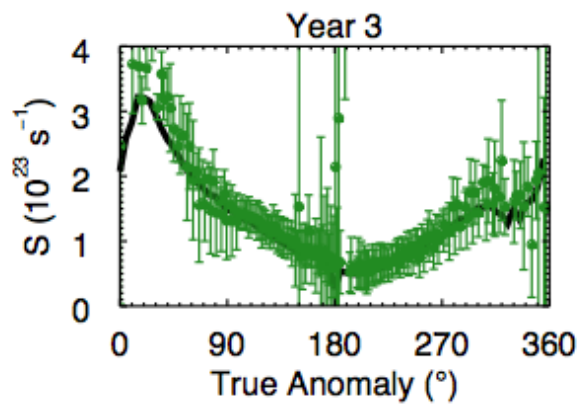
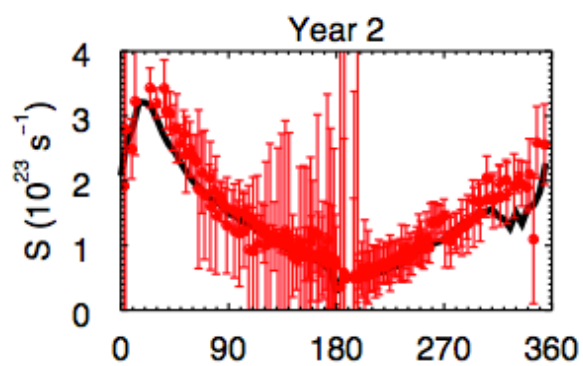
- Strong seasonal variation in source rate
- Year-to-year variability is small
 - 89% of points w/in 1σ of the black line
 - 97% of points w/in 2σ of the black line

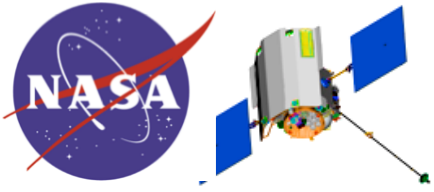
All data

Burger et al. (2014)

Killen & Hahn (2015)

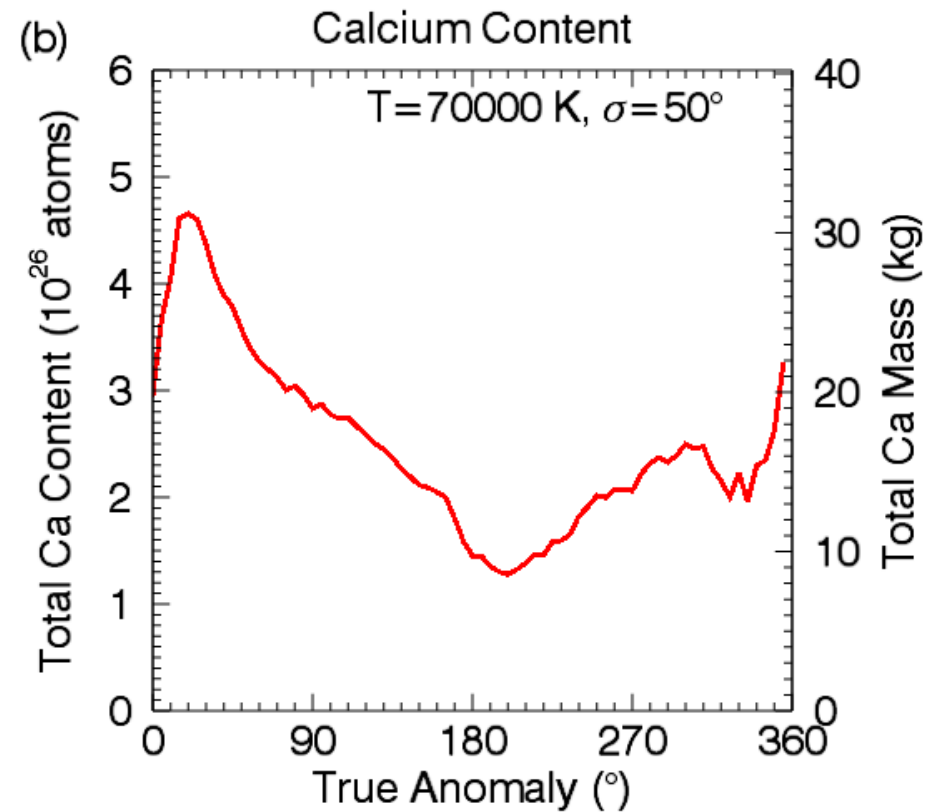
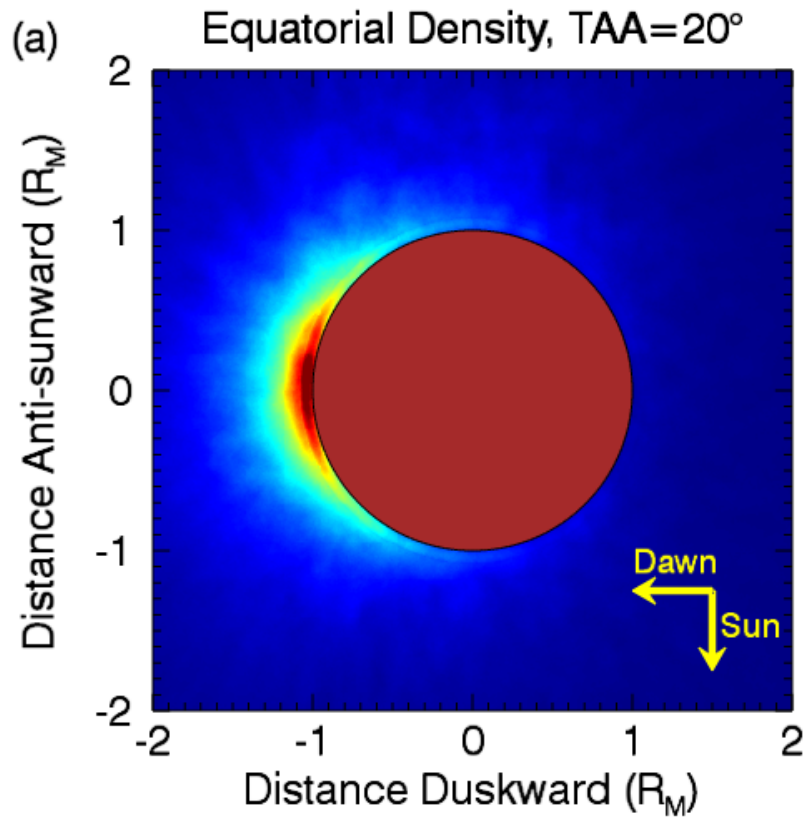


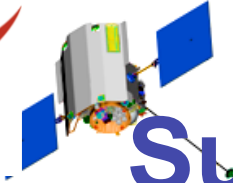




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8.5-31 kg Ca



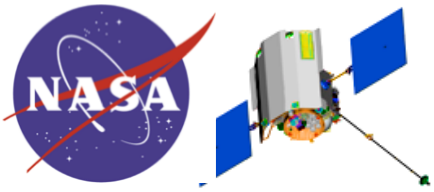


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Summary of Observations



- The calcium source is concentrated on the dawn hemisphere
 - may move around a little bit, but not much
- Comes off very hot ($T > 50,000$ K)
 - Nominal model has $T = 70,000$ K
- Source size and temperature don't change much if at all
- Source strength varies with Mercury's true anomaly (position in orbit)
 - Source strongest just after perihelion
 - Source weakest just after aphelion

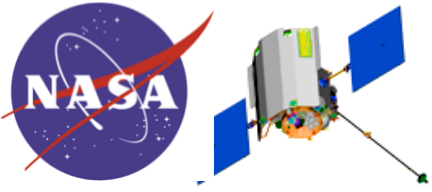


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What is the Source?



- **Not related to the surface geology**
 - **Source is (approximately) fixed in local time and does not rotate with Mercury**
- **Not related to the magnetosphere (ion sputtering or electron stimulated desorption)**
 - **Magnetosphere is highly variable**
 - **Wouldn't produce a source at dawn**
 - **There are possible sporadic sources that may be associated with the magnetosphere (still trying to understand these)**
- **Not related to Ca freezing on the nightside and vaporizing as it moves into sunlight**

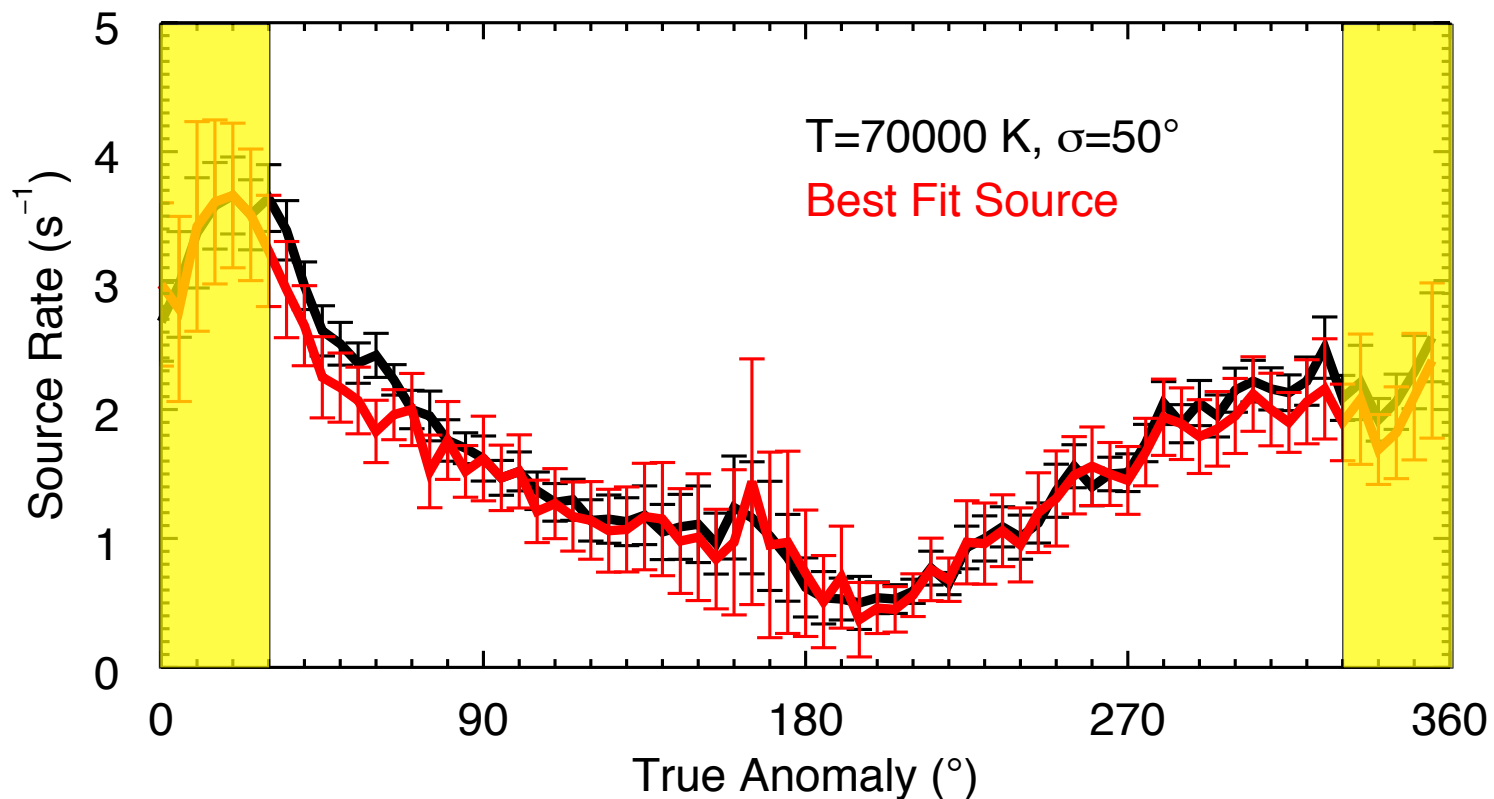


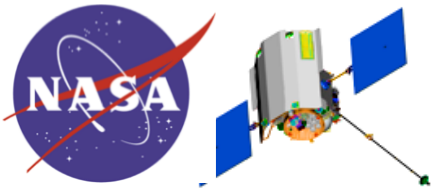
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Sun Reverses in Sky



- The motion of the Sun in Mercury's sky reverses near perihelion, where the source is strongest





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Current Hypothesis



- **Micrometeoroid impact vaporization and molecular dissociation**
 - Interplanetary Dust bombards the surface producing vapor plumes containing CaO and Ca(OH)_2 at ~ 5000 K (Berezhnoy 2013)
 - Ca-bearing molecules quickly dissociate producing hot Ca that escapes Mercury (Killen et al. 2005)
- **Pros:**
 - Models at Earth suggest dust impacts peak at dawn (Janches et al. 2006, Pifko et al. 2013)
 - CaO and Ca(OH)_2 are more likely to be produced in plumes than atomic Ca (Berezhnoy & Klumov 2008; Berezhnoy 2013)
 - Source rate consistent with dust disk + comet stream (Killen & Hahn 2015): See Rosemary's talk later
- **Cons:**
 - No evidence of a dawn/dusk asymmetry in impacts at Mercury
 - Plume chemistry is very uncertain
 - Not clear that dissociation produces the $>50,000$ K Ca the model requires