

Methodology for extracting exoplanetary atmospheres using high-resolution transmission spectroscopy

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Outline:

- ◆ Telluric correction
- ◆ Differential extinction correction
- ◆ Stellar master spectrum
- ◆ Extraction of planetary spectrum
- ◆ Search for atomic/molecular features
 - Direct detection method
 - Cross-correlation method



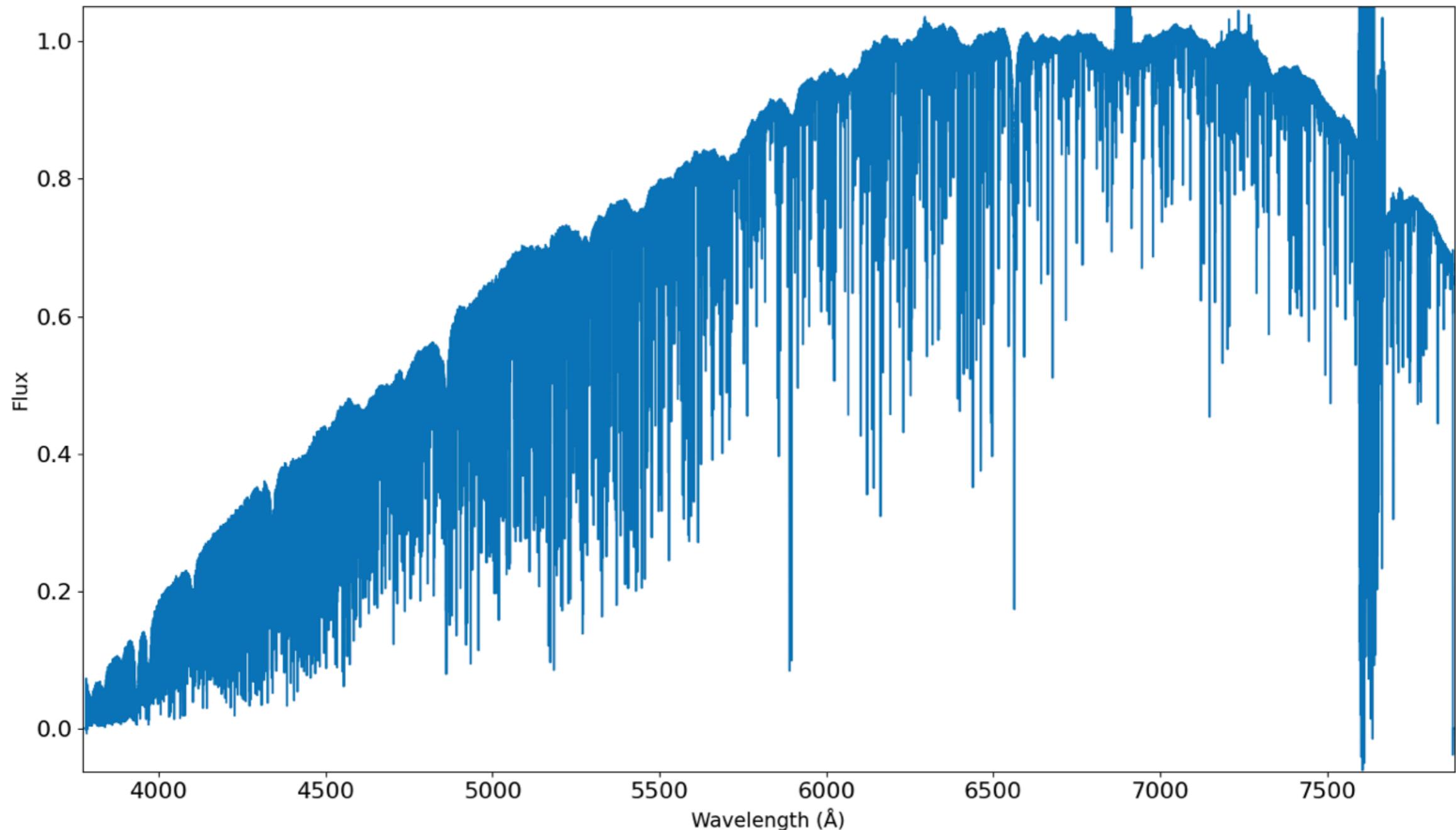
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ESPRESSO stellar spectrum



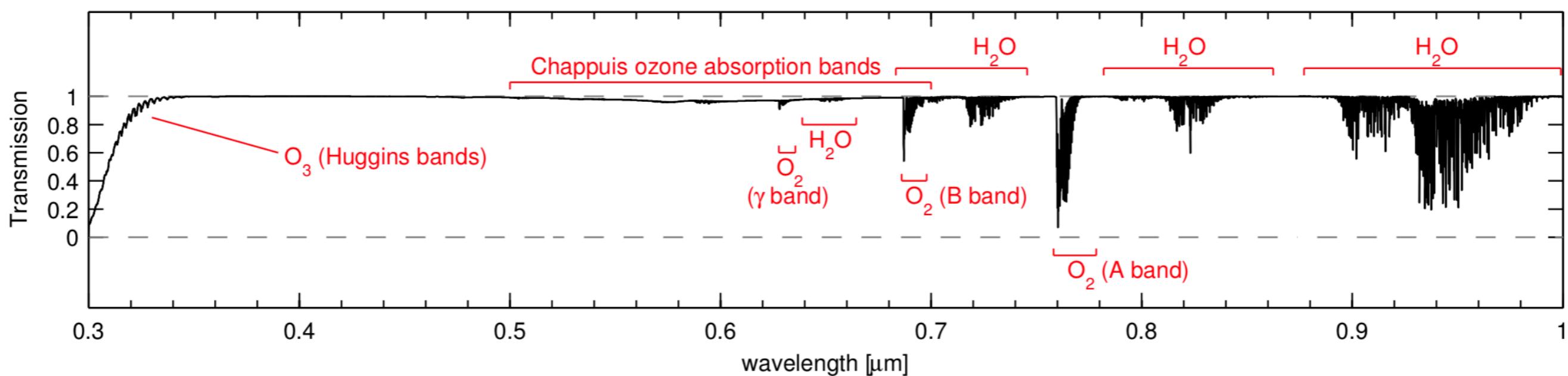
Wavelength range: 3782 Å - 7887 Å

R = 138000

Spectral type: F8

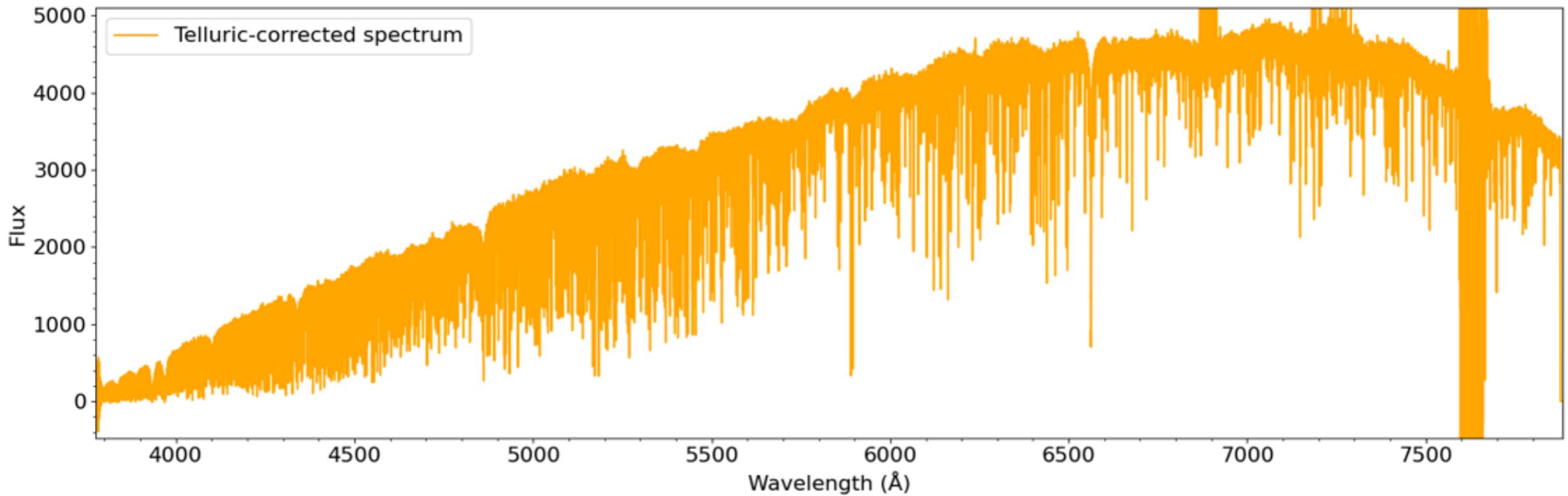
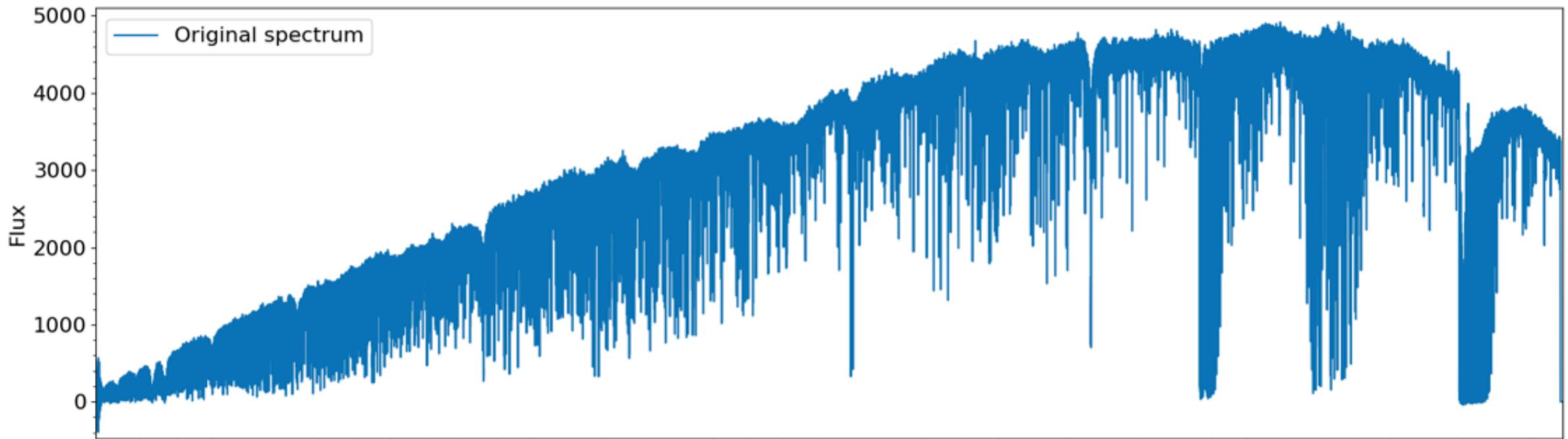
Teff = 6100 K

Telluric correction



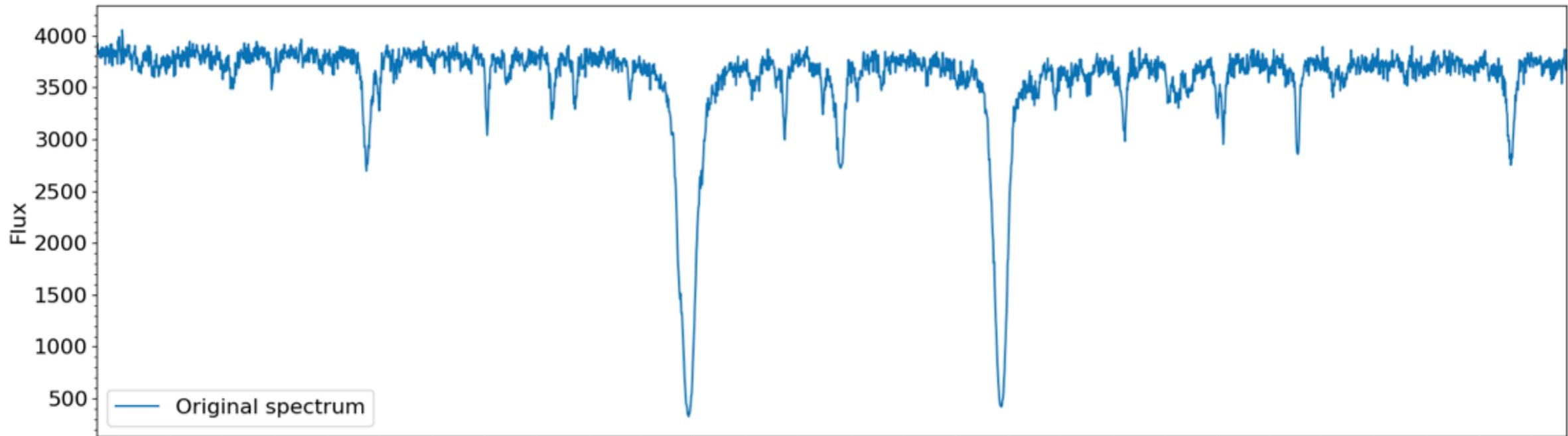
Credit: Smette et al. 2015

Telluric correction

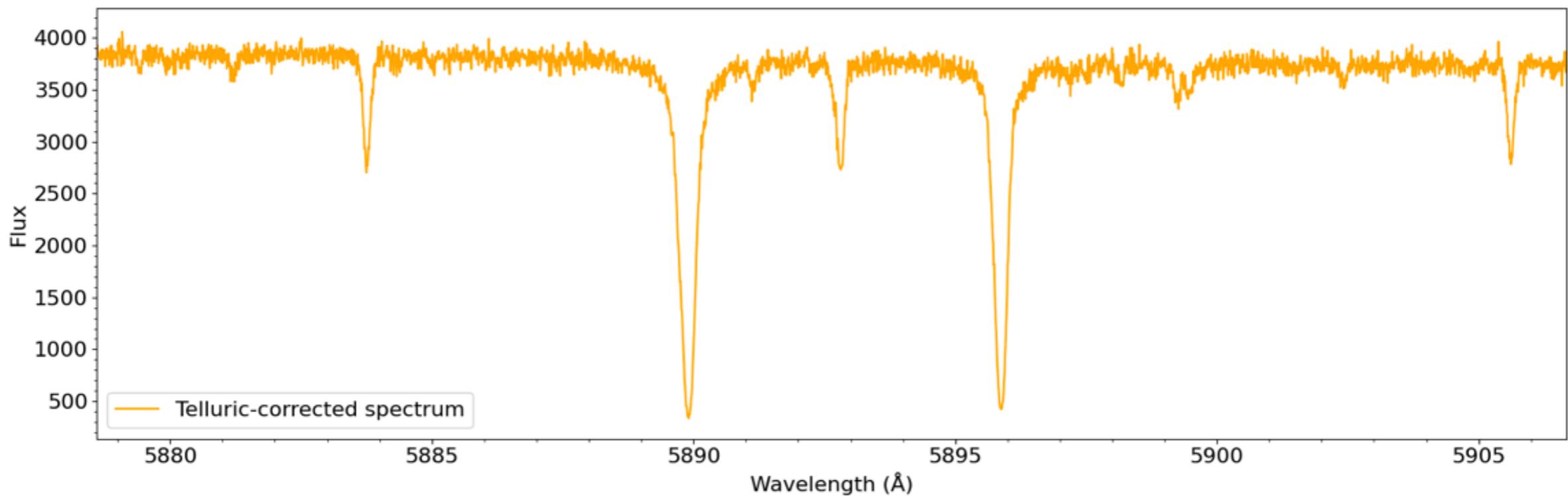


Telluric correction performed with Molecfit (Smette et al. 2015; Kausch et al. 2015)

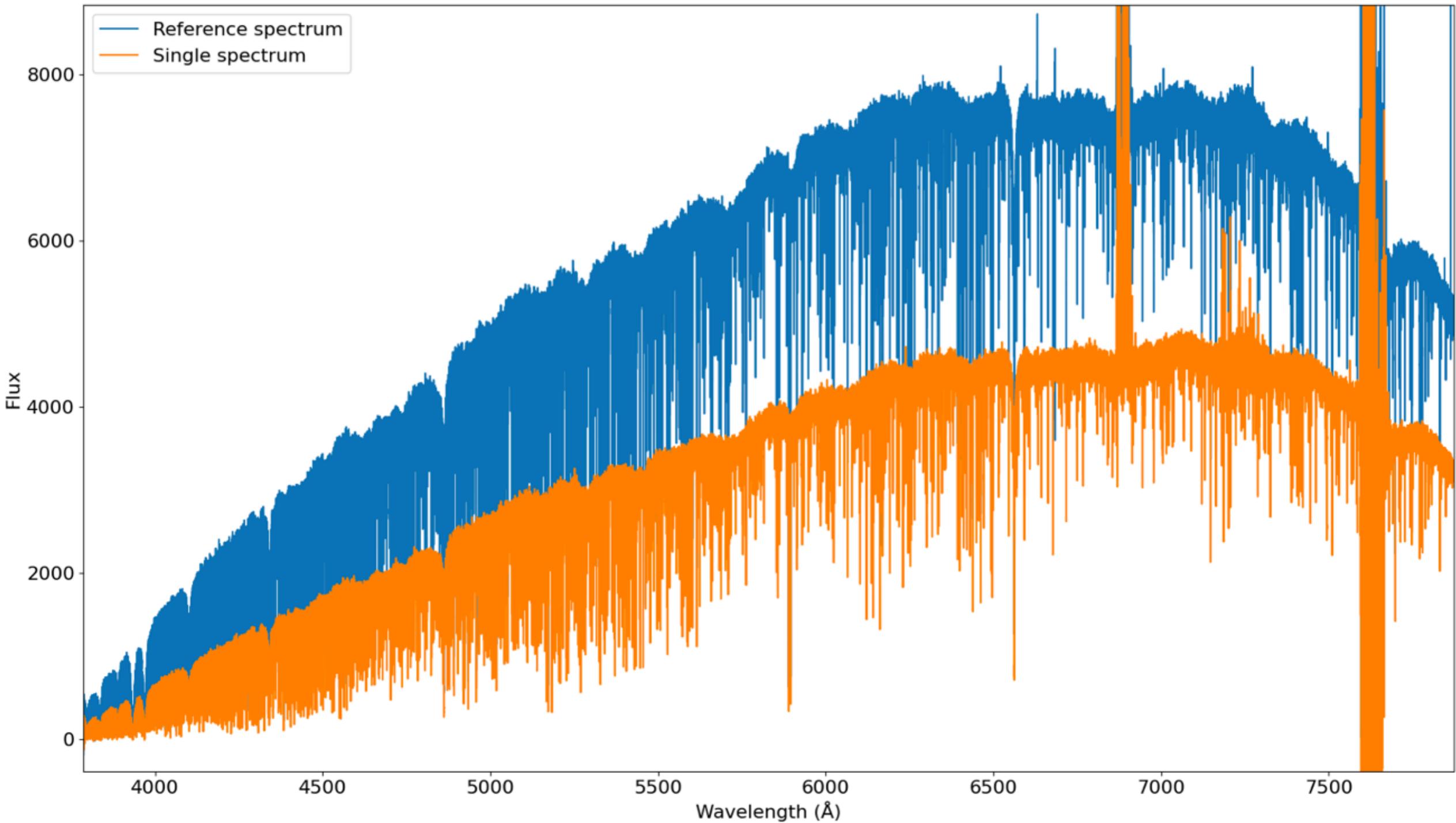
Telluric correction



Na I doublet

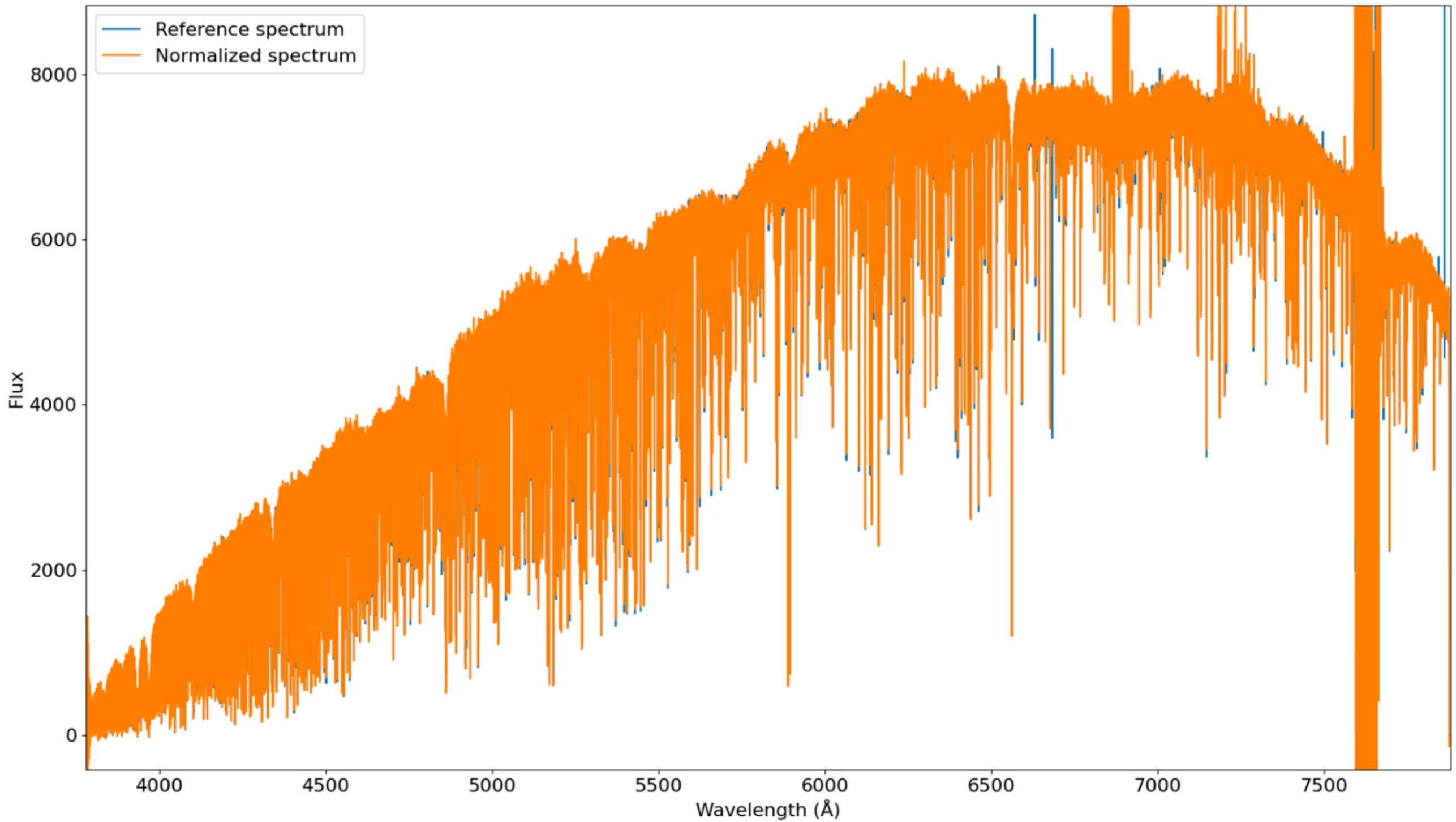


Differential extinction correction

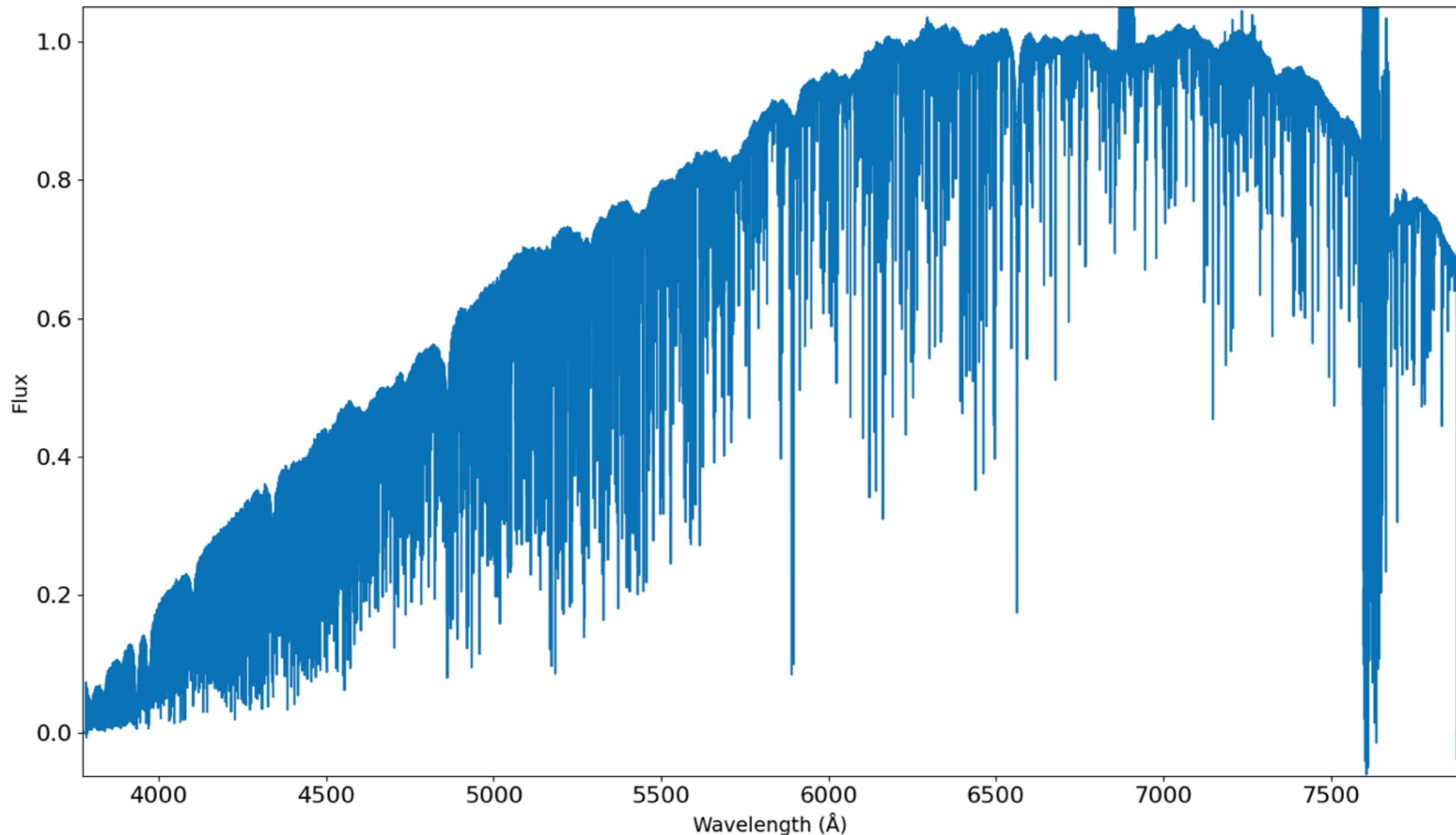


Reference spectrum: exposure with min airmass

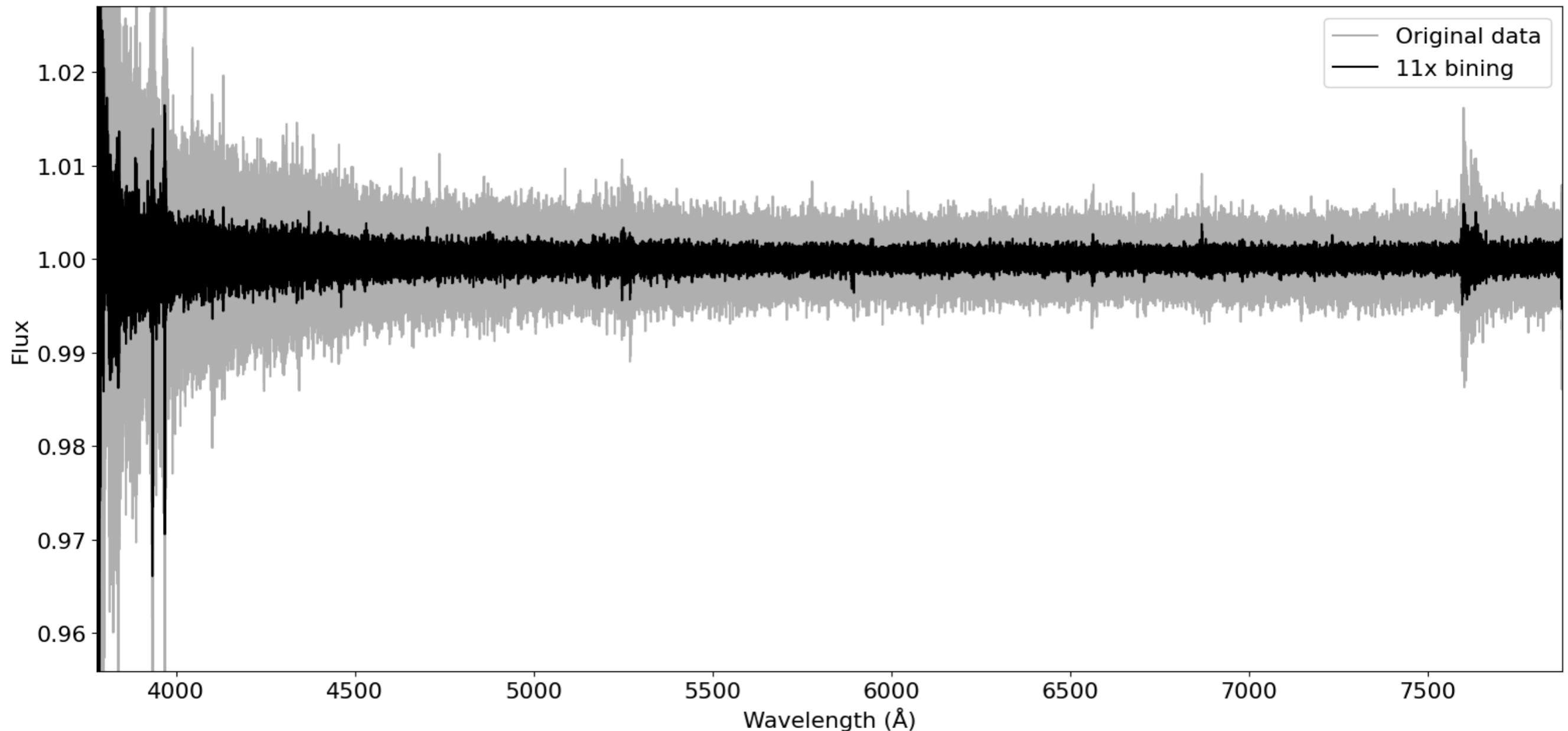
Differential extinction correction



Stellar master spectrum using out-of-transit data



Extraction of planetary spectrum



WASP-76 b

$T_{\text{eq}} = 2200 \text{ K}$

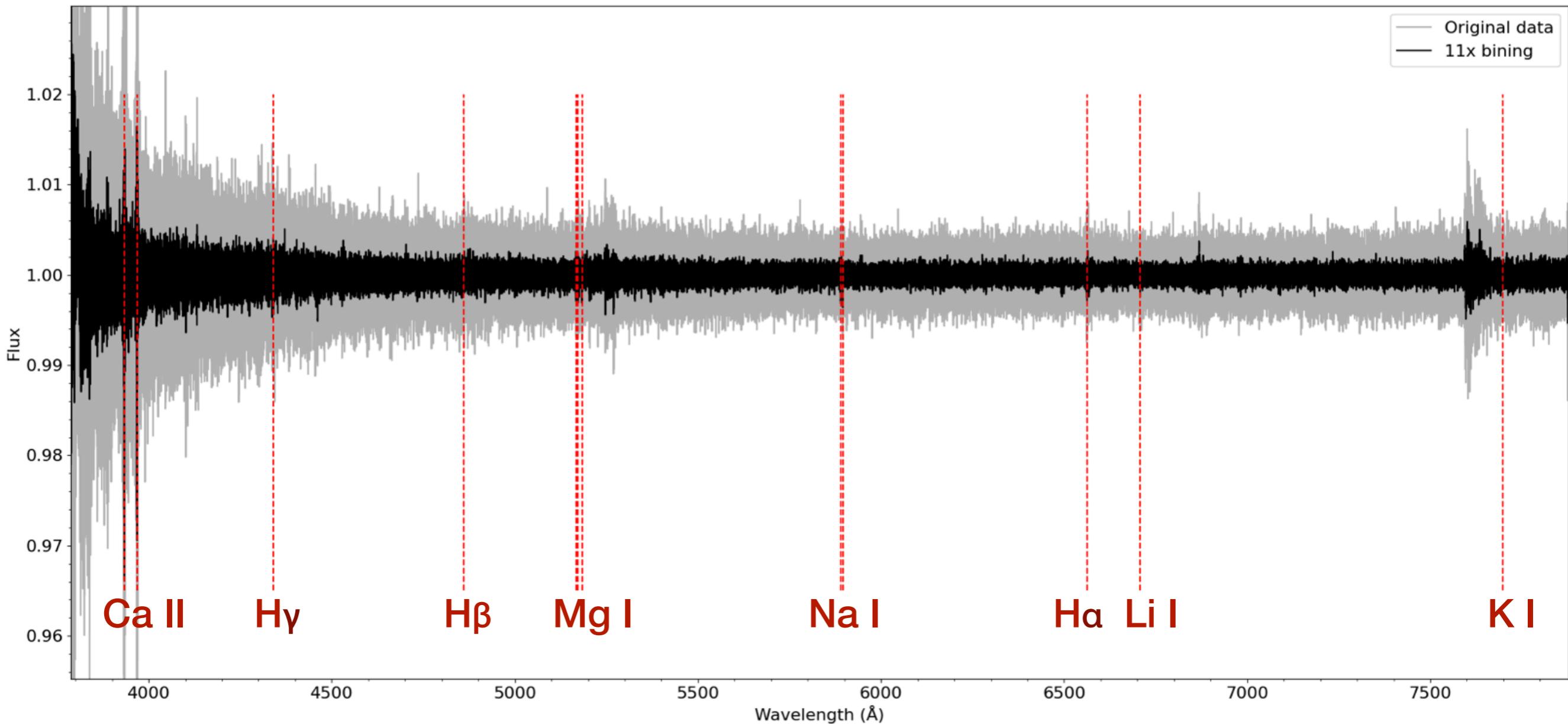
$M_p = 0.92 R_{\text{Jup}}$

$R_p = 1.83 R_{\text{Jup}}$

$\rho = 0.201 \text{ g/cm}^3$

Tabernero et al. 2021

Identification of individual atomic lines



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$M_p = 0.92 R_{\text{Jup}}$

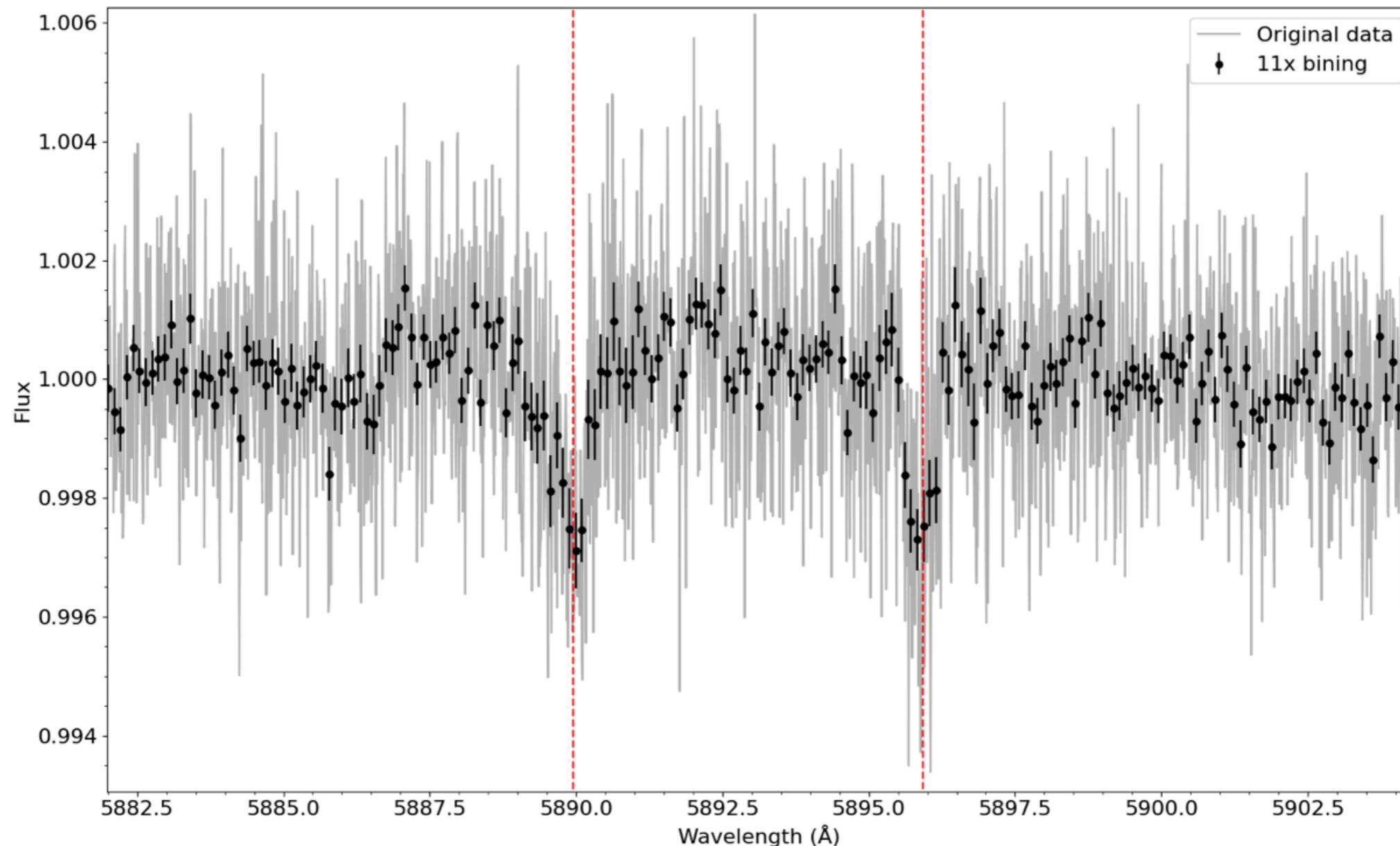
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Identification of individual atomic lines

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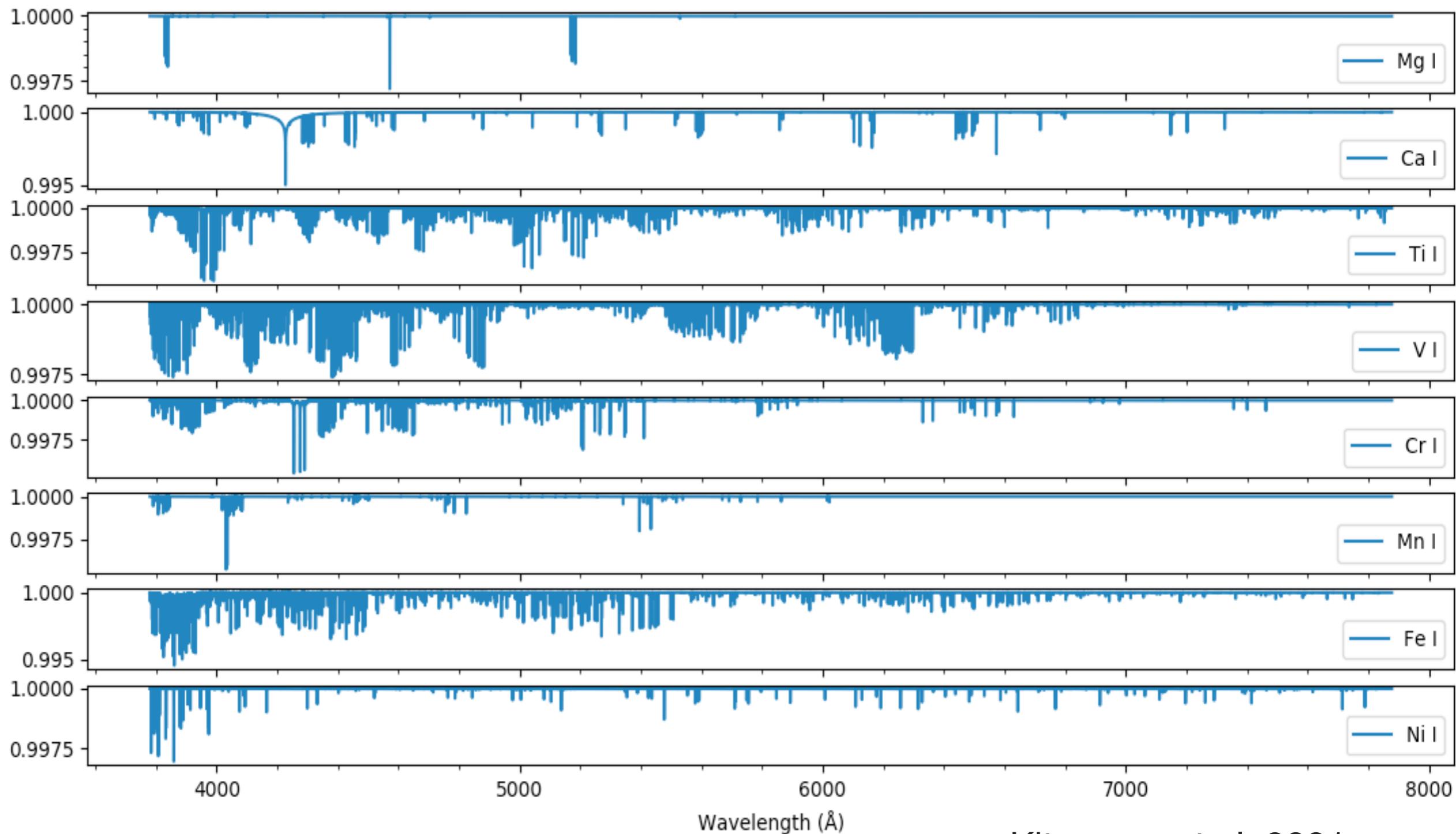


WASP-76 b $T_{\text{eq}} = 2200 \text{ K}$ $M_p = 0.92 R_{\text{Jup}}$
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Tabernero et al. 2021

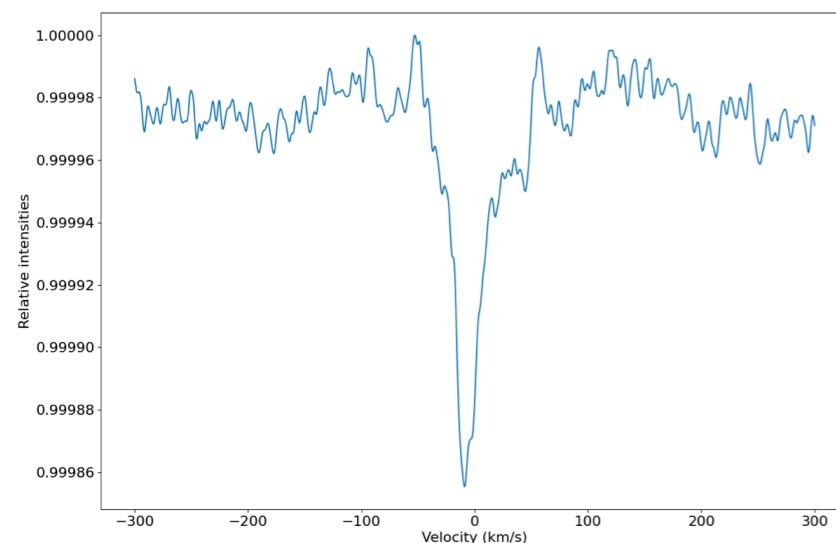
Cross-correlation method

- ◆ Planetary spectrum
- ◆ Reference spectrum:
 - Combs / masks
 - Synthetic spectra

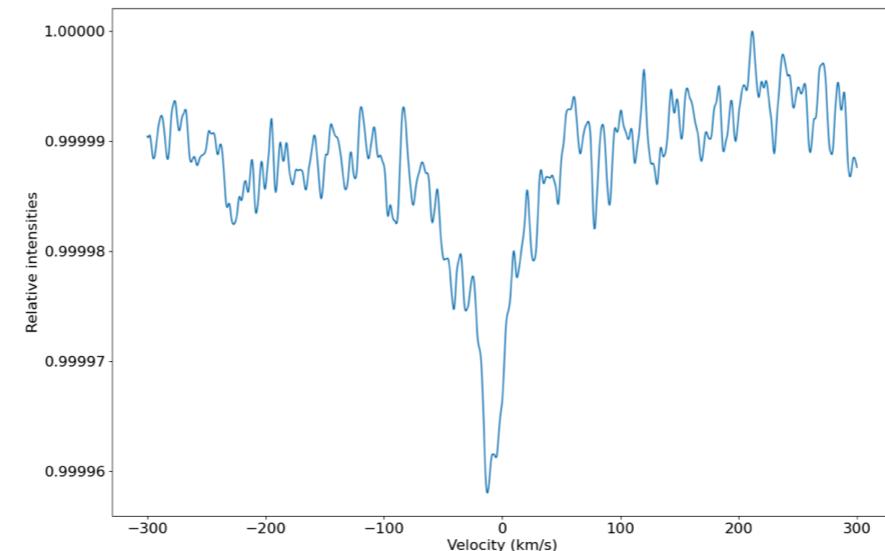


Cross-correlation method

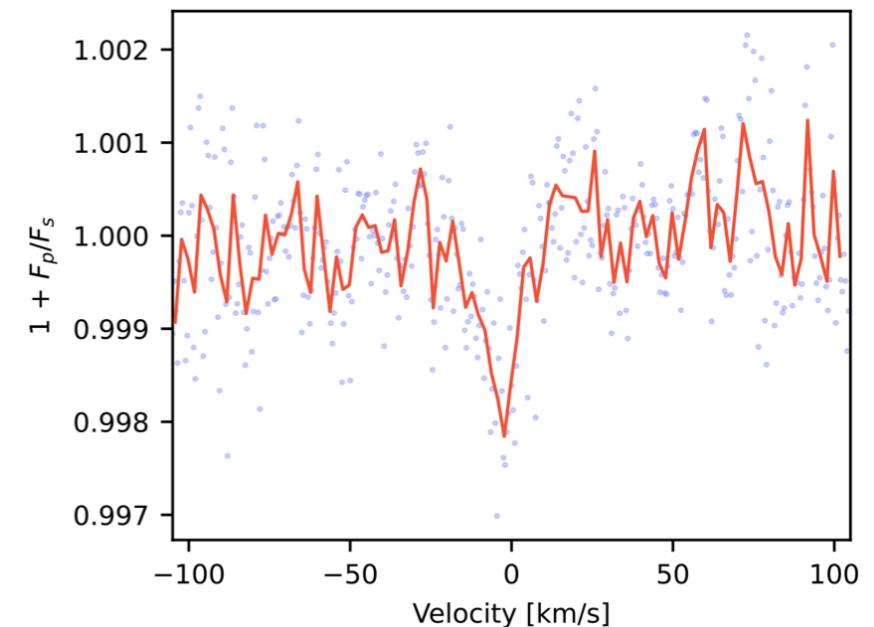
Fe I



Cr I



Ba I



CCFs computed with iSpec (Blanco-Cuaresma et al. 2014)

Azevedo Silva et al. 2022

petitRADTRANS (Mollière et al. 2019) synthetic spectra are used as templates

WASP-76 b

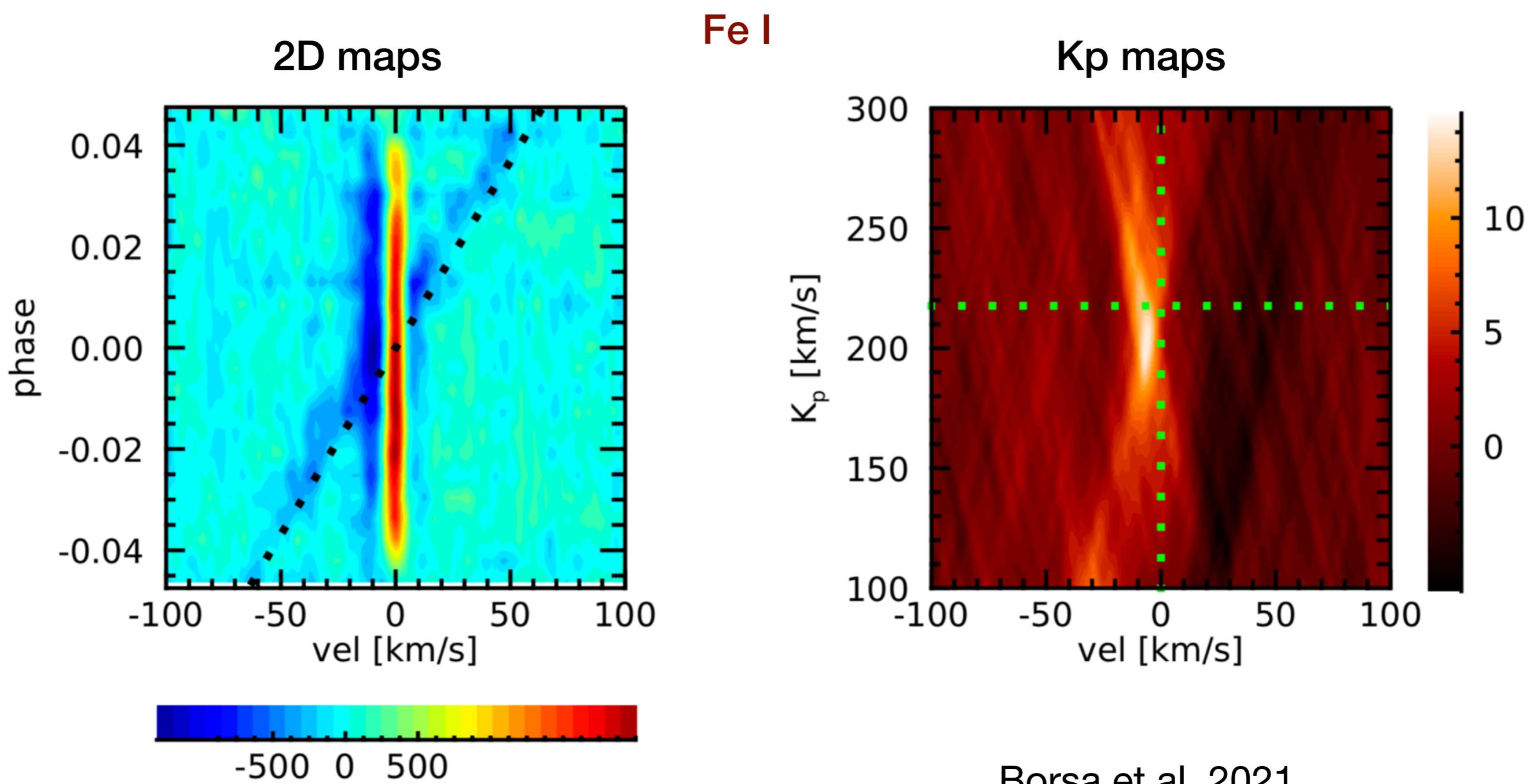
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