UFOs in (all?) high-z QSOs

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Outline

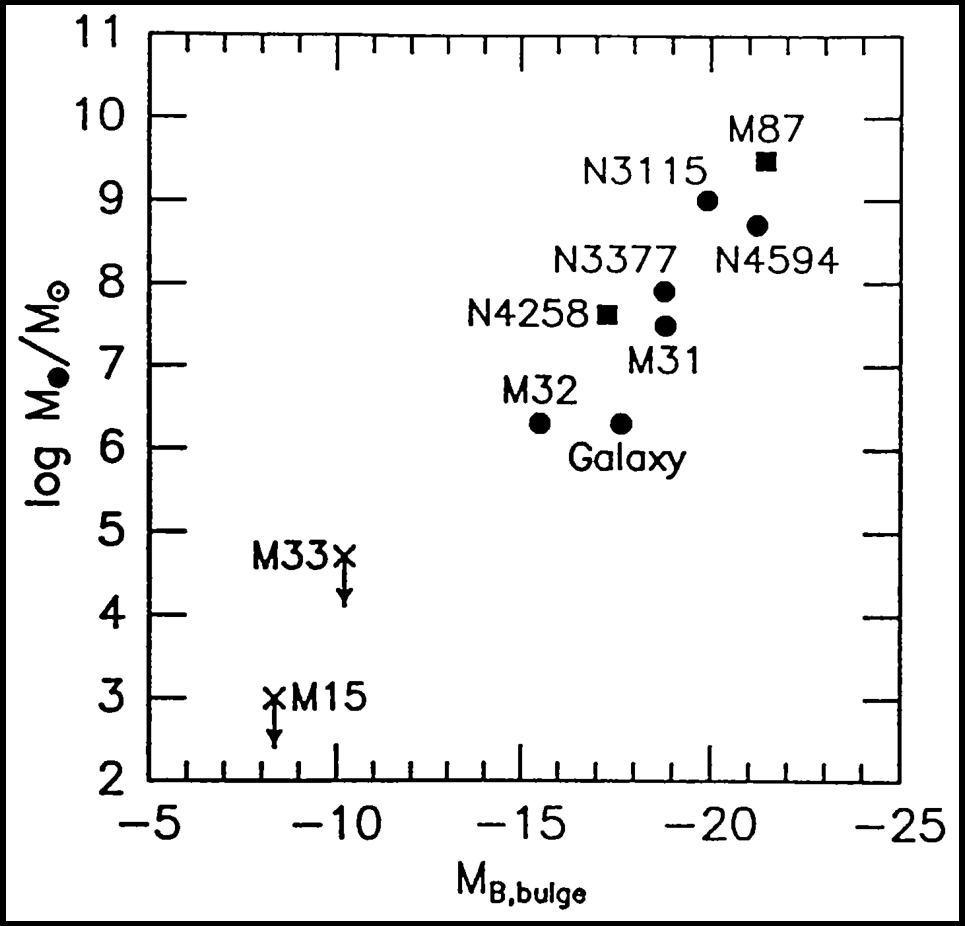
i. Motivation(s)

ii. UFOs in (all?) high-z QSOs

Main Collaborators: M. Dadina, G. Chartas, J. Kaastra, J. Kriss, C. Vignali, G. Lanzuisi, F. Tombesi, M.Giustini, J. Reeves, M. Gaspari, J. Gofford, B. DeMarco, G. Ponti, V. Braito

Framework (i/iv): Co-evolution of galaxies

First unexpected "revolution" in extragal. astrophysics: not only most (all?) galaxies have SMBHs (MDOs) in their centers, these also correlate with bulge properties



Annu. Rev. Astron. Astrophys. 1995. 33:581–624

INWARD BOUND—THE SEARCH FOR SUPERMASSIVE BLACK HOLES IN GALACTIC NUCLEI

John Kormendy[1]

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Douglas Richstone

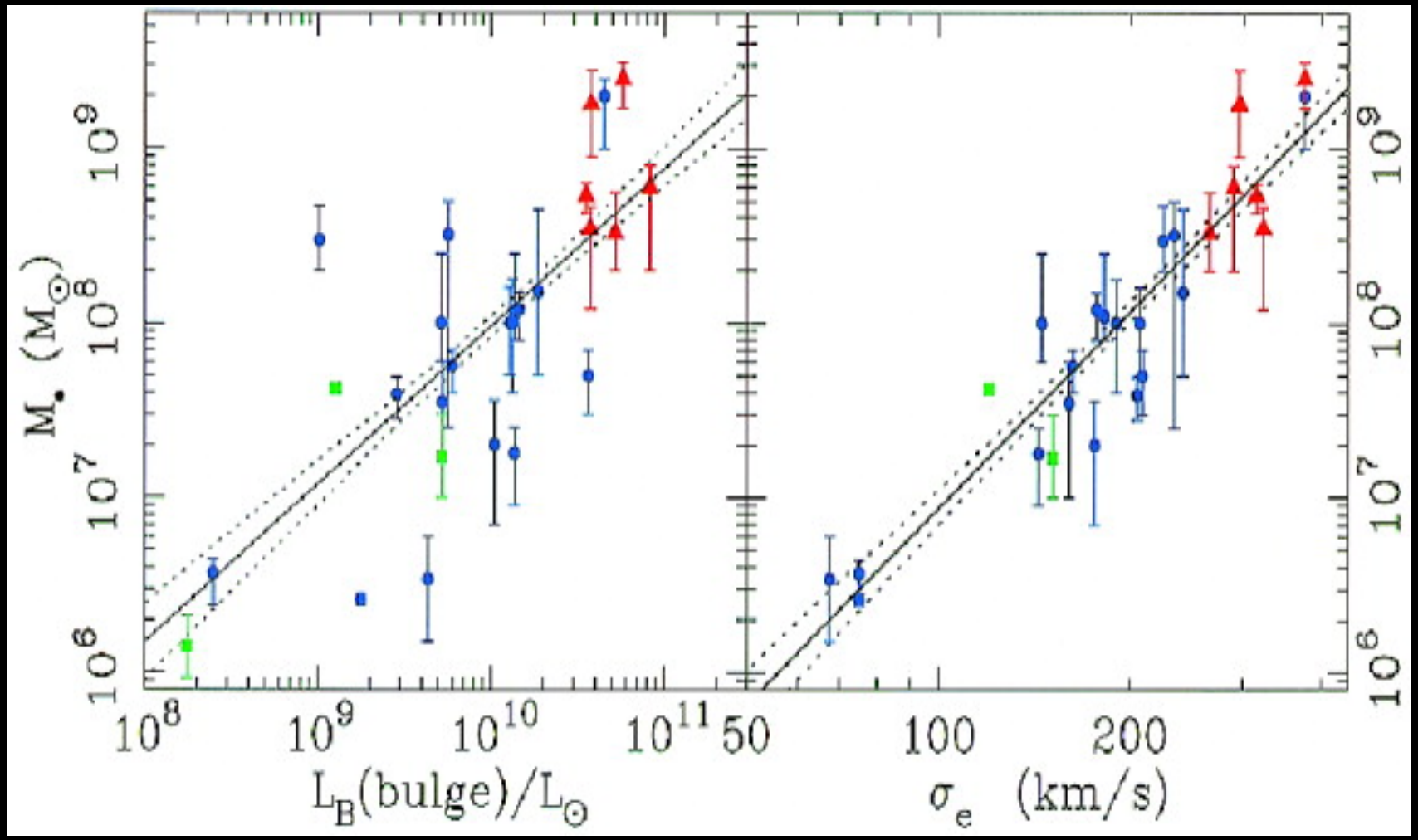
Department of Astronomy, University of Michigan, Dennison Building, Ann Arbor, Michigan 48109

A statistical survey finds BHs in ~20% of nearby E–Sbc galaxies, consistent with predictions based on quasar energetics. BH masses are proportional to the mass of the bulge component. Most candidates are inactive; in some cases, the abundance of fuel is not easily reconciled with BH starvation. Flashes caused by the

Kormendy & Richstone, 1995, ARA&A

Framework (ii/iv): Feedback in the co-evolution of galaxies

⇒evidence for feedback mechanism between SMBH(AGN) and its' host galaxy?



$M_{bh} \sim \sigma^4$

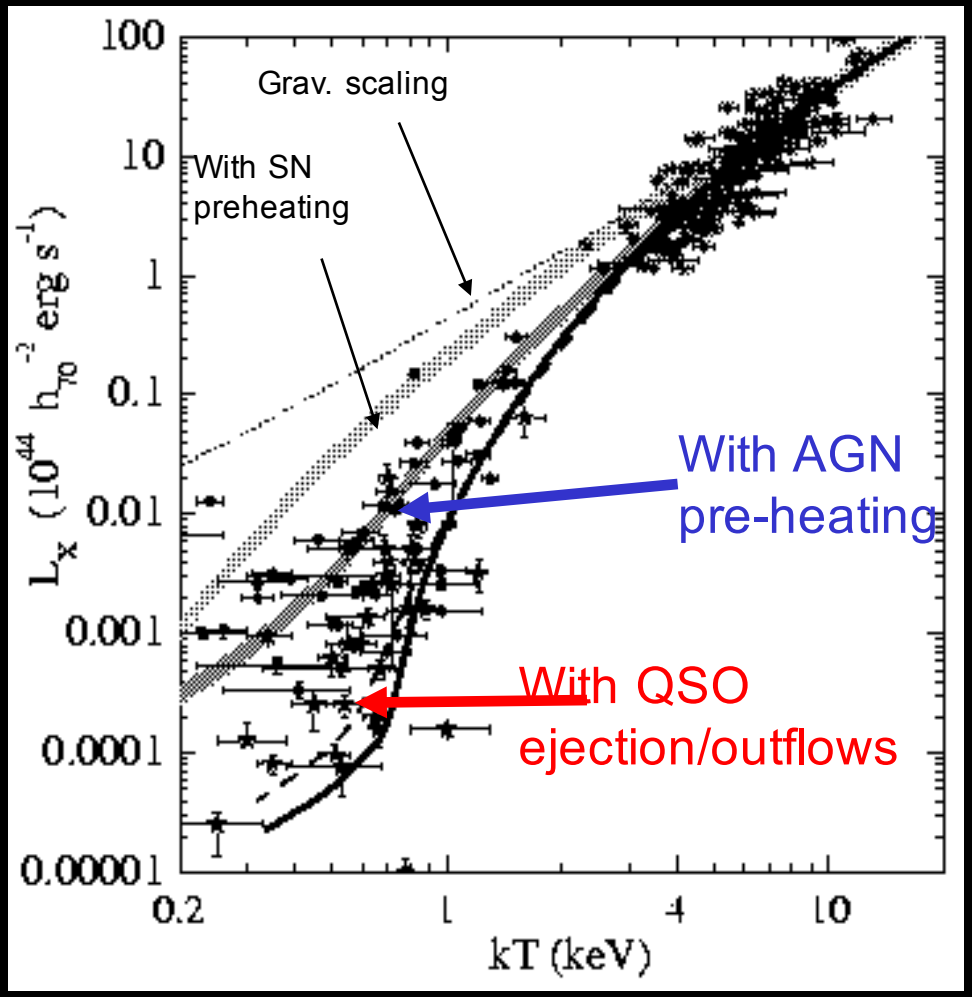
Magorrian et al. '98
Tremaine '02; Gebhardt '02...etc
(see e.g. King and Pounds '03, Crenshaw, Kraemer & George '03, ARA&A)

Framework (iii/iv): (P)re-heating of groups and clusters of galaxies

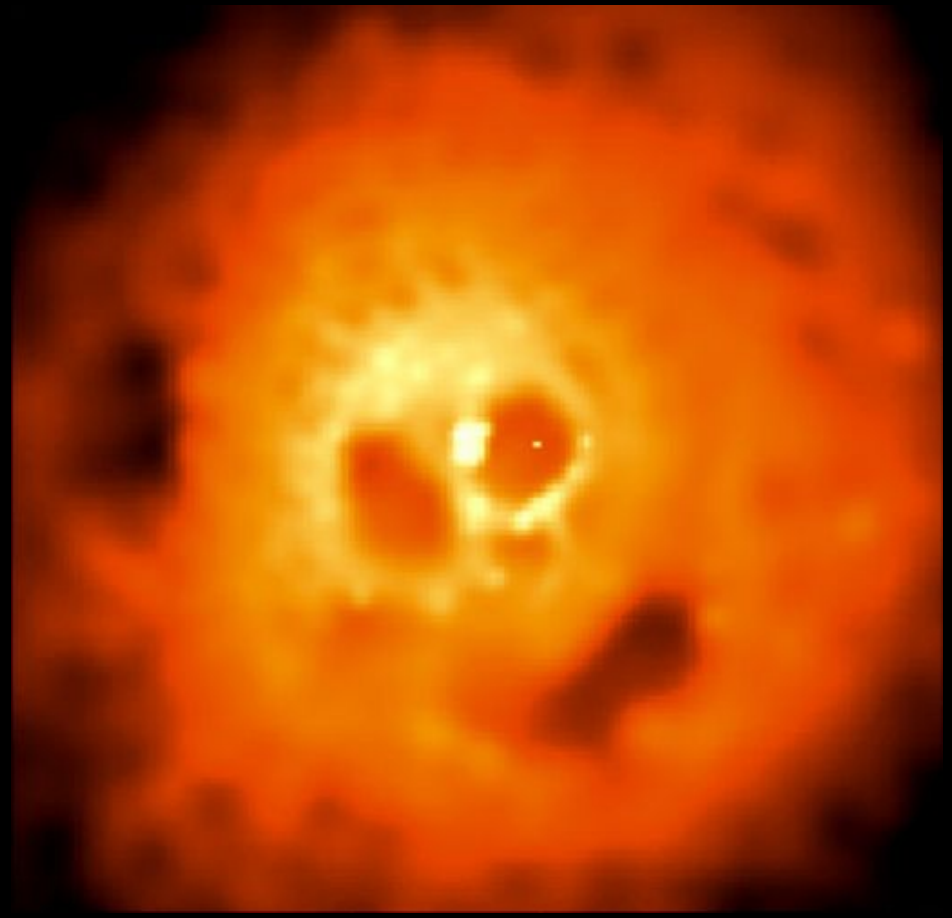
Second unexpected "revolution" in extragal. astrophysics:

need preheating to recover L-T relations & cooling flows extra-heating

⇒ Energy feedback from AGNs/QSOs in groups&clusters?



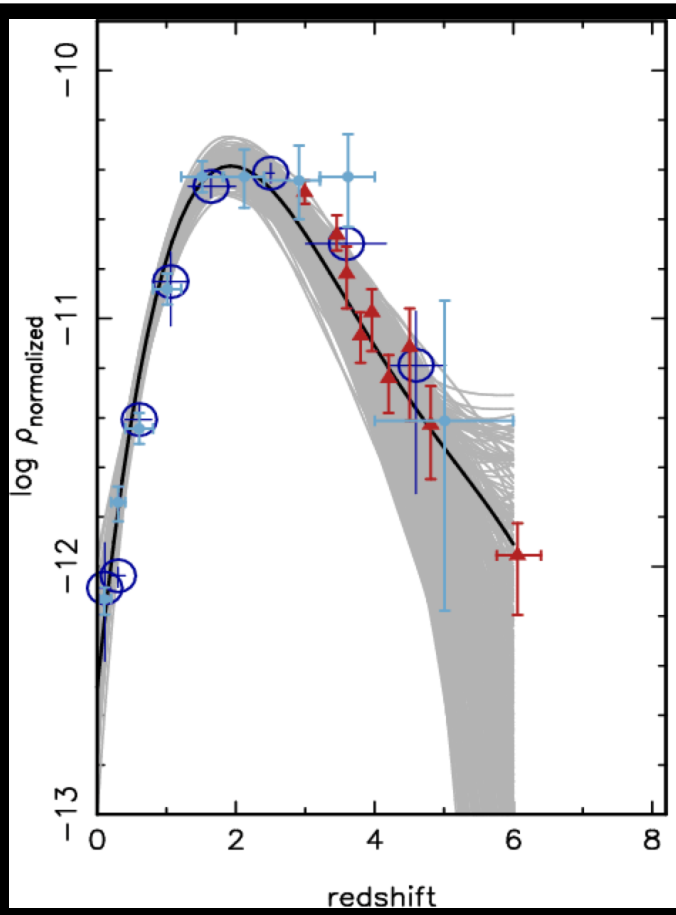
Lapi, Cavaliere & Menci, '05



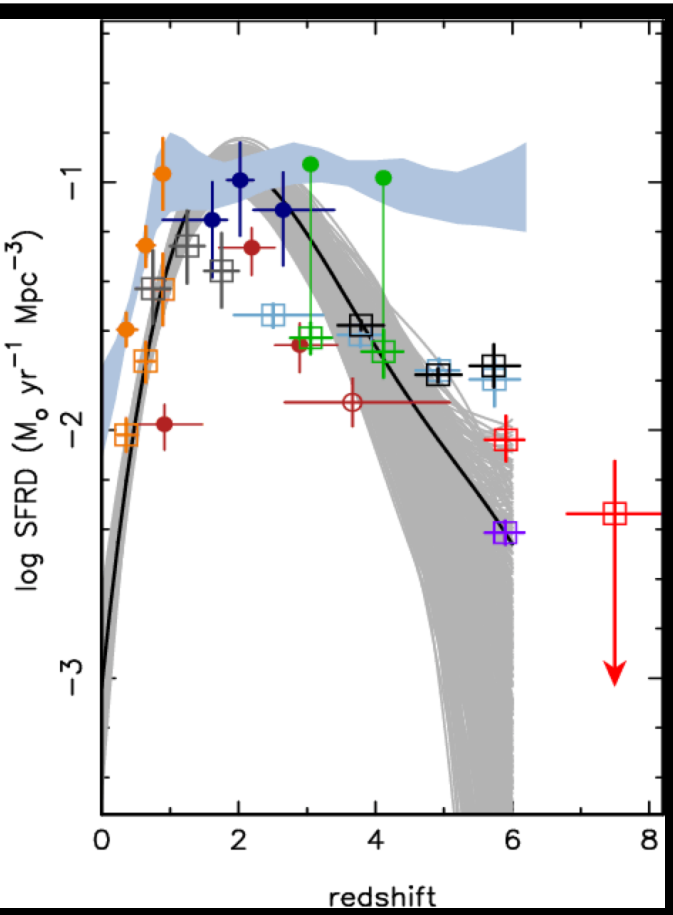
Perseus Cluster

Fabian et al. '05

Framework (iv/iv): MBH vs SFR, which arrived first at z~2-3?



QSO space density



SFR space density

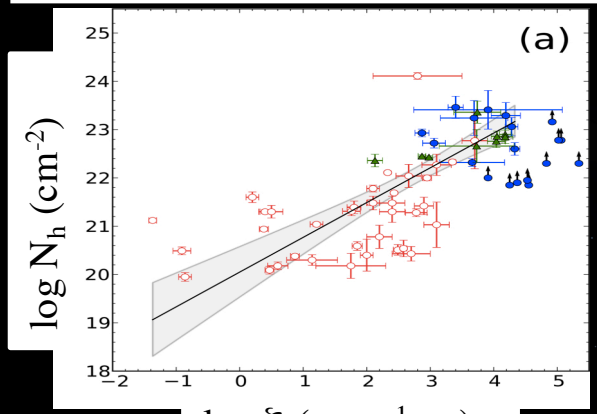
Wall et al. '05
Madau et al. '96

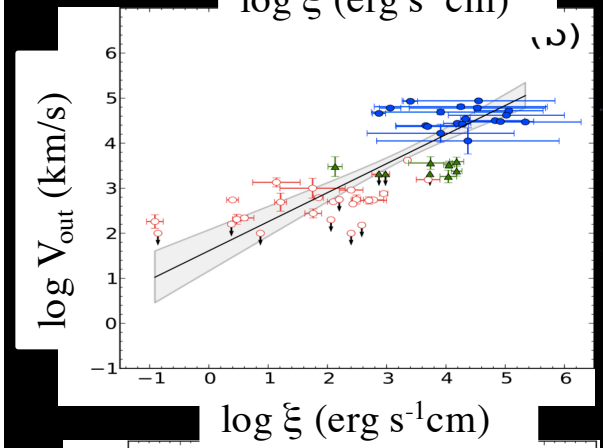
M_{bh}-σ relation, AGN-gal coevolution,
L-Tx relations, Heating cooling flow,
Galaxies colors & sizes

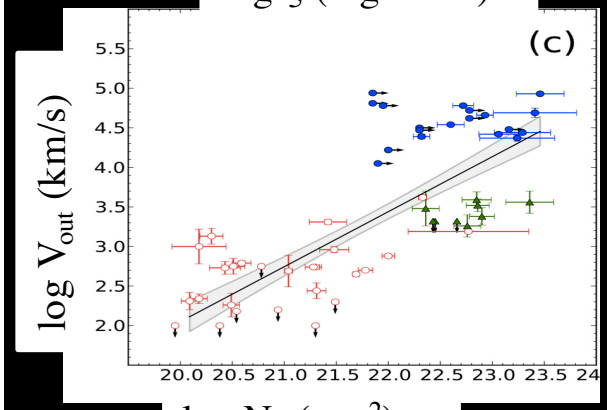
BUT HOW?
(Jet, Winds/UFOs, L_{AGN}, mix?)

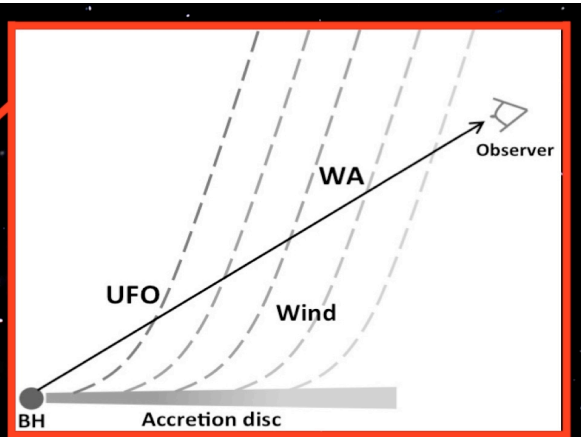
AGN Feedback !

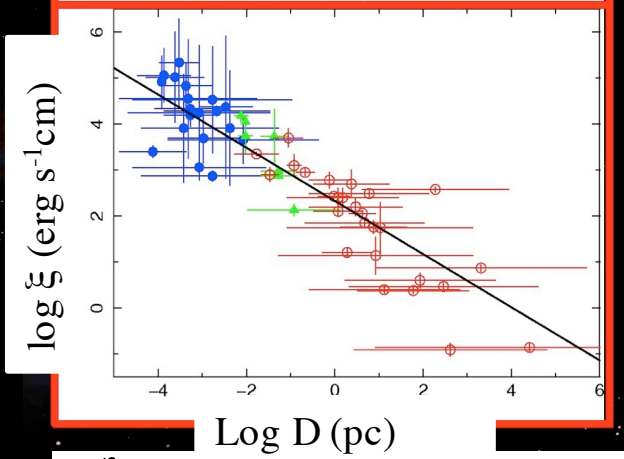
At low-z: A possible (unifying) X-ray view of UFOs and non-UFOs (WAs)



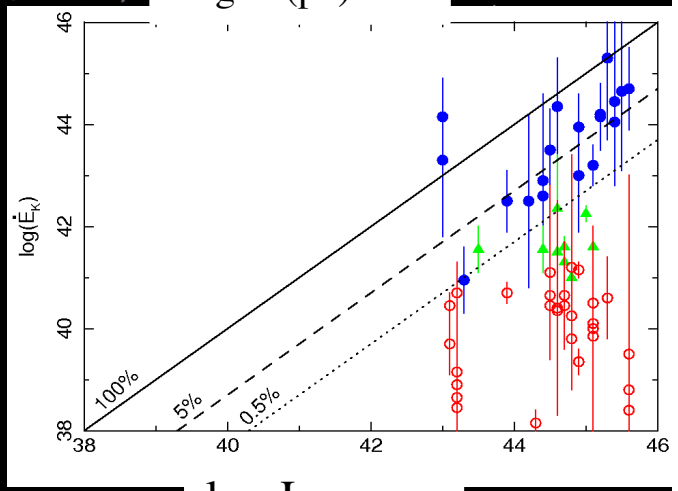








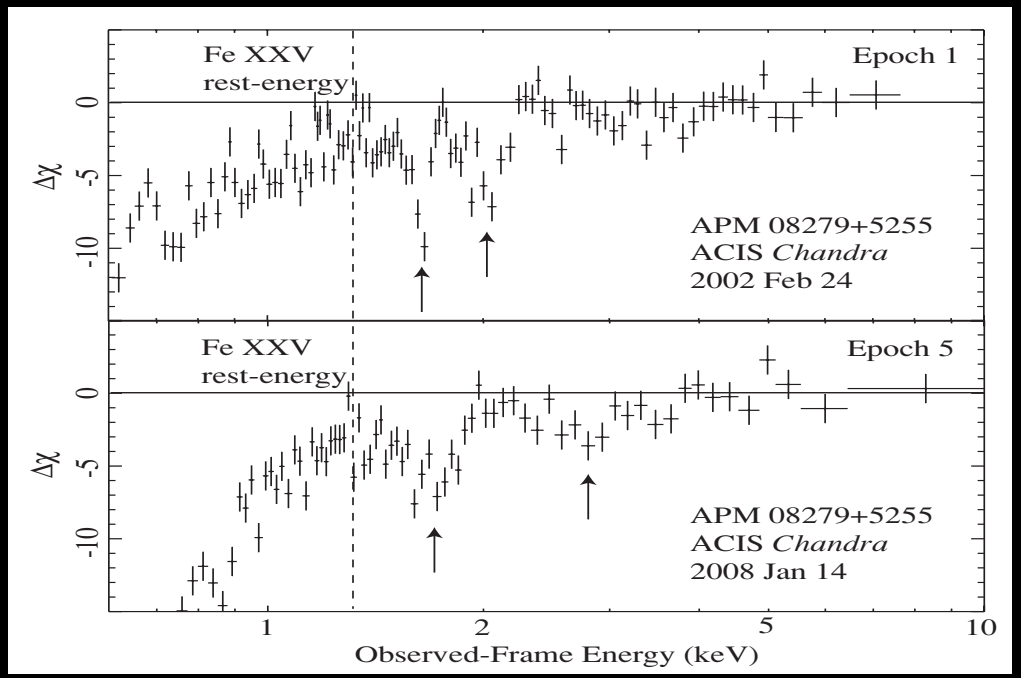
Tombesi, MC et al., '12b, '13



→ UFOs kinetic energy >1% of Lbol
→ Feedback (potentially) effective!

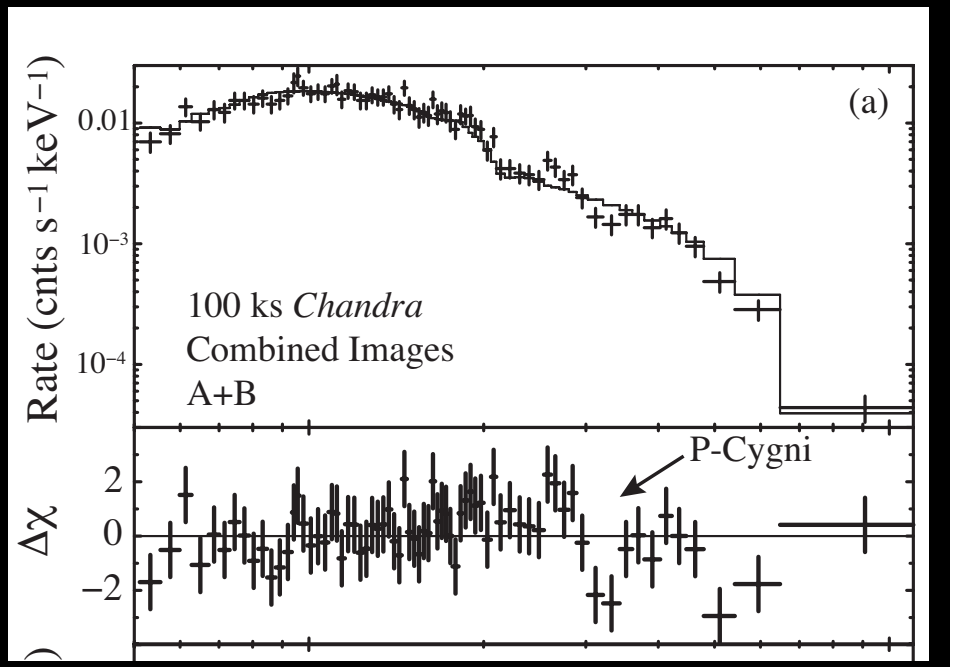
At high-z: UFOs and/or FeK complex features seen in lensed high-z QSOs

APM 08279+5255 (z=3.91)
V_{out}~0.2-0.76 c



Chartas et al. 2002, 2009

HS0810+554 (z=1.5)
V_{out}~0.1-0.4 c

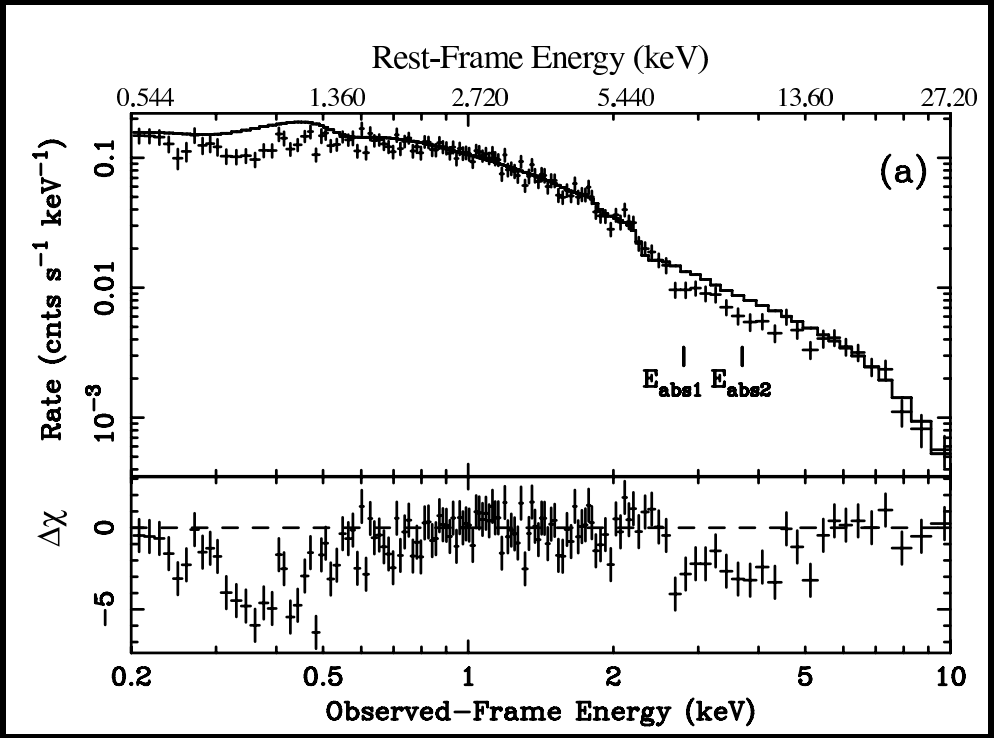


Chartas, MC, et al. 2014, 2015

→ Ubiquitous complex (i.e. ionized and/or partially covering), and massive absorption in high-z QSOs?

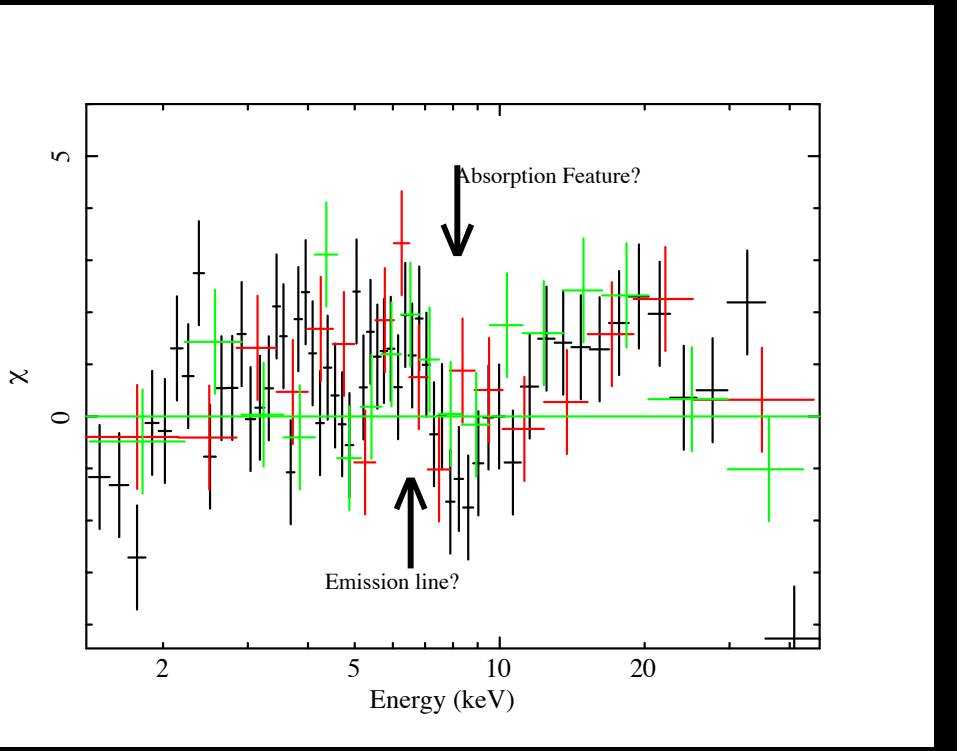
At high-z: UFOs and complex absorbers measured in lensed high-z QSOs

PG1115+080 (z=1.7)
V_{out}~0.1-0.34 c



Chartas et al., '07

B1422+231 (z=3.6)
Ionized absorber, V_{out}=?

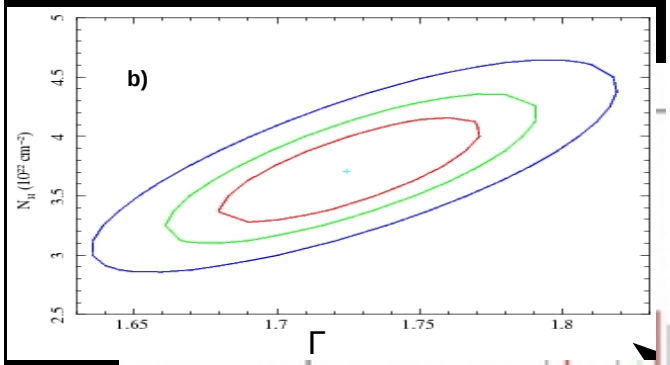


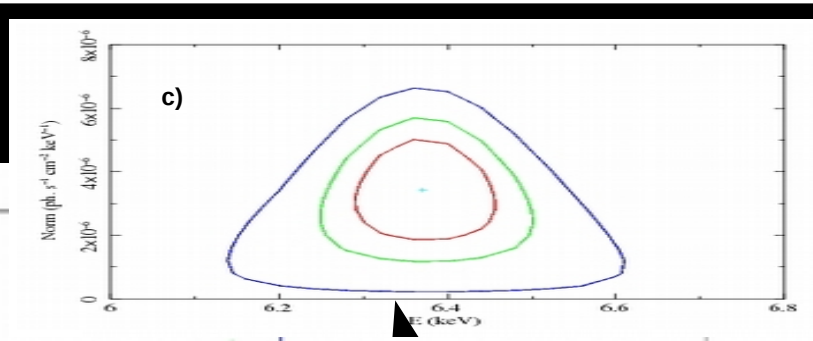
Dadina, MC, et al., '16

The new X-ray view: UFOs and complex absorbers measured in lensed high-z QSOs

0414+0534 (z=2.6)

0.32 c

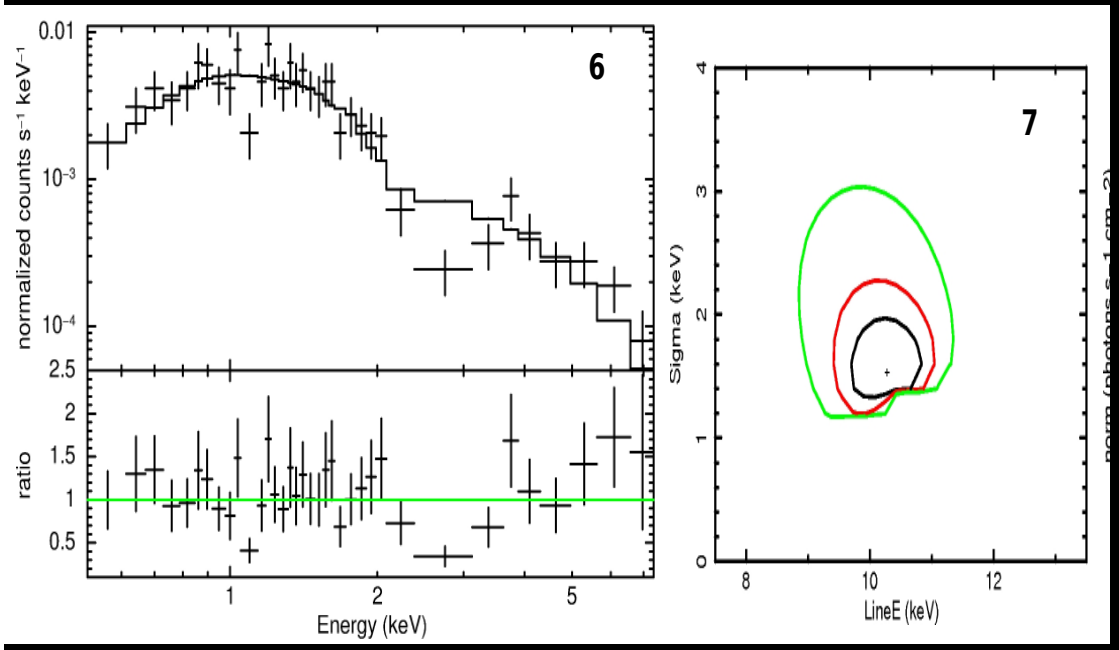




Dadina, MC, et al.,see poster

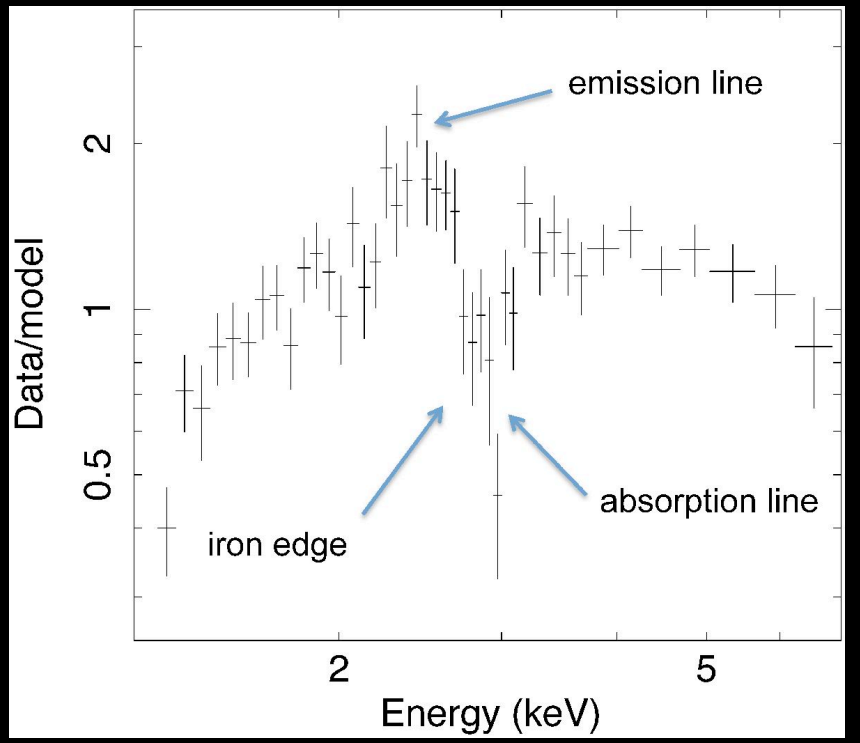
The new X-ray view: UFOs seen also in non-lensed high-z QSOs

(z=2.73) high-z RQ (NAL) QSO HS1700+6416,
V_{out}~0.12-0.6c



Lanzuisi et al., ’12

PID352 (z=1.6)
V_{out}~0.15c

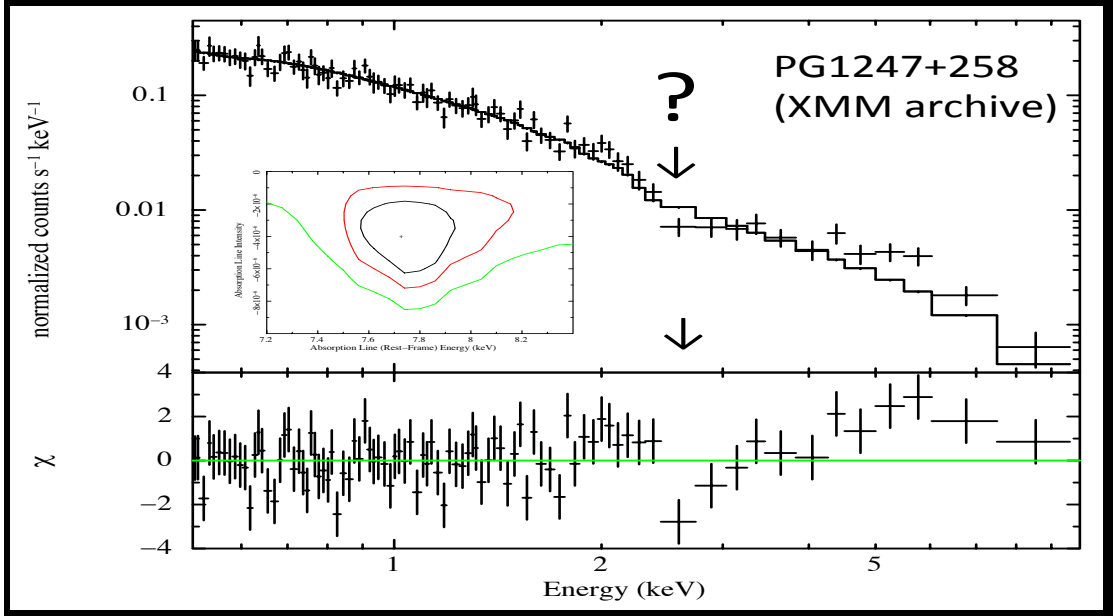


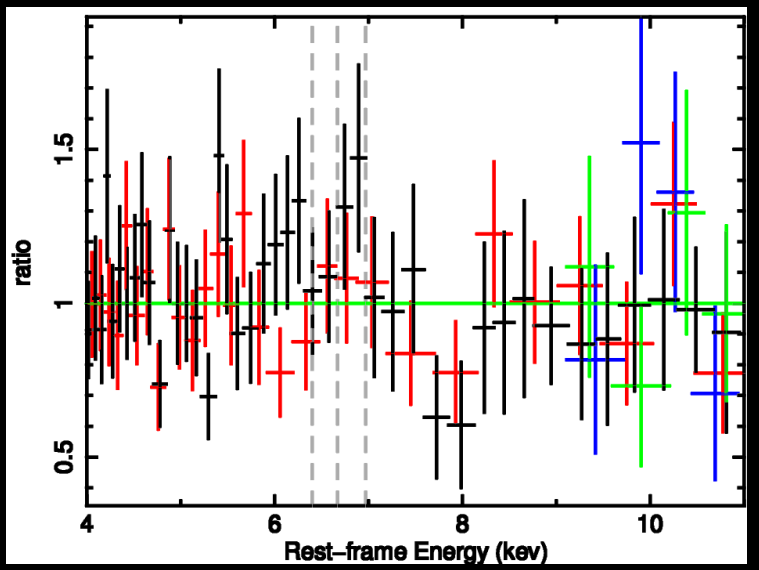
Vignali et al., ’15

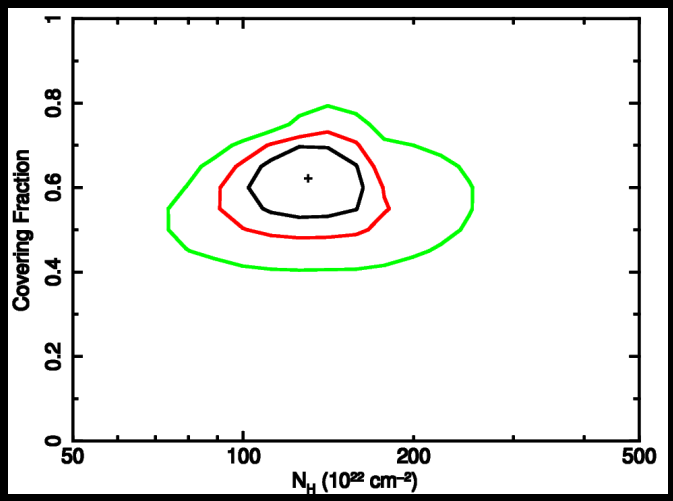
→ Ubiquitous complex (i.e. ionized and/or partially covering) absorption?

UFOs seen also in non-lensed high-z QSOs

(z=2) PG1247+268
V_{out}~0.15c







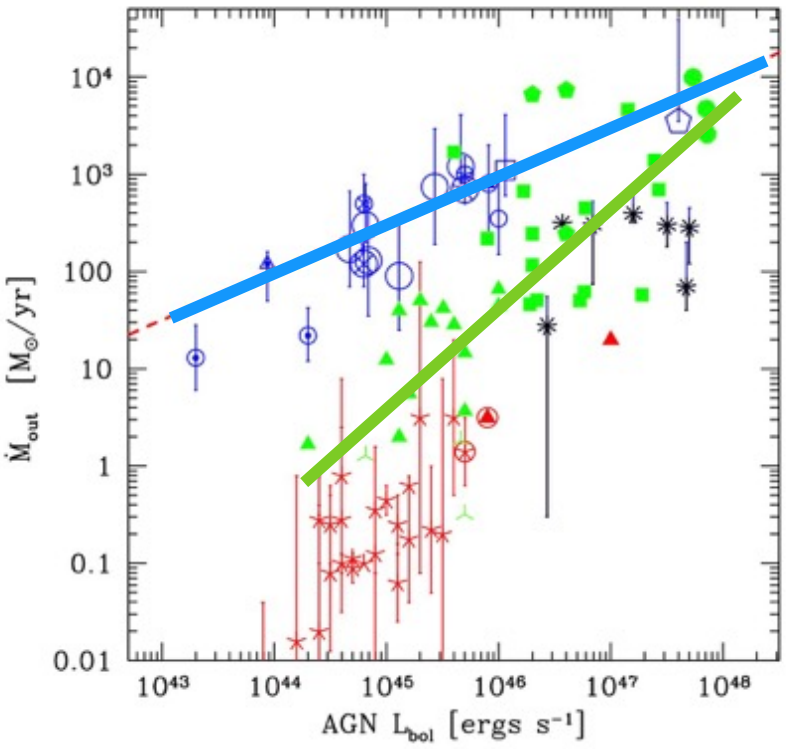
Another high-z UFO candidate…

Lanzuisi et al., ’16

→ Ubiquitous complex (i.e. ionized and/or partially covering) absorption?
→ Desperately need more and longer XMM observations on high-z QSOs

Just were approved a ~450 ks of XMM time to observe 4 non-lensed QSOs at z~2 (PI: MC)

Need X-ray and multi-ni coverage of a representative sample of high-z QSOs.

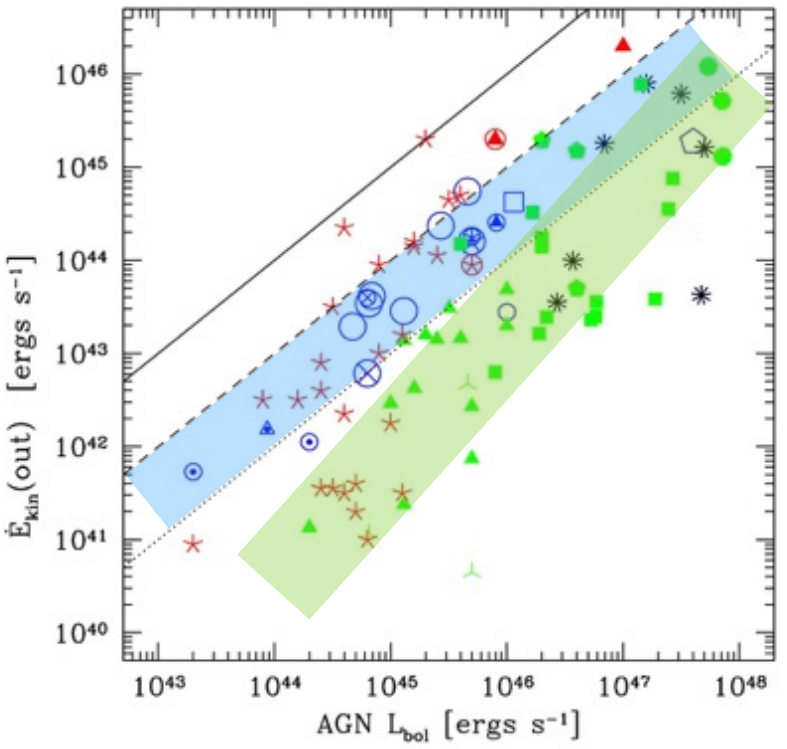


Molecular
small=nucl
large=gal.

Ionized

BAL

X-ray



Remarkable correlation between wind mass outflow rate and AGN bolometric luminosity: $M_{out} \sim L_{bol}^{0.5}$ for molecular winds $M_{out} \sim L_{bol}$ for ionized winds

E_{kin}(out) = 1-10% L_{bol} (molecular) E_{kin}(out)=0.1-10% L_{bol} (UFOs, BALs)

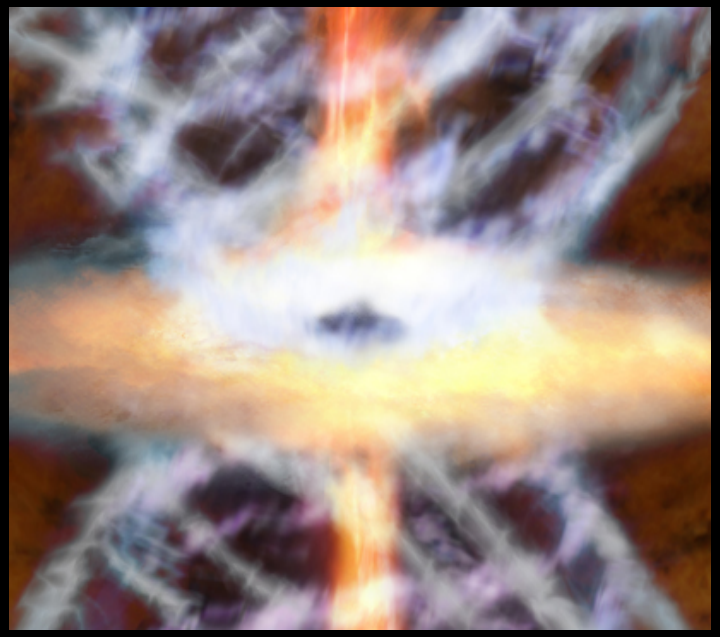
E_{kin}(out) = 0.1-1% (ionized low L_{bol}) = 1-10 % (ionized high L_{bol})

Fiore et al., ’15

Summary:

- Science Case (outflows)
  - *Recognized importance, and "pathfinder" to future missions/observatories (from ground based Obs. ALMA, MUSE, SINFONI to Athena).*
  - *Important implications for both astrophysics of winds/outflows formation and acceleration, and the cosmological impact/feedback of AGN winds.*

- Cosmological impact/feedback:
  - Few decent high-z QSOs spectra available, ALL show UFO-like features in their X-ray spectra
  - Need to have good quality (>10000 cts) X-ray spectra for a representative sample of (30-40) high-z QSOs to characterize and measure the frequency of massive and energetic outflows in high-z QSOs (for z~0-2, L~0.1-Ledd). Need multi-ni coverage to obtain full outflow energetics. Multi-ni would also "guarantee" more publications per XMM-ks, like experience in low-z AGNs.
  - The future: from XMM (LPs and VLPs) to Athena (core science)



Thank you very much
for your attention