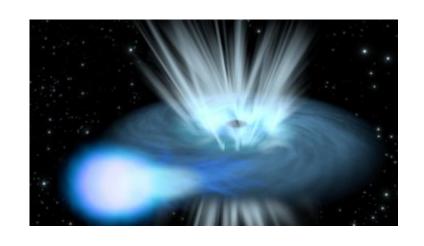


A new catalogue of Universultraluminous X-ray sources (and more!)



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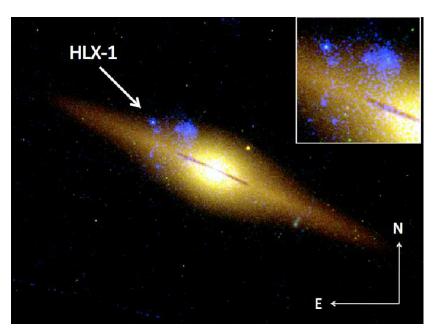
Ultraluminous X-ray sources

Kaaret, Feng & Roberts (2017)

☐ Eddington limit is

$$L_{Edd} = 1.3 \times 10^{38} (M/M_{Sun}) \text{ erg s}^{-1}$$

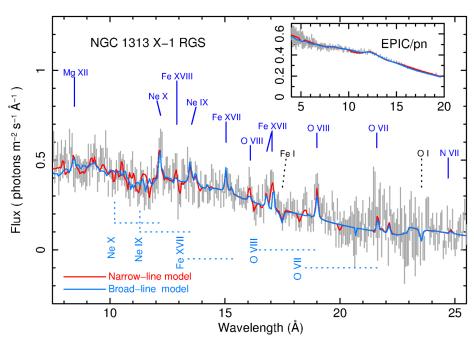
- □ **ULXs**: point sources in nearby galaxies with $L_x > 10^{39}$ erg s⁻¹
- ☐ Implies big black holes; or super-Eddington accretion
- ☐ IMBHs now thought to be a minority of ULXs



The optical/UV counterpart of ESO 243-49 HLX-1, a good IMBH candidate (Mapelli et al. 2013)



Super-Eddington accretion



Detection of an outflowing, fast (v = 0.2 c) wind in NGC 1313 X-1 (Pinto et al. 2016)

Observational evidence includes:

- Optical nebulae (e.g. Pakull & Mirioni 2003)
- X-ray behaviour (e.g. Gladstone et al. 2009; Sutton et al. 2013)
- Outflowing material in wind (Middleton et al. 2015; Pinto et al. 2016, 2017)
- Pulsations 3
 neutron star ULXs!
 (e.g. Bachetti et al. 2014; Israel et al. 2017)



ULX catalogues

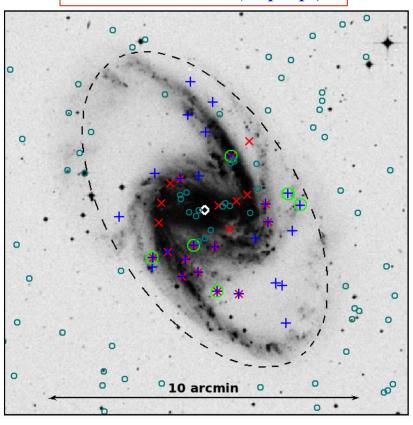
- □ Long heritage
 - ROSAT HRI (e.g. Roberts & Warwick 2000, Colbert & Ptak 2002)
 - Chandra ACIS (e.g. Swartz et al. 2004, Liu et al. 2011)
 - **XMM-Newton EPIC** (e.g. Walton et al. 2011)
- ☐ Critical enabler for science
 - Statistical samples population studies
 - Comparative source properties
 - Outliers (e.g. brightest, most variable...)



A new catalogue

- □ Take 3XMM-DR4 and cross-correlate with RC3 & Catalogue of Neighbouring Galaxies (Karachentsev et al. 2004)
 - ~2000 galaxies covered
- ☐ Follow method of Walton et al. (2011), with some improvements

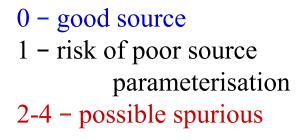
Earnshaw et al. (in prep.)

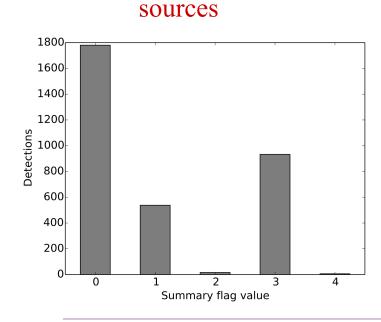


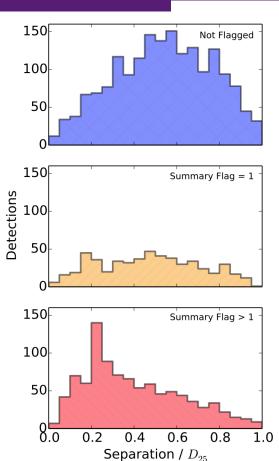
NGC 1365. Cyan circles – 3XMM detections; green circles – ULXs; blue plusses – unflagged detections; red crosses – flagged detections

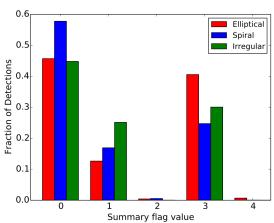


Source flags







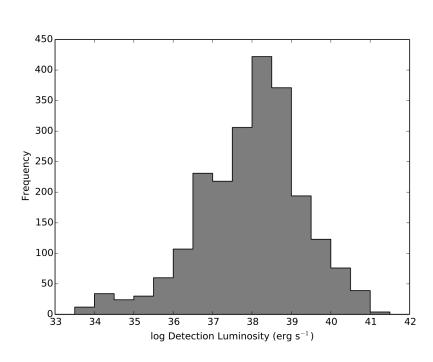


Keep 0 & 1

Remove 2-4; affects sources near centres of galaxies and/or in elliptical galaxies most.



New catalogue: numbers

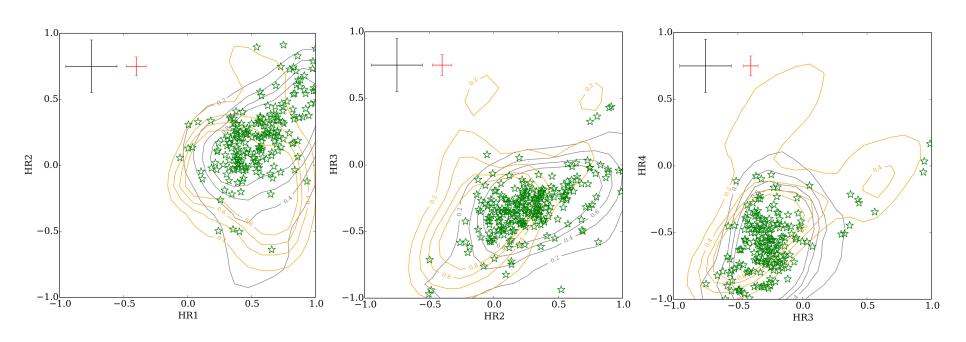


Luminosity distribution of detections

- 2256 detections of 1464 sources in 349 galaxies
- □ 539 detections of 372ULXs in 248 galaxies
- Complete subset:298 detections of 227ULXs in 122 galaxies
- □ Contamination ~ 20%



X-ray colours

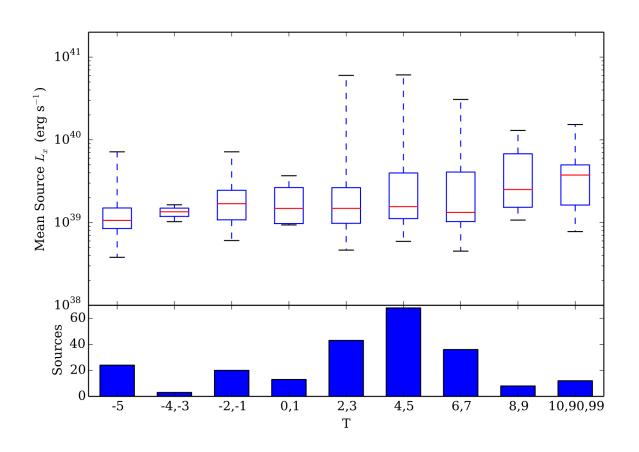


Standard 3XMM bands (0.2-0.5; 0.5-1; 1-2; 2-4.5; 4.5-12 keV) converted into colours using HRn = (RATE_b - RATE_a)/(RATE_b + RATE_a).

Complete ULX sample is green stars; grey contours from sub-ULX sample; orange contours from nuclear sources (AGN candidates)



Luminosity versus host type



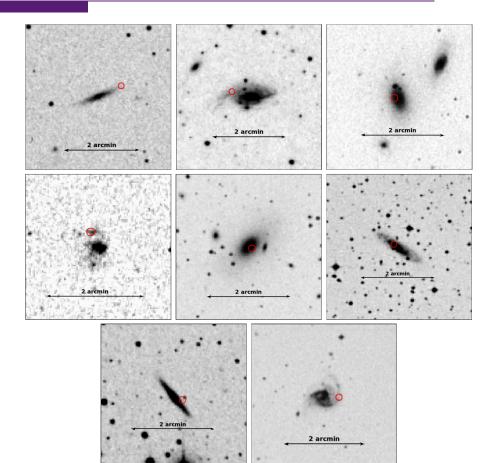
Tentative upwards trend in luminosity towards later type galaxies?

Cf. established theoretical & observational links to metallicity (e.g. Prestwich et al. 2011, Linden et al. 2010, Marchant et al. 2017) – more XRBs, and bigger BHs.



New IMBH candidates

- □ 22 sources with ≥ 1
 detection above 5 ×
 10⁴⁰ erg s⁻¹
- ☐ 4 well-known (incl. IMBH & PULX); 10 rejected as known or suspected AGNs
- 8 new IMBH candidates (see right, on DSS images)





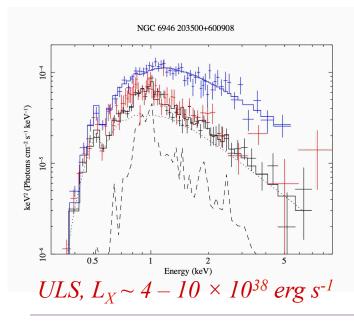
The Eddington Threshold

- □ Star-forming galaxy XLFs flat so emit most (>80%) of energy (radiative & mechanical) in HMXBs with luminosities above 10³⁸ erg s⁻¹.
- □ Bulk of output in $1 30 \times 10^{38}$ erg s⁻¹ regime (*Eddington Threshold*)
- ☐ (Far) more binaries created in early Universe
- □ Feedback impacts on young galaxies
- Cosmologically important regime!

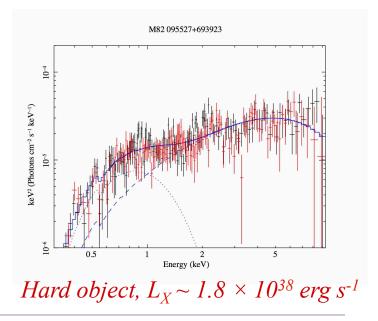


Important – but poorly understood

- Observational constraints poor
- ☐ *Athena* science; pilot studies with *XMM-Newton*
 - E.g. Earnshaw & Roberts (2017) 4 objects



Stacked spectra (Roberts et al. in prep.)





Summary

- ULXs are heterogeneous but super-Eddington emission appears common
- ☐ ULX catalogues an important resource
- □ Derived a new catalogue from 3XMM-DR4, including 372 ULXs & > 1000 less luminous objects in nearby galaxies
- Science: X-ray colours, luminosity as a function of host, extreme objects, Eddington Threshold, etc...