# Dependence of the broad Fe K $\alpha$ line on the physical parameters of AGN

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#### Abstract

The dependence of the broad Fe Kα line on the physical parameters of AGN is investigated by applying the X-ray spectra stacking method to a sample of Type I AGN. A broad line is detected in the stacked spectra of the high  $\lambda_{Edd}$  sub-sample (log $\lambda_{Edd} > -0.9$ ). The profile of the line can be well fitted with relativistic broad line model. We found hints that the Fe K line becomes broader as the  $\lambda_{Edd}$  increases. A broad line might be present in the narrow-line Seyfert I (NLSI) galaxies. Our results indicate that the detection/properties of the broad Fe K $\alpha$  line may strongly depend on  $\lambda_{Edd}$ , which can be explained if the ionization state and/or truncation radius of the accretion disc changes with  $\lambda_{Edd}$ . The non-detection of the broad line in the BLS1 sub-sample can be explained if the the average EW of the relativistic Fe Kα line is weak or/and the fraction of sources with relativistic Fe Kα line is small in BLS1 galaxies.

#### Introduction

- $\rightarrow$  Broad Fe K $\alpha$  line is significantly detected in ~50% of AGN (e.g. Nandra et al. 2007).
- → It is important in measuring the BH spin, understanding the growth of BH as well as the accretion history.
- The reason for the lack of apparent relativistic broad Fe K $\alpha$  line in the X-ray spectra of some AGN is still unclear.

### Sample & Method

➡ 8862 Type I AGN selected from SDSS-DR4 data (Dong et al. 2012).

### Results I: stacked spectra of $\lambda_{Edd}$ sub-samples



- $\bullet$  A significant broad Fe K $\alpha$  line is detected in the high  $\lambda_{Edd}$  sub-sample.
- $\bullet$  The line width of the Fe K $\alpha$  line increases with  $\lambda_{Edd}$ , consistent with lnoue et al. 2007
- ➡ 156 sources are detected by XMM-Newton with high signal-to-noise ratio.
- The sample is divided into different subsamples, see Fig. I
- The X-ray spectral stacking method (Corral) et al. 2008) is applied to each sub-sample.



## Results II: stacked spectra of NLSI/BLSI sub-samples



 $\bullet$  A broad Fe K $\alpha$  line might be shown in the NLS1 sub-sample, consistent with Liu et al. 2015.

No broad line is shown in the BLSI sub-sample.

Conclusion

9 6 8  $\log M_{\rm BH} (M_{\odot})$ 

Fig I. Distribution of the sources in the  $M_{BH} - \lambda_{Edd}$ plane. The dashed lines mark the boundaries of

different sub-samples.

• Our results indicate a dependence of properties of the broad FeK  $\alpha$  line on the  $\lambda_{Edd}$ .

 $\star$  lonziation state of the accretion disk changes with  $\lambda_{Edd}$ ?

 $\star$  Disk truncation?

• A broad line is shown in the NLSI sub-sample, though with low significance.

#### References

- I. Dong X-B et al., 2012, ApJ, 755, 167
- 2. Corral A. et al., 2008, A&A, 492, 71
- 3. Nandra K. et al., 2007, MNRAS, 382, 194
- 4. Liu Z. et al., 2015, MNRAS, 447, 517
- 5. Inoue H., Terashima Y., Ho L. C., 2007, ApJ, 662, 860

Liu Z. et al., 2016, MNRAS, 463, 684; doi:10.1093/mnras/stw2042

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