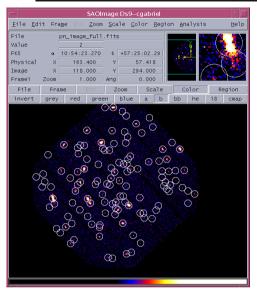
EPIC source catalogues

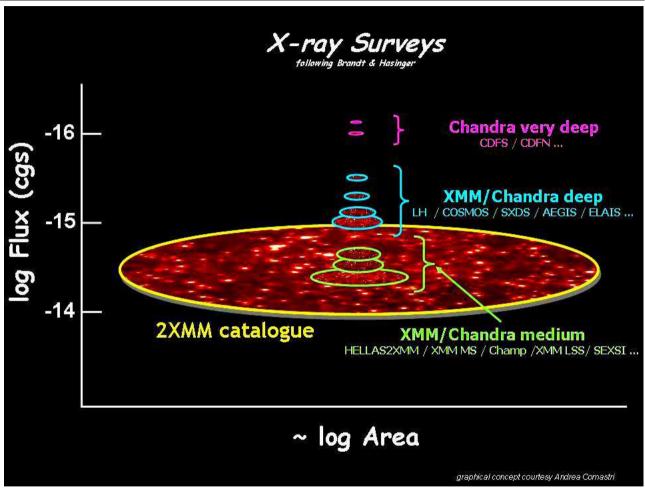
Richard Saxton

June 2014



3XMM – serendipitous sources cat







3XMM: Creation process

- Uniform pipeline processing of all public observations
- Combine detection of sources in all EPIC cameras simultaneously with ML method to maximise sensitivity
- Visual screening to remove obvious spurious sources
- Cross-correlation with standard catalogues (USNO, Rosat 2mass etc.)
- Merge all observations into one catalogue
- First released in August 2007 (increments every year(ish), currently DR4)
- Source-specific products (spectra, lightcurves) produced for brighter sources (>100 counts in combined image)

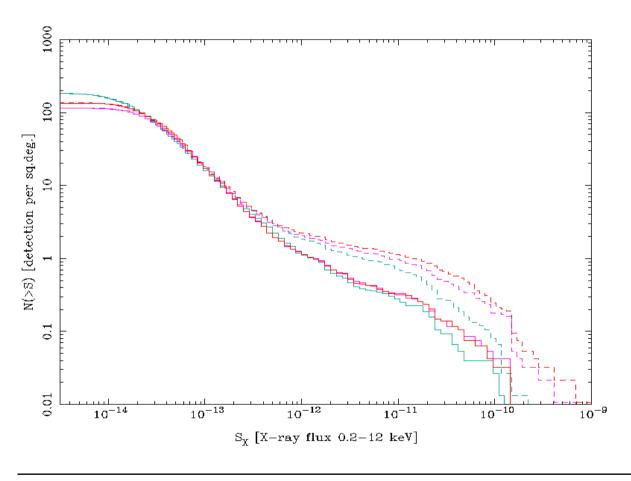


3XMM-DR4: statistics

- 7427 observations.
- ~531000 detections in total.
- ~373000 unique sources
- ~11000 X-ray sources probably extended
- median flux is $F_{2-12}=1.3x10^{-14}$ ergs s⁻¹ cm⁻²
- **2** 20% have $F_{0.2-12}$ <1x10⁻¹⁴ ergs s⁻¹ cm⁻²
- Minimum flux ~1.0x10⁻¹⁵ ergs s-1 cm-2
- Total area independently covered = 794 deg²
- Astrometry generally good to ~1 arcsec
- 248000 sources have spectral products extracted



Source flux distribution

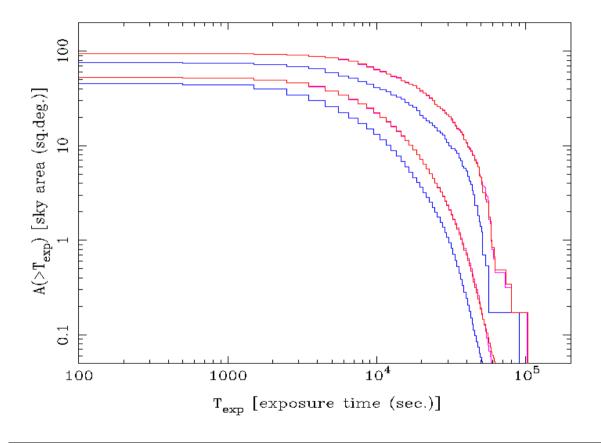


Log n - Log s curves for pn (red), Mos-1 (purple) and Mos-2 (grey).

Dashed curves include the observation targets.



Sky Area



Area covered by catalogue: upper curves are the nominal values for MOS and EPIC-pn; lower curves correct for field overlaps and vignetting.



Important columns

3XMM J010230+640102

- IAU source name

RA_CORR, DEC_CORR, LII,BII - standard coords.

EP DET ML

- maximum likelihood for source

EP_FLUX, EP_FLUX_ERR - 0.2-12 keV flux

M1_n, M2_n, PN_n

- count rate in band 'n' in each camera

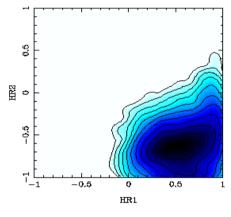
CAT NAME 1

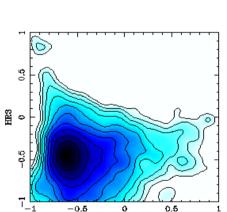
- cross-correlation results

Band	Energy (keV)
1	0.2 - 0.5
2	0.5 – 1.0
3	1.0 – 2.0
4	2.0 – 4.5
5	4.5 – 12.0

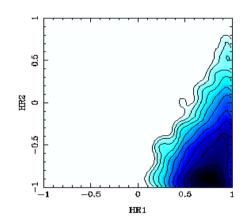


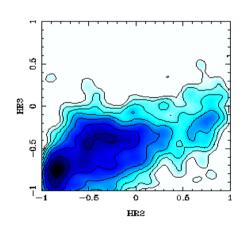
Hardness ratios





HR2





The bulk spectral properties of the EPIC_pn catalogue detections.

$$HR1 = (F2 - F1) / (F2 + F1)$$

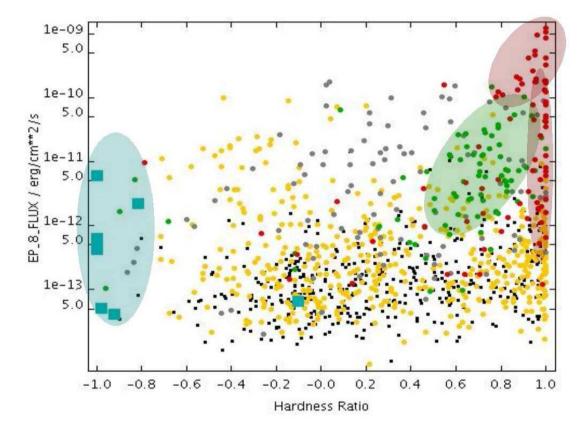
$$HR2 = (F3 - F2) / (F3 + F2)$$

$$HR3 = (F4 - F3) / (F4 + F3)$$

Left panels are for high galactic latitude, right panels are for low galactic latitude.



Hardness ratios of variable sources in 3XMM



HR = (F2 - F1) / (F2 + F1)

F1 = 0.2 - 1 keV fluxF2 = 1 - 12 keV flux

EP_8_FLUX is 0.2-12 keV

Red = X-ray binaries

Green = CVs

Grey = AGN

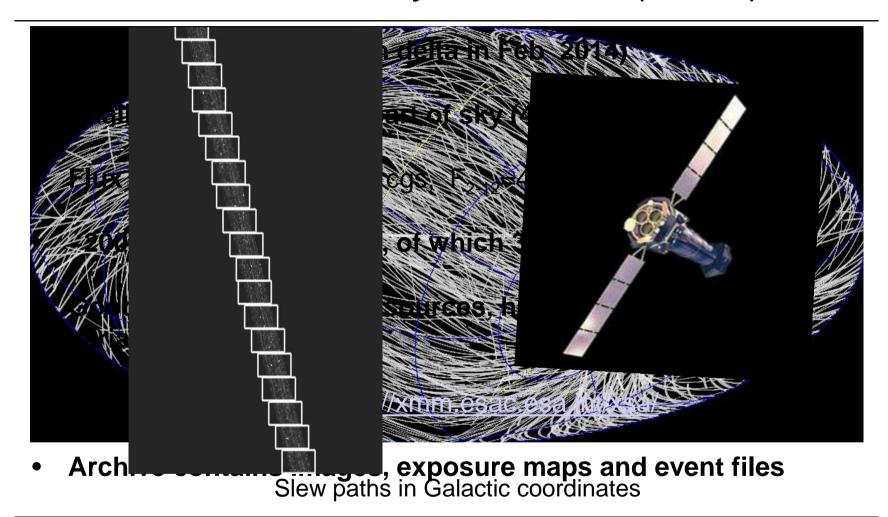
Yellow = Stars

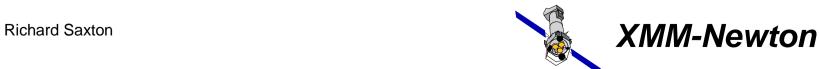
Turquoise = SSS

Gosling et al. 2008

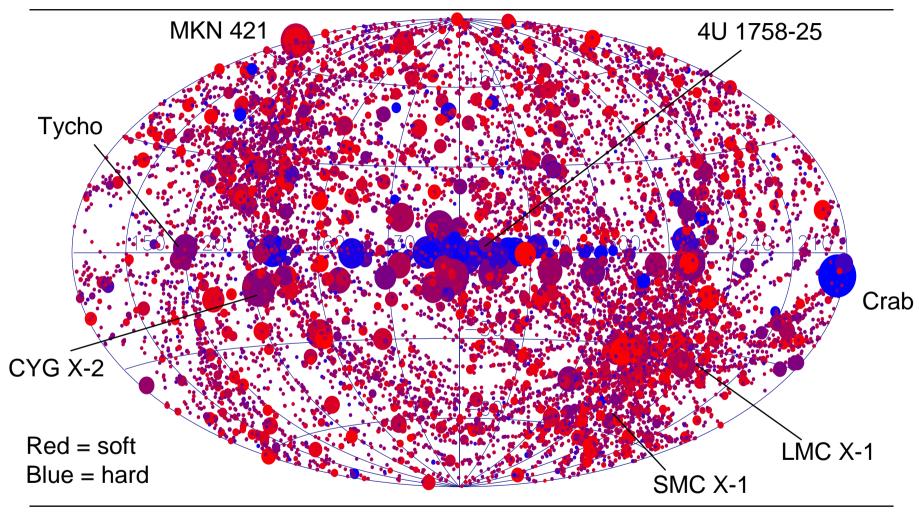


XMM slew survey: XMMSL1(Delta-6)





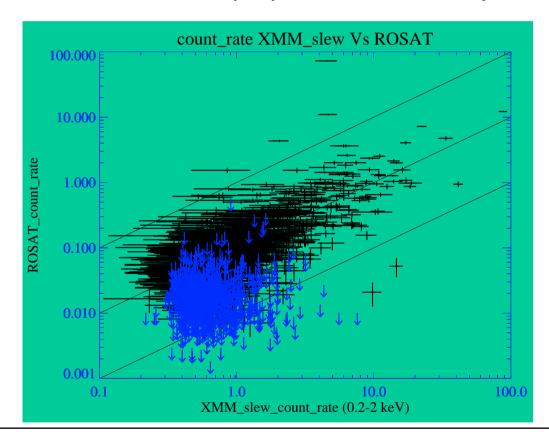
Slew sources



XMM-Newton

Correlations with ROSAT

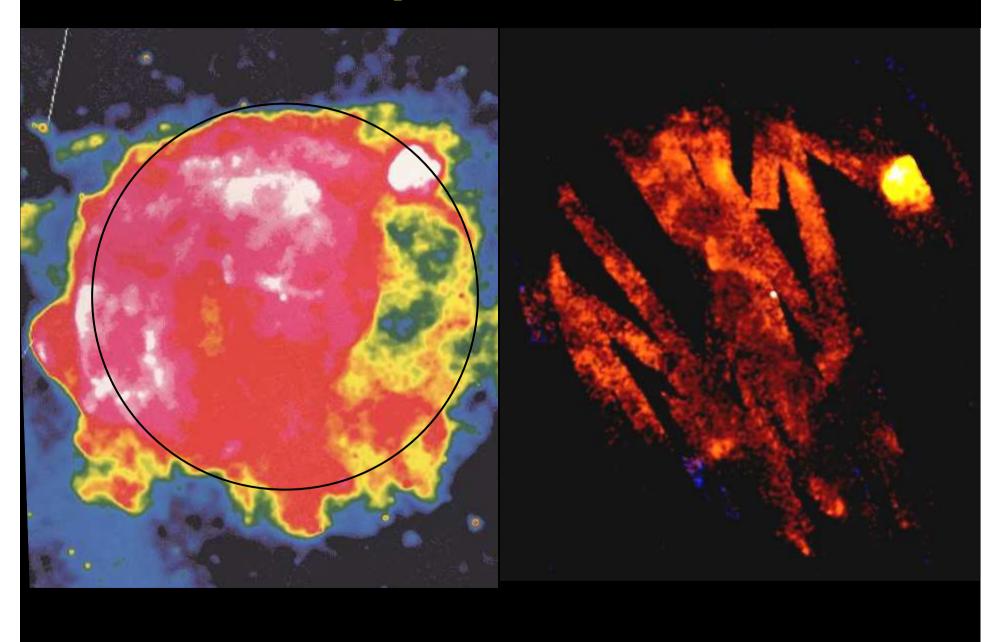
- 55% of sources correlate with ROSAT (non-extended, DET_ML>10)
- 1% show variability by > factor of ten (5% of AGN)



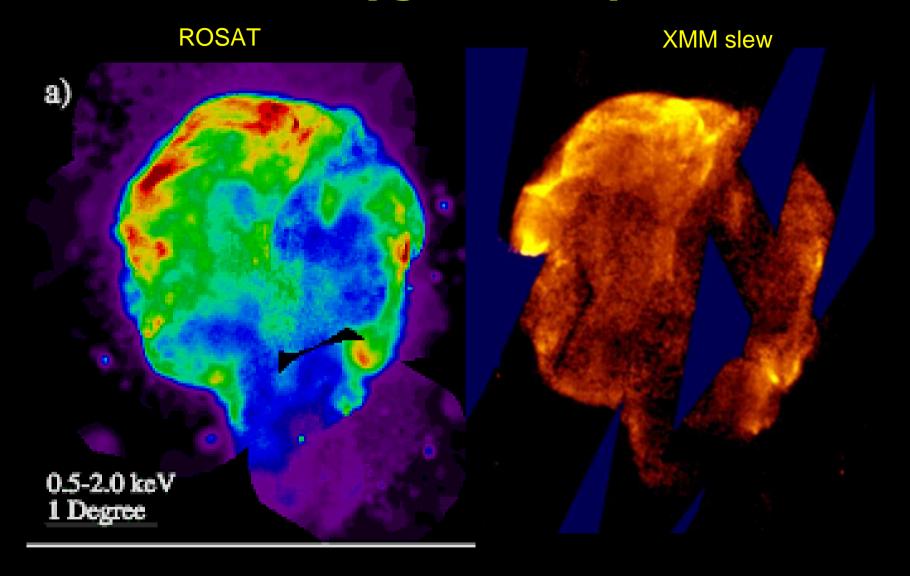
Mean rate: XMM-slew/ROSAT~7



Vela Supernova Remnant

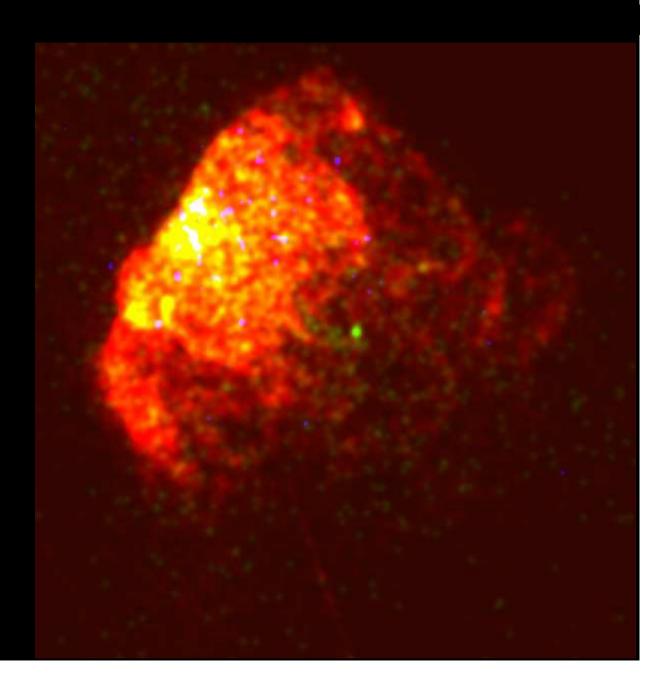


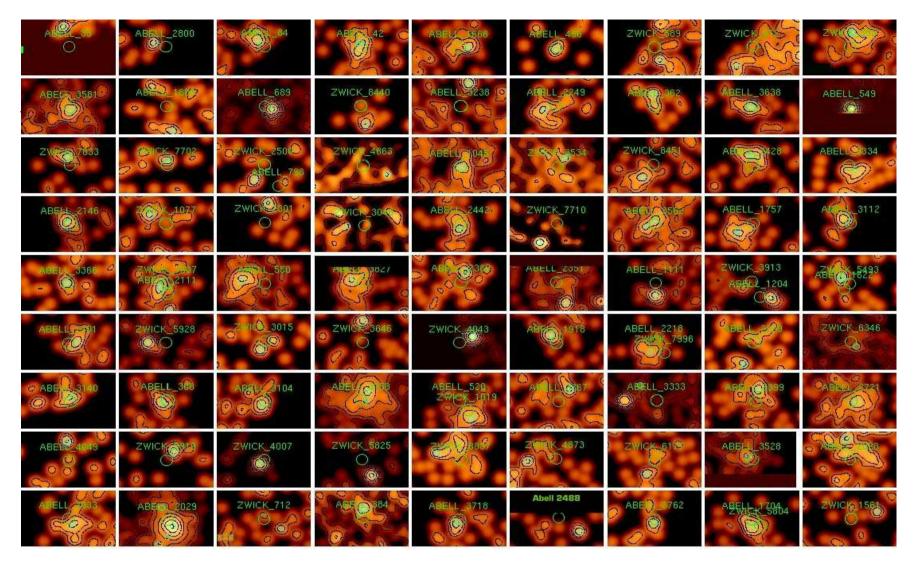
Cygnus Loop



Analysing extended structures- Puppis SNR

XSA contains slew images, exposure maps and event files. Can perform spectral analysis of extended structures.





>200 clusters of galaxies detected



Summary

- > 3XMM, XMMSL1 relatively unexplored at present.
- Data are public
- ➢ Slew survey catalogue currently covers 65% of sky. Addition of 4000 deg^2 per year.
- > 3XMM covers 2% of sky 500k sources



How to access the catalogue

- o Fits file from http://xmmssc-www.star.le.ac.uk/
- o The XSA http://xmm.esac.esa.int/xsa/
- o SCI-DB http://xcatdb.u-strasbg.fr/xcat-db/
- o User documentation: http://xmmssc-

www.star.le.ac.uk/Catalogue/UserGuide_xmmcat.html

o LEDAS http://ledas-www.star.le.ac.uk

Current ref: Watson et al. 2009, A&A 493, 339-373 (2XMM) Saxton et al. 2008, A&A 480, 611 (XMM slew)

