# The geometric origin of quasi-periodic oscillations in black hole binaries

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# **Black Hole Binaries**

- Unique laboratories for relativity
- Too small to directly image
- Must infer geometry from spectral & timing properties



# **Truncated Disk Model**







e.g. Done, Gierlinski & Kubota (2007)

# **Truncated Disk Model**





### Quasi-periodic oscillations



### Quasi-periodic oscillations



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A spinning black hole **distorts** space and time The satellite's motion is **influenced** by the spin of the black hole



# FLOW DISK



Solid body precession at average LT frequency

Viscosity aligns inner regions with the BH and outer regions with the binary partner

Bardeen & Petterson (1975)

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Fragile et al (2007)



Ingram, Done & Fragile (2009)

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Tell-tale sign of relativistic precession: a rocking iron line





# 11/22 Phase Resolving Energy × Flux Energy

#### Geometric origin $\rightarrow$ Line EW varies

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#### Geometric origin $\rightarrow$ Line EW varies



Non-geometric origin -> Line EW in phase with continuum



Non-geometric origin -> Line EW in phase with continuum

#### Periodic function: constant phase difference



Ingram & van der Klis (in prep)

#### Periodic function: constant phase difference



Ingram & van der Klis (in prep)

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Quasi-periodic function: changing phase difference



...but does the phase difference vary randomly or around a well defined mean?

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Quasi-periodic function: changing phase difference



Split long light curve into many segments and measure the phase difference  $\psi$  for each segment

Phase difference varies around a mean: there is an underlying waveform N° of segments (normalised) 0.5 1 0.5 2 1.5 0

ψ/π

Ingram & van der Klis (in prep)

Phase difference varies around a mean: there is an underlying waveform Obs 1 Obs 2 N° of segments (normalised) 0.5 1 Obs 1:  $v_{qpo} \sim 0.46 Hz$ Obs 2:  $v_{qpo} \sim 2.25 Hz$ 2 0.5 1.5 0

ψ/π

Ingram & van der Klis (in prep)

Phase difference varies around a mean: there is an underlying waveform Obs 1 Obs 1: 5000  $v_{\text{qpo}} \sim 0.46 \text{Hz}$ Counts/s 4000 Obs 2 Obs 2:  $v_{qpo} \sim 2.25 Hz$ 3000

Phase (QPO cycles)

0.5

0

Ingram & van der Klis (in prep)

2

1.5

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Enables waveform fitting of QPO models



Can now reconstruct waveform in each energy band using phase lag with the full band



Ingram & van der Klis (in prep)



Ingram & van der Klis (in prep)

**Observation 1** 

**Observation 2** 



Ingram & van der Klis (in prep)



Ingram, van der Klis & Done (in prep)



Ingram, van der Klis & Done (in prep)

# Conclusions

- The QPO does have an underlying waveform
- Waveform fitting will provide a diagnostic tool
- Modulations seen in spectral parameters: photon index, disk temperature & Fe line equivalent width
- Variation in Fe line equivalent width is strong evidence for geometric QPO origin
- Need better data to see a rocking iron line shape





### Testing precession











Ingram, van der Klis & Done (in prep)

### Quasi-periodic oscillations

