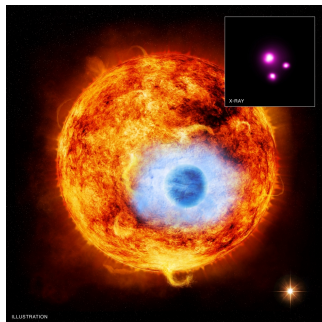
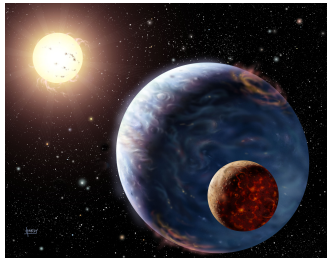
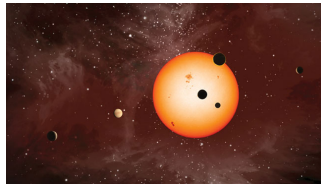
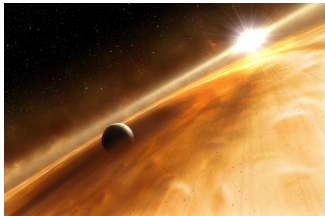


Extrasolar planets and their hosts:  
Why exoplanet science needs  
X-ray observations

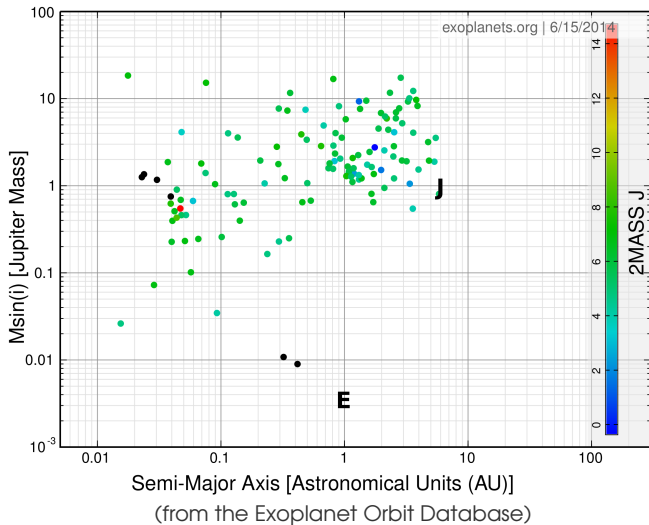
**Dr. Katja Poppenhaeger**  
Sagan Fellow

Harvard-Smithsonian Center for Astrophysics

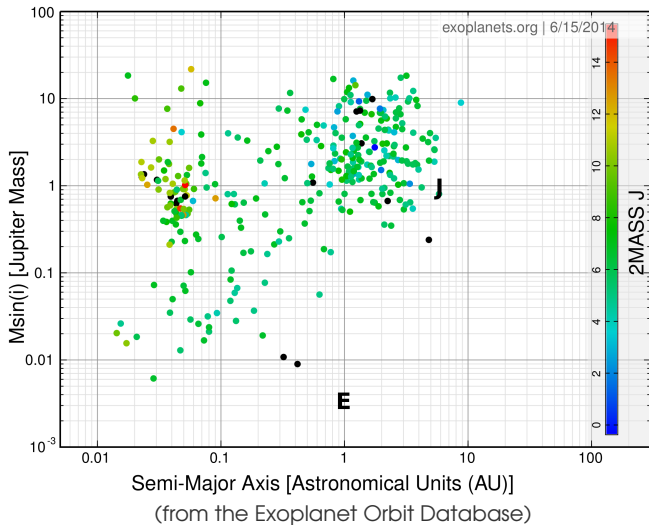
# Exoplanets



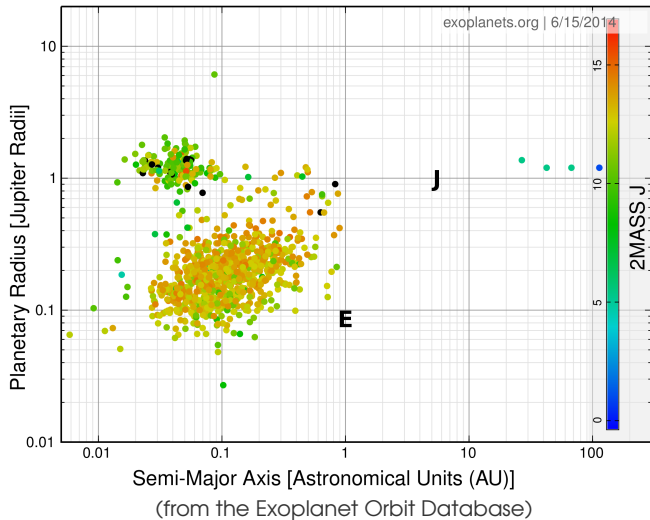
# Exoplanets in 2005



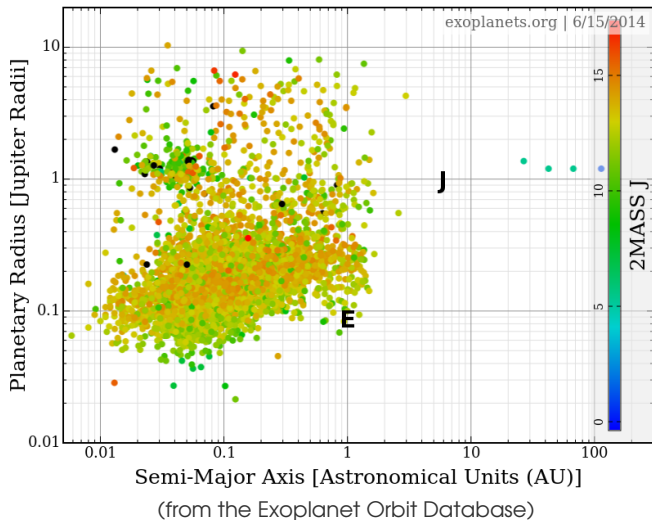
# Exoplanets in 2010



# Exoplanets in 2014



# Exoplanets in 2014, including Kepler candidates



# Exoplanets at AAS

Winter meetings of the American Astronomical Society:  
press releases:

[exoplanets](#) vs. everything else

AAS 2011 Seattle:	<a href="#">11%</a> (4/38)
AAS 2012 Austin:	<a href="#">12%</a> (4/34)
AAS 2013 Long Beach:	<a href="#">29%</a> (13/45)
AAS 2014 National Harbor:	<a href="#">31%</a> (9/29)

# Exoplanets at AAS

Winter meetings of the American Astronomical Society:  
press releases:

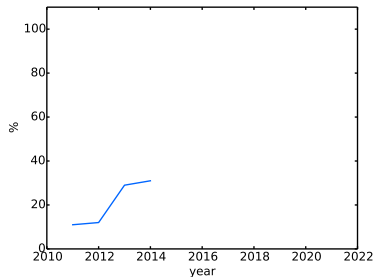
exoplanets vs. everything else

AAS 2011 Seattle: 11% (4/38)

AAS 2012 Austin: 12% (4/34)

AAS 2013 Long Beach: 29% (13/45)

AAS 2014 National Harbor: 31% (9/29)





# Exoplanets at AAS

Winter meetings of the American Astronomical Society:  
press releases:

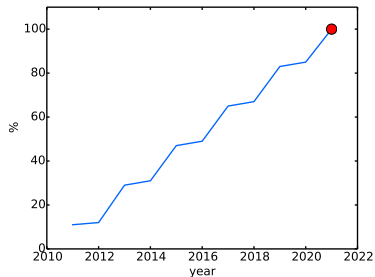
exoplanets vs. everything else

AAS 2011 Seattle: 11% (4/38)

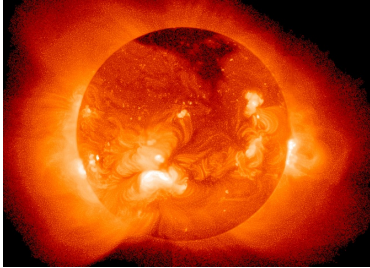
AAS 2012 Austin: 12% (4/34)

AAS 2013 Long Beach: 29% (13/45)

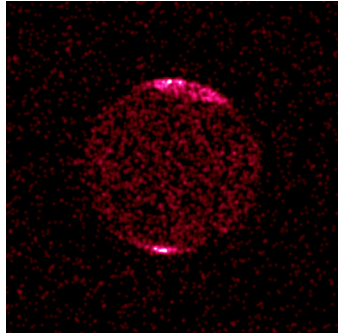
AAS 2014 National Harbor: 31% (9/29)



# X-ray emission in exoplanet systems

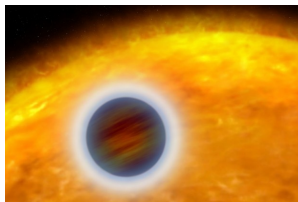


central star (if cool star):  
magnetic activity, flares



planets:  
aurorae, charge exchange

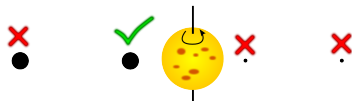
# Stellar high-energy emission



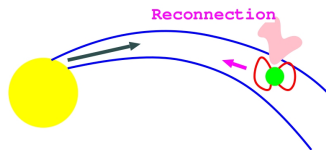
planetary **atmospheres**:  
use star as background  
candle



planetary **mass loss**:  
stellar high-energy  
emission is driver for  
evaporation

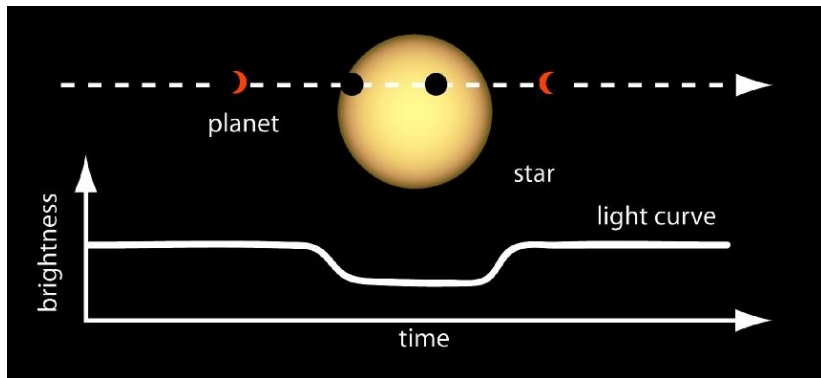


stellar **magnetic activity**  
biases planet detections



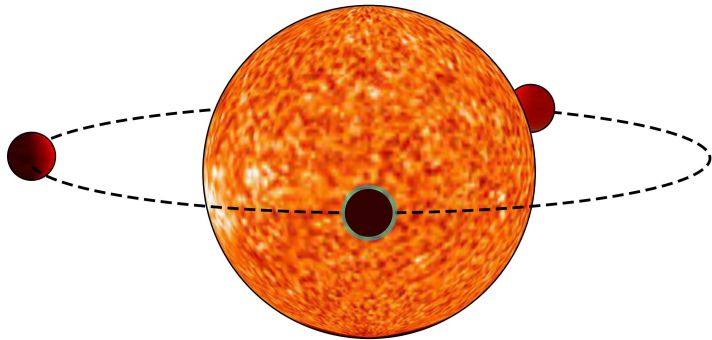
**interactions** of planets with  
stellar rotation/activity

# Exoplanets: transits



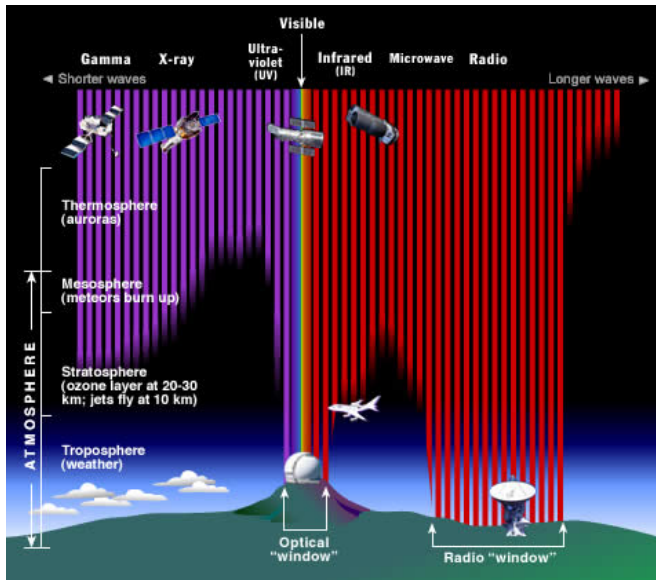
picture credit: NASA

# Exoplanet atmospheres: transmission spectroscopy



K. B. Stevenson

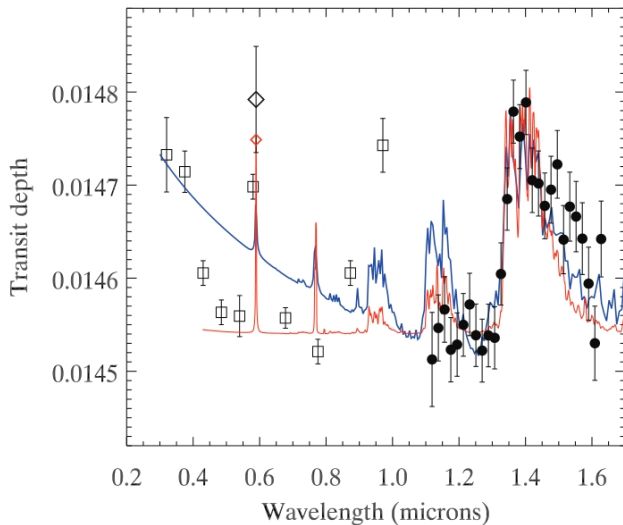
# Exoplanet atmospheres: transmission spectroscopy



picture credit: NASA

# Exoplanet atmospheres: transmission spectroscopy

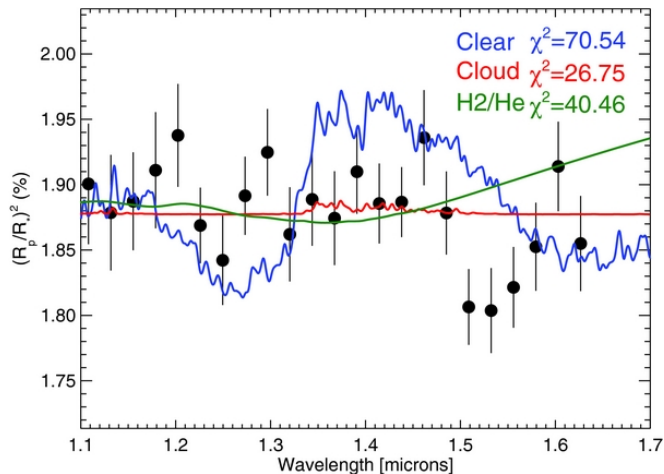
HD 209458b (hot Jupiter)



Deming et al. 2013

# Exoplanet atmospheres: cloud layers

HAT-P-12b (warm Saturn)

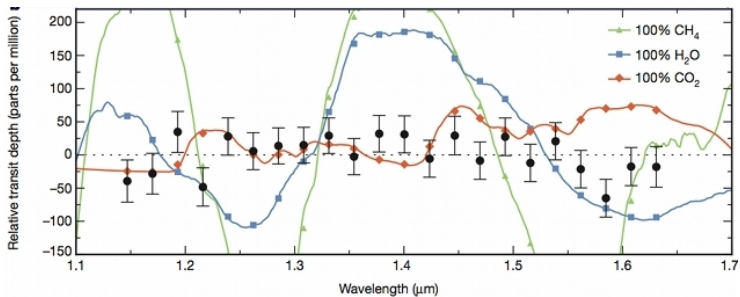


Line et al. 2013



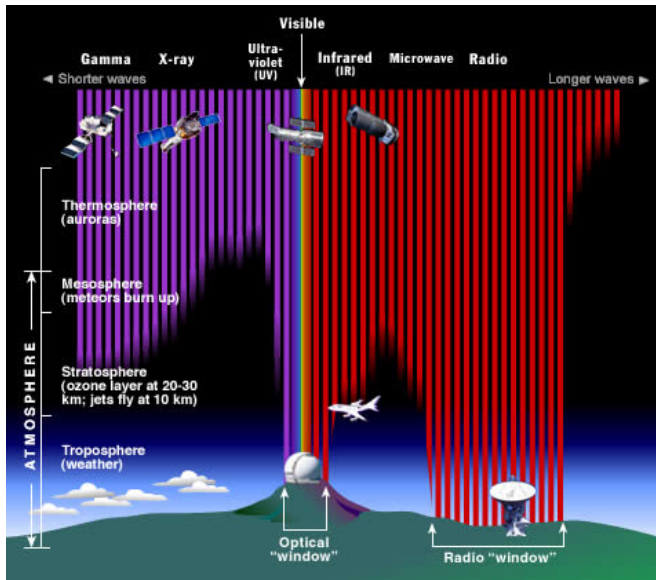
# Exoplanet atmospheres: cloud layers

GJ 1214b (super-earth)



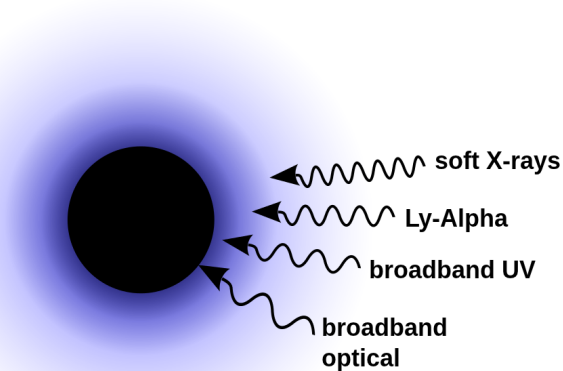
Kreidberg et al. 2014

# Exoplanet atmospheres: transmission spectroscopy



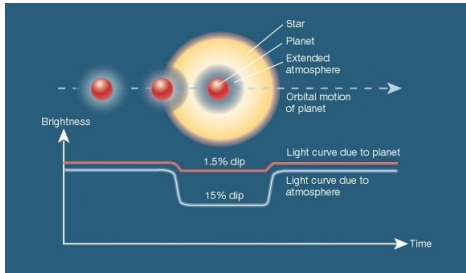
picture credit: NASA

# Exoplanet atmospheres: outer layers



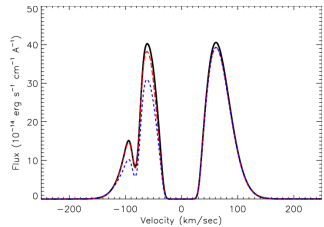
Poppenhaeger et al. 2013

# Exoplanet atmospheres: outer layers



picture credit: D. Charbonneau / Nature

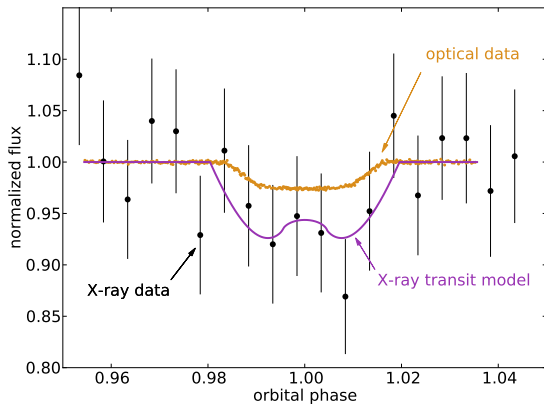
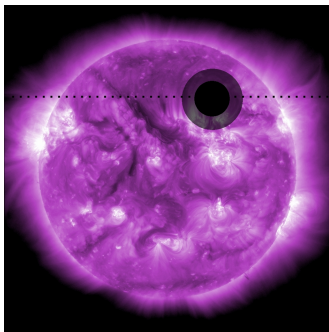
deeper transits



Lecavelier des Etangs et al. 2010

moving (evaporating)  
atmosphere: line profiles

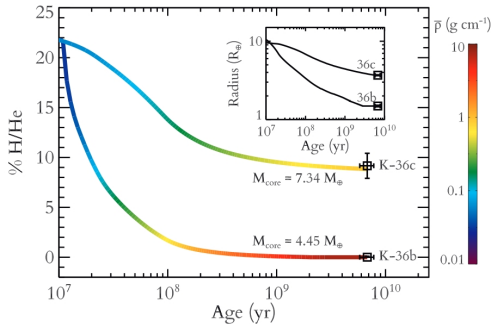
# X-ray transits (hot Jupiter HD 189733b)



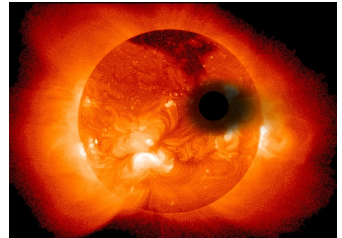
5 X-ray observations co-added, 0.2-2 keV  
Poppenhaeger et al., ApJ 2013

# Atmospheric evaporation, driven by X-rays and UV

## Evaporation of gaseous envelope

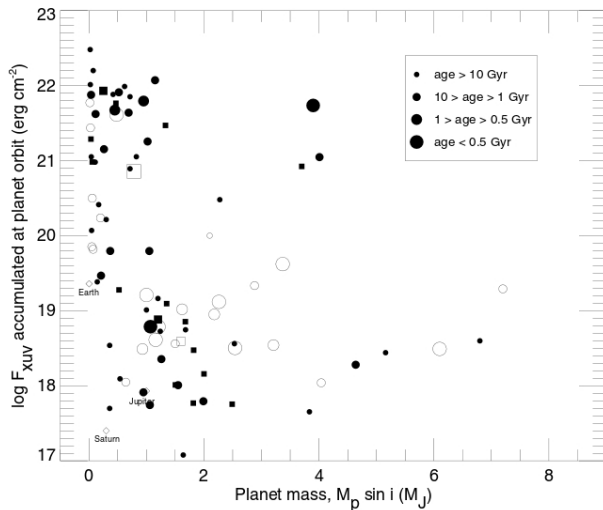


Lopez et al. 2013



YOHKOH/modified by K.P.

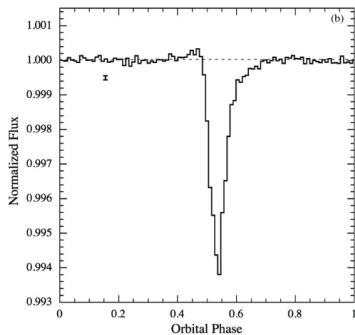
# Imprints of evaporation in observed exoplanet population



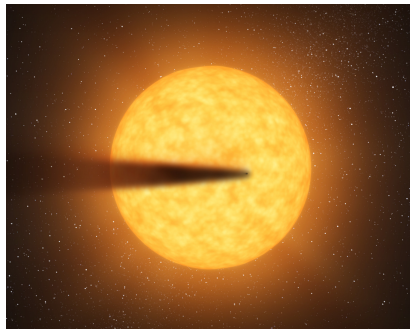
Sanz-Forcada et al. 2011

# Evaporating planets - transit profiles

KIC 12557548, dusty tail?



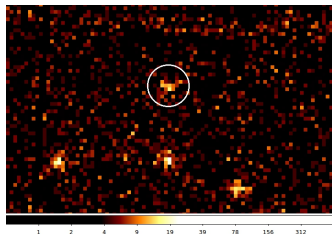
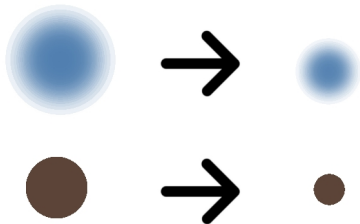
Rappaport et al. 2012



picture credit: NASA/JPL/Caltech



# Small planets can lose lots of mass



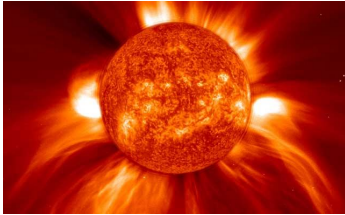
XMM detections of host stars:

water-planet GJ 1214 b (Lalitha et al. 2014 submitted),  
rocky planet CoRoT-7 b (Poppenhaeger et al. 2012)

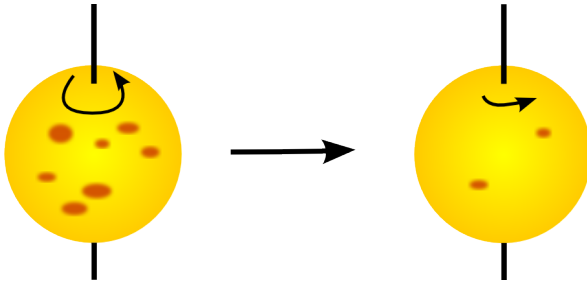


→ mixed systems like KOI-314?

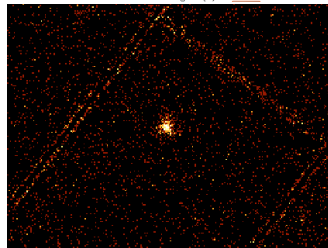
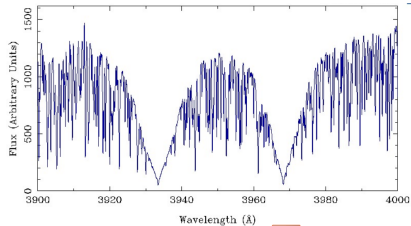
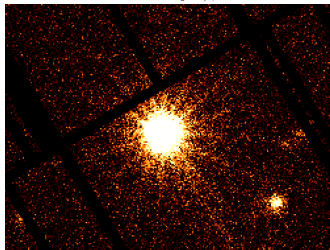
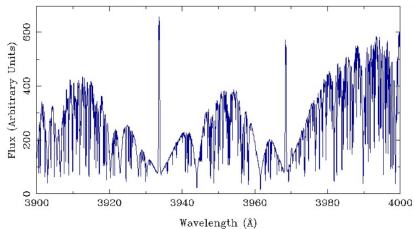
# Decline of magnetic activity: magnetic braking



loss of angular  
momentum through  
stellar wind



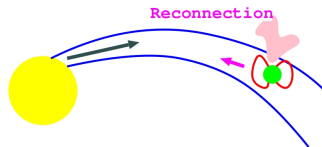
# Activity decline with stellar age



# Star-planet interaction

(template: star-star interaction in close binaries)

2 basic scenarios:



magnetic

discovery papers:  
Shkolnik et al. 2005, 2008  
for 2 individual systems

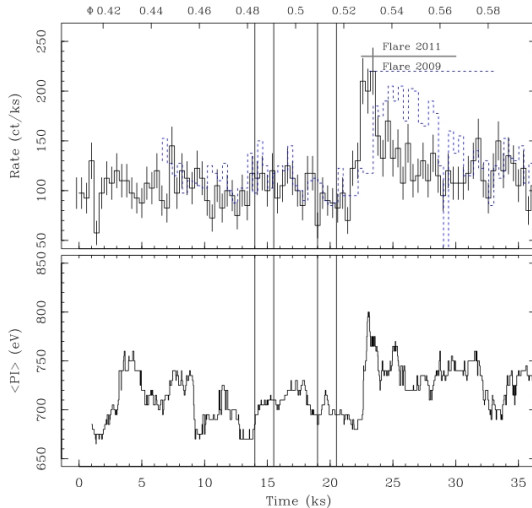


tidal

spin-up (inhibited  
spin-down) of host star;  
stronger for thick outer  
convection zones

# Individual planet-host stars: Flare triggering?

HD 189733b (hot Jupiter)



Pillitteri et al. 2011

# Stellar samples: planet-induced activity enhancements?



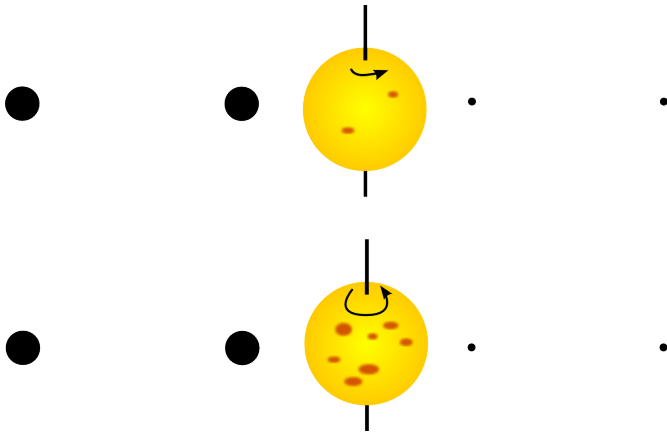
high activity?



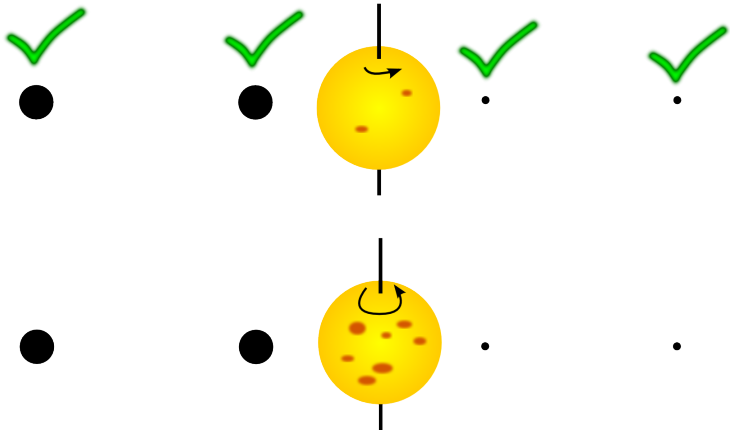
low activity?

- ▶ searching for trends in samples of planet-hosting stars: Kashyap et al. 2010, Poppenhaeger et al. 2010, 2011, Lanza 2011, Shkolnik 2013, Miller et al. 2012, 2013, and others
- ▶ caveat: stellar activity biases against planet detection!

# Dealing with selection effects

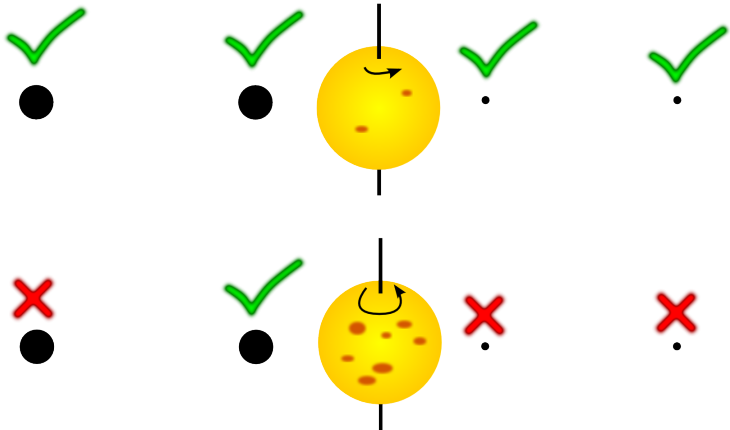


# Dealing with selection effects

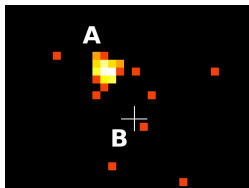




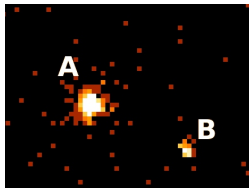
# Dealing with selection effects



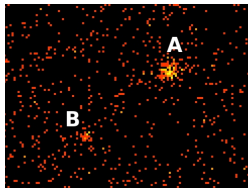
# Systems with a built-in negative control: Planet-hosting stars with stellar companions



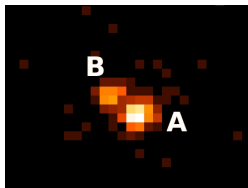
CoRoT-2 AB



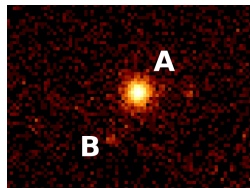
HD 189733 AB



55 Cnc AB

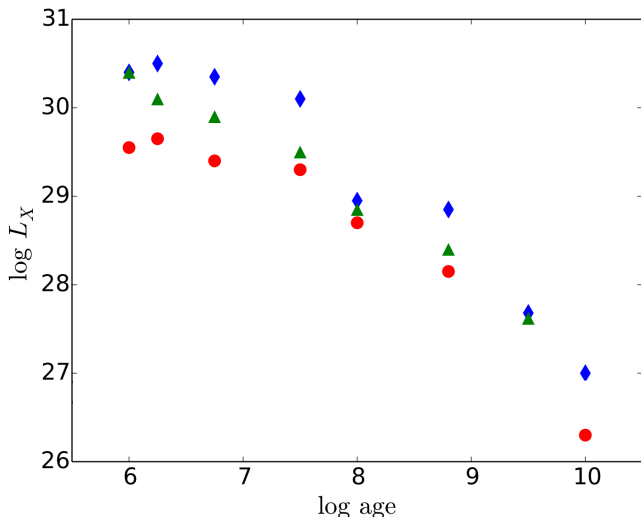


$\tau$  Boo AB

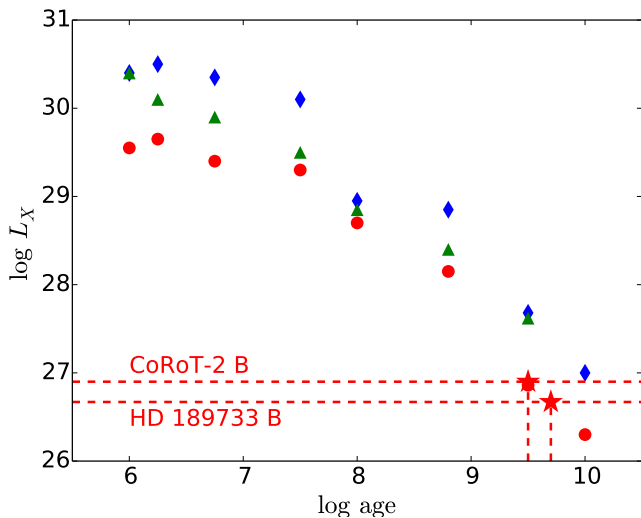


$\nu$  And AB

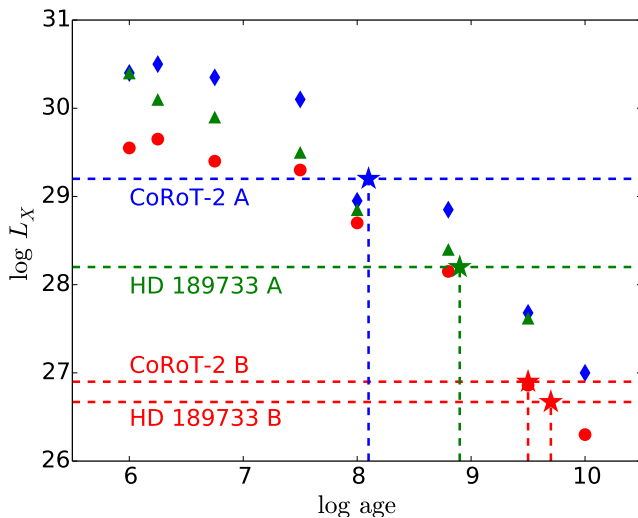
# Stellar age and X-ray luminosity



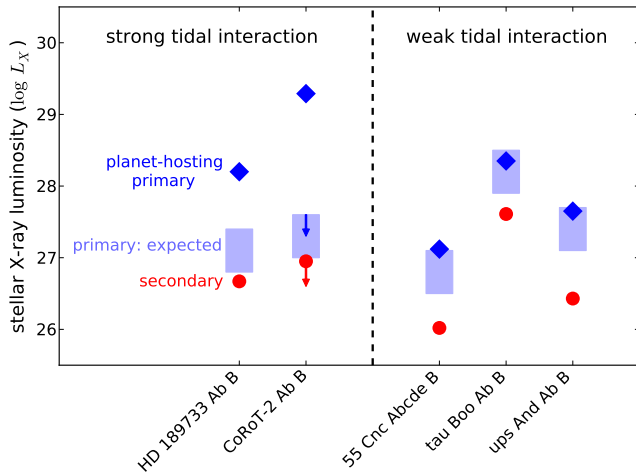
# Stellar age and X-ray luminosity



# Stellar age and X-ray luminosity

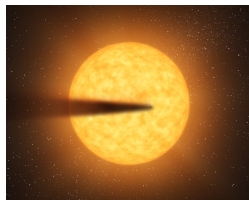
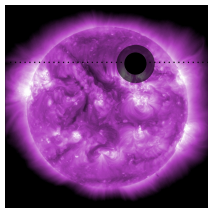
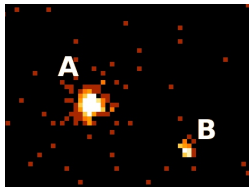


# X-ray activity for 5 systems



Poppenhaeger et al. 2014, A&A Letters

# X-rays: exoplanets & host stars



- ▶ stars: activity biases in exoplanet samples
- ▶ stars: tidal / flare triggering interactions of planets and stars
- ▶ exoplanets: short wavelengths probe outer atmospheres, via transit depths or line profiles
- ▶ exoplanets: X-ray/UV-driven evaporation
- ▶ upcoming opportunities: Athena+, Astro-H, eRosita
- ▶ Arcus: X-ray grating!

# NASA SMEX proposal: X-ray spectrograph Arcus

