

Predicting Jets in Cygnus X-2

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Cygnus X-2

Luminous neutron star LMXB: L > 10³⁸ erg s⁻¹ (> L_{Edd})

 Z-track source : Cyg X-2, GX 5-1, GX 340+0, Sco X-1, GX 349+2, GX 17+2 and LMC X-2 Transient XTE J1701-462

- Long-term variability: non-periodic: 50 90 days
- Radio emission shows relativistic jets: observed only in certain states of the source
- Need observations of > 2 days: X-ray + radio –
 For movement between states jets may not be seen
- Can we find a trigger that jet launch imminent







Low Mass X-ray Binaries



Nature of Island/ Banana states have not been understood

Nature of HB, NB and FB have not been understood

Extended ADC model for Z-sources (*)

- Based on realization that accretion disk corona is extended in LMXB (dipping studies and work of Schulz on line emission of the ADC)
- Investigations of the Z-sources: Cyg X-2, GX 340+0, GX 5-1

Sco X-1, GX 17+2 GX 349+2

based on the extended ADC

provide a convincing physical explanation

- Mdot increases on NB (decreases on HB)
 Radiation pressure launches jet
- Flaring Branch = Unstable Nuclear Burning

Mdot > Mdot_{crit} --- Stable burning Mdot < Mdot_{crit} --- Unstable burning





High radiation pressure launches jet



Jets:

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large increase in $T^4 =>$ high radiation pressure \rightarrow disk disruption \rightarrow jet



Cygnus X-2 in the RXTE ASM

- I6 years of data (February 1996 December 2011)
- We have superimposed all pointed data on the ASM data
- Aim is to recognize the 3 states of the Z-track
- In particular, the state that produces radio jets



- 8% of ASM data simultaneous with RXTE/PCA pointed observations
- Duration of the XTE pointed observations: 50 days

Cygnus X-2: Longterm variability



Power spectrum = Porb + harmonics (Boyd & Smale 2004)

Dark points: ASM data simultaneous with the pointed observations

Cygnus X-2: RXTE/PCA Pointed Data



Case A: Complete Z-track

The Soft Apex in the RXTE/ASM



Only 3 points (green) at real Soft Apex (I \sim 30 count s⁻¹)

Other points (red) - in many obs. the Soft Apex never reached (I > 35 count s⁻¹) Reason: Mdot rarely falls to Mdot_{crit}

> Only at minima of long-term variability This clearly suggests that long-term variability is due to change of Mdot

The Hard Apex in the RXTE/ASM



Majority of Hard Apex points: **ASM intensity > 40 count s⁻¹ Width of ASM track = Z-track motion on NB**: 20 – 45% in intensity Superimposed on long-term variations

The Hard Apex can be at various count rates depending on the level of the long-term variability

Cygnus -2: Radio detections in the RXTE era



- 2013 Only observation to catch switch-on of jets: simultaneous X-ray (Swift) and eVLBI Spencer, Rushton, Balucinska-Church, Paragi, Schulz, Wilms, Pooley, Church; MNRAS (2013)
- 2009 simultaneous X-ray (XMM) and eVLBI; only radio flux upper limit Balucinska-Church, Schulz, Wilms, Gibiec, Hanke, Spencer, Rushton, Church; A&A (2011)
- 2005 simultaneous X-ray (RXTE) and Merlin observations (Rushton PhD)

Simultanous X-ray/radio observations of Cygnus X-2: Swift and European VLBI network

(Spencer, Rushton, Balucinska-Church, Paragi, Schulz, Wilms, Pooley, Church; MNRAS (2013)



Observations: 22 - 23 February 2013 (Hard Apex / Horizontal Branch) X-ray luminosity (1 - 30 keV): $1.3 - 1.7 \cdot 10^{38}$ erg/s (9 kpc)



Simultanous X-ray/radio observations of Cygnus X-2: Swift and European VLBI network

Spencer, Rushton, Balucinska-Church, Paragi, Schulz, Wilms, Pooley, Church; MNRAS (2013)





Peak flux density/rms:

22 Feb 2013: 523/20 microJy beam⁻¹ 23 Feb 2013: 583/30 microJy beam⁻¹

23 Feb, 2013: Core – SE : ~ 5 mas

jet velocity v/c = 0.33 +/- 0.12 weak detection of a counter jet

Jet launching criterion



MAXI for jet launching ?

Jets are seen at > 40 count s⁻¹ in the ASM i.e. when Cygnus X-2 gets to the Hard Apex by ascending the NB of the Z-track



40 count s⁻¹ in the RXTE ASM is equivalent to 1.8 count s⁻¹ in MAXI

Thus low count statistics make jet prediction more difficult

However, the longterm variability can still be monitored to detect the periods of high luminosity

A Unified Model of LMXB

- We want to predict jets for *more certain triggering* of observations
- More important to understand the physics of jet formation
- We have a model of the Z-track sources predicting jet formation

 We now have model of Island and Banana states in Atolls predicting that jets may be seen in Atolls

Combining these gives a Unified Model of LMXB (Church et al. 2014)

Unified Model of LMXB* MNRAS 438, 2784-2797 (2014)



Thermal equilibrium:

electron temperature of Comptonizing region equals NS temperature for *all LMXB* with $L > 1 \cdot 10^{37}$ erg s⁻¹

* See Poster B01

The Unified Model: Atoll sources



Banana State: the basic state of Atolls: $L > 1 \cdot 10^{37}$ erg s⁻¹ thermal equilibrium $kT_{NS} = kT_{ADC}$ motion on banana due to Mdot variations (like NB of Z-sources)

Island state: $L < 1 \cdot 10^{37}$ erg s⁻¹ thermal equib lost: $kT_{ADC} >> kT_{NS}$ heating of ADC by unknown mechanism

Radiation pressure

measured values can be high suggesting jet formation

The Unified Model

Atoll sources:

basic state Mdot varies	
ADC heated producing hard spectrum	
may also be formed by high radiation pressure	
Z-track sources:	
Mdot increases as in Banana state	
Unstable nuclear burning when	

Mdot falls below critical value for Unstable Nuclear Burning on NS (as in Bildsten theory 1998)

Horizontal Branch: recovery after jet launch

Flaring Branch: not possible in Atolls – far from critical value Island Branch: not possible in Z-sources as $kT_{ADC} = kT_{NS}$ always

CONCLUSIONS

- ASM data or similar can be used to predict jet launching
- Jets are formed at the Hard Apex of the Z-track strongly suggesting that the high, measured radiation pressure launches the jets
- High radiation pressure under certain conditions in Atolls => jets