XMM-Newton Future Operational Ground Segment

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Introduction

- ith the majority of visibility from the southern hemisphere

- Thermal tensitivity of Instruments may require immediate reaction to on-shoard event Currently XMM is supported routinely by 6 fifteen metre antennas located in Australia (Perth and Dorgarea), South America (Keurou and Santiago de Chile) and Spain (Villafrence and Maspalanas)
- provide additional support in the case of unavailability of Perth or Dongara due to LEOPs or outages



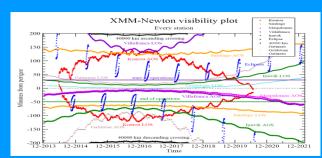
Operational Ground Segment Evolution

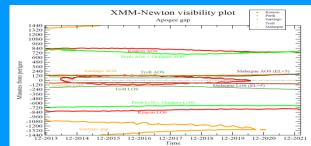
Figure 1 : ESTRACK Network

- NMM-Newton has undertaken a number of lifetime enhancing measures in the last few years 4 Wheel Drive : Slews are now performed using all loar instead of three reaction wheels on-board This has resulted in fiel savings of approximately 50%.

 - Anti-Reaction Wheel Coging Strategies Upgrade of the Mission Control System to modern virtual machines running on Solaris 10
 - Pleixible-Perigee passage optimizing the spacecraft attitude to allow more efficient slewing to first scientific target
- The end result is a dramatic improvement in fuel consumption (\sim 50%) and robolescence and/or catastrophic failure extending mission life to potentially ~2028

- Orbital evolution during the next 6 years leads to the opening of a large gap in visibility Mission extension past end-of-2015 decommissioning of Perth





Ground Station Requirements

- Antenna Sensitivity: -170dBW/som

| Station | | nt-r(mdeg/s) | Azr(mdeg/s) | Elr(mdeg/s) |
|----------------|-----|--------------|-------------|-------------|
| DONGARA (13m) | | | | |
| PERTH (15m) | | | | |
| SANTIAGO (13m) | | | | |
| KOUROU (15m) | | | | |
| VILLA2 (15m) | | | | |
| GOLDSTNE (34m) | 4.2 | | | |

If the tracking rate is too high key-holing can occur effectively limiting the tracking capability of the antenna

Another key concern is the availability of the station, XMM requires continuous uninterrupted coverage from its reas Villafranca2 has a higher loading due to utilization by Cluster and INTEGRAI

| ne final ground station(s) should meet the | following Technical Spo |
|--|-------------------------|
| RX/TX frequency ratio 240/221 in | Requiren |
| coherent mode | RX/TX Frequ |
| Polarisation predominantly RHC | Uplink Transmiss |
| | Downlink Transmi |
| Pointing accuracy sufficient for search pattern | Modulation T |
| | Polarisatio |
| Timing critical for science | EIRP |
| observations (order of magnitude less than millsecs required) | Ranging To |
| | Pointing Accu |
| Ranging critical for orbit | Station Time Assurement |

Ground Segment Requirements

Table 2 : Station Specifications

Ground Segment Architecture Overview

ust be guaranteed to maximise scientific return

| CASE | Coverage | 2016 | 2017 | 2018 | 2019 |
|------|---|-------------|-------------|-------------|-------------|
| 1 | Full Orbit Support | 44:00 hours | 44:00 hours | 44:00 hours | 44:00 hours |
| 2 | Full Perth Replacement | 24:00 hours | 24:24 hours | 24:36 hours | 23:40 hours |
| 3 | Partial Perth Replacement (NNO available) | 15:00 hours | 14:00 hours | 16:30 hours | 12:20 hours |



TMTCS - Handles TC & TM frames across the SLE IFMS – Modulates TC onto carrier and demodulates TM



Conclusions

- XMM-Newton in excellent shape with all units both payload and service module operating
- Successful implementation of 4WD greatly extends potential operational life

- **Robust Requirement defined for Public Tender**
- operation and reduce mission overheads
- Winter eclipse season



References



Figure 6 : New Norcia Deep Space Network Station

- aring Noise Detection, Modelling and Mitigation Measures on ESAs N-Ray Observatory XMM-Newton, Kirsch, Proceeding of 9th GNNC Conference 2014, AAS Rocky Mountain Section 4WD Implementation for XMM_Newton, F. Schmidt, Proceeding of 13th SpaceOps Conference 2014, JPL
- Proceeding of 13th SpaceOps Conference 2014, JPI
- eOps Conference 2014, JPI