Planck Cosmological Parameters

Steven Gratton, on behalf of the Planck Collaboration Ferrara, 1st Dec 2014

Changes for parameters, level 1

- Data from full mission
- Understanding of "4K" line systematics
- Increase in sky area used
- Corrections for low levels of correlated noise between detectors (though negligible for usual parameters)

2013 Survey1 x Survey2 vs 2013 detset nominal



2013 Year1 x Year2 extended sky vs DS nominal & DS full extended sky



2013 nominal to 2013 Y1xY2 extended

Parameter	2013N	2013CYE
100 θ _{MC}	1.04131 ± 0.00063	1.04121 ± 0.00048
$\Omega_{\rm b}h^2$	0.02205 ± 0.00028	0.02230 ± 0.00023
$\Omega_{c}h^{2}$	0.1199 ± 0.0027	0.1188 ± 0.0022
H ₀	67.3 ± 1.2	67.8 ± 1.0
n _s	0.9603 ± 0.0073	0.9655 ± 0.0062
Ω _m	0.315 ± 0.017	0.308 ± 0.013
σ ₈	0.829 ± 0.012	0.828 ± 0.011
τ	0.089 ± 0.013	0.094 ± 0.013

2013: nominal to Y1xY2 extended

Parameter	Δparam (% of 2013N σ's)	σ (% of 2013N σ's)
100 θ _{MC}	-16	76
$\Omega_{\rm b}h^2$	89	82
$\Omega_{c}h^{2}$	-41	81
H ₀	42	83
n _s	71	85
Ω _m	-41	76
σ ₈	-8	92
τ	38	100

Changes for parameters, level 2

- Improved TOI processing
- Calibration and beams
- Alternative implementation of likelihood used
- Improved foreground handling, including dust emission at all frequencies
- Cross-half-mission, not detsets, now the default
- Polarization now an option in the likelihood

So now, with 2014 TT



2013 Y1xY2 E (CamSpec) to 2014 cross-half-mission (Plik)

Parameter	2013CYE	2014 CHM Plik
100 θ _{MC}	1.04121 ± 0.00048	1.04086 ± 0.00048
$\Omega_{ m b}h^2$	0.02230 ± 0.00023	0.02222 ± 0.00023
$\Omega_{c}h^{2}$	0.1188 ± 0.0022	0.1199 ± 0.0022
H ₀	67.8 ± 1.0	67.26 ± 0.98
n _s	0.9655 ± 0.0062	0.9652 ± 0.0062
Ω _m	0.308 ± 0.013	0.316 ± 0.014
σ ₈	0.828 ± 0.011	0.830 ± 0.015
τ	0.094 ± 0.013	0.078 ± 0.019
$10^9 A_s e^{-2\tau}$	1.831 ± 0.011	1.881 ± 0.014

So, how did we do for 2013?

Parameter	Δparam (% of 2014 σ's)
100 θ _{MC}	94
$\Omega_{\rm b}h^2$	-74
$\Omega_{c}h^{2}$	0
H _o	4
n _s	-79
Ω _m	-7
σ ₈	-7
τ	58
$10^{9}A_{s}e^{-2\tau}$	-320

How are we doing for 2014?

Parameter	CamSpec-Plik (% of 2014 σ's)
100 θ _{MC}	17
$\Omega_{\rm b}h^2$	13
$\Omega_{c}h^{2}$	-23
H ₀	22
n _s	48
Ω _m	-21
σ ₈	-7
τ	5
$10^{9}A_{s}e^{-2\tau}$	-43

Now we have 2014 TE & EE...











2014 TT to TT, TE, EE

Parameter	2014 Planck TT + lowP	2014 Planck TT,TE,EE+lowP
100 θ _{MC}	1.04086 ± 0.00048	1.04073 ± 0.00032
$\Omega_{ m b}h^2$	0.02222 ± 0.00023	0.02224 ± 0.00015
$\Omega_{c}h^{2}$	0.1199 ± 0.0022	0.1199 ± 0.0014
H ₀	67.26 ± 0.98	67.22 ± 0.64
n _s	0.9652 ± 0.0062	0.9639 ± 0.0047
Ω _m	0.316 ± 0.014	0.316 ± 0.009
σ ₈	0.830 ± 0.015	0.831 ± 0.013
τ	0.078 ± 0.019	0.079 ± 0.017
$10^{9}A_{s}e^{-2\tau}$	1.881 ± 0.014	1.883 ± 0.012



 $\Omega_{\rm k}$, $m_{\rm v}$ & $N_{\rm eff}$...



Summary

• 2014 TT a significant evolution over 2013

 Improved parameter constraints, driven by larger sky area used and full vs nominal data

- Polarization data now included, though not as well-tested as temperature
- ACDM still seems to fit well, now both the temperature and polarization data, though
 - Marginal tensions (curvature, A_lens... as 2013)
 - low-*l* anomalies

The scientific results that we present today are a product of the Planck Collaboration, including individuals from more than 100 scientific institutes in Europe, the USA and Canada.

