

Automated Multi-Dataset Analysis (AMDA): An online database and analysis tool for heliospheric and planetary plasmas

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DASH meeting, ESAC, 16 October 2024

CDPP, the French national data centre for natural plasmas of the solar system

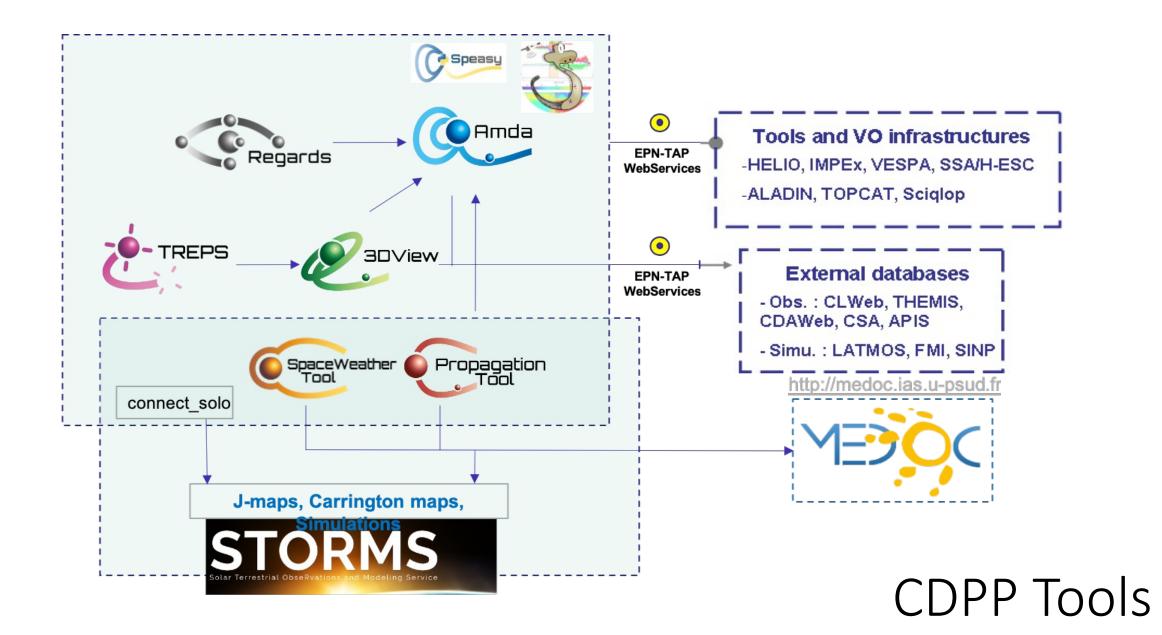




 Created in 1998 jointly by <u>CNES</u> and <u>INSU</u>, the CDPP assures the long term preservation of data obtained primarily from instruments built using French resources, and renders them readily accessible and exploitable by the international community.

Objectives of CDPP

- Data: continue the long-term archiving activity of data obtained by experiments with French participation but also (i) make available recent data on which the community mobilizes and (ii) make them easily usable (extraction in standard formats)
- Services: develop attractive Value-Added Services that save users time and energy and promote an increase in the scientific return on data exploitation
- Interoperability and Virtual Observatories (VO): continued investment by the CDPP in the development of standards and in future VO projects.



http://amda.cdpp.eu/

- A data analysis tool in your browser
 - physical parameters not files !
- Data are
 - replicated from ESA/Cluster Science Archive, NASA/PDS
 - or accessed remotely : CDAWeb, simulation and model databases, ...
 - public or restricted to communities
 - can be exported in companion tools (SAMP)
 - or uploaded by the user
 - can be accessed via web-services (SOAP/REST)
- Sessions are saved (so it's better to register !)
 - register at <u>amda@irap.omp.eu</u>
 - Public access w/o registration also available



AMDA reference paper

Automated Multi-Dataset Analysis (AMDA): An on-line database and analysis tool for heliospheric and planetary plasma data

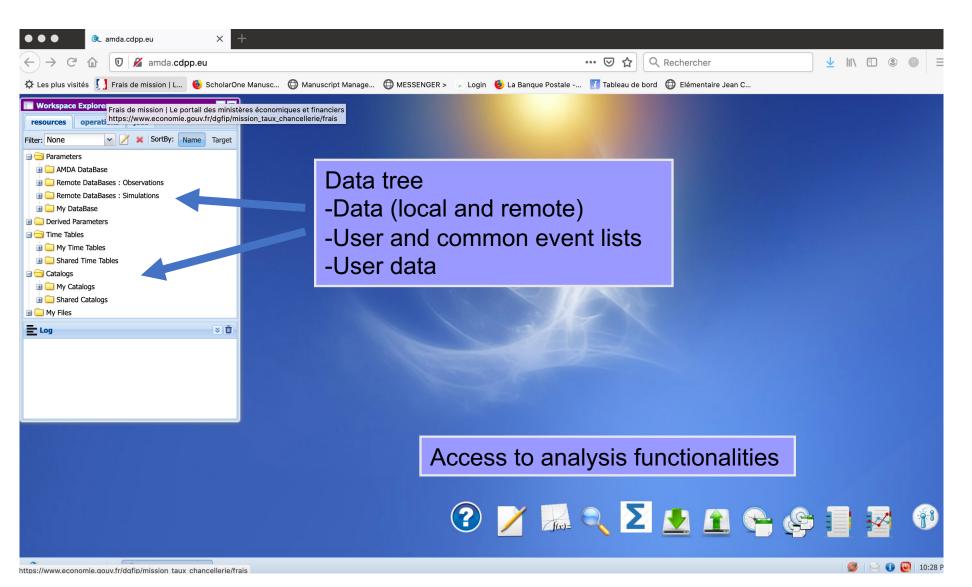
Vincent Génot, E. Budnik, C. Jacquey, M. Bouchemit, B. Renard, N. Dufourg, N. André et al.

Planetary and Space Science, Elsevier, 2021, 201, pp.105214.

https://doi.org/10.1016/j.pss.2021.105214

http://amda.cdpp.eu/





AMDA Database

- More than <u>80 missions</u> (heliosphere, magentosphere, planets, comets, ...) and <u>800 datasets</u> (electromagnetic fields, plasmas, particles, waves, ...)
- AMDA has been integrated in the Space Weather Service Network of the European Space Agency.
- AMDA data can also be accessed through our HAPI server.
- AMDA data can also be accessed through our python tool Speasy.
- AMDA data can also be accessed through EPN-TAP.

Datasets available in the online tool CDPP/AMDA

Space missions data

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Heliosphere

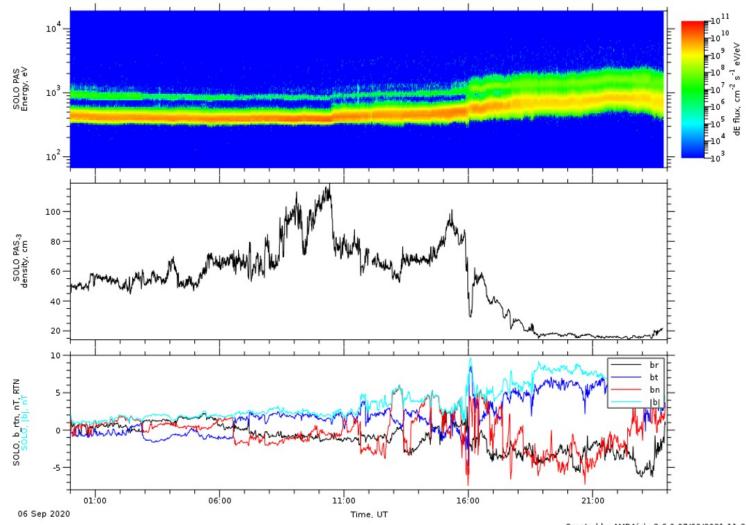
Planets

Comets

Earth

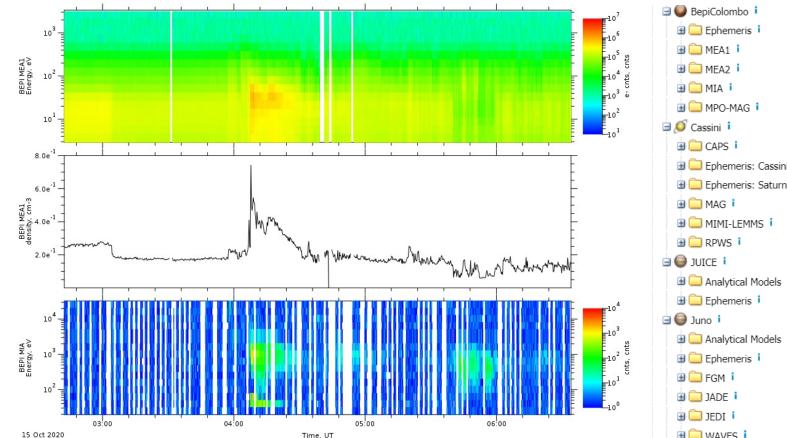
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Solar Orbiter (Solar wind)



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BepiColombo (Venus flyby)

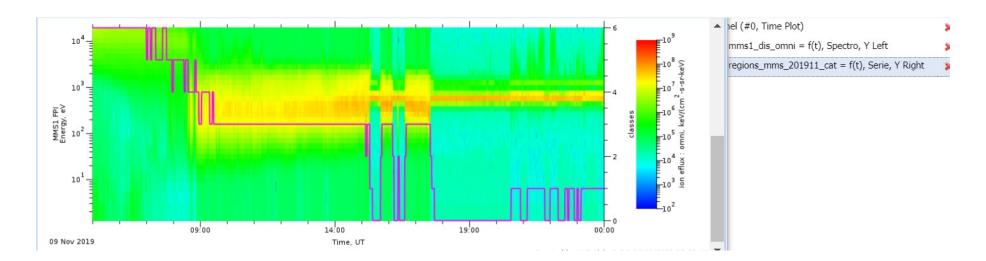


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Catalogues

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- 🔄 INSTRUMENTS	Creation date: Intervals:		Start Time	Stop Time	Duration (Min)			
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- 20 SOLAR_WIND	Stop Time: 2007-02-26T00:28:46	5	2007-02-25T04:24:33	2007-02-25T04:27:22	2.82	3		
- 😑 SUN		6	2007-02-25T09:32:51	2007-02-25T09:39:52	7.02	3		
- O VENUS	Contact:	7	2007-02-25T10:00:31	2007-02-25T10:05:33	5.03	1		
🖽 🧰 BepiCoord_WG	renard	8	2007-02-25T10:22:41	2007-02-25T10:35:19	12.63	1		
SPACECRAFT_ORBIT_NUMBERS		9	2007-02-25T10:53:05	2007-02-25T10:59:10	6.08	3		
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		16	2007-02-25T22:32:02	2007-02-25T22:36:15	4.22	3		
		17	2002-02-25123-16-01	2007-02-25723-22-41	6.67	4		

- shared catalogs
- Time formats for ASCII exports



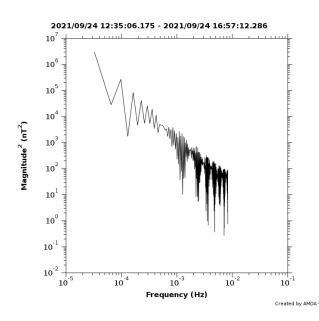
New functionalities

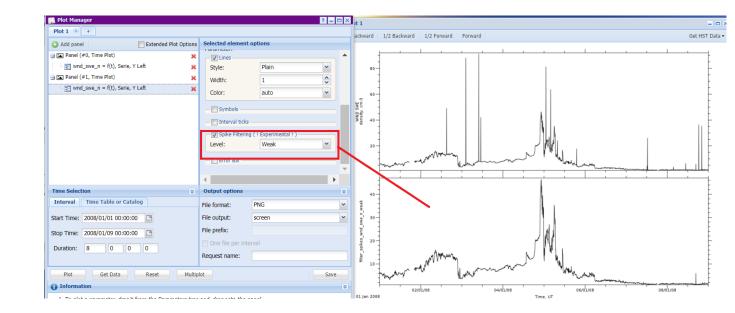
Spike filtering

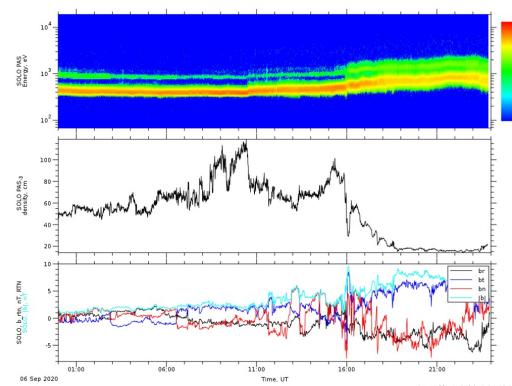
The user can remove spike in time-series data.

FFT

The user can perform FFT on selected time windows.

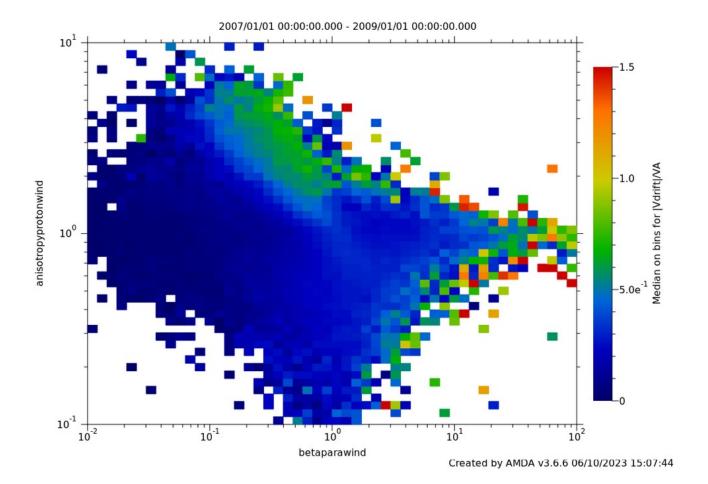




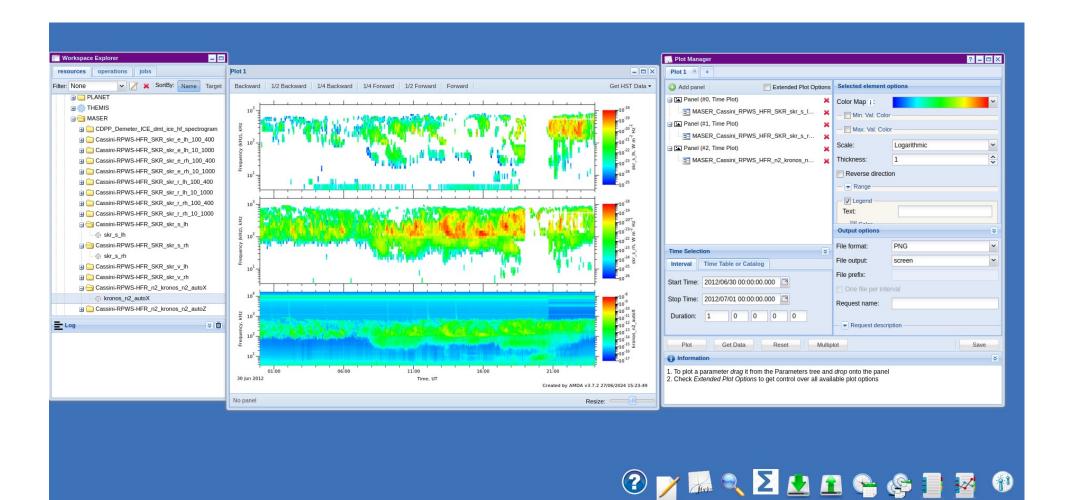


Created by AMDA(c) v3.6.0 07/09/2021 11:35:21

1D/2D Histograms



Connexion AMDA-MASER (radio data)

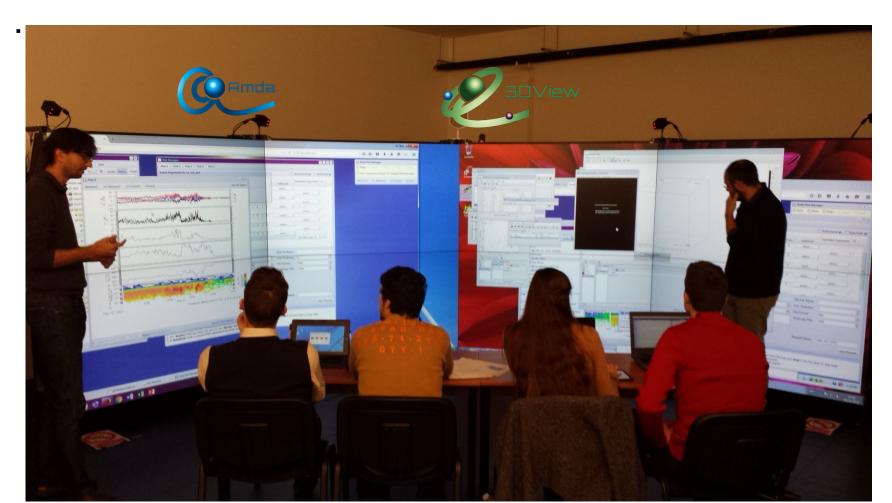


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Hands-on sessions for students

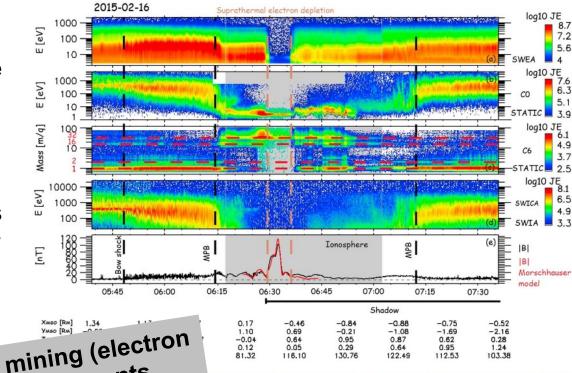
Tools for education in space sciencesAt the Masters Degree level, in summer schools, ...



Use of the PDS/MAVEN data from AMDA

Altitude dependence of nightside Martian suprathermal electron depletions as revealed by MAVEN observations, *M. Steckiewicz et al., GRL, 2015*

- MAVEN observes on almost each periapsis in the nightside ionosphere suprathermal electron depletions
- Observed depletions are populated by 6 eV electrons resulting from absorption by CO2 and by 3 eV O2
- The geographical 05:45 06:00
 distribution of nightside suprathermal electron
 depleti
 AMDA was used for data mining (electron depletion) and selection of events depletion) and selection of events field intensity (measured by MAG in black)

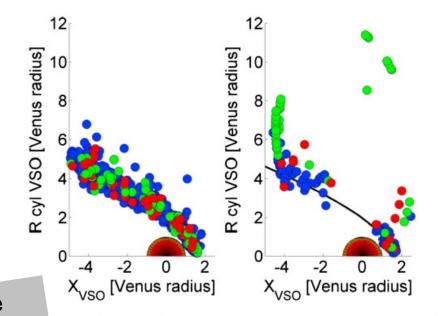


nofevence 10, 2015. (a) SWEA energy-time spectrogram of omnidirectional electron energy flux, (b) STATIC energycome spectrogram of omnidirectional ion energy flux (C0 mode), (c) STATIC mass-time spectrogram of omnidirectional ion energy flux (C6 mode), (d) SWIA energy-time spectrogram of omnidirectional ion energy flux (SWICA mode), and (e) magnetic field intensity (measured by MAG in black and calculated from the model of *Morschhauser et al.* [2014] in red) versus time. The grey shading highlights the ionosphere. The shadow corresponds to solar zenith angle (SZA) larger than 100°.

Use of Venus Express multi-datasets from AMDA

Space weather effects on the bow shock, the magnetic barrier, and the ion composition boundary at Venus, *Vech et al., JGR, 2015*

- Statistical study of the ICME-Venus interaction
- Analysis of solar wind and magnetic barrier conditions during ICME passages
- Decreased altitude of the nightside ionosphere during ICME passages



AMDA was used to search extreme events in long time series (9 years)

Locations of the bow shock crossing for all the 42 investigated events. Since green marks show the bow shock crossings (including the multiple bow shock crossings as well) on the day when the ICME arrived at Venus, the red dots represent the bow shock location during the following day, and the rest of the days are marked with blue dots. (left) Thirty-six cases when the interaction between the magnetic cloud and the induced magnetosphere was not observed and (right) six cases when the signature of the passing magnetic cloud was detected.