## The polarized Galaxy at 43 and 95 GHz as seen by *QUIET*

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Tone Melvær Ruud for the QUIET collaboration

Institute of Theoretical Astrophysics University of Oslo

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# Q/U Imaging ExperimenT

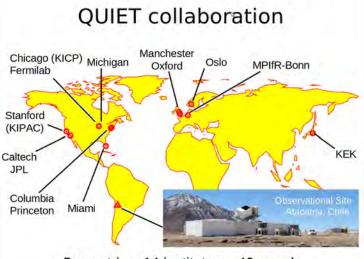
- ► Ground-based polarimeter experiment (Stokes Q,U)
  - ▶ 19-element Q-band (43 GHz) array (2008 Oct 2009 Jun)
  - ▶ 90-element W-band (95 GHz) array (2009 Aug 2010 Dec)
- Main goals:
  - Detect polarized CMB; probe for B-modes
  - Demonstrate capabilities of MMIC detectors ("radiometer on a chip")
- ► Chajnantor Plateau, Atacama Desert, Chile (5080 m.a.s.l.)
  - ► Large accessible sky fraction, easy access, low PWV
  - Natural sky rotation suppresses systematic errors



QUIET vitals Earlier results



QUIET vitals Earlier results

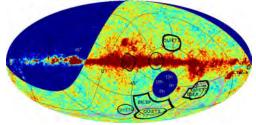


5 countries, 14 institutes, ~40 people

### Analyses so far

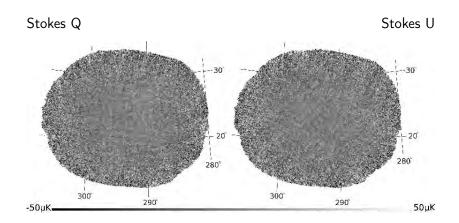
10 000 hrs of observation, 4 CMB fields, 2 foreground fields

- ► First Season QUIET Observations: Measurements of CMB Polarization Power Spectra at 43 GHz in the Multipole Range 25 ≤ ℓ ≤ 475, ApJ 741:111, 2011
- Second Season QUIET Observations: Measurements of CMB Polarization Power Spectrum at 95 GHz, ApJ 760:145, 2012
- ► The QUIET Instrument: ApJ 768:9, 2012
- ► The Q/U Imaging ExperimenT: Polarization Measurements of Radio Sources at 43 and 95 GHz, ApJ 806:112, 2015



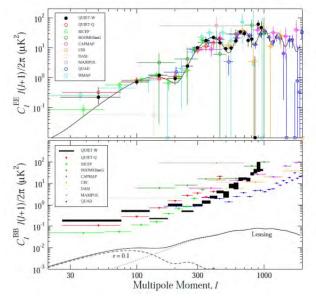
QUIET vitals Earlier results

### QUIET W-band Maximum-Likelihood CMB map (2012)



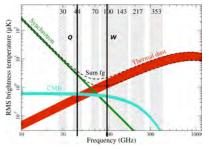
QUIET vitals Earlier results

### Polarization power spectra 2012

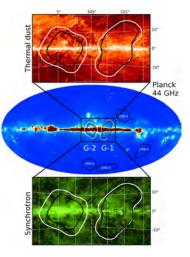


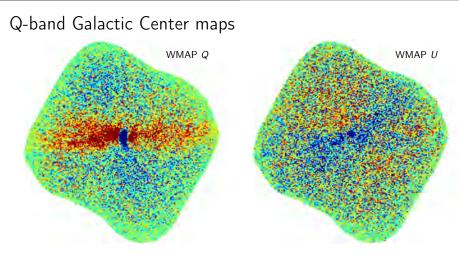
# QUIET foreground analysis (2015)

- The Q/U Imaging ExperimenT: Polarization measurements of the galactic plane at 43 and 95 GHz: In press; http://arxiv.org/abs/1508.02778
- QUIET has the highest reported Q-band sensitivity in the literature



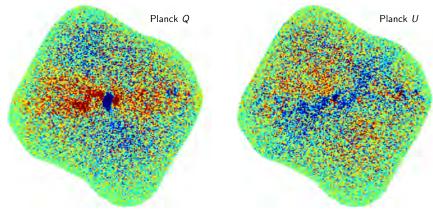
Planck Collaboration X (2015)



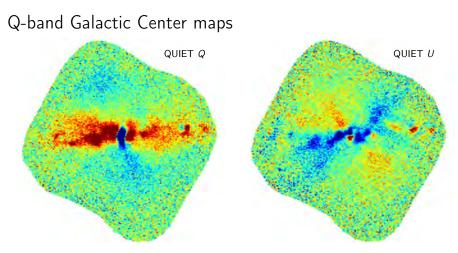


- The deepest large-area Galactic polarization maps published to date at these frequencies
  - 2.3-6 times deeper than WMAP and *Planck* in terms of instrumental noise, depending on frequency and field

## Q-band Galactic Center maps

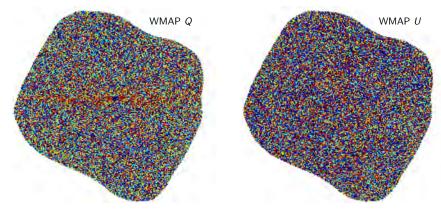


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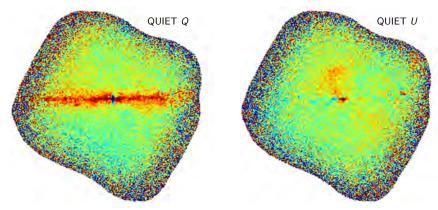
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### W-band Galactic Center maps



- ► Excellent agreement between QUIET and WMAP in all maps
  - No indication of significant residual systematic errors in either experiment

### W-band Galactic Center maps

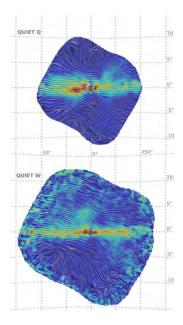


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#### The Galaxy as seen by QUIET Comparison with *Planck*

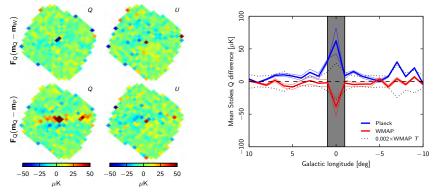
### Co-added QUIET+WMAP maps

- Given the good agreement, we combine our maps with WMAP by inverse-variance-weighting to produce optimal all-scale maps for further analysis
  - ► Full-field polarized synchrotron spectral index of -3.01 ± 0.01 (G-2) and -2.91 ± 0.01 (G-1)
  - ▶ Robust detection of polarized synchrotron spectral index steepening of ≃ 0.2 off- vs. on-plane
  - Detection of Faraday rotation at Galactic center
- All quantities consistent with earlier measurements



The Galaxy as seen by QUIET Comparison with *Planck* 

### Comparison of Q-band Galactic Center maps



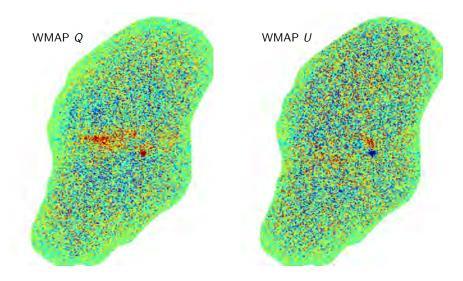
- Planck 44 GHz data deviates from QUIET and WMAP in the Galactic Plane
  - Spatial profile suggests residual temperature-to-polarization leakage at 0.2% level, consistent with reported uncertainties for this channel
  - The new QUIET maps provide unique opportunity to improve understanding of *Planck* data

# Summary

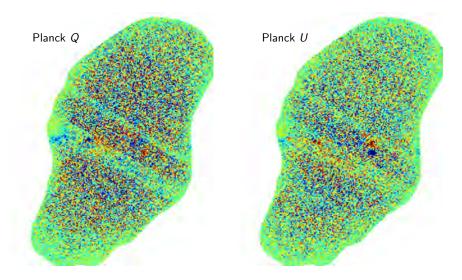
- QUIET has published the deepest large-area Galactic polarization maps to date at Q- and W-band http://arxiv.org/abs/1508.02778
- Comparison to Planck 44 GHz map indicates residual Planck systematic errors consistent with reported values
  - New QUIET maps can improve this situation
- Raw QUIET and co-added QUIET+WMAP maps are publicly available at Lambda together with all necessary ancillary information, including full pixel-pixel covariance matrices http://lambda.gsfc.nasa.gov/product/suborbit/quiet\_prod\_table.cfm

Thank you for your attention!

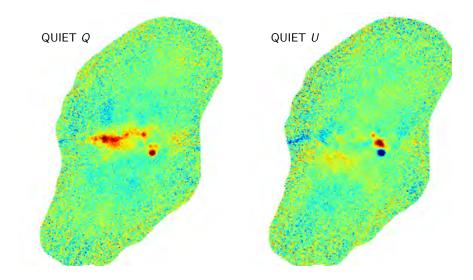
### Q-band Galactic field G-1 maps



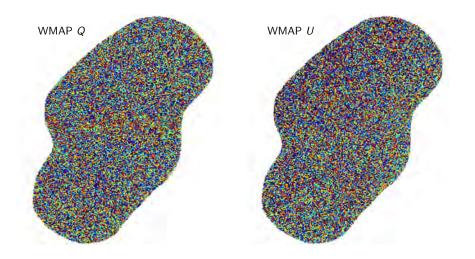
### Q-band Galactic field G-1 maps



### Q-band Galactic field G-1 maps



### W-band Galactic field G-1 maps



### W-band Galactic field G-1 maps

