



# Planck unveils the Cosmic Microwave Background





# **IAU GA Focus Meeting 5**

# **The Legacy of Planck**

# Honolulu, 11-13 August 2015

planck

The aims of FM5 are to:

- publicize the wide variety of current and potential uses of Planck data
- foster exchanges between users of Planck data
- get feedback to optimize the last delivery of Planck data products

FM5 is sponsored by Div H (Interstellar Matter and Local Universe) and Div J (Galaxies and Cosmology)

Scientific Organizing Committee

Chair: J. Tauber - European Space Agency - The Netherlands
Co-Chair: N. Mandolesi - IASF - INAF - Italy
Co-Chair: J. L. Puget - Institut d'Astrophysique Spatiale – France
B. Barbuy, M. Bersanelli, F. R. Bouchet, J. Dunkley, G. Efstathiou, E.
Falgarone, W. Freedman, M. Hazumi, J. M. Lamarre, C. Lawrence, P.
Martin, H.U. Norgaard-Nielsen, T. Souradeep, R. Sunyaev, A. Zacchei





1. Warning: last-minute modifications to the programme may not be included in the online programme

Last talk of FM5 (Thursday @ 12:15): Rashid Sunyaev on "Planck and spectral distortions of the CMB"

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2. Speakers:

*please make sure that your talks are uploaded BEFORE your session !* 

- a. Contributed talks: <15 mins
- b. Invited talks: <30 mins

please allow 2-3 mins for questions !

- 3. Take time to visit all the posters !
- 4. Visit the ESA booth direct access to the Planck Legacy Archive is available, with expert support







# Introduction The Planck products

J. Tauber Planck Project Scientist European Space Agency on behalf of ESA and the Planck Collaboration

# **Plan for this Introduction**

- 1. A little background on Planck
- 2. Overview of the Planck data products
- 3. Overview of the Planck Legacy Archive





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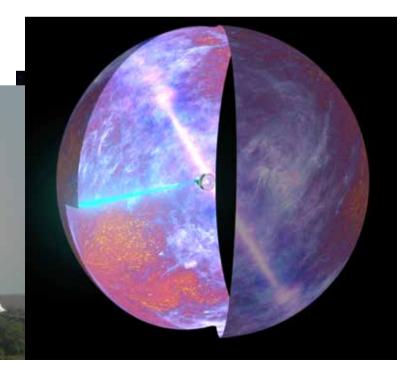
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# A little background



- First proposals COBRAS & SAMBA in 1993
- Start of Planck in 1995
- Launch in May 2009
- Operations Aug 2009- Oct 2013
- First data release March 2013
- Second data release Feb July 2015



# Planck Legacy Archive

- The Planck data products are generated by the LFI and HFI Data Processing Centres, under the supervision of the Planck Science Team
- The Planck Legacy Archive is the main distribution means of Planck data
  - It contains all the data released by Planck since 2011
    - ALL the data acquired by Planck is now available !

- <u>http://pla.esac.sci.int</u>
- Visit the ESA booth in the Exhibition Hall to get direct access and support !
- A partial mirror of the PLA exists at IPAC
  - <u>http://irsa.ipac.caltech.edu/data/Planck/docs/index.html</u>



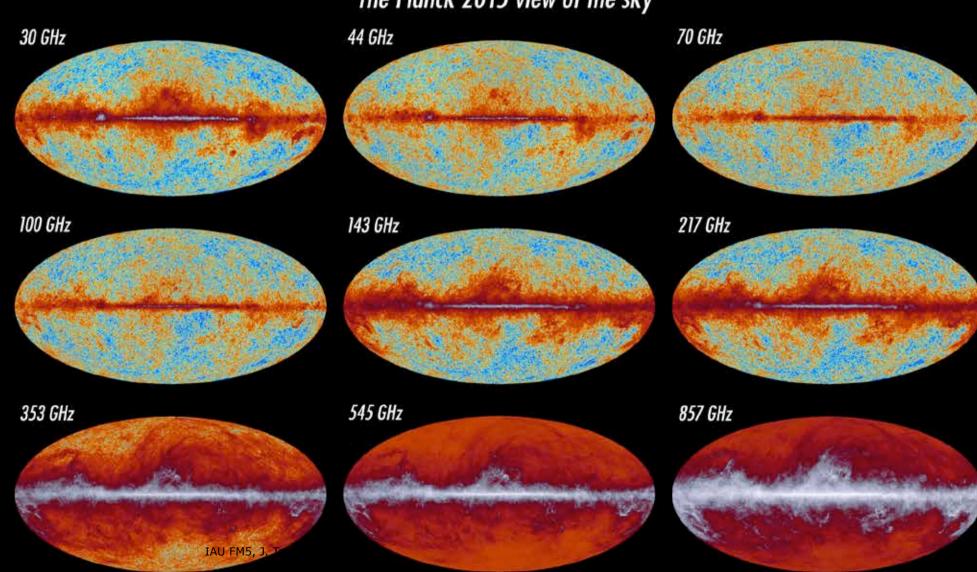
# Planck data releases

- 2011: The Early Release Compact Source Catalogue
  - Intended as a "quick" product to enable follow-up of interesting sources, mainly with Herschel
- 2013: the first major release of data
  - Contained data products based on the first 15 months of observations, calibrated on the WMAP solar dipole
    - All-sky Temperature maps by frequency
    - physical component maps and catalogues
- 2015: the first complete release of data
  - Data products using ALL the data acquired by Planck, calibrated on the orbital dipole
    - All-sky Temperature and Polarization maps by frequency
    - Physical component maps and catalogues
    - Timelines of cleaned and calibrated data
- 2016: the "final" release of data
  - Data products with improved handling of systematic effects
  - "semi-raw" timelines



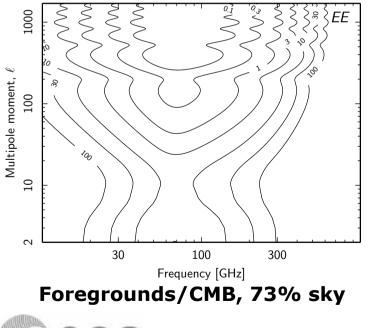
# The main Planck products: 2015 temperature maps

The Planck 2015 view of the sky



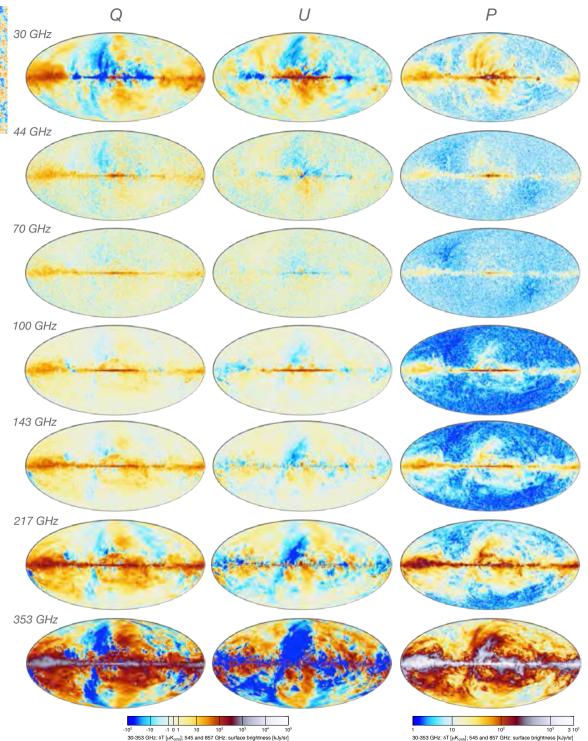
# The 2015 polarization maps

Ratio of amplitude of polarized foregrounds to CMB is <1 for ell>40 at ~70 GHz



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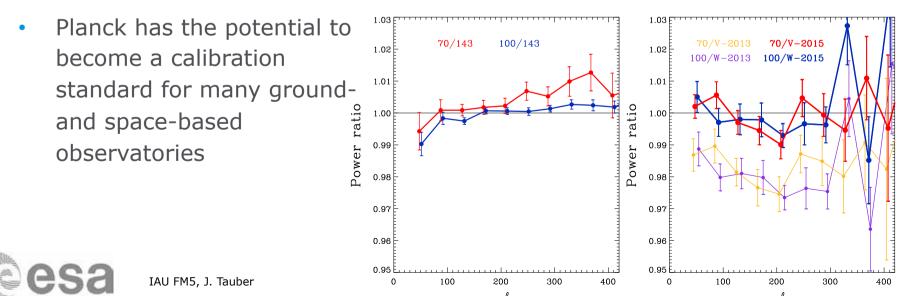
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# Characteristics of the maps



- The Planck temperature frequency maps are the most exquisitely calibrated maps you have ever seen in this wavelength range !
  - Absolute calibration accuracy is as low as ~0.1% (100-143 GHz)
  - The Solar dipole amplitude is determined by Planck to 0.06% (0.02° in direction); worst residual 0.2% at 217 GHz
  - Relative calibration between 70 and 217 GHz <0.3% (0.5% when including 353 GHz)</li>
  - The 545 and 857 GHz maps are calibrated on planets
    - The relative calibration is  $\sim 1\%$ , but the absolute calibration is driven by the model uncertainty of 5%

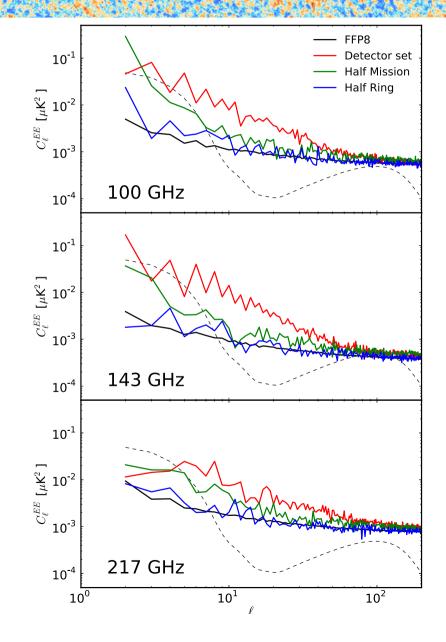


# Low-ell systematics

- At large angular scales (>10°), the HFI polarized maps (100-217 GHz) still contain significant levels of systematics
- Large angular scales (*l* < 30) should NOT be used for cosmological analysis
- The polarized CMB maps produced using these maps and described in Planck papers have been high-pass-filtered accordingly
- We already know how to reduce the level of systematics very significantly – the 2016 products will incorporate this

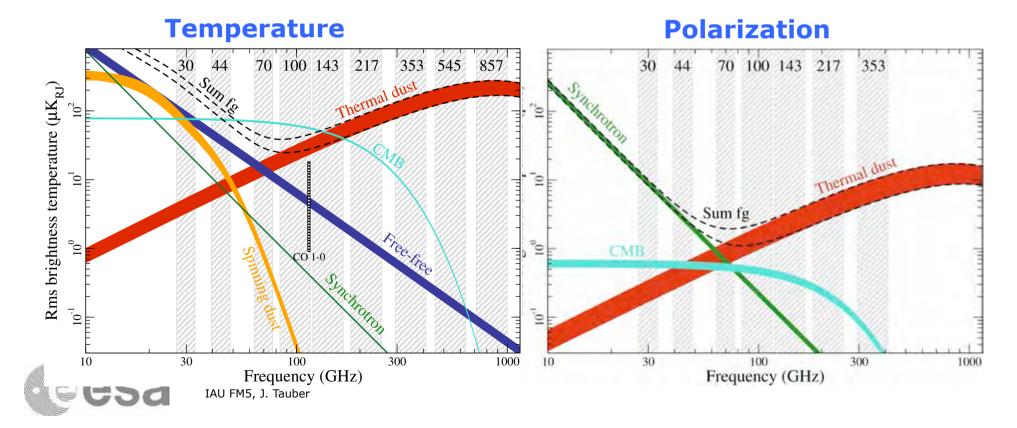
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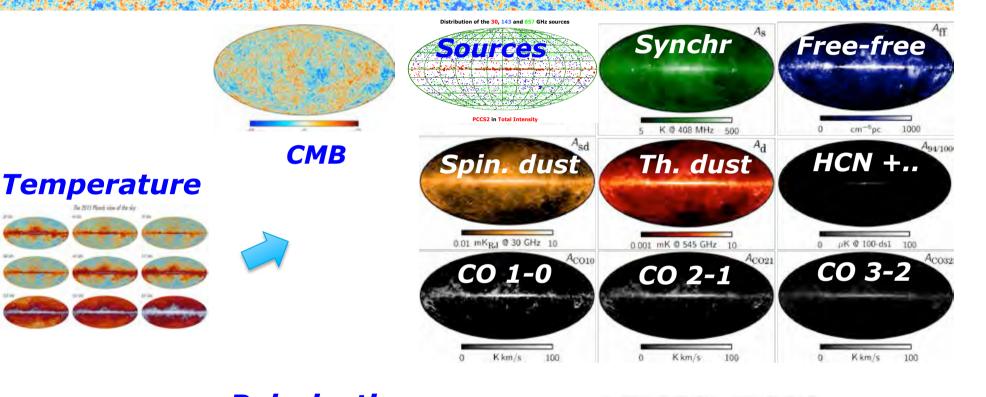


# Component separation

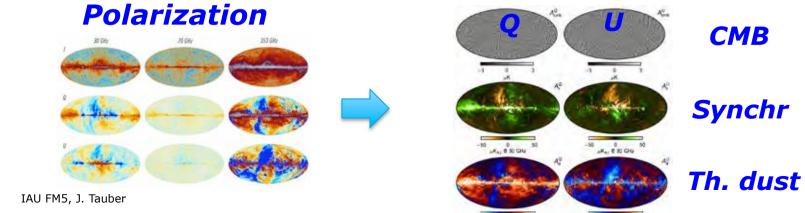
- planck \*
- The main tool to separate physical components are their spectral characteristics
- Separating CMB and foregrounds are different processes: 4 methods are used for the CMB, 1-2 methods to resolve foregrounds
- Separating temperature and polarization are independent processes



# **Component** separation

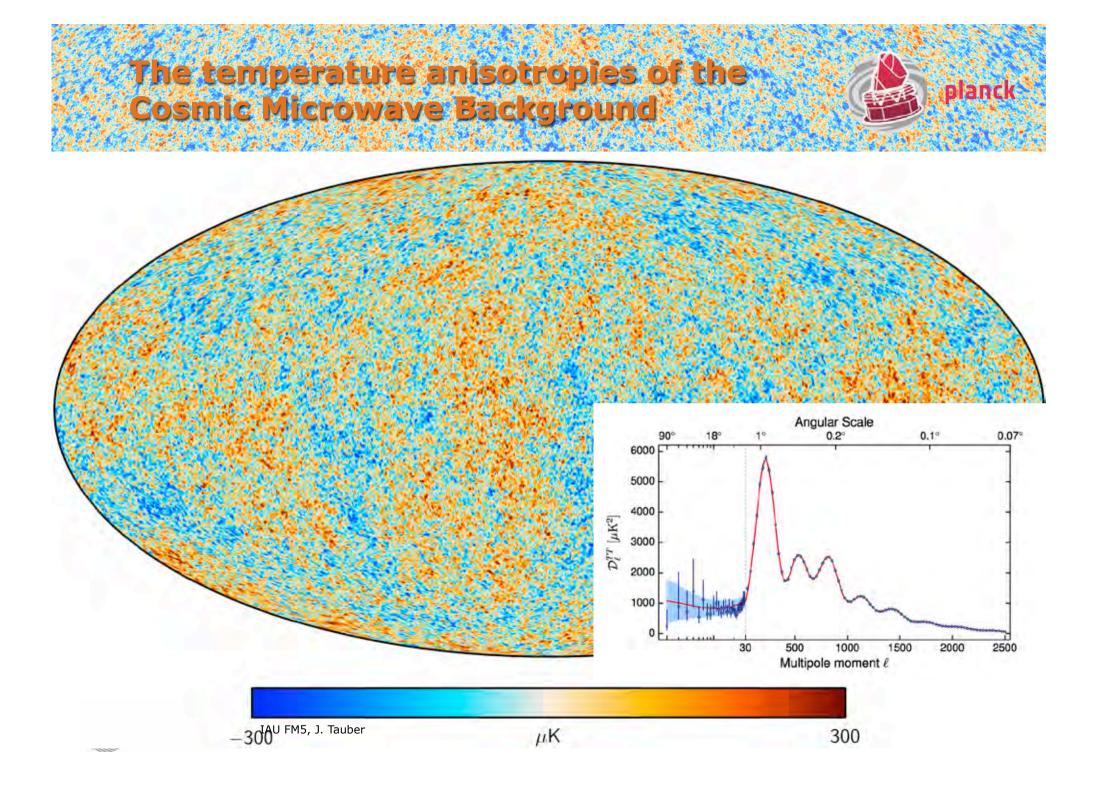


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100 -10 0 10 10 100 PKa 0 153 GHz 100 -30 0 is its pKat \$ \$53 Gets

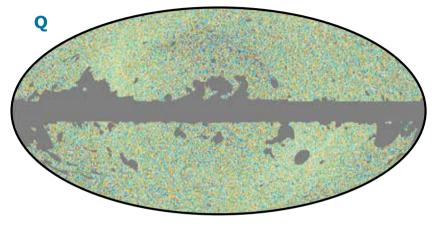


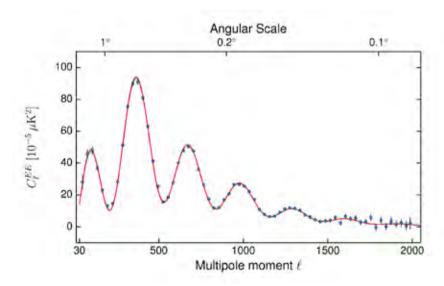


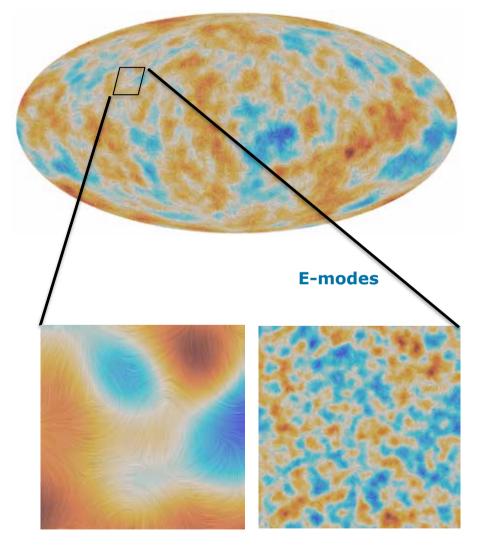


# The polarization anisotropies of the Cosmic Microwave Background

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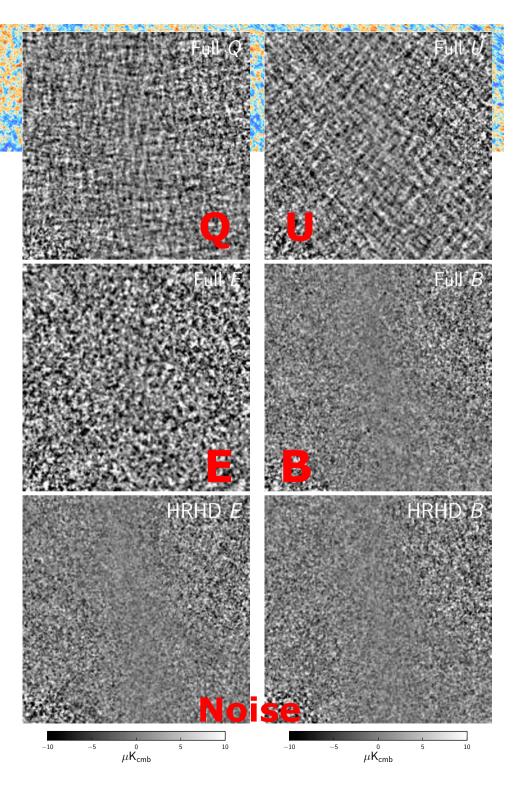




# **Polarized CMB**

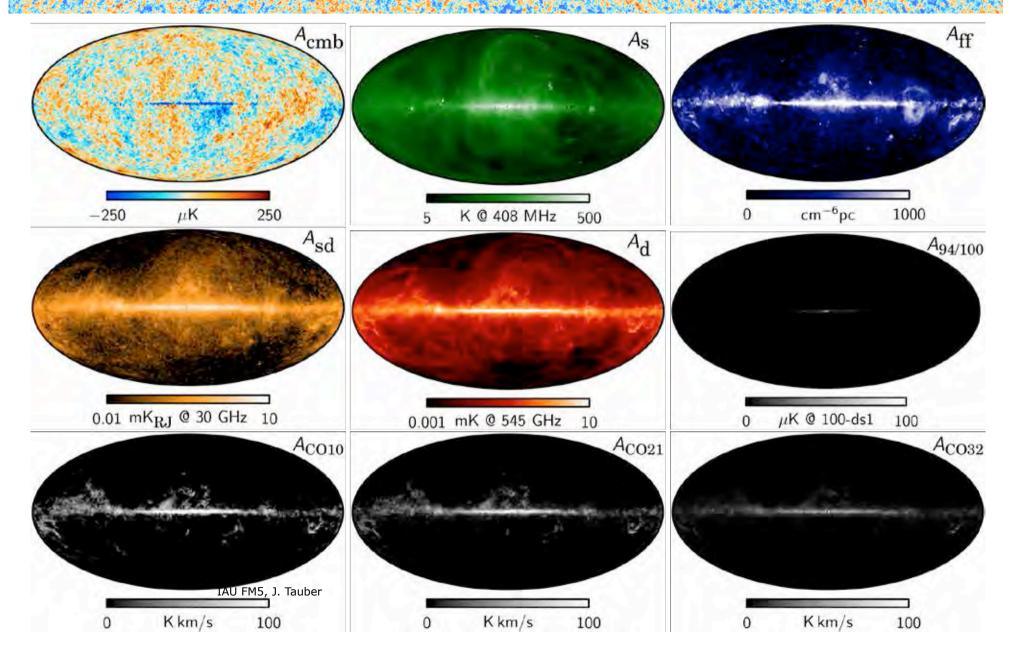
Zooming into a 20x20 deg patch (NEP) shows:

- Characteristics (+ and x) patterns for U and Q maps
- High signal-to-noise for Emodes
- Very low signal-to-noise for B-modes





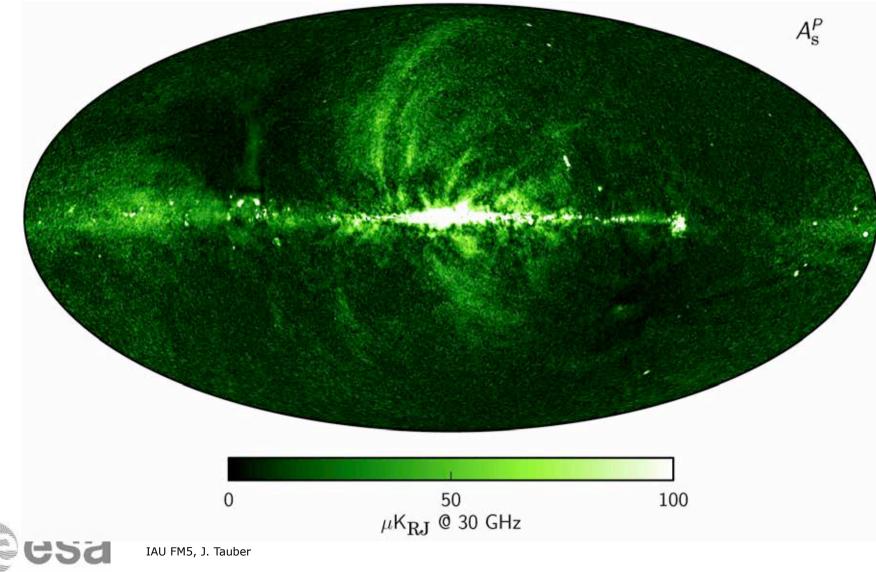
# Astrophysical foregrounds from Planck: a complete view of the submm sky





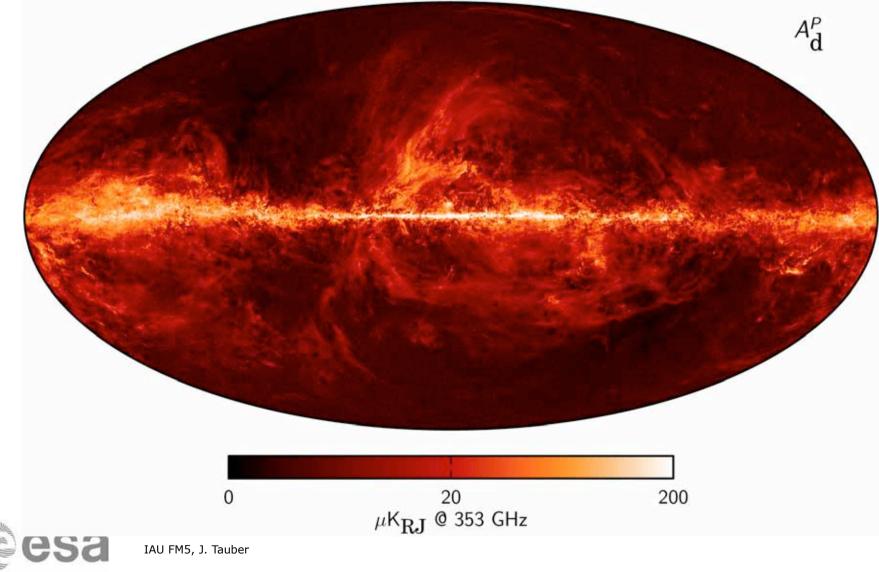
Contraction of the







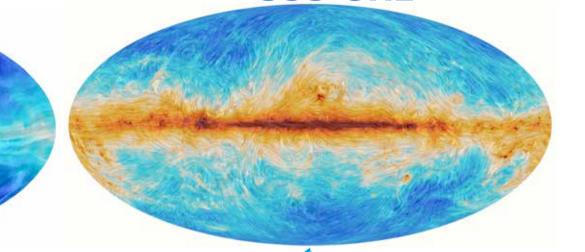


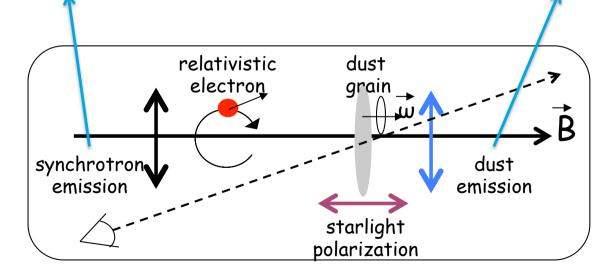


# Tracing the Galactic magnetic field









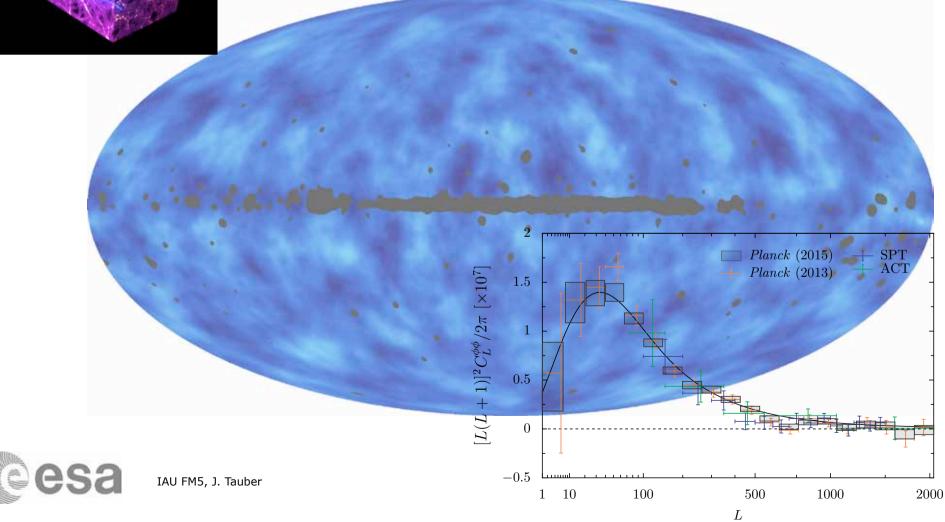


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# Lensing of the CMB

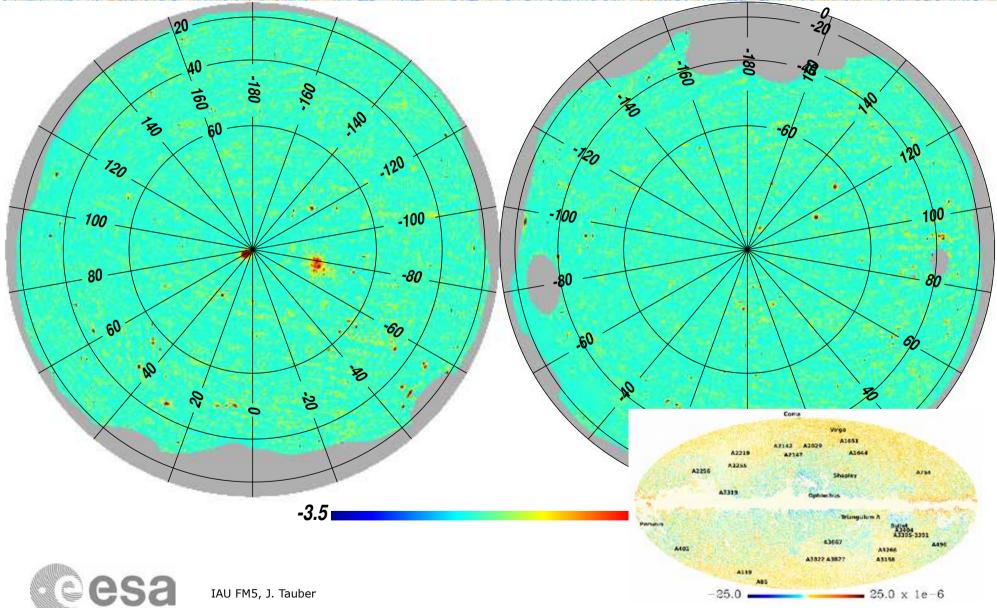






# Diffuse Sunyaev-Zeldovich emission



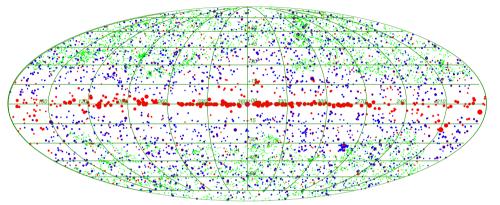


## Planck source catalogues

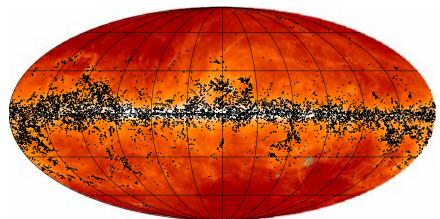


### **Radio and dusty galaxies**

Distribution of the 30, 143 and 857 GHz sources



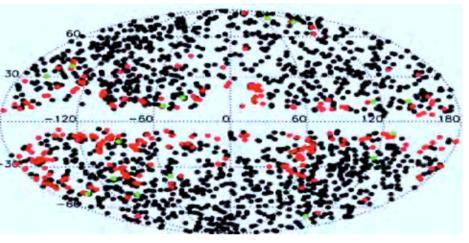
# PCCS2 in Total Intensity Cold galactic clumps



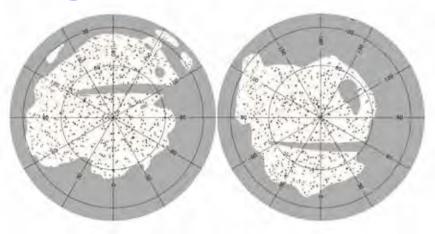


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### **Galaxy clusters via the SZ effect**



### **High-redshift sources**



# The Planck Legacy Archive

- Contains all the Planck data products
  - Timelines
  - Frequency maps
    - Full and Partial data maps
    - Instrument characterisation data
    - Beams
  - Physical component maps
    - CMB: 4 different versions
    - Commander foreground products
  - Likelihood code
  - Cosmological data (power spectra, parameters, etc)
  - Helpful software, e.g. unit conversion, color correction etc
  - Online documentation (the Explanatory Supplement)
- In the future: the Planck Sky Model...



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# PLA product downloads

- Most downloaded products:
  - 1. Likelihood code
  - 2. Frequency maps
  - 3. Catalogues
  - 4. CMB maps
  - 5. Component maps
  - 6. Simulations
  - 7. ...





- 120 papers published by the Planck Collaboration
  - 81 of these accepted by a refereed journal (by 30 Jul)

- Cited by ~6000 individual papers (5600 refereed)
  - About 10% of these actually use Planck data
    - (remaining 90% "only" use the Planck best-fit cosmology)
    - Breakdown by topic:
      - 55% CMB cosmology
      - 16% SZ & clusters
      - 16% external galaxies
      - 11% interstellar medium



# Summary



- Planck provides a very complete view of both the near and the very distant Universe
- It will remain for many years a unique source of data to address a wide range of problems, from cosmology to astrophysics
- The Planck Legacy Archive already provides all the data that Planck has acquired
  - Every cosmologist and astronomer will find something useful in it
  - USE IT !!!
- Look out for the 2016 release
- Visit the ESA booth to get hands-on support on how to use the PLA





### PCI I



And in case

### Planck pre-launch status: The Planck mission

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