Twenty Third Meeting of the Hipparcos Science Team ESOC, 11-12 December 1989

Attendance:

HST: Prof. P.L. Bernacca, Dr M. Crézé, Dr. M. Grenon, Prof. M. Grewing. Prof. E. Høg, Prof. J. Kovalevsky, Dr F. van Leeuwen, Dr L. Lindegren, Dr H. van der Marel, Mr C.A. Murray, Mr R.S. Le Poole, Dr C. Turon.

ESTEC: M.A.C. Perryman, R. D. Wills

ESOC: J. van der Ha, A. Schütz, P. Davies; O. Ojanguren, D. Heger, C. Sollazzo (part time)

MATRA: I. Asseman, J.P. Gardelle

Apologies for absence were received from F. Donati.

The agenda given in Annex I was adopted.

1. Introduction

Perryman reported on the November SPC meeting, and Turon reported on the letter of intent she had submitted to ESA on behalf of the consortia leaders and the HST, concerning Hipparcos 2. Høg reported that he had been approached by Dr. Khromov, Vice President of the Astronomical and Geodetic Society of the USSR, offering a free launch of Hipparcos 2.

2. Satellite Status

Heger reported on the satellite status (Annex II). One of the multiplexers had failed on the payload RTU, leading to a loss of telemetry data on day 336 (2 December, 1989). No effect had been observed on the satellite.

On the solar array degradation, no clear picture had yet emerged—voltage degradation had been steep over the last week, even though no excessive solar activity had been reported over this interval. HST members were not happy that (a) errors on data points seemed

to be unknown, (b) no explanation of the degradation mechanism had been presented. Le Poole stressed that little effort seemed to be given to estimating the lifetime of the satellite, which was a major cost driver.

Heger summarised the operations which had been conducted since early November (Annex II). Previous results had led to about 31 per cent data recovery. Presently, with Kourou, 56 per cent data recovery is being achieved.

3. Calibration Results

Davies updated the calibration status since the last mini-HST meeting of 16 November (Annex III):

M3: typically $M_3 < 0.01$.

Piloting improvements: Matra had modified the CCCM, and the transverse offset residuals are now < 0.5 arcsec, but longitudinal offsets are still around 2 arcsec. This could be further improved once the grid rotation (and axis) are better known.

IFOV profile: the average profile is very close to that specified. Beyond 180 arcsec there is great difficulty separating the response from the background effects.

Chromaticity: Wills presented results of the chromaticity calibration (see Annex IV) ($C_c = -0.27 \pm 0.06$ milli-arcsec, $C_v = 0.79 \pm 0.06$).

Refocusing Strategy: Davies proposed that refocusing monitoring is done with a certain frequency, and adjustments are made when the grid is two steps from the optimum focus, moving to two steps in the new direction, i.e. changing focus every 16–20 days. After discussions, it was agreed that the focusing would be monitored every week; on the basis of the history, movements would be made so that the grid would always be within ± 0.5 steps of best focus. It was agreed to do this around perigee ± 1 hour, so that RGC's would not have a changing scale factor during the RGC reductions. After the meeting, Davies provided further details of the focusing requirements—these are included in the status report ESTEC.STATUS.05.

Basic Angle Stability Check Points: The identified dates for verification of the basic angle stability were: 25 November - 10 January; 4 December - 25 February; 12 December - 17 February. None of these pairs can be used presently, so basic angle stability checks could only be made using this procedure after the eclipse hibernation phase.

Goldstone: van der Ha reported that NASA will support the Hipparcos project through use of the Goldstone antenna (Annex V). Of the 70m, two 34m and the 26m antennae at Goldstone, it is the latter that will be used for Hipparcos. Datation will be performed at the Goldstone station, before propagation of the data to ESOC via three satellite links.

The station should be operational by end-March. The plan is to keep the Kourou station until further notice, principally because of the perigee coverage.

4. Results of Consortia Evaluation

Van Leeuwen presented results of the RGO analysis so far (Annex VI).

Basic Angle: 58 deg 00 arcmin 31.212 arcsec was found by RGO (31.254 arcsec found by Utrecht). The basic angle had been determined to an accuracy of 6–7 milli-arcsec from 7 hours of data. The star mapper rotation was found to be 5 arcmin 59 ± 1 arcsec. Attitude reconstitution is good to 0.1-0.2 arcsec for the fainter stars.

IDT reductions had gone well for the data interval covering about three satellite revolutions. For the phase shift between the first two harmonics, RGO find (very provisionally) 15.9 milli-arcsec for the preceding FOV, 14.1 milli-arcsec for the following FOV. M2/M1 = 0.343 for the preceding FOV, 0.329 for the following FOV, cf the ESOC value of about 0.304 (but note that ESOC do not correct for the presence of background). A tape is expected to be sent to Copenhagen (for the GRC step) around 15 December.

Kovalevsky reported on the attitude and IDT analysis in FAST (Annex VII). The same grid rotation angle had been found as RGO. After star position correction, the smoothed attitude is good to 80-150 milli-arcsec. For the main grid, IDT reductions had been made on a 5×5 matrix, with preliminary results consistent with those from ESOC (Annex VII). For the photometry, the in-orbit results are consistent with the on-ground predictions.

Van der Marel reported on the great-circle work carried out at Utrecht. Grid rotation and basic angle correction derived so far have been determined from the star mapper data and may not be final values. Instrumental terms have been determined by Hans Schrijver at Utrecht (except for y-dependence and basic angle) from the IDT analysis (the following results were updated during the meeting on 12 December).

Component	P	F
y	2428	2800
y^2	9	7
y^2 y^3	20	32
x	-874	-872
xy	0	-5
xy^2 x^2	40	44
	4	5
x^2y x^3	-3	-1
x^3	23	26

with an rms error of 15 milli-arcsec, and a basic angle as given above.

Concerning feedback to ESOC, van der Marel/Schrijver will forward updated parameters (scale factor, basic angle, single grid rotation) by 15 December to be updated by ESOC and verified through CCCM calibration.

Tycho: Høg reported on the Tycho analysis at Heidelberg. Grewing reporting on the analysis work at Tubingen. He presented results on the transits of Venus. A document was passed to ESTEC/MATRA on the potential impact of ghost images. (Results were presented at the TDAC meeting the following day which showed that the data has been propagated through the PGC and photometry tasks).

Spikes: TDAC (Grewing/Høg) will send details of spikes found in Tycho data (counts, duration etc.) to be sent to ESTEC/MATRA (Action 6).

Strategy for Choice of ν_0 : Crézé presented results of his studies on uniformity of sky scanning. Presently, a value of $\nu_0 = 40^\circ$ would be replaced by a value of $\nu_0 = 80^\circ$ after the first hibernation period.

5. DDID Discussions

van Leeuwen requested clarification of d (see Action 8). Provisional tape 4 went out on Wednesday 6 December, using day 315, with three ground stations. All errors should have been corrected on this tape. Some further comments had been received from Pieplu, and these will be studied by ESOC. Tape 5, with good focus, would be sent out by 15 December. Tape 6 will also be sent out based on one week of data. Feedback from tapes 4–6 is expected from the DRC before the data tapes are sent out routinely. Courier will continue to be used for series 4-6. In future, until further notice, ESOC will send tapes by courier to ESOC and to RGO.

6. Input Catalogue

Schütz had received 109 900 updates: 54 stars suppressed, 6 stars added due to alternate observation strategy, and 179 changed due to revised positions. Turon reported that updated catalogues had been sent to CNES, RGO, Utrecht, and one would be sent to Lund. The catalogue was put into operation at ESOC on 1 December (IC6). The relevant changes to the Annexes were made at the same time.

Modulation of the observation strategy will not be carried out until at least after the eclipse period, as agreed between Crézé and Schütz.

Minor planets: Schütz/Davies will check a few minor planets for signal modulation and send reports to Perryman (Action 4). Similarly, Davies will put into action the reporting of variable star magnitudes.

Publication of Input Catalogue: Turon still foresees this for the end of 1990.

Charts Status: Grenon now has optical disk reader in Geneva, and will be ready to prepare charts in April. Approximately 10 000 stars will be included: all faint stars in the catalogue $(V>10.5~{\rm mag})$, plus stars throughout the catalogue with 'confusing' neighbours, using laws similar to those used for flagging veiling glare. 15×15 arcmin fields would be used, with approximately 24 charts per page.

Turon reported that the INCA data base computer and software will be changed in mid-1990.

7. Hibernation Policy

A strategy/policy paper wass requested from ESTEC (Action 9).

Photometry: use of TDAC to be considered by Høg for the next meeting (Action 10).

Comparison: will be done on Tape 5—Perryman to coordinate.

Next HST meeting: subsequently confirmed for 21–22 February 1990 at ESOC.

M.A.C. Perryman

14 December 1989

Distribution: Participants, H. Hassan, K. Van Katwijk, K. Clausen, B.G. Taylor.

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	Signatures				

Twenty Third Meeting

of the

HIPPARCOS SCIENCE TEAM

ESOC, 11-12 December 1989

(Portacabin 1, 14:00 hrs on 11 December VIP Room, 09:00 hrs on 12 December)

AGENDA

- 1 Debrief of SPC Meeting (Perryman) and 'revised mission' status
- Satellite status (Heger) (orbit, payload status, power degradation, ground station coverage)
- Calibration results (Davies, Wills)
 - review of all calibration results
 - CCCM (transverse and longitudinal offset)
 - chromaticity results
 - IFOV profile
 - refocusing strategy
 - basic angle stability 'check points'
- 4. Results of Consortia evaluation:
 - Utrecht first-look
 - FAST/NDAC/TDAC reports
 - strategy for choice of nu0, omega0
 - 'spikes'
- 5. DDID discussions:
 - comments on data tape
 - status of DDID 6/7 (for nominal mission)
 - tape sending policy
- Input Catalogue aspects (Schutz/Turon)
 - updated catalogue
 - minor planets/variable stars
 - modulation strategy
 - publication and charts status
- 7. Miscellaneous:
 - hibernation policy and data reductions
 - comparison activities
 - next HST meeting
 - photometry aspects

PAYLOAD 11.12.1989

P/L SUBSYSTEM NOMINAL WITH FOLLOWING EXEPTION:
- FAILURE OF MUX 8 OF IMX2 OF THE P/L RTU
ON DAY 336 AT 13:16:40z (AR 17)

P/L THERMAL CONTROL IS NOMINAL IDT CHAIN 2 SELECTED TYCHO CHAIN 2 SELECTED MDE1, TCE1, AND INVERTER 1 ON

CALIBRATIONS SCHEDULED AND PERFORMED:

- REFOCUSING
- ISP
- GRM
- TRANSVERSE OFFSET
- LONGITUDINAL OFFSET
- SINGLE SLIT RESPONSE
- JITTER
- STRAYLIGHT
- IFOV
- CHROMATICITY

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TABLE 1 (Before Anomaly)

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J252	MDE1	TYC:SHUTT	ER	CLOSED			K252	MDE2	TYC1SHUT	TER	INTERMEDIATE	
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TABLE 2 (After Anomaly)



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DOC N° ISSUE N° DATE

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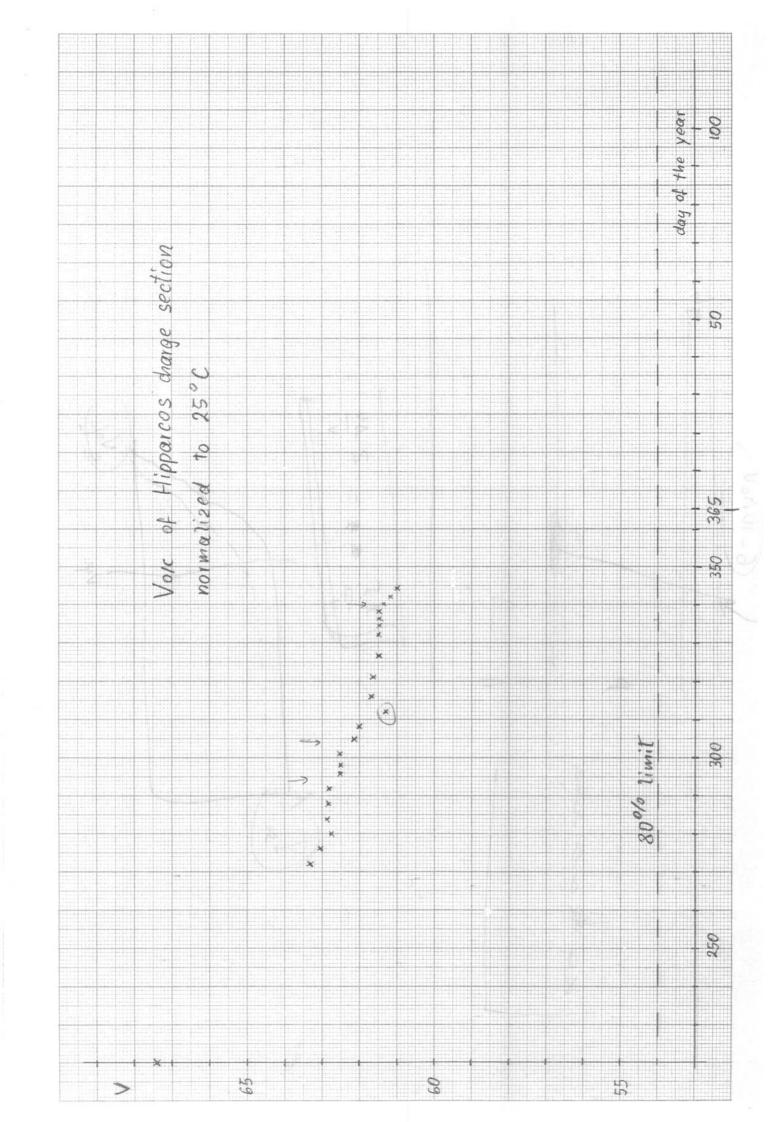
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31.08.87

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		3	232 233 234 235	P/L,INV	I002 I004
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		2	244 245 246 247	DEB2	H001 H002 H006 H014
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TABLE 3.5.4.4 : CONTINUED



SOLAR ARRAY DEGRADATION

DAY	Vo/C
272	63.29
276	63.04
280	62.78
284	62.88
288	62.86
292	62.82
296	62.51
298	62.51
301	62.53
305	62.13
308	62.01
312	61.27
316	61.73
321	61.64
327	61.50
332	61.52
334	61.50
336	61.51
338	61.51
340	61.37
342	61.18
344	61.00

SUMMARY OF HIPPARCOS OPERATIONS

FOR THE PERIOD DAY 313 TO 329

- DAY 313 9.11.89
 UPDATE OF STAR MAPPER THRESHOLDS LAMBDA MAX/MIN
 (REJECTION CRITERIA AS A FUNCTION OF STAR MAGNITUDES)
- DAY 319 15.11.89
 CBS/ACS TRANSFER DISABLE/ENABLE CHANGED FROM 22K TO 18K
 GYRO 3 DESTORAGE
- DAY 321 17.11.89 CHANGE UOP FROM 90 mins. TO 10 mins

OPEN SHUTTER 1 min 30 sec AFTER OCCULTATION CLOSE SHUTTER 1 min 30 sec BEFORE OCCULTATION NO PSF DURING SHUTTER CLOSURE TO AVOID 'FALSE' STAR IDENTIFICATIONS

- DAY 322/323 18/19.11.89
 MISSION PLAN (AUTO COMMAND) OPERATIONAL
- DAY 324 20.11.89
 ACS/CBS TRANSFER ALWAYS ENABLE DURING PERIGEE
 (NO ALTITUDE CONSTRAINS)
- DAY 325 21.11.89
 INNOVATION THRESHOLD FOR EXTENDED WINDOW 10 arcsec.
 INNOVATION THRESHOLD FOR NARROW WINDOW 5 arcsec.
- DAY 326/327 22./23.11.89

 TIME-TAG BUFFER FULL (117 CMD'S) AND STILL CMD'S
 TO BE LOADED.

 AS A CONSEQUENCE THE IDT SHUTTER REMAINED OPEN
 DURING AN EARTH OCCULTATION. OBSERVED COUNT RATES
 WERE IN THE ORDER OF 10**4 COUNTS /SEC.
 (SIRIUS COUNT RATES ARE ABOUT 10**7)

FINAL STAR CATALOGUE RECEIVED ON DAY 327 15:00 L (OVER 90% CHANGES), BUT NOT IMPLEMENTED YET.

DAY 328 24.11.89
INNOVATION THRESHOLD FOR EXTENDED WINDOW 30 arcsec.
INNOVATION THRESHOLD FOR NARROW WINDOW 10 arcsec.
(REASON: AFTER OCCULTATIONS RTAD HAS BEEN LOST SEVERAL TIMES)

NARROW WINDOW TO BE USED ALWAYS, UNLESS FD REQUESTS A CHANGE; i.e. PRIOR TO T-B ANGLES/DRIFTS UPLINK

'NEW' SINGLE COMMANDS FOR SHUTTER OPEN/CLOSE BACK ONTO THE SYSTEM (MISSION PLAN)

DAY 329 25.11.89 GYRO 5 DESTORAGE

SUMMARY OF HIPPARCOS OPERATIONS

FOR THE PERIOD DAY 330 TO 345

DAY	330	26.11.89 START OF ROUTINE SCIENCE DATA COLLECTION
DAY	336	02.12.89 FAILURE OF MUX 8 OF IMX 2 OF P/L RTU (AR 17)
DAY	339	05.12.89 KOUROU TAKEN OUT OF SUPPORTING NETWORK DUE TO REED SOLOMON DECODER MALFUNCTIONING
DAY	341	07.12.89 UPLINK OF NEW CALCULATED COEFFICIENT TORQUE PARAMETERS
DAY	344	10.12.89 TEST OF KOUROU TO SUPPORT DOWNLINK TELEMETRY
DAY	345	11.12.89 KOUROU SCHEDULED FOR ACTIVE SUPPORT

HIPPARCOS DATA RECOVERY COMPARED TO GEOSTATIONARY ORBIT AND POSSIBLE DEGREE OF SCIENCE EXPLOITATION.

DAY 310 TO 313 (incl.) 6-11-89 TO 9-11-89 (incl.)

THE PASS PATTERN REPEATS EVERY FOUR DAYS WITH ONLY SLIGHT VARIATIONS. THE PERIOD DAY 310 TO 314 WAS CHOSEN SINCE IT INCLUDES KOUROU FOR THE FIRST TIME.

GEOSTATIONARY		96 HOURS	100%
PRESENT DATA RECOVERY 16 KOUROU, PERTH, ODENWALD	.25 Н	79.75 Н	83%
DATA LOST DUE TO: - KOUROU SPOT PASS TDF DATA FLOW	6.2 H	73.55 Н	76.6%
- NOT SCHEDULED - LATE A.O.S. (LINES) OR BAD LINES - STATION OVERLAP - GROUNS STATION OR HDCS PROBLEMS	6.2 Н	67.35 H	70.1%
NO SCIENCE DATA DUE TO 21 OCCULTATIONS	5.3 н	62.05 H	64.6%
BECAUSE OF ACTIVE VAN ALLEN FURTHER SCIENCE DATA AVAILA	H [[[다양양] 6명 2 5.[[] 12 12 12 12 12 12 12 12 12 12 12 12 12	1 0 1 M	
RTAD IN CLOSE LOOP (CBS/ACS TRANSFER ENA.) > 22 K ALTITUDE 1	0 н	52.05 H	54.2%
DURING REFERENCE PERIOD NO CONVERGENCE HAS BEEN ACHIEVED ON DAYS 310 & 311, 3 PASSES; AND CONVERGENCE FOR THE 5 REMAINING PASSES ON DAYS 31 312 AND 313 HAS TAKEN PLACE AFTER 1 H AVERAGE AND > 22 K ALTITUDE 22		30 H	31.25%
RUNNING AVERAGE			
RUNNING AVERAGE			31.25%

THE FOLLOWING IS A SUMMARY OF ADDITIONAL OPERATIONAL ACTIVITIES WHICH HAVE BEEN UNDERTAKEN TO TRY AND IMPROVE THE QUANTITY OF USEFUL SCIENTIFIC DATA.

- VAN ALLEN BELT ACTIVITIES (ACS/CBS TRANSFER ENABLED/DISABLED AT 22K, 18K, 0)
- RTAD INITIALISATION AFTER LOS (BEST ACHIEVEMENT TO DATE 30 TO 45 MINS.)
- IMPROVEMENT TO GROUND RTAD MONITORING S/W
- UPDATING OF A.O.S./L.O.S. PROCEDURES
- LAMBDA MAX/MIN VALUES
- INNOVATION THRESHOLD CHANGES
- REFOCUSSING
- SINGLE SLIT RESPONSE FOR A FOCUSSED TELESCOPE
- MISSION PLAN (AUTOMATIC COMMAND IMPLEMENTATION)

SOME OF THE ABOVE ACTIVITIES ARE STILL UNDER INVESTIGATION, AND FUTURE CHANGES MIGHT BE REQUIRED.
IN ADDITION TO THE ABOVE, OTHER WAYS OF IMPROVING THE PROPORTION OF SCIENCE DATA GATHERED IS BEING INVESTIGATED.

DAY 334 TO 337 (incl.) 30-11-89 TO 03-12-89 (incl.)

GEOSTATIONARY		96 HOURS	100%
PRESENT DATA RECOVERY KOUROU, PERTH, ODENWALD	17.86 Н	78.14 H	81,39%
DATA LOST DUE TO: - KOUROU SPOT PASS TDF DATA FLOW	н	Н	,-%
- NOT SCHEDULED - LATE A.O.S. (LINES) 3 H OR BAD LINES - STATION OVERLAP - GROUNS STATION OR HDCS PROBLEMS	4.68 H	73.46 H	76.5%
BECAUSE OF ACTIVE VAN ALLI FURTHER SCIENCE DATA AVAIL			
RTAD IN CLOSE LOOP (CBS/ACS TRANSFER ENA.)	н	73.46 Н	76.5%
NO SCIENCE DATA DUE TO 17 OCCULTATIONS	5.95 H	66.54 H	69.31%
DURING REFERENCE PERIOD NO CONVERGENCE HAS BEEN ACHIEVED	12.5 H	54.04 H	56.29%



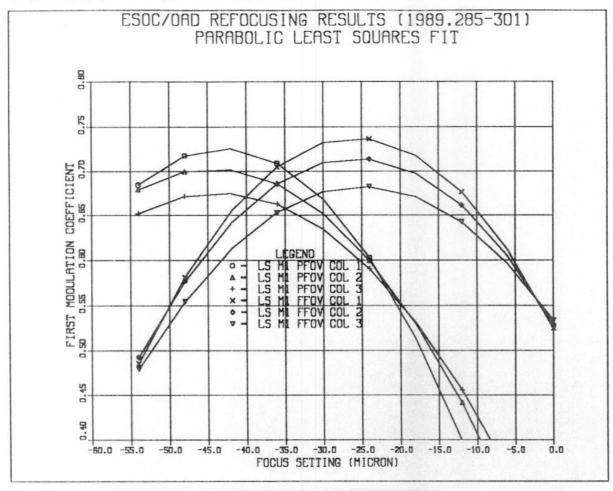
Calibration Status on 8-DEC-1989

- Status presented previously to special calibration meeting held at ESOC on 16-NOV-1989
- Results were presented on the following topics:
 - Detector Noise
 - SSR
 - IDT Piloting focusing, M_1 , M_2 , $\Delta \phi$
 - Transverse Offset
 - Longitudinal Offset
- New results mainly concerned with:
 - M₃ determination
 - Transverse-/Longitudinal- offset corrections
 - IFOV Profile
 - Chromaticity

Brief Summary of 'Old' Results

- Detector Noise
 - Acceptable for SM data above 20000 km (7.5 hours/orbit)
 - Very high for SM during solar flares
 - Fairly low for IDT in all cases
- SSR Filter computed and now used on-board
- IDT Processing
 - M1 values for 10 focus settings fitted by parabolae for 3 colours and 2 FOV (M1=0.69, M2=0.21 for B-V=0.5)
 - . $\Delta\phi$ between 1st and 2nd harmonics approximately 21 mas and fairly independent of colour and position
 - Differential defocus approximately $\pm 10 \mu m$
 - Focus changes by about 1 step/ 4 days
- Transverse Offset
 - Mean over PFOV=-3.8 arcsec, FFOV=-4.3 arcsec
 - Residual offset dependent on position in FOV ∈ (-3.1, +3.2)
- Longitudinal Offset
 - Mean over PFOV=-1.0 arcsec, FFOV=-0.4 arcsec
 - Residual offset dependent on position in FOV ∈ (-3.2, +1.7)
 - Total piloting budget gave bias < 3.4 arcsec everywhere.
 Now improved.

TELESCOPE REFOCUSING



- Collect 45 minutes IDT Data per focus /setting
- Compute M1 for 3 colour groups in each field-of-view
- Fit results using least squares to a parabola
- Differential defocus of +- 9 microns

IDT 3rd harmonic

- Calibrated over central portion of grid for 3 colours and 2
 FOV
- Stars used brighter than B=6, 92 observations reduced using a weighted least squares 7 parameter estimator
- Results distributed asymmetrically i.e. M3 is low and M3>0
- Phase change not well determined but all means < 22 mas

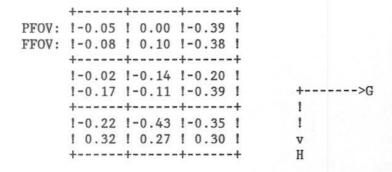
	PFOV			FFOV	
COL1	COL2	COL3	COL1	COL2	COL3
0.009	0.009	0.008	0.010	0.009	0.005

Piloting Improvements

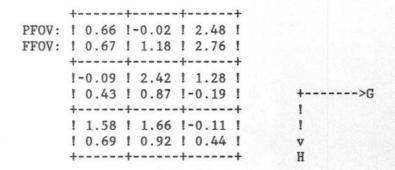
- CCCM modified to empirically remove calibrated residuals (MATRA recommendation)
- Transverse offset residuals now < 0.5 arcsec
- Longitudinal offset residuals still biased by about 1 arcsec (in opposite direction) but always less than 2.8 arcsec. However, these results are not well determined as not many observations were made.
- Overall rms bias of piloting residuals is about 1.1 arcsec
- Grid rotation and drift angle may explain most of the original residuals, therefore, when we get an accurate grid rotation we should:
 - 1. uplink the grid rotation to the spacecraft
 - 2. uplink a CCCM that is NOT empirically corrected to the spacecraft
 - 3. repeat the transverse and longitudinal offset calibrations as a verification the the piloting is good

Piloting Residuals

Transverse Offset (determined to 0.07 arcsec)



Longitudinal Offset (determined to 0.50 arcsec)



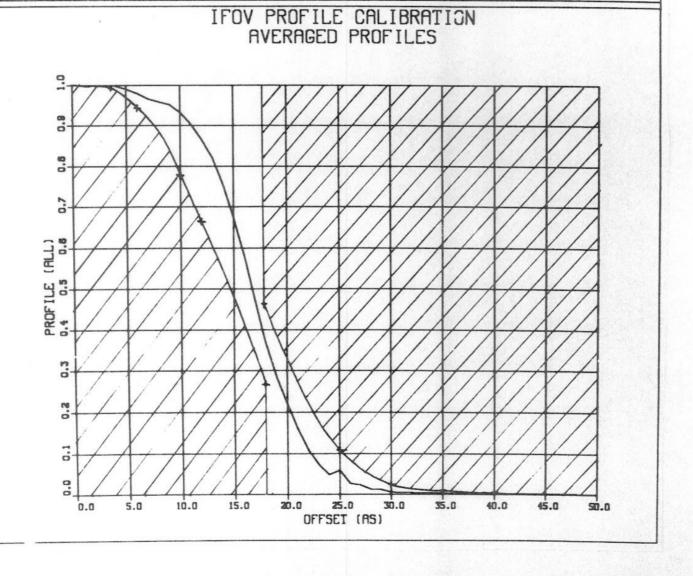
Piloting Residuals

IFOV Profile

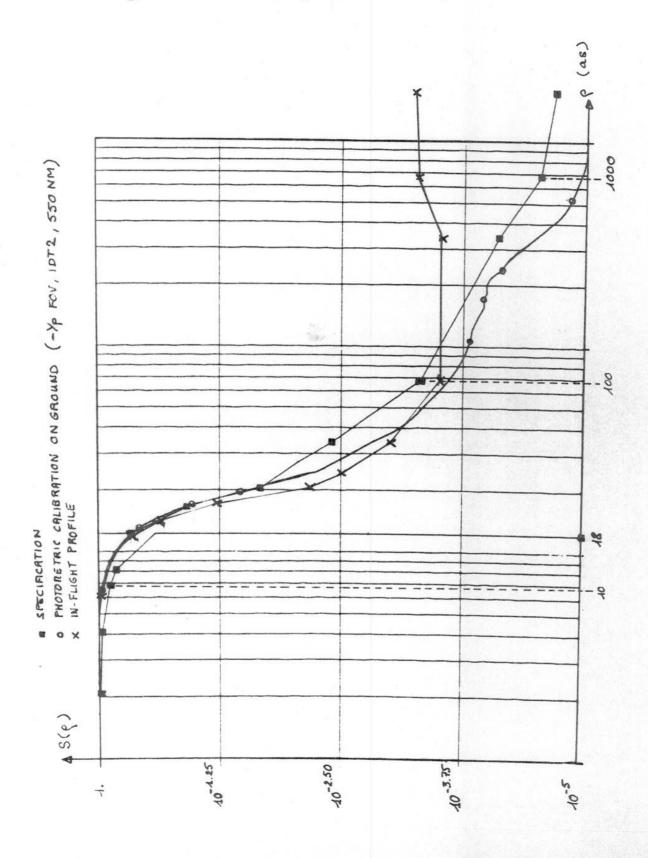
- Calibrated over 12 days from 9-NOV to 21-NOV i.e. a lot of telemetry
- Some cases (large orisets) not well covered
- Individual observations highly sensitive to RTAD convergence, this is a problem as we can't easily tell when RTAD is converged
- Following profiles computed:
 - 15 individual profiles (5 zones, 3 colours)
 - 5 profiles averaged over colour for each zone
 - 3 profiles averaged over zone for each colour
 - 1 profile averaged for zone and colour
- Needs some work to remove statistical 'outliers' when RTAD was not properly converged
- On ground results better for large offsets

Averaged Profile Compared with Specification

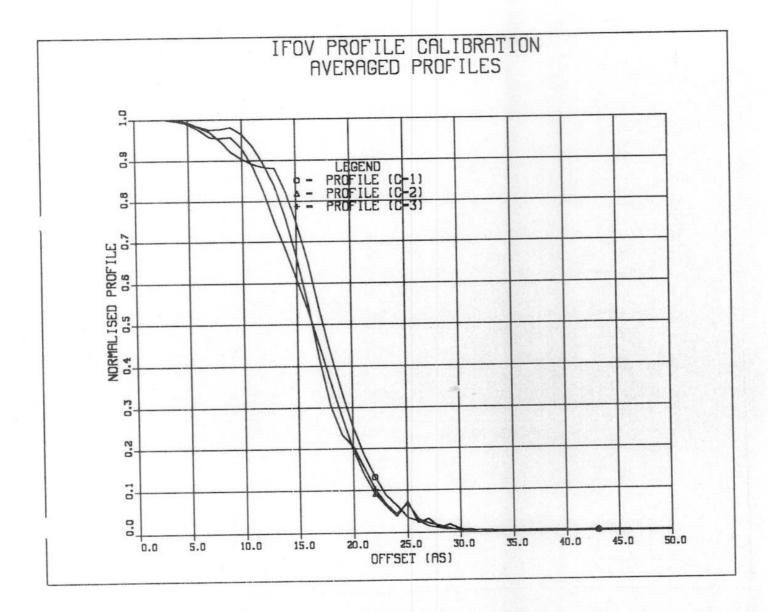
	E EEE	
PROFILE (01) 1.0 0.99829 0.96741	0.95235 0.90575 0.82165 0.75424 0.67294	0.46865 0.36637 0.28°42 0.21981 0.15970 0.07335 0.02753 0.02753 0.02364 0.01315 0.00653 0.00032 0.00032 0.00032 0.00032
OFFSET (arcsec)	9.0 11.0 14.0 15.0	17.0 17.0 18.0 20.0 22.0 23.0 25.0 25.0 25.0 25.0 25.0 30.0 32.0 35.0 100.0 2500.0



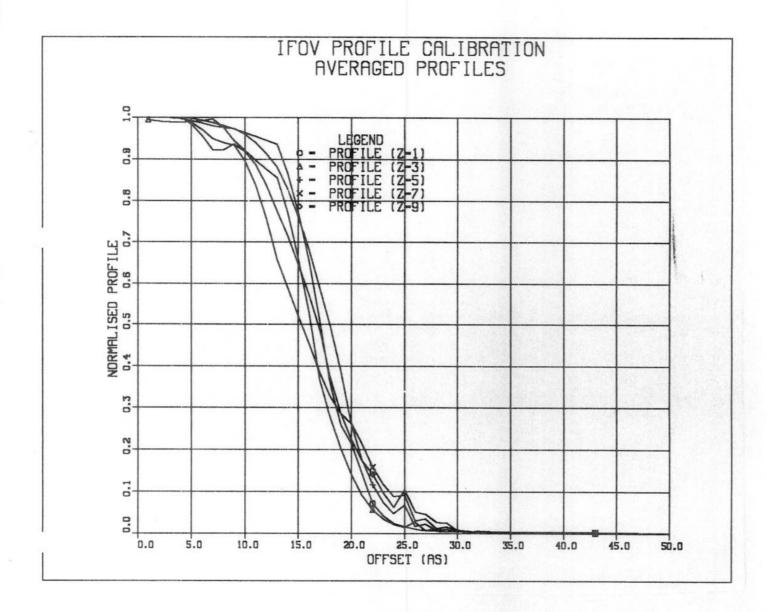
Averaged Profile Compared with Specification



Profiles by Colour Group



Profiles by Zone



Profiles by Zone

Coordinates of spots (are see):

Re	Blue	
(pav)	(pav)	
(-16, +13)	(+43, -8) (+43, -8)	Predicted
(-14, -13)	(+44, +8)	Measures

Measured Chromaticity from 961 observations

	1 0 FE 0 3 FE 0 1 FE 0	₹0.07	18.07
FFOV	±0.72	£0.75	-0.74 =0.24
	£2.0∓ 21.1+	+0.84	+0.42 ±0.23
		5	
			·I
	10.07	-2.03	-1.32 ±0.22
PFOV	+0.17 ±0.28	-1.24	±0.33
	+0.03 ±0.29	40.25	₹0.30

"Content chromoticity" $C_c = -0.27 \pm 0.06$ mas $\begin{cases} spatial anomytes only. \end{cases}$



Annex

GOLDSTONE SUPPORT FOR HIPPARCOS RECOVERY MISSION

89-12-01

Jozef C. Van der Ha

HIPPARCOS Ground Segment Manager ESOC, Darmstadt, FRG

GOLDSTONE SUPPORT for HIPPARCOS RECOVERY MISSION

GROUND STATIONS VISIBILITY STATISTICS

SINGLE STATIONS:

Odenwald (ODN): 34.2 % (ratio of visibility relative to full)

• Perth (PER): 35.2 %

• Kourou (KRU): 39.0 %

TWO & THREE STATIONS:

• ODN + PER: 61.8 %

ODN + PER + KRU: 81.1 %

PERTH IS FULLY OPERATIONAL (since mid September)

KOUROU IS FULLY OPERATIONAL (since begin November)

Maximum Non-Visibility:

..... ODN + PER + KRU: 11 Hrs

ADDING GOLDSTONE as FOURTH STATION:

1. Total Coverage: 93.6 %

2. Maximum Non-Visibility: 1 hr; 18 min

COVERAGE CHARACTERISTICS OF NETWORK OPTIONS

Network	Coverage Per- centage	Maximum Gap
ODN + PER + KRU	81.1	almost 11 hrs
ODN + PER + GDS	91.4	1 hr 27 min
ODN + PER + KRU + GDS	93.8	1 hr 13 min

Table 2. COMPARISON OF COVERAGE CHARACTERISTICS OF NETWORK OPTIONS

ACTUAL COVERAGE SCHEDULE STATISTICS (4 DAY PERIOD)

DAY	ODN	PER	KRU
1	11.5	0.5	-
2	9	3.5	6
3	5.5	8	10
4	7.5	13	2.5
Total	33.5	25	18.5

Table 3. ACTUAL TOTAL STATION COVERAGE INTERVALS (HRS) IN 4-DAY PERIOD

IN TOTAL: ABOUT 77 HOURS OUT OF 96 HOURS (i.e. 80 %)

GOLDSTONE will Add:

About 11 Hours (Single Coverage)

+ About 21 Hours (Redundant Coverage)

GOLDSTONE (SINGLE) VISIBILITY INTERVALS OVER YEAR 90

PERIOD	DAY 1	DAY 2	DAY 3	DAY 4
Mid April	10 - 14	11 - 11.5	-	6.5 - 17
Begin June	2 - 3 & 6 - 12.5	-	-	5 - 9.5
End July	0.5 - 2 & 4.5 - 11	-	-	3.5 - 7
End Sept	1.5 - 8.5	-	-	2 - 4 & 22.5 -24

Table 4. Intervals (in GMT) of Goldstone SINGLE Coverage over Year

Band of Goldstone Visibility Shifts Backward in Day over Year:

from:

2 - 16 (in April)

to:

20 - 8 (in Sept)

Trend:

- 15hr / Month

- 3 min /day

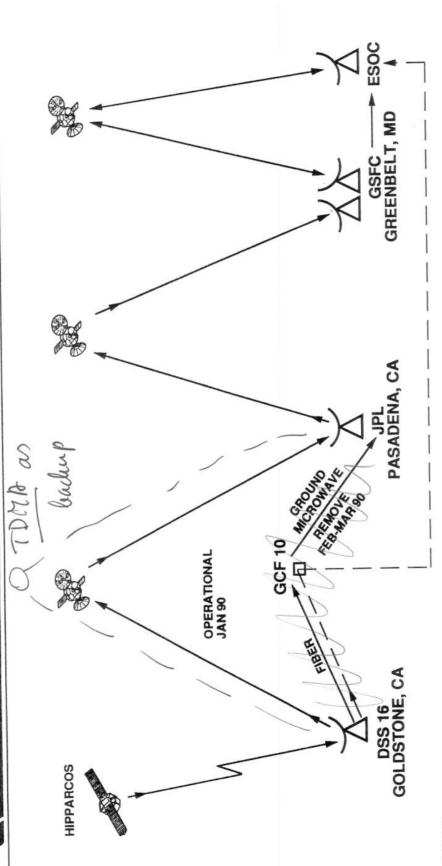
GOLDSTONE (REDUNDANT) VISIBILITY INTERVALS OVER YEAR 90

PERIOD	DAY 1	DAY 2	DAY 3	DAY 4
Mid April	4.5 - 9.5	1 - 11	3.5 - 7.5 & 10.5 - 12	12 - 15
Begin June	3 - 7	0 - 10	1.5 - 6	9.5 - 12
End July	2 - 4.5 & 21.5 -24	0 - 8	0 - 4.5	7 - 9.5
End Sept	0.5 - 2 & 20 - 24	0 - 6 & 20 - 24	0 - 3	4 - 7

Table 5. Intervals (in GMT) of Goldstone REDUNDANT Coverage over Year

HIPPARCOS 56 KB/S DATA ROUTING





LEGEND:



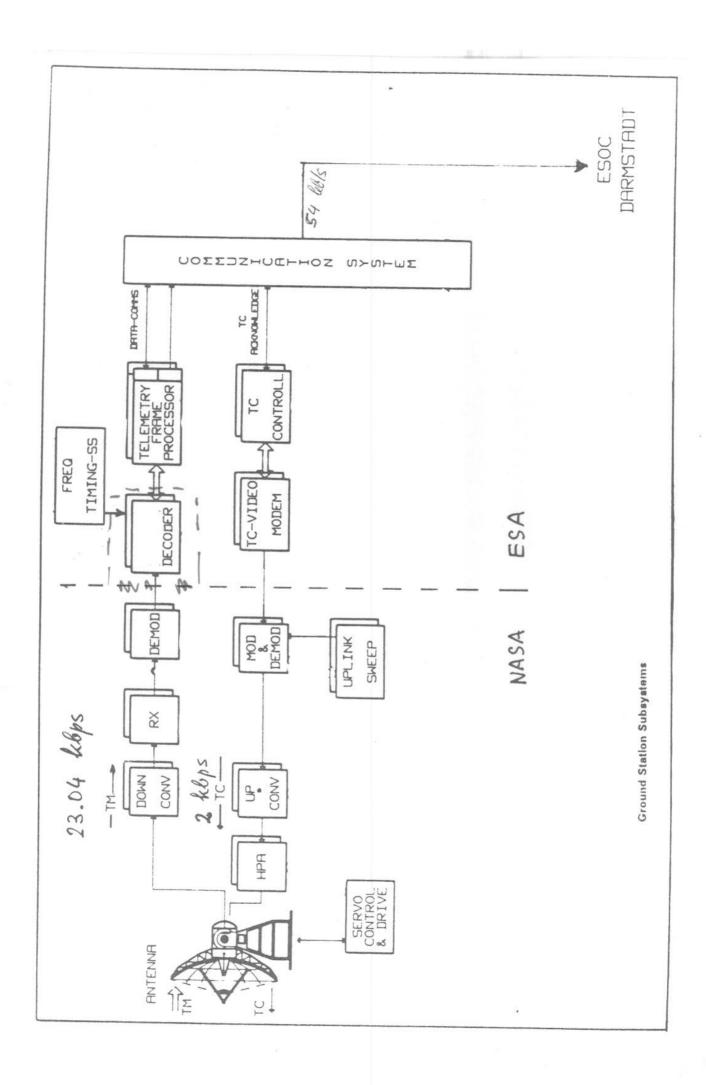
COMMUNICATION SATELLITE

EXPECTED FINAL ROUTING (COM SAT)

— TECHNICALLY POSSIBLE ROUTING (ESOC TO GCF 10 DIRECT)

CARRIER TERMINAL INTERFACE

JPMc 12-6-89



Schedule of Implementation

until end Nov: Finalisation of Implementation Concept

begin Dec: Discussion and Agreement with NASA/JPL

Dec: Procurement of Necessary Equipment

Dec/Feb: Tailoring/Integration of Equipment (at ESOC)

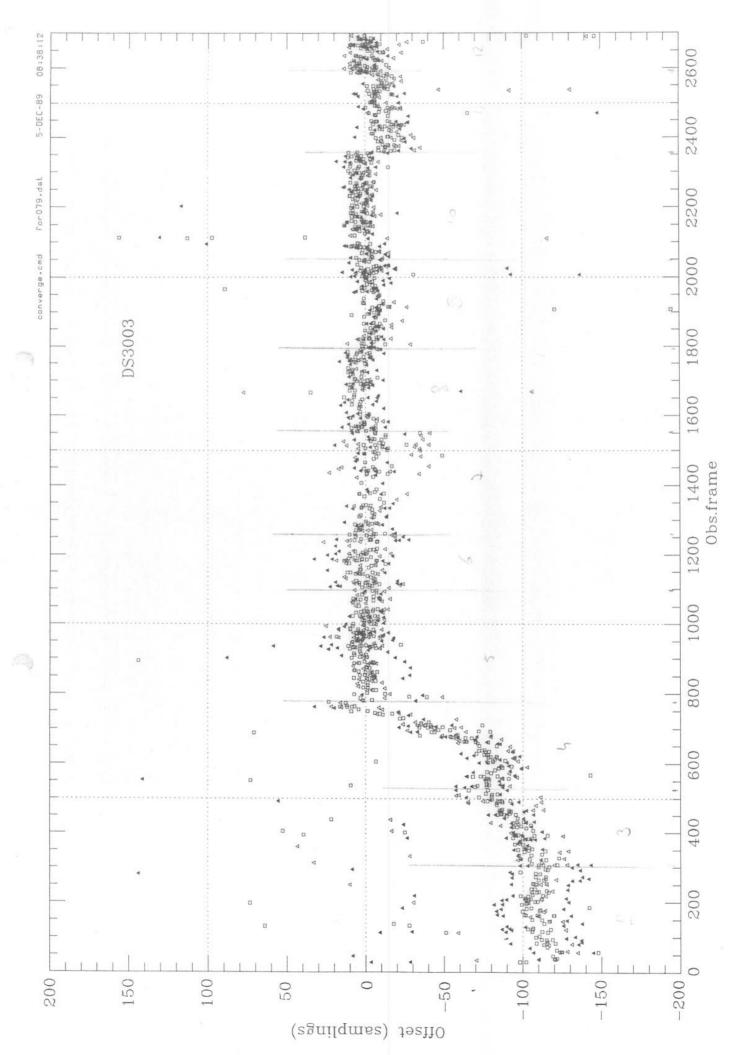
Begin Feb: Dismantling of ODN Redundant Chain

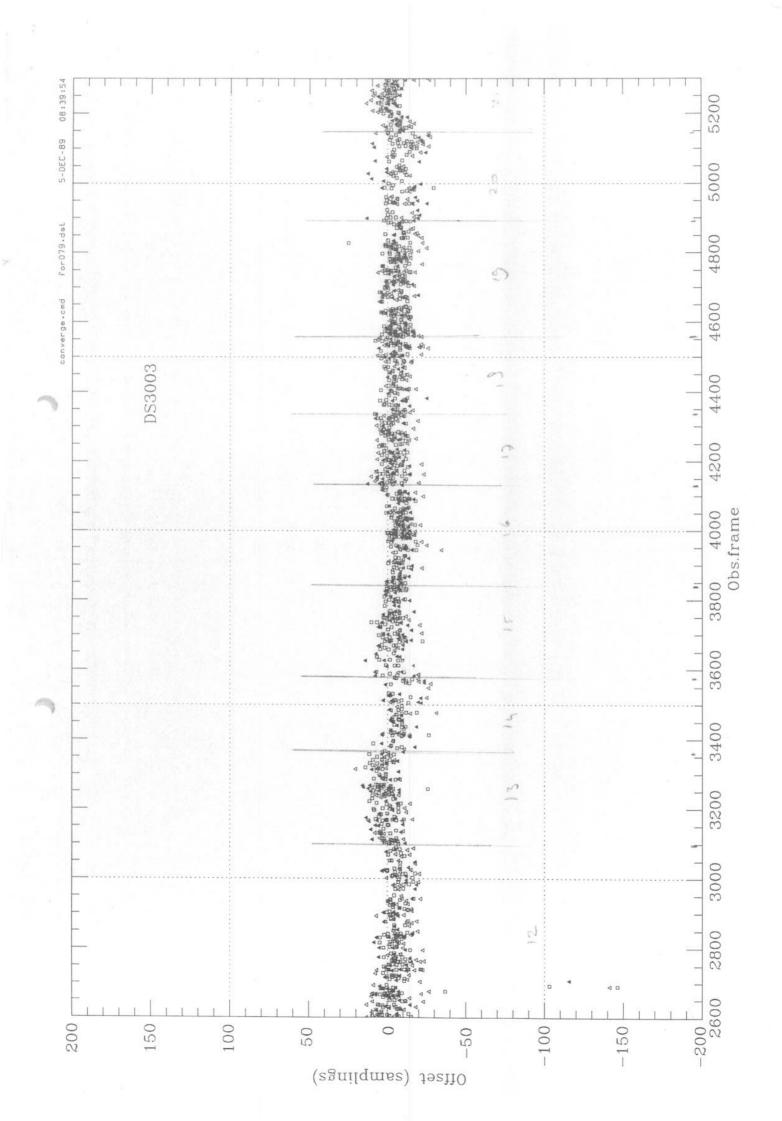
End Feb: Transport of Equipment to USA

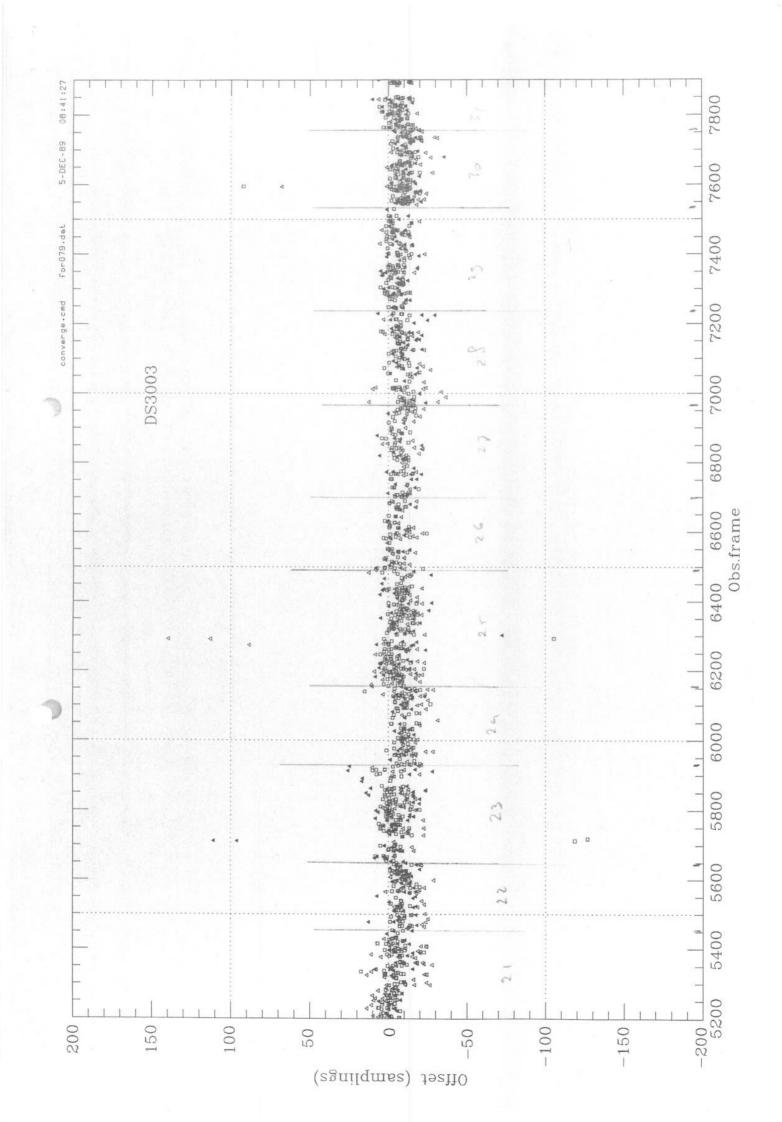
March: Integration/Testing at Goldstone

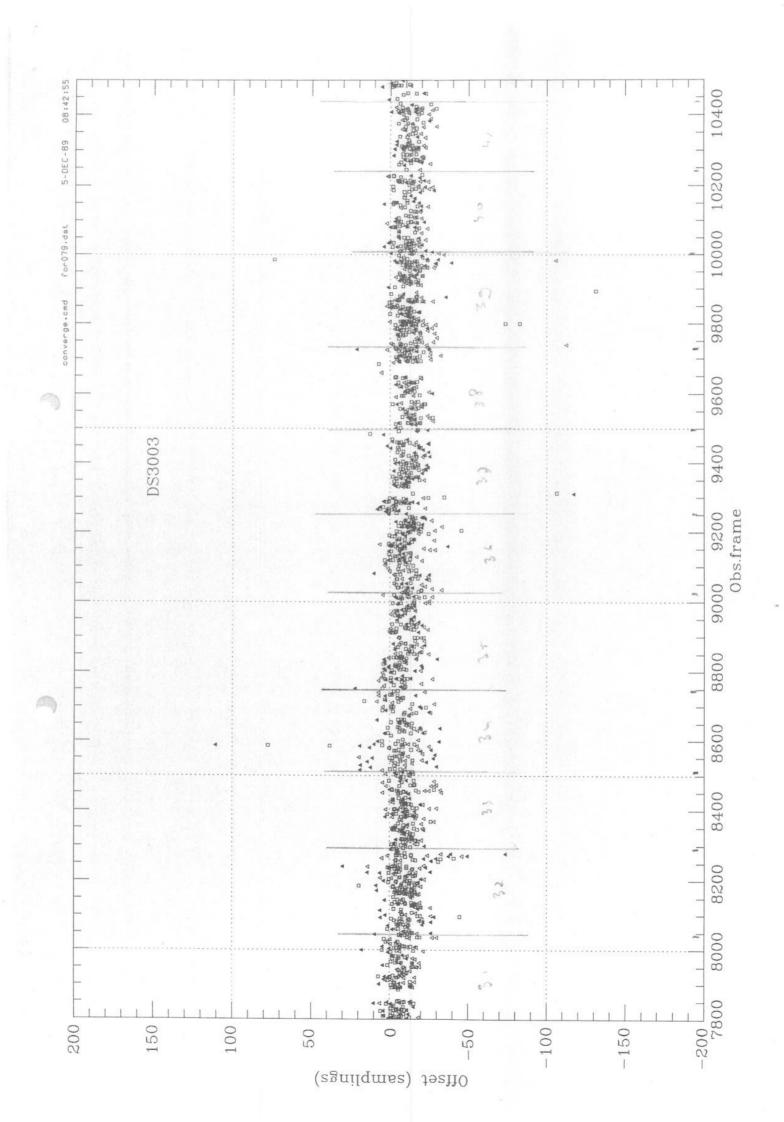
End March: Goldstone Operational

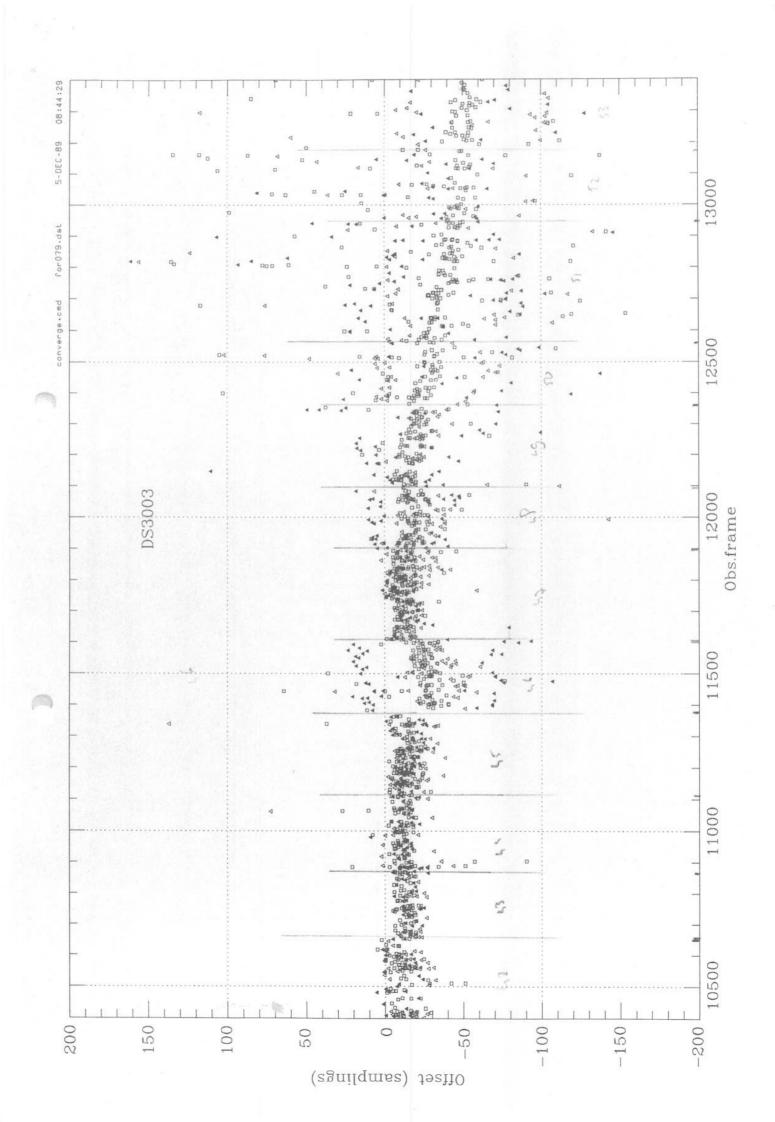
	200.00		
		1990	
	iber December	January	March
ACTIVITIES	13 20 27 4 11 18 25	1 8 15 12 29 5 12 19 26	5 5 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6 6
		JETABT ODN BEMOVA	4STN. OPERATIONAL
MAJOR MILESTONES IMPLEMENTATION CONC.	♣AGREEMENT WITH	WITH NASA	START INTEGRATION GO
56 Kbps LINK	↓ ORDER 2 LINKS	COMMUNICATIONS	JLINK OPERATIONAL
MULTIPLEXERS (4 UN.)	ORDER: ESOC-MAD-NASCOM	\$000 CS	
VOICE LINE;		CLINE AVAIL. 2 MONTHS	SHI
VITERRI DECODER	NGC O	TECHNICAL IMPLEMENTATION	
יויייי דוניייי	CORDER NEW UNIT		
TIME CODE GENERATOR	NDIU FROM MASP	SPA_OMAS	
PCM PREPROC.SPARES	LUNIT TAKEN FR	FRQM ODN	USPARE GTYS
			AND DELIVERY DATES ARE TBD
TC CONTROLLER-SPARES	UNIT TAKEN FROM ODN		
TPCU	ITEM		, at
	MODS, ODW	DDW REMOVAL, ESOC INTEGRATION & TESTING	
TAILORING 6 INTEGRAT			37 60
ODN 2ND CHAIN REMOV.		♦START ESOC	INTEGR.
TRANSPORT TO GOLDSTO	TRANSPORT CHECK US	TRANSPORT, GOLDSTONE INSTALLATION AND TESTING CHECK US EXPORT REDS	
TATE A GOSTAL			COLDSTONE INTEGR.
0			
N progress	current plan	target plan	4 planned mil
ESOC/OPS ISSUE: DRAFT REV.I ORIG.: T.MORENO(SPSD/PCS)	SELO -T MORERO CERED ABORD	00 11 00 1702	THE PROPERTY OF THE PROPERTY O











CERGA. FIRST CALIBRATION OF MA

MEAN POLLOWING FIELD 0.413 ±0.007

MEAN VALUES OVER THE FIELD

H

- G+

0,705	0,701	0,698	0,695	0,693	0,698
0,695	0,690	0,687	0,682	0,675	0,686
0,687	0,687	0,682	0,677	0,672	0,681
0,679	0,681	0,678	0,681	0,671	0,678
0,672	0,679	0,681	0,677	0,674	0,677
0,688	0.688	0,685	0,682	0,677	

RMS pa Square 2: 0030

CERGA. FIRST CALIBRATION OF MZ

MEAN PRECEDING FIELD 0,219 ±0.004

MEAN POLLOWING FIELD 0,213 ±0.004

M2/M1 = 0.297 Puc 0.326 fl

MEAN VALUES OVER THE FIELD

5+

HI

0,224	0,227	0,221	0,223	0,225	0,229
0,215	0,214	0,211	0,210	0,209	0,212
0,216	0,213	6,207	0,208	0,210	0,211
0,214	0,214	0,210	9211	0,214	0,213
0,218	0,218	0,214	0,219	9225	0,219
0.217	0,217	0,213	0,1.14	0.217	

RMS per Square ~ 0.024

CERGA FIRST CALIBRATION OF 4-92

(m.a.s)

MEAN PRECEDING FIELD -19.2 ±5,6

MEAN FOLLOWING FIELD -20,3 ± 4.4

					>6+
1 —					MEAN
-20.7	-19,4	-18,5	-17,8	-18.6	19.0
-21,5	-20.5	-20,3	- (8,4	-18,1	19.8
-21,9	-21.3	-20,2	-18,9	-16.8	19.8
-21,5	-20,9	-19.9	-18,8	-16,0	19.4
-19,6	-19,3	-20,1	-17.5	-16,9	18.7

RMS per square ~ 12 mas

