



Mathematical Best Fit Algorithm for Energy Distributions in VO Context

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Project Setup

Name

Mathematical fitting algorithms for Energy Distributions in VO context

Aim

To implement an algorithm that looks for the theoretical model that best fits a given Spectral Energy Distribution (SED)

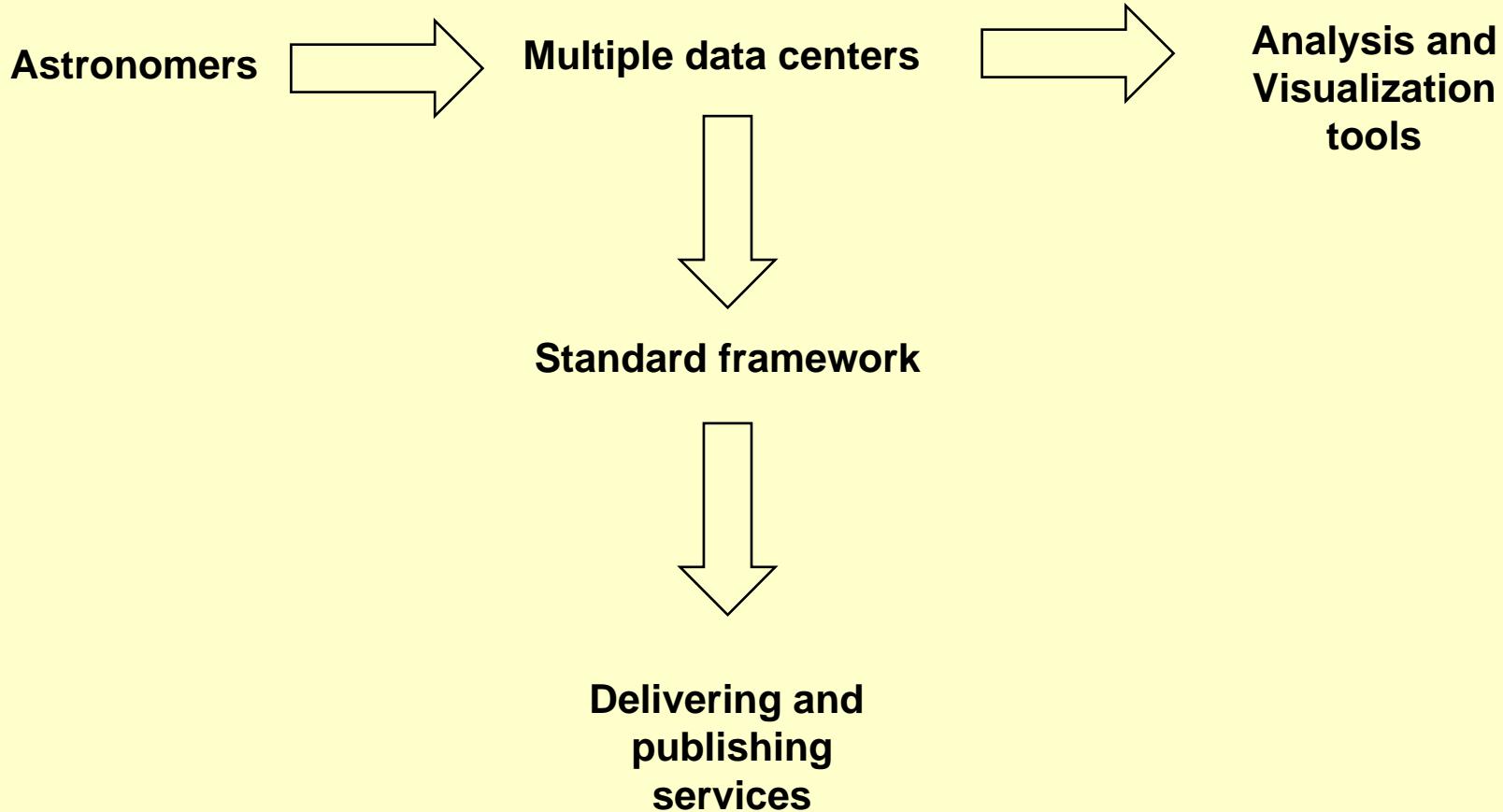
Tutors

**Pedro Osuna (ESA)
Jesus Salgado (INSA)**



Virtual Observatory

- ❑ Its aim is to offer seamless access to astronomical data worldwide





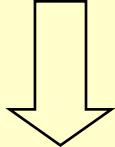
Virtual Observatory (cont.)

- In order to fulfill this aim the International Virtual Observatory Alliance (IVOA) was created in June 2002

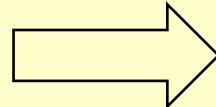
- Data
 - management
 - analysis
 - distribution
 - interoperability



- Data Access Layer (DAL)
 - working group within IVOA
 - standards for remote observational data access



Simple Spectral Access Protocol (SSAP)



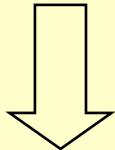
To define a uniform interface
to spectral data



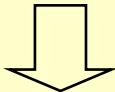
Virtual Observatory (cont.)

- Theory Group

- interest group within the IVOA
- formed in January 2004
- define requirements needed for a full interoperability between observations and models



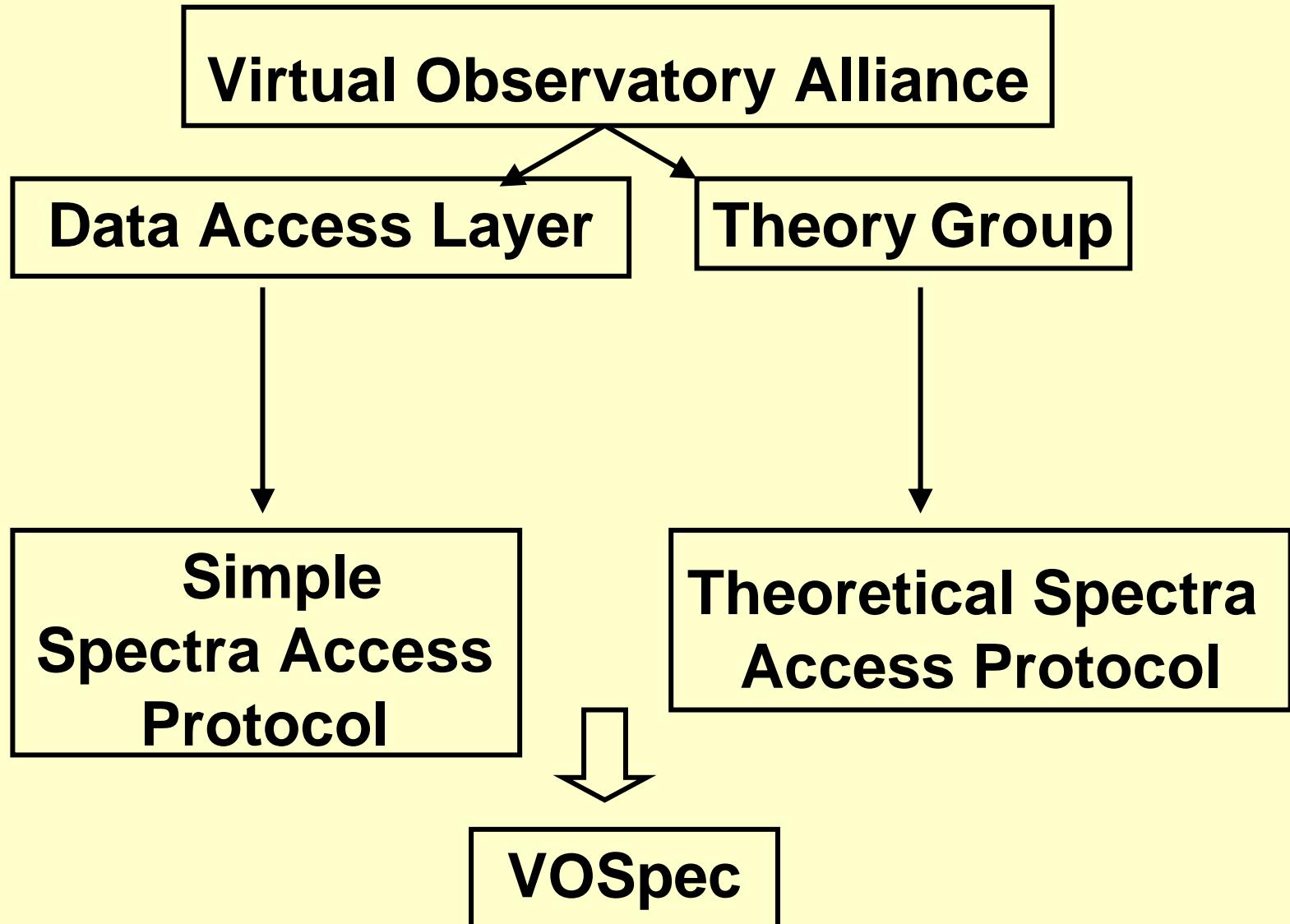
Theoretical Spectrum Access Protocol (TSAP)



Incorporate theoretical models to the VO



Project environment





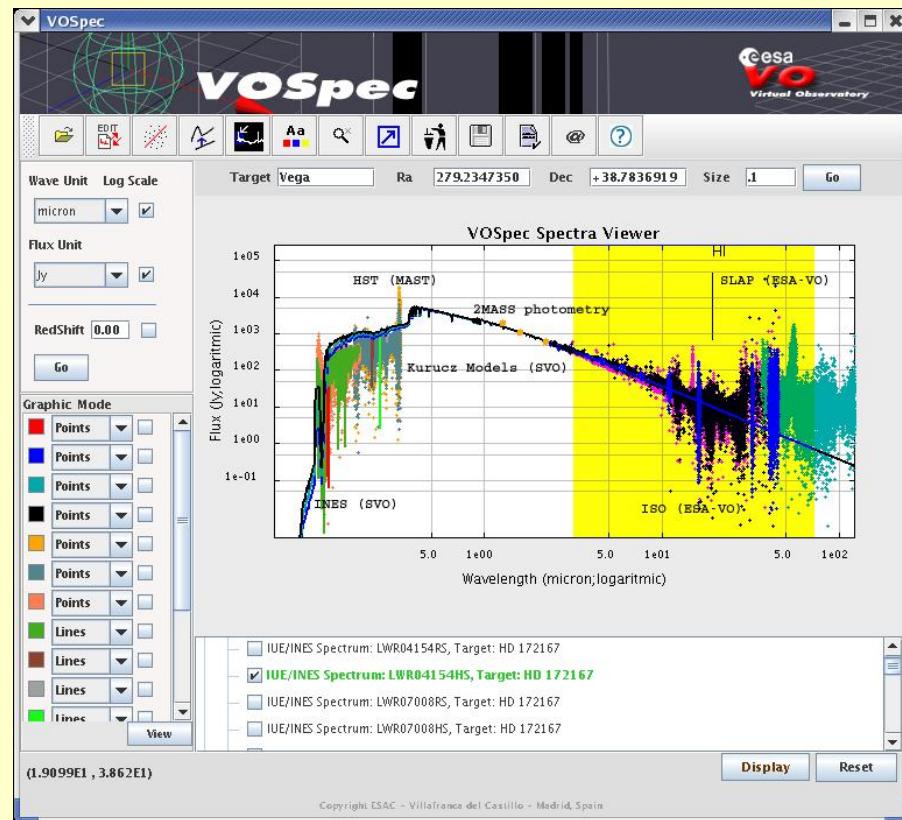
VOSpec

❑ Tool for handling

- VO compliant spectra through **SSAP**
- Theoretical models through **TSAP**

❑ Handling spectra with VOSpec:

- Display and superimpose spectra
- Automatic unit conversion through dimensional analysis
- Multi-wavelength analysis
- Polynomial/Black Body/Gaussian fitting





Why Best Fit?

SVO: ATLAS9 Kurucz ODFNEW/NOVER models (Castelli et al.)

-TSA Service Description & Options

SVO Theoretical Data Access Service: ATLAS9 Kurucz ODFNEW/NOVER models (Castelli et al., 1997, AA, 318, 841)

teff_min	3500	description	phys.temperature.effective
teff_max	3500	description	phys.temperature.effective
logg_min	3750	description	phys.gravity
logg_max	4000	description	phys.gravity
meta_min	4250	description	phys.abund.Fe
meta_max	4500	description	phys.abund.Fe

Select Close

Currently to find the model that best fits a given SED:

- Choose parameters manually
- Inspect result visually (“Chi-by-eye”)
- Modify parameters manually
- Reinspect visually
- Loop.....

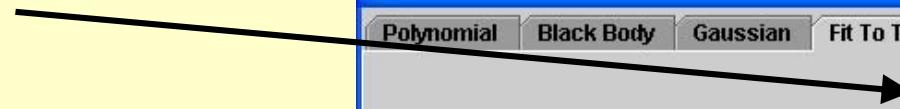


Best Fit finds the model automatically



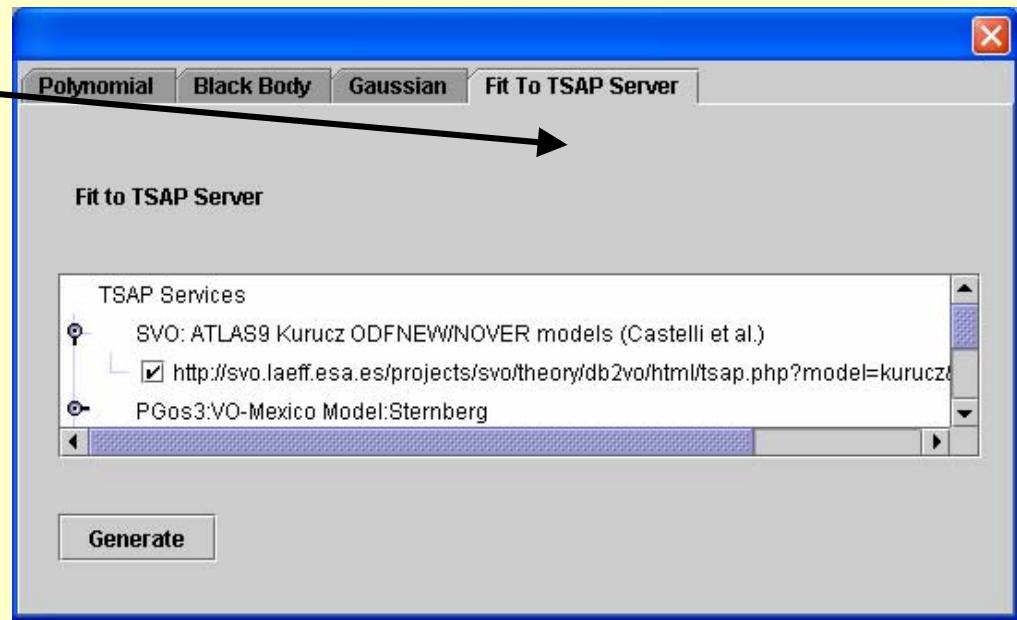
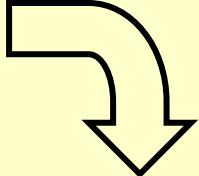
Best Fit brief description

- ❑ It is going to be integrated as a new fitting utility in **VOSpec**



- ❑ Java language

- ❑ Aim



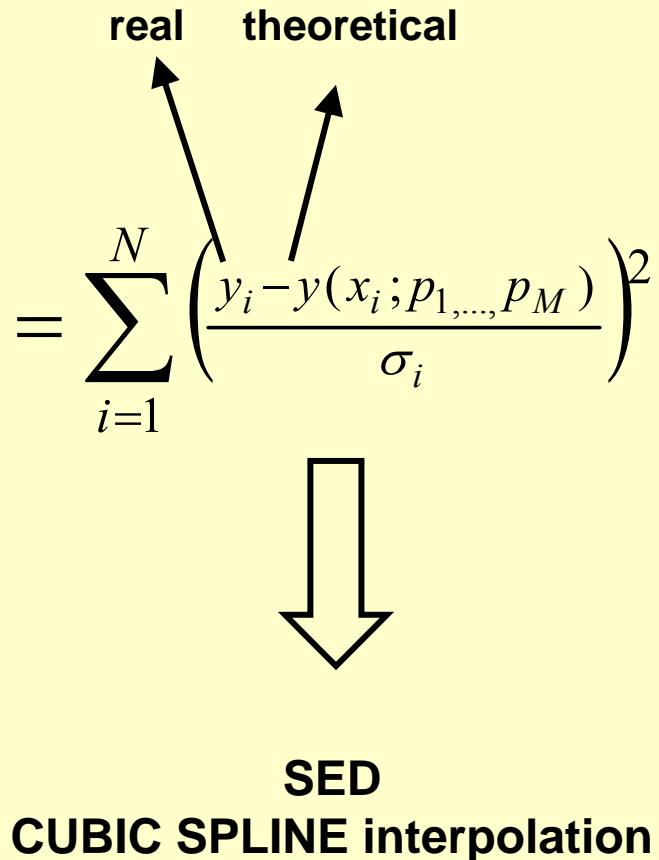
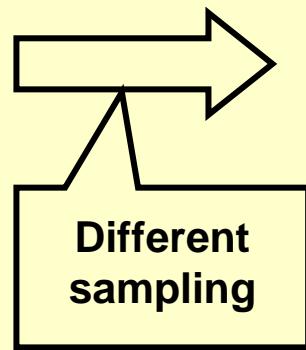
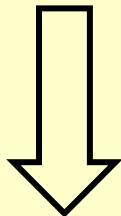
Find the theoretical model
that best fits the SED



Best Fit brief description (cont.)

□ To fulfill this aim:

- Chi-square function measures the agreement between real and theoretical spectrum
- The best fit spectrum is the one that achieves a minimum in the Chi-square function
- Problem in minimization in many dimensions

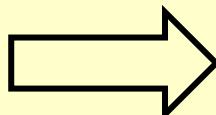


Levenberg-Marquardt method



Levenberg-Marquardt method

- ❑ Iterative method
- ❑ Standard technique for nonlinear least-square problems



Modified for discrete case

- ❑ Combination of
 - ↗ Steepest descent method
 - ↗ Gauss-Newton method

❑ Steepest descent method

- Initial estimate $\xrightarrow{\hspace{1cm}}$ \mathbf{x}_0
- Successive estimate $\xrightarrow{\hspace{1cm}}$ $\mathbf{x}_{i+1} = \mathbf{x}_i - c \cdot \nabla \chi^2(\mathbf{x}_i)$
- Iterations stop when $\nabla \chi^2(\mathbf{x}_i)$ is sufficiently small
- Hard to determine the step size
- Slow, but guaranteed convergence



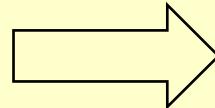
Levenberg-Marquardt method (cont.)

□ Gauss-Newton method

- Taylor expansion

$$\chi^2(\mathbf{x} + \Delta\mathbf{x}) = \chi^2(\mathbf{x}) + \nabla \chi^2(\mathbf{x}) \cdot \Delta\mathbf{x} + \frac{1}{2} \cdot \Delta\mathbf{x}^T \cdot \mathbf{H}(\mathbf{x}) \cdot \Delta\mathbf{x}$$

- Set the derivative of the previous expression equal to 0
- Quadratic convergence



Iterative equation

$$\mathbf{x}_{i+1} = \mathbf{x}_i - \mathbf{H}^{-1} \cdot \nabla \chi^2(\mathbf{x}_i)$$



Levenberg-Marquardt method (cont.)

□ Steepest descent and Gauss-Newton methods are complementary in the advantages they provide



Levenberg Marquardt Method

$$\mathbf{x}_{i+1} = \mathbf{x}_i - c \cdot \nabla \chi^2(\mathbf{x}_i)$$

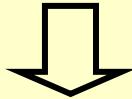
$$\mathbf{x}_{i+1} = \mathbf{x}_i - \mathbf{H}^{-1} \cdot \nabla \chi^2(\mathbf{x}_i)$$

$$\mathbf{x}_{i+1} = \mathbf{x}_i - (\mathbf{H} + \lambda \mathbf{I})^{-1} \cdot \nabla \chi^2(\mathbf{x}_i)$$



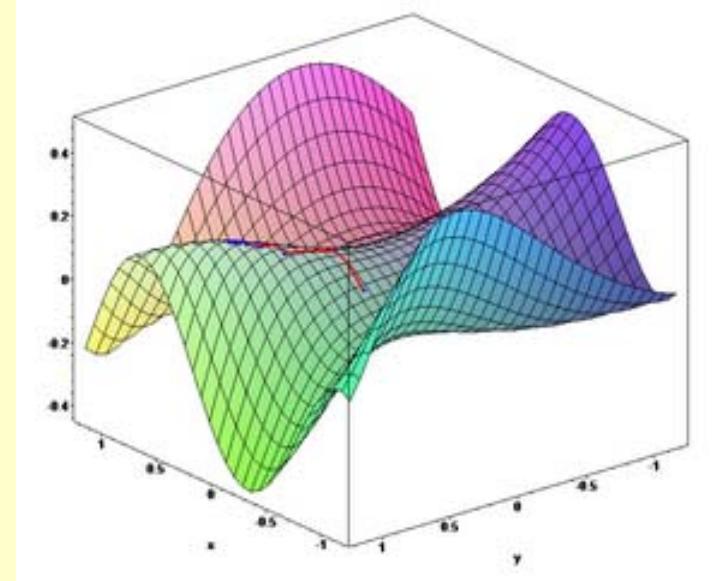
Levenberg-Marquardt method (cont.)

- Far from the minimum
 - Linear approximation with steepest descent method to the closest minimum
 - Adds a measure of the step in the steepest descent method



Close to the minimum
Gauss-Newton convergence

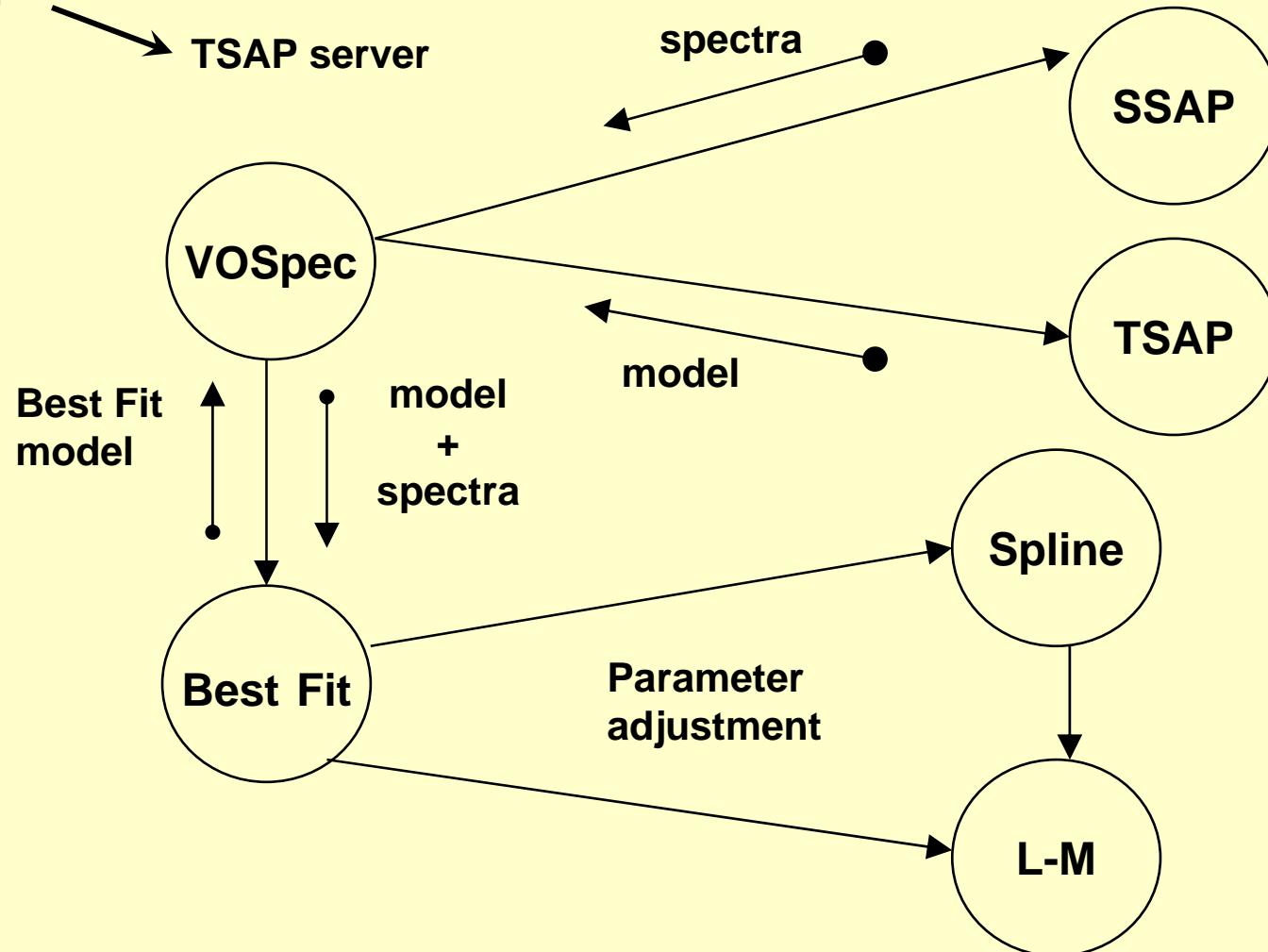
- L-M method works well in practice





Best Fit diagram

- ❑ Given
 - Real spectrum from SSAP server
 - TSAP server





Conclusions

- IVOA standards are nearly mature enough to do real science
- BF algorithm implementation using TSAP protocol increases VO functionality
- Nonlinear least-square methods (LM) are very appropriate for the implementation of the algorithm
- Very easy integration with ESA VOSpec tool