## The Cosmic Battery in X-ray binaries

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GRS 1915+105 (Mirabel 1994)







## Where does the magnetic field come from?

- The Cosmic Battery (C, Kazanas 1998, ApJ)
- The Cosmic Battery Revisited (C, Kazanas, Christodoulou 2006, ApJ)
- Simulations of the Poynting-Robertson Cosmic Battery in Resistive Accretion Disks (Christodoulou, C, Kazanas 2008, ApJ)



 $-\frac{L\sigma_T}{4\pi r^2 c}\frac{v_\phi}{c}$  $F_{PR} =$ 

#### protons electrons

• The Cosmic Battery (Contopoulos, Kazanas 1998)

$$m_e \left(\frac{d\mathbf{v}_p}{dt}\right) = -e \left(\mathbf{E} + \frac{\mathbf{v}_e \times \mathbf{B}}{c}\right) + m_e \mathbf{g} + \mathbf{F}_{rad}$$
$$m_p \left(\frac{d\mathbf{v}_p}{dt}\right) = +e \left(\mathbf{E} + \frac{\mathbf{v}_p \times \mathbf{B}}{c}\right) + m_p \mathbf{g}$$
$$\mathbf{E} + \frac{\mathbf{v}_e \times \mathbf{B}}{c} = \frac{\mathbf{F}_{rad}}{e}$$
$$\frac{d\mathbf{B}}{e} = -c\nabla \times \mathbf{E}$$

$$dt = \nabla \times (\mathbf{v}_e \times \mathbf{B}) - \nabla \times \left(\frac{c\mathbf{F}_{\text{rad}}}{e}\right)$$

• The Biermann Battery (Biermann 1950)

$$n_{e}m_{e}\left(\frac{d\mathbf{v}}{dt}\right) = -n_{e}e\left(\mathbf{E} + \frac{\mathbf{v} \times \mathbf{B}}{c}\right) - \nabla p_{e} + n_{e}m_{e}\mathbf{g} + \mathbf{F} \cdot \mathbf{O}\mathbf{h}\mathbf{m'}$$
$$\mathbf{E} + \frac{\mathbf{v} \times \mathbf{B}}{c} = -\frac{\nabla p_{e}}{n_{e}e} + \frac{\mathbf{F} \cdot \mathbf{O}\mathbf{h}\mathbf{m'}}{n_{e}e}$$

$$\frac{d\mathbf{B}}{dt}$$

$$= -c\nabla \times \mathbf{E}$$

$$= \nabla \times (\mathbf{v} \times \mathbf{B}) + \nabla \times \left(\frac{c\nabla p_e}{n_e e}\right)$$

$$= \nabla \times (\mathbf{v} \times \mathbf{B}) - c\frac{\nabla n_e \times \nabla p_e}{n_e^2 e}$$

$$-\nabla \times \left(\frac{c\mathbf{F} \cdot \mathrm{Ohm'}}{n_e e}\right)$$

• Vorticity and magnetic fields

$$\begin{split} \frac{d\mathbf{v}}{dt} &+ (\mathbf{v} \cdot \nabla)\mathbf{v} = -\frac{1}{\rho}\nabla p - \mathbf{g} \\ \frac{d\omega}{dt} &= \nabla \times (\mathbf{v} \times \omega) + \frac{\nabla \rho \times \nabla p}{\rho^2} , \quad \omega \equiv \nabla \times \mathbf{v} \\ \frac{d\mathbf{B}}{dt} &= \nabla \times (\mathbf{v} \times \mathbf{B}) - \frac{\nabla n_e \times \nabla p_e}{n_e^2 e} \end{split}$$

**OBJECT PARAMETERS** 

Object	$r_{in}$ (cm)	$L (L_{\odot})$	$M (M_{\odot})$	$\frac{\tau_{eq}}{(yr)}$	$\frac{B_{\rm eq}}{(\rm G)} \equiv \frac{\Psi_{\rm eq}}{r_{\rm in}^2}$
Black Hole Neutron Star White Dwarf Protosun AGN Galaxy	$3 \times 10^{6}$ $10^{6}$ $10^{8}$ $7 \times 10^{9}$ $10^{13}$ $10^{22}$	10 <sup>4</sup> 10 <sup>4</sup> 10 332 10 <sup>12</sup> 10 <sup>10</sup>	10 1 10 <sup>-1</sup> 10 <sup>8</sup> 10 <sup>10</sup>	$7 \times 10^{-4}$ $10^{-3}$ $2 \times 10^{4}$ $2 \times 10^{9}$ $7 \times 10^{7}$ $6 \times 10^{30}$	$5 \times 10^{7}$ $2 \times 10^{7}$ $2 \times 10^{4}$ $8 \times 10^{3}$ $5 \times 10^{3}$ $6 \times 10^{-5}$

- The Biermann Battery (Biermann 1950)
- The Cosmic Battery (Contopoulos, Kazanas 1998):
  - Rotation of solar coronal layers (Cattani, Sacchi 1966)
  - Rotation of protogalaxies with respect to the cosmic microwave background radiation (Harrison 1970; Mishustin, Ruzmaikin 1972)
  - Ring currents and poloidal magnetic fields in nuclear regions of galaxies (Lesch et al. 1989)
  - A Cosmic Battery Reconsidered (Bisnovatyi-Kogan, Lovelace, Belinski 2002)
  - Disk accretion onto a black hole at subcritical luminosity (Bisnovatyi-Kogan, Blinnikov 1977)





 $E_r \neq 0$  $\begin{array}{c} \rightarrow \rightarrow \rightarrow \rightarrow \\ \rightarrow & E_r = 0 \end{array}$  $\rightarrow$  $E_r \neq 0$ 





#### Numerical simulations



Contopoulos & Kazanas 1998

#### Numerical simulations



#### Contopoulos, Kazanas & Christodoulou 2006

#### Numerical simulations

- The Poynting-Robertson effect in GR: Monte Carlo ray-tracing (Koutsantoniou, Contopoulos 2013)
- Kinematic integration of the induction equation
- 2D/3D GR resistive MHD simulations with Cosmic Battery source term
- Where is the ISCO? (Contopoulos, Papadopoulos 2012, 13)

#### Jet formation

Evidence so far indicates that the production of Poynting-flux dominated relativistic jets requires a steady large-scale dipolar field near the black hole

(McKinney, Tchekhovskoy, Blandford 2012)

The large-scale dipolar field required for jet formation is produced through the Poynting-Robertson drag on plasma electrons at the inner edge of the accretion disk around the central black hole

(Contopoulos, Kazanas 1998)

#### Observational evidence

- Extragalactic jets:
  - The Invariant Twist of Magnetic Fields in the Relativistic Jets of Active Galactic Nuclei (C, Kazanas, Christodoulou, Gabuzda 2009, ApJ)
  - Evidence for Helical Magnetic fields in kpc-Scale AGN Jets and the Action of a Cosmic Battery (Gabuzda, Christodoulou, C, Kazanas 2012, JPCS)
- X-ray Binaries
  - Formation and destruction of jets in X-ray binaries (Kylafis, C, Kazanas, Cristodoulou 2012, A&A)
  - The HID in X-ray binaries (Kazanas, Kylafis, C 2013)









sources. Adding this information to the above indicates that the 3C303 jet has an average current within  $\simeq 400$  pc of its axis, which is  $7 \times 10^{17}$  ampères, and directed away from the central BH (AGN) (Ji *et al.*, 2008). It is the largest electrical current ever measured.

This result is of particular interest in the context of poynting flux energy flow models for extragalactic jets, as we shall discuss in the next section.

#### The Cosmic Battery in X-ray binaries: the Hardness-Intensity Diagram (HID)

#### Spectral transitions



Grebenev et al. 1993

#### Spectral transitions



Fender 2001



Belloni 2010



Kording et al. 2008

### Jet formation<sup>01</sup> and destruction



RXTE PCA count s<sup>-1</sup>

1(

### Jet formation<sup>0,01</sup> and destruction



#### Jet formation and destruction



Kylafis, Contopoulos, Kazanas, Christodoulou 2012; Koutsantoniou, Contopoulos 2013

#### Other observations

- AGN MHD disk winds with  $B \sim 1/r$ ,  $n \sim 1/r$ (Fukumura et al. 2012)
  - Not compatible with disk solutions (Fereira, Li '90s)
  - The magnetic field is not advected in, it diffuses out (Sikora, Begelman 2013; Lovelace et al.)

## Thank you for your attention





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