

The ALMA Science Archive

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ALMA

- 66 antennas at 5000m elevation in the Atacama Desert
- Interferometry with baselines up to 16km
- Wavelength range from 3mm to 400μ m (84 to 702 GHz)
- Average data rate in full operations: 200TB/yr=6.6Mbytes/s
- Built by ESO, NRAO, NAOJ in cooperation with Chile

- Credits: the whole ALMA Archive team
- Acknowledgement: long discussions with the (former) HST teams at CADC and ESO

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The ALMA Work Flow



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ALMA Science



Credits: ALMA (ESO/NAOJ/NRAO). Visible light image: the NASA/ESA Hubble Space Telescope, ALMA (ESO/NAOJ/ NRAO)/M. Maercker et al., *A.C. Boley et al., ALMA (ESO/NAOJ/NRAO), T.A. Rector (University of Alaska Anchorage). Visible-light image: ESO*

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Science Archive Rationale

- As the success of ALMA is measured by the science output of the community, the goal is to render the user-experience for PIs and archival researchers as perfect as possible
- Creating a Science Archive that allows archival rearchers to discover and retrieve the data they can use, helps maximise the scientific return of ALMA

Great return-for-investment ratio



Science Archive Design

- speak the language of the scientists (query by physical concepts)
- provide only relevant information, keep hurdles small
- help users to quickly decide whether or not the data returned in a given search is relevant for them

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- complete, correct, consistent, homogenized
- fast responses
- programmatic access to metatada and data
- one small set of optimized database tables
- use existing standards and technologies
- iterative development and evolution

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Science Archive Policies

- All public data and metadata is accessible **anonymously**
- There will be **no private metadata** in the ASA
- PIs are normal archive users with more access rights
- Authors are required to acknowledge the use of ALMA data using a standard data tag
- Metadata will be available when the first data has been ingested into the archive. Titles, PI/CoI names, abstracts, project codes directly after Science Assessment.



Current version (almascience.org/aq)

Atacama Large Millimeter/Submillimeter Array In search of our Cosmic Origins

You are here: Home > ALMA Data > Archive Query

ALMA Science Archive Query

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rm Result Table

Search Reset

Query Help

Position	Energy	Time	Polarisation
Source name (Sesame)			Polarisation type
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Source name (ALMA)	Source name to be resolved with Sesame.	V* R Scl	
RA Dec			
Search radius	Description	Coordinates (RA Dec)	
0:10:00	We use the external resolver Sesame to obtain coordinates for	01:26:58.09 -32:32:35.4	
	the source. Sesame gueries the	Object type	
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Water vapour	Example	Resolver	Results view
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Observe target	NGC3375		
			Release status
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Still very limited

Only raw data searches possible

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You are here: <u>Home</u> > <u>ALMA Data</u> > Archive Query			
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Add/remove displayed columns

Drag & drop columns above or below the red bar, move the red bar itself or click on the checkboxes.

der columns

drop the columns or drag & drop the column headers directly in the results table

Atacama Large Millimeter/S

In search of our Cosmic Origins

all columns Reset column order Order alphabetically

✓ project_code	Project code, in the form YYYY.NNNNN.C.AAA, where:			
	Name of the source as registered in the ASDM. Partial matches through wildcards (?, *), and			
V SOORCE_NAME	boolean OR expressions (" "), can be used.	More columns		
🗹 RA	Right Ascension of the field pointing. (Default unit 'deg')	resolution		
☑ DEC	Declination of the field pointing. (Default unit 'deg')			
BAND	ALMA receiver band.	/s 🗸)		
✓ integration	Aggregated integration time for the field in the ASDM. (Default unit 's')	58.00		
RELEASE_DATE	RELEASE_DATE	58.00		
	Estimated velocity resolution from all the spectral windows, from frequency resolution. (Default	58.00		
	unit ' m/s ')	58.00		
		59.00		
	Estimated frequency resolution from all the spectral windows, using median values of channel	58.00		
freq_resolution	widths. (Default unit ' kHz')	58.00		
POL_PRODUCTS	Polarisation products provided.	58.00		
□ start_date	to_char(asa_science.start_date, 'YYYY-MM-DD HH24:MI:SS')	58.00		
□ PI_NAME	case-insensitive partial match over the full PI name. Wildcards can be used	76		
	Estimated precipitable water vapour from the XML_CALWVR_ENTITIES table. (Default unit	76		
PWV	'mm')			

Similar to HST interface (now at ESA!)

Code reuse from CADC (VO technology)

2011.0.00131.S

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Current version

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• ESO's SAF and ALMA use the same Request Handler



Evolution: Division of Labour







Evolution of Astronomy

- ALMA will produce about the same amount of data in one year as ESO has produced in its first 50 years
- LSST, LOFAR, SKA, PanStars, Euclid, Gaia, ELTs
- T. Tyson: Astronomy is transformed from being a datastarved science to one where data is overabundant
- Multi-wavelength science: less time per wavelength regime
- Astronomers do not scale: bytes/astronomer grow exponentially
- Now: astronomers compete for observing time
 Future: observatories will compete for astronomers to work with their data



Conclusion

- The goal of ALMA is to help the scientists wherever possible from proposal preparation over the sciencegrade data products and data-reduction to archival research and to work towards a great end-to-end userexperience (by "turning around" and looking at the whole workflow from a user's point of view)
- Part of this effort is to construct a powerful Science Archive which helps maximising the scientific return of the observatory

