

# The detection of transiting extrasolar planets with CoRoT -Light curves analysis-

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# The CoRoT mission

- Launched December 2006
- Find extrasolar planets using the Transit Method
- Continuously monitoring ~12.000 stars per observing field with magnitudes [11.5-16]
- Mission duration: 2.5 years



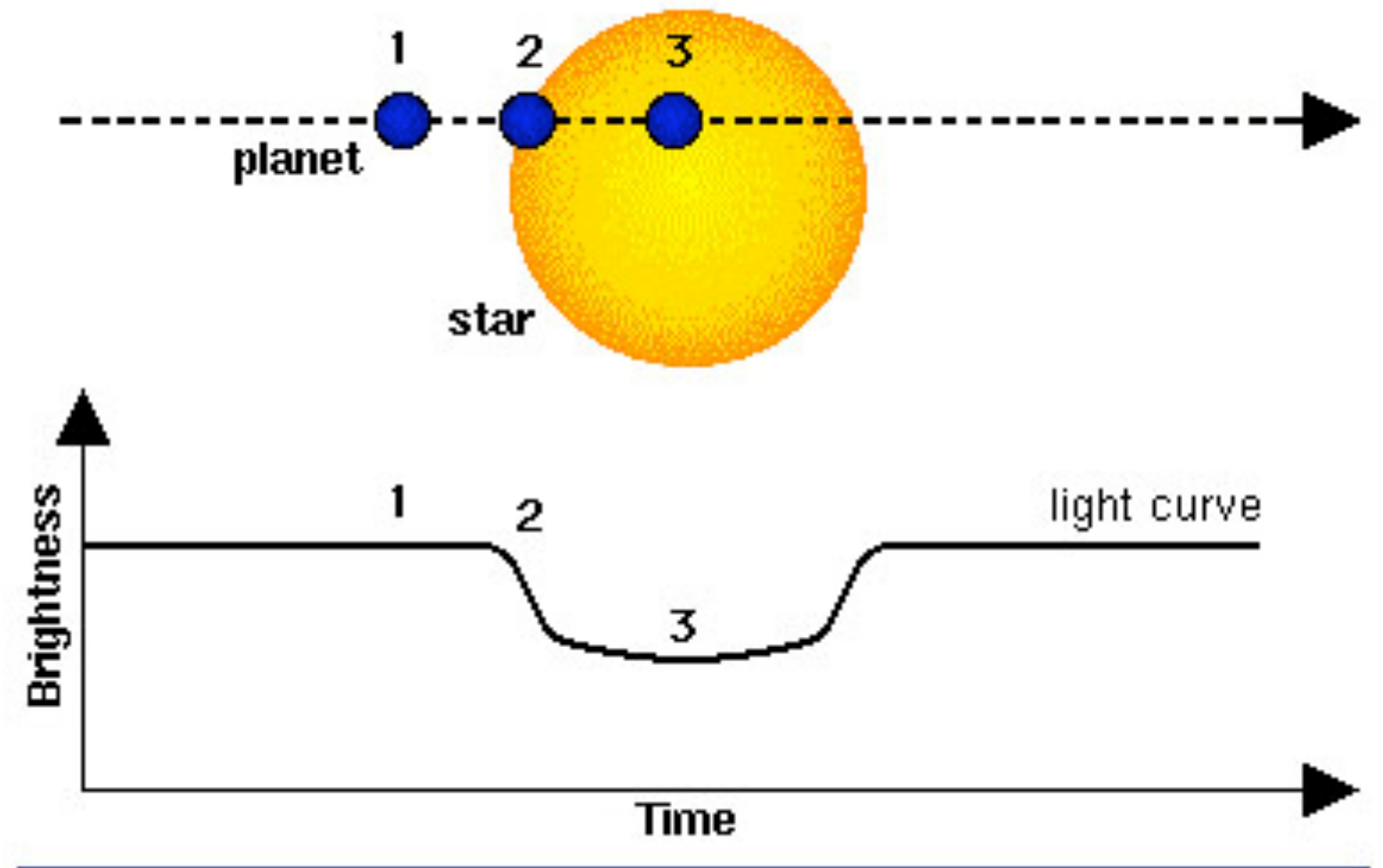
# What's an extrasolar planet?

- An extrasolar planet (or exoplanet) is a planet orbiting a star different from the Sun
- Its mass must be below the limiting mass for thermonuclear fusion of deuterium ( $\sim 13 M_{\text{jup}}$  for solar metallicity)
- Its minimum mass/size required is the same as that used in our Solar System
- 307 exoplanets known so far (01/08/2009)

# The Transit Method

- When star pass in front of its parent star, it causes a drop in brightness
- Transit depth:  
~1% for Jupiter-size  
~0.01% for Earth-size

$$\frac{\Delta F}{F} = \left[ \frac{R_p}{R_*} \right]^2$$





# Advantages/Disadvantages

- If more than one transit is visible, the orbital period is known
- Simultaneous monitoring of large number of sources
- Direct measurement of planet/star size ratio
- But transit probability very small and method sensitive to large planets

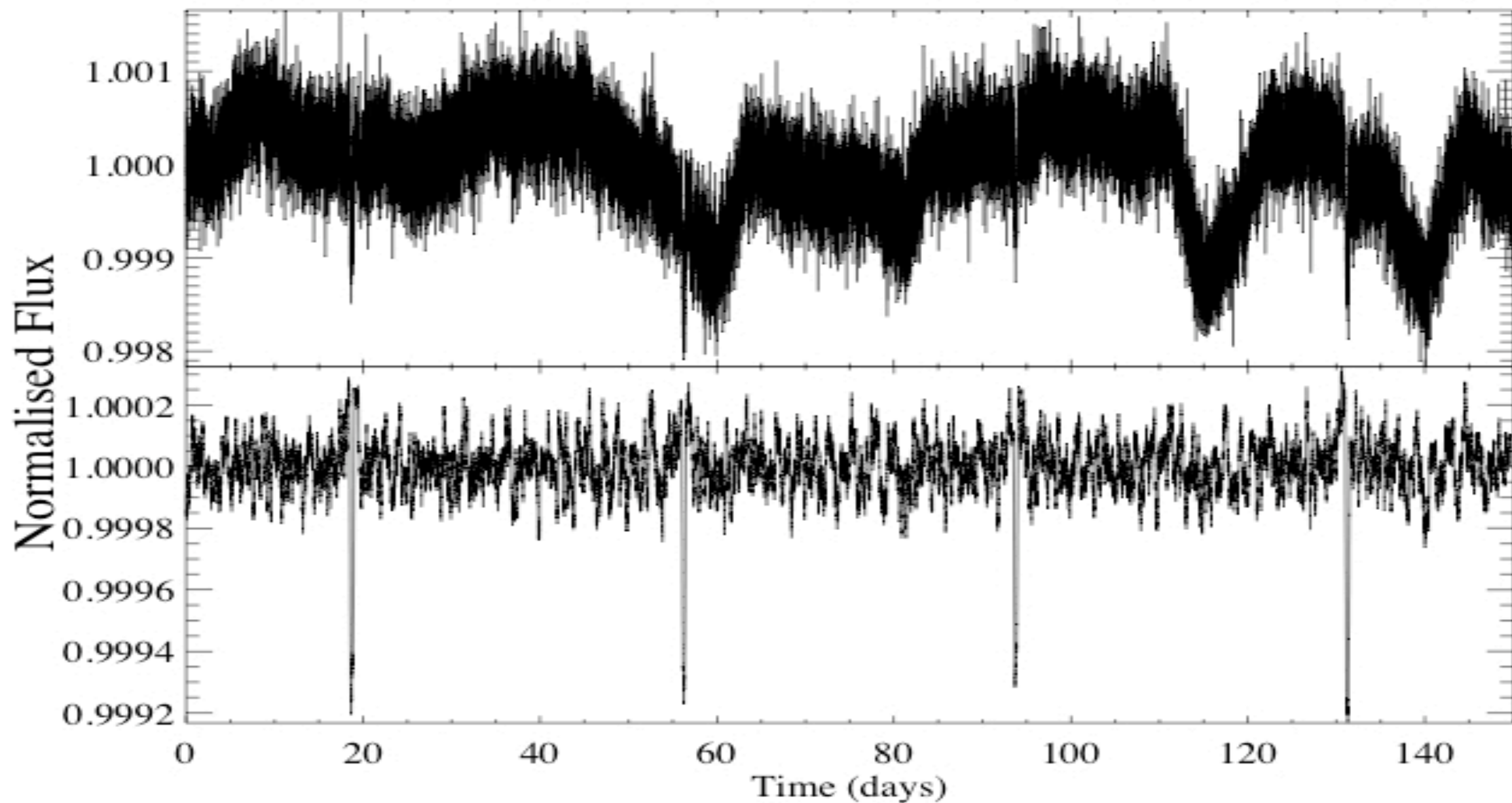
# CoRoT fields

- Year divided into two 6-months of observation toward galactic center and galactic anti-center.
- Each of these 6 months is divided between a short run (20 days) and long run (150 days)
- Observation runs already performed:

	IRa01	LRc01	SRc01	LRa01
run type	Initial run	Long run	Short run	Long run
direction	anti-center	center	center	anti-center
length (days)	60	150	20	150

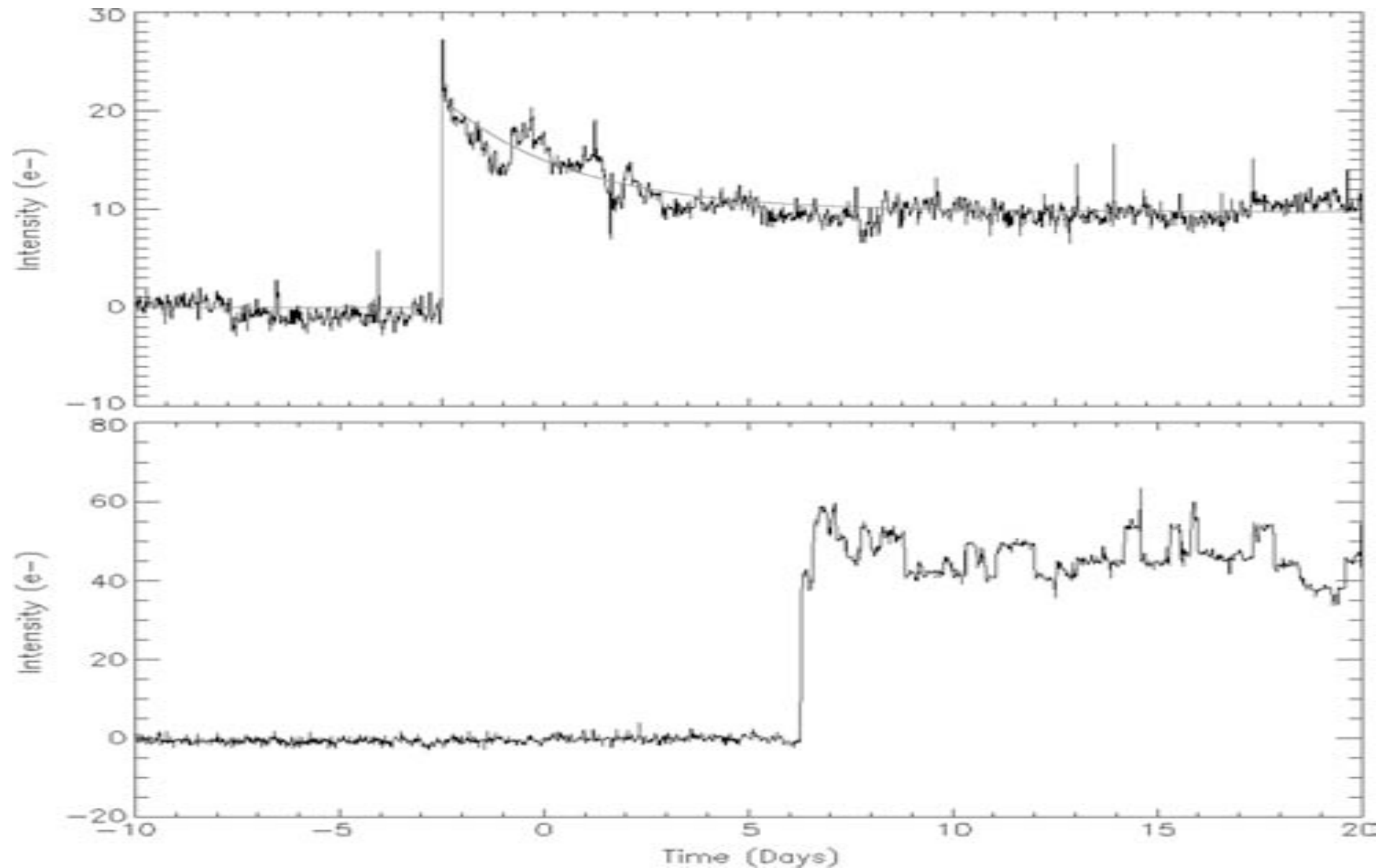
# What do we expect to see?

(stellar noise and photon noise only)



Simulated light curve using SOHO data (Carpano & Fridlund 2008)

# Extra-noise: “Hot-pixels”

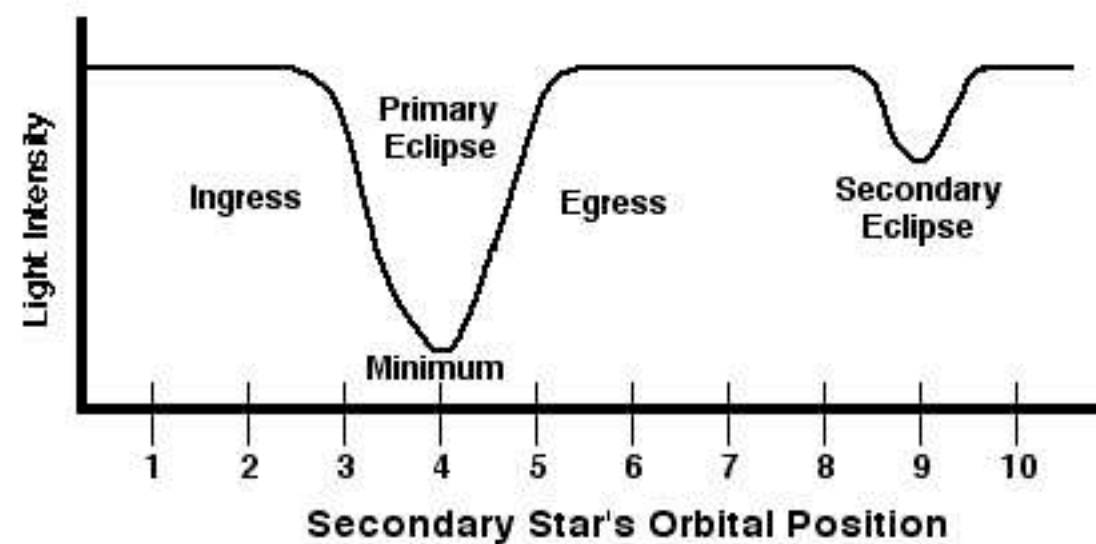
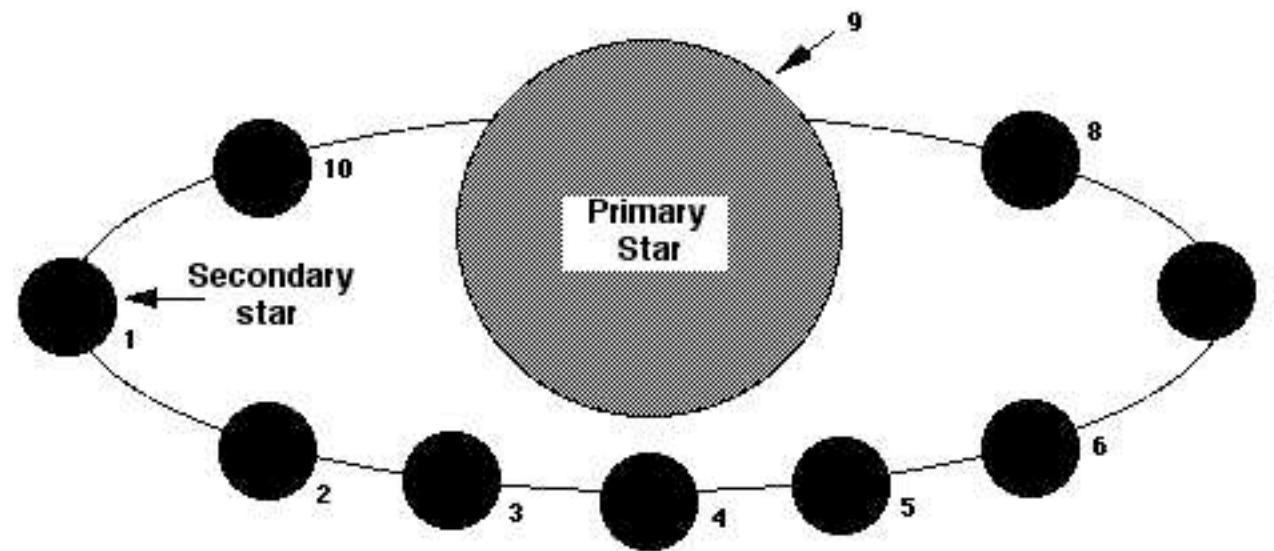


Generation of bright pixel after particle impact (Pinheiro da Silva et al. 2008)

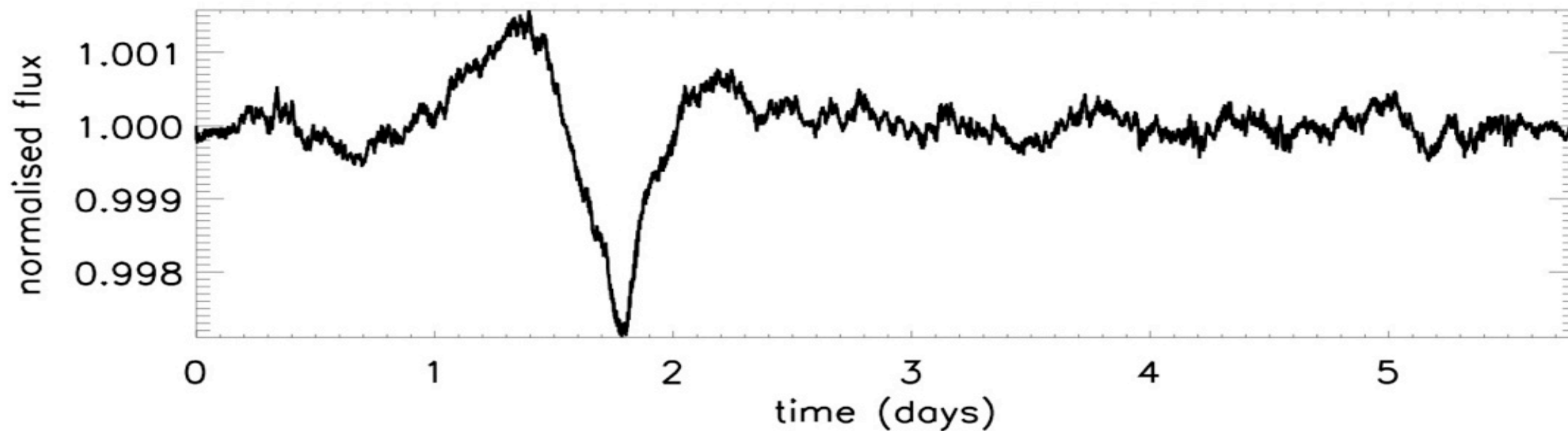
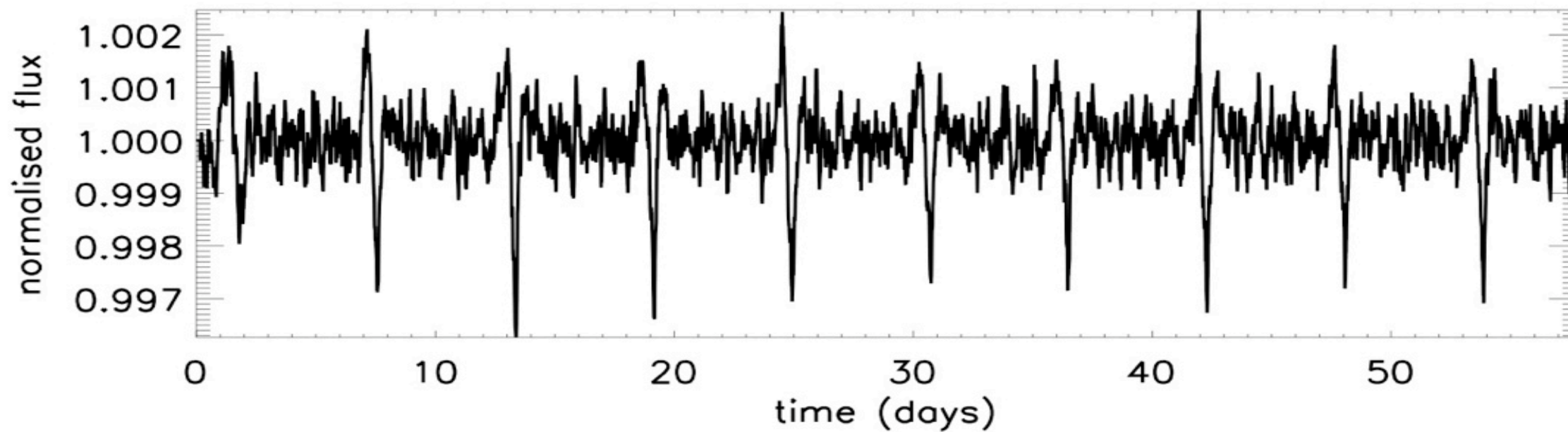


# Looks like a planet... but it's not!

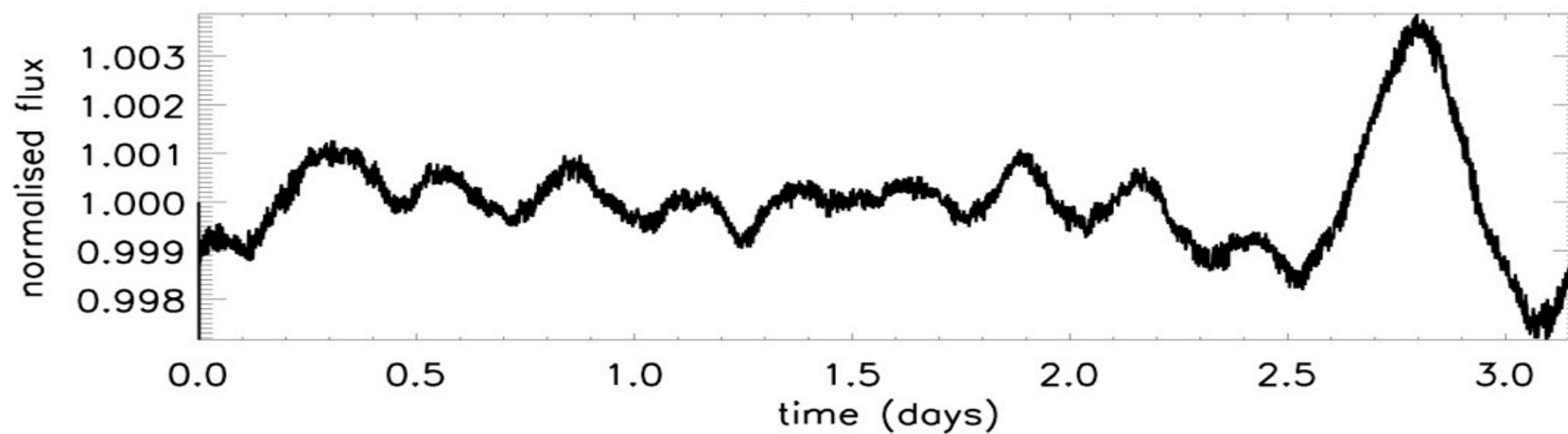
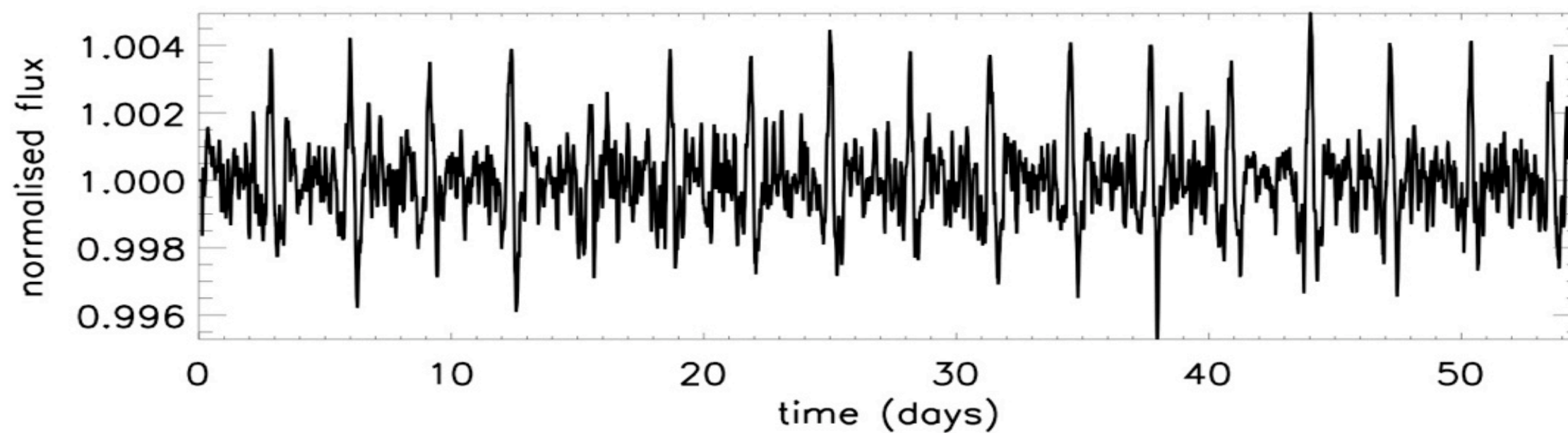
(eclipsing binaries at grazing angle)



# Not a planet... but interesting!



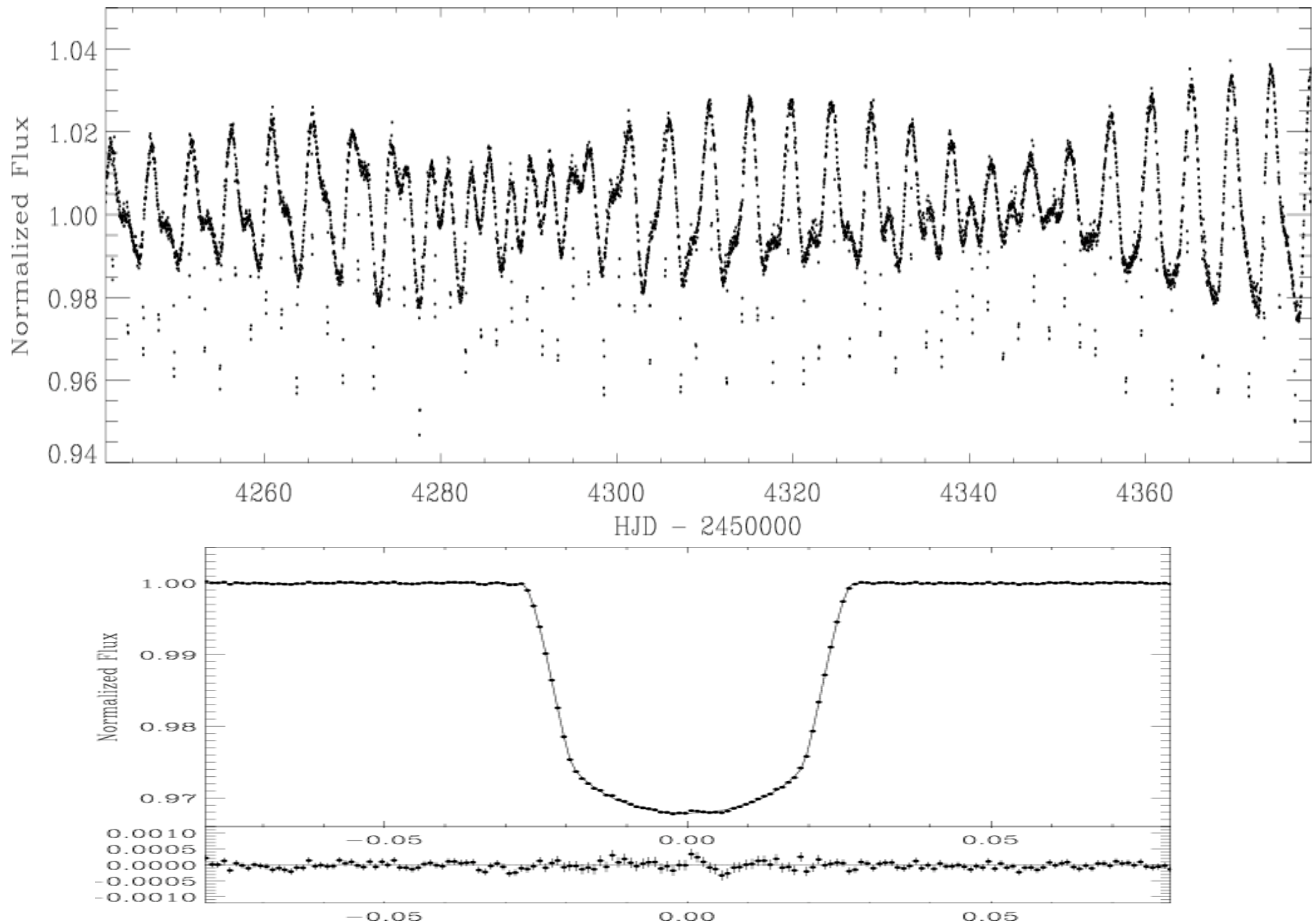
# Oscillations?



# Planets discovered with CoRoT

- CoRoT-Exo-1b (published):  $R=1.5R_{\text{jup}}$ ,  $M=1M_{\text{jup}}$ ,  $P=1.5\text{days}$
- CoRoT-Exo-2b (published):  $R=1.5R_{\text{jup}}$ ,  $M=3.3M_{\text{jup}}$ ,  $P=1.7\text{days}$
- CoRoT-Exo-3b (submitted) is a peculiar Brown dwarf:  $R=1R_{\text{jup}}$ ,  $M=21M_{\text{jup}}(!)$ ,  $P=4.3\text{days}$
- CoRoT-Exo-4b (accepted):  $R=1.2R_{\text{jup}}$ ,  $M=0.75M_{\text{jup}}$ ,  $P=9.2\text{days}$
- +CoRoT-Exo-5b (in preparation)

# Example: CoRoT-Exo-2b





# Conclusions

- ~30.000 light curves analysed so far
- 1% are possible planet candidates (from CoRoT data only)
- 1/2 of the candidates are/will be soon followed up (ground-based FU)
- Results: 5-7 confirmed so far (+few other good candidates)
- About ~15 planets are expected from theory

Detection of planets with CoRoT  
(especially small ones) is very challenging!