

**The role of cosmic-ray nuclei and
electrons in the LISA test mass
charging rate**

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LISA SUN DISTANCE RANGE: $149 \div 152 \cdot 10^6$ km
(0.9933 AU \div 1.0133 AU)

LISA LATITUDINAL RANGE: $0.7 \div 1.0$ degree

What does this mean for LISA?

SOLAR MINIMUM

A radial correction factor of the order of 10^{-3} %

A latitudinal correction factor of the order of 0.33%

Measurements near the Earth OK!

SOLAR MAXIMUM

No latitudinal correction is needed

COSMIC-RAY COMPOSITION

Protons : 90%

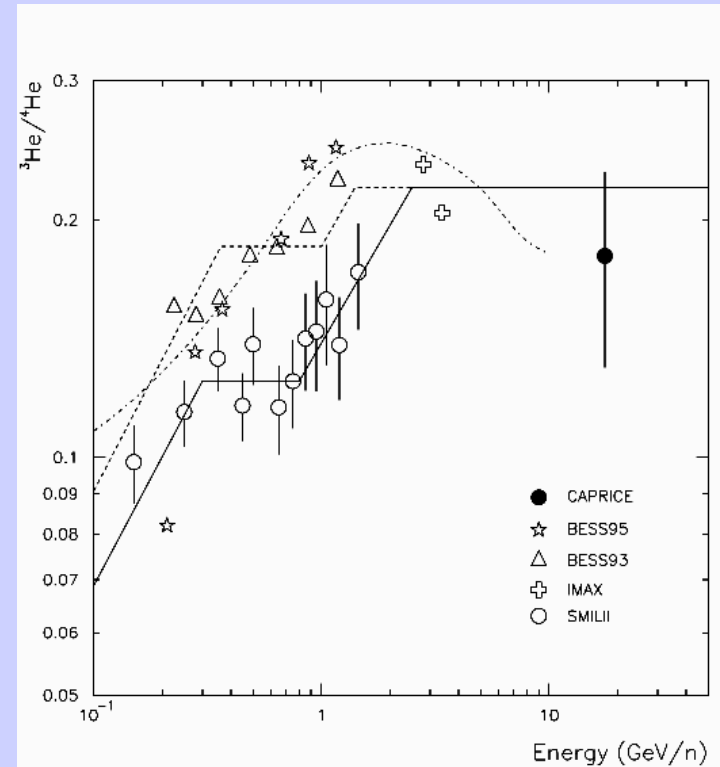
Helium : 8%

Heavy Nuclei: 1%

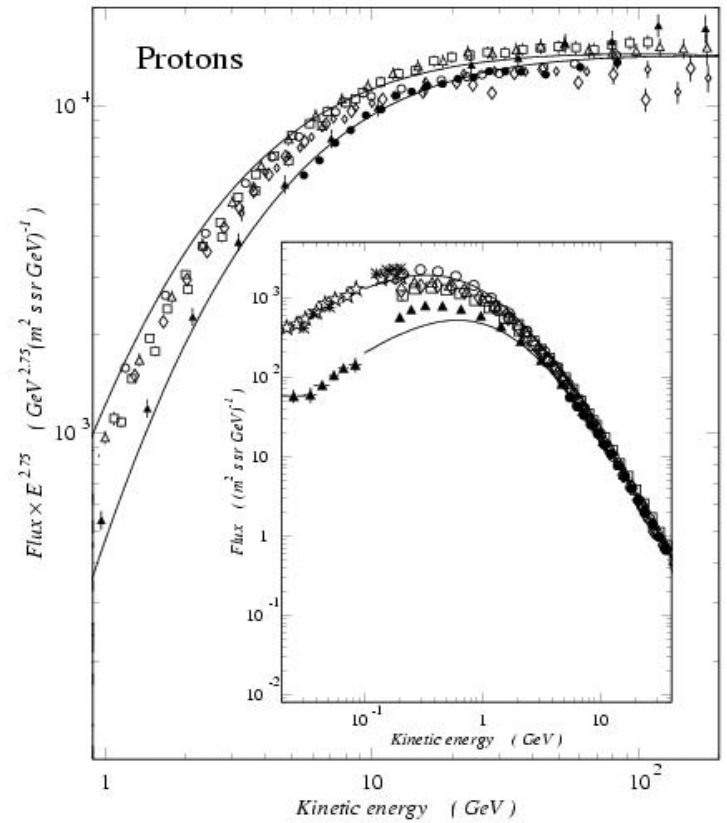
Carbon
Oxygen
Nitrogen
Magnesium
Silicon
Iron

Electrons: 1%

$^3\text{He}/^4\text{He}$ Ratio



Proton Differential Flux



Helium Differential Flux

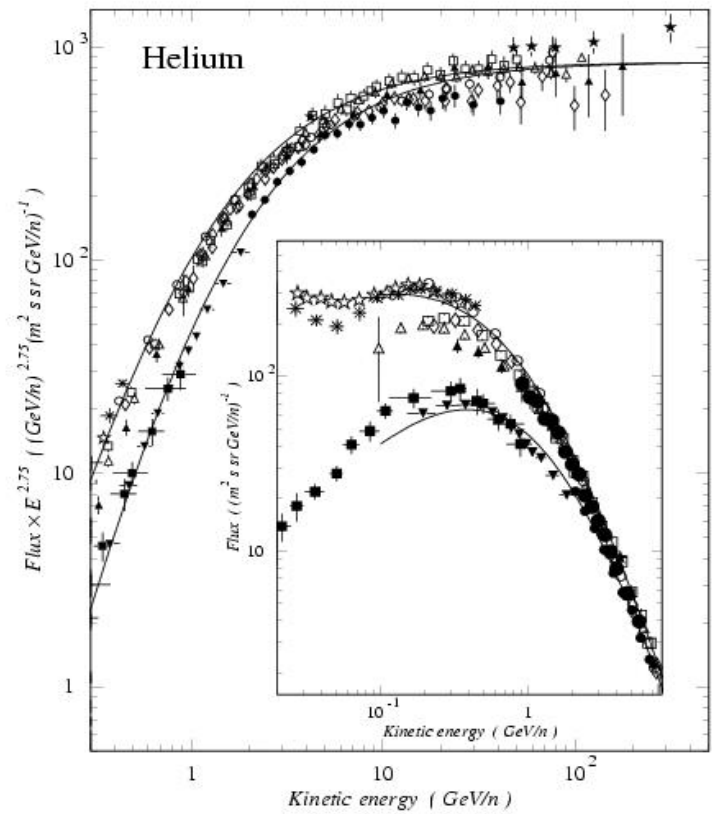
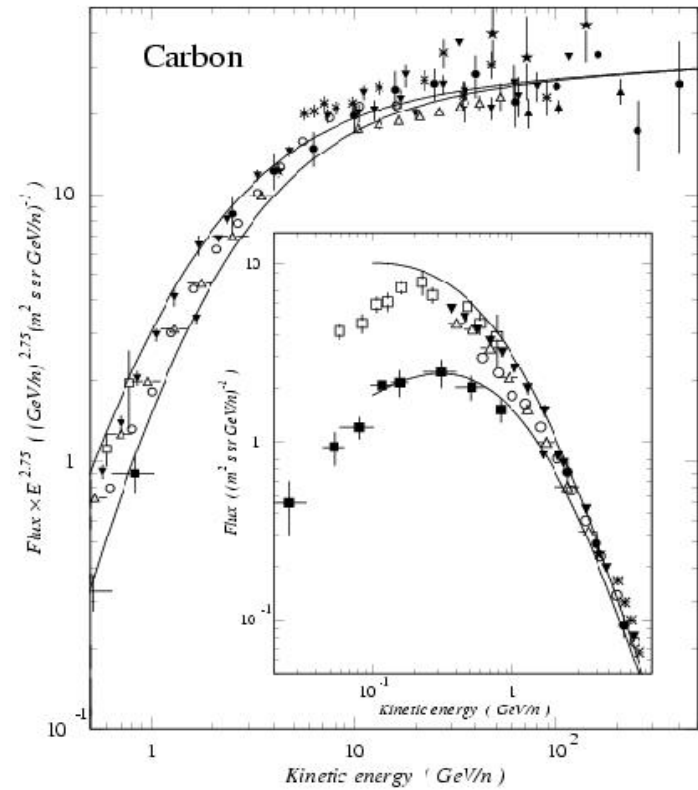
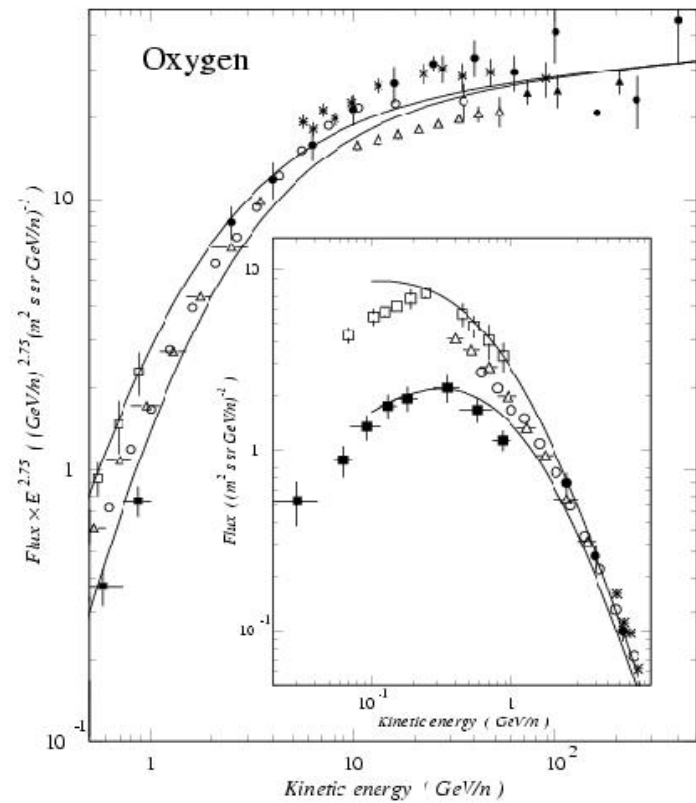


Figure from C. Grimani et al., Class. Quant. Grav., 21, S629, 2004

Carbon Differential Flux



Oxygen Differential Flux



Primary cosmic-ray interpolation function

$$F(E) = A(E + B)^{-\alpha} E^{\beta} \quad \text{Part}/(\text{m}^2 \text{ sr s GeV}/n)$$

Where E is GeV/n

Solar minimum

| | A | B | α | β |
|----|-------|------|----------|---------|
| p | 18000 | 1.09 | 3.66 | 0.87 |
| He | 850 | 0.99 | 3.10 | 0.35 |
| C | 23 | 0.95 | 3.00 | 0.29 |
| O | 21 | 0.95 | 3.00 | 0.32 |

Solar maximum

| | A | B | α | β |
|----|-------|------|----------|---------|
| p | 18000 | 1.55 | 3.90 | 1.11 |
| He | 850 | 1.25 | 3.60 | 0.85 |
| C | 23 | 1.22 | 3.40 | 0.69 |
| O | 21 | 1.22 | 3.40 | 0.72 |

Most Abundant Heavy Nuclei

| Element | Z | Abundance (% O) |
|----------------|----------|----------------------------|
| N | 7 | 28% |
| Ne | 10 | 16% |
| Mg | 12 | 20% |
| Si | 14 | 14% |
| Fe | 26 | 9% |

Simpson, J. A., *Ann. Rev. Nucl. Phys. Sc.*, 33, 323, 1983

Stephens, S. A. and Streitmatter, R. E., *Adv. Sp. Res.*, 27, 749, 2001

Heavy nucleus interpolation function

$$F(E) = A(E + B)^{-\alpha} E^{\beta} \text{ Part}/(\text{m}^2 \text{ sr s GeV}/n)$$

| Element | A | B | α | β |
|---------|-----|------|----------|---------|
| N | 5.9 | 0.95 | 3.00 | 0.32 |
| Ne | 3.4 | 0.95 | 3.00 | 0.32 |
| Mg | 4.2 | 0.95 | 3.00 | 0.32 |
| Si | 2.9 | 0.95 | 3.00 | 0.32 |
| Fe | 1.9 | 0.95 | 3.00 | 0.32 |

Electron Flux Measurements

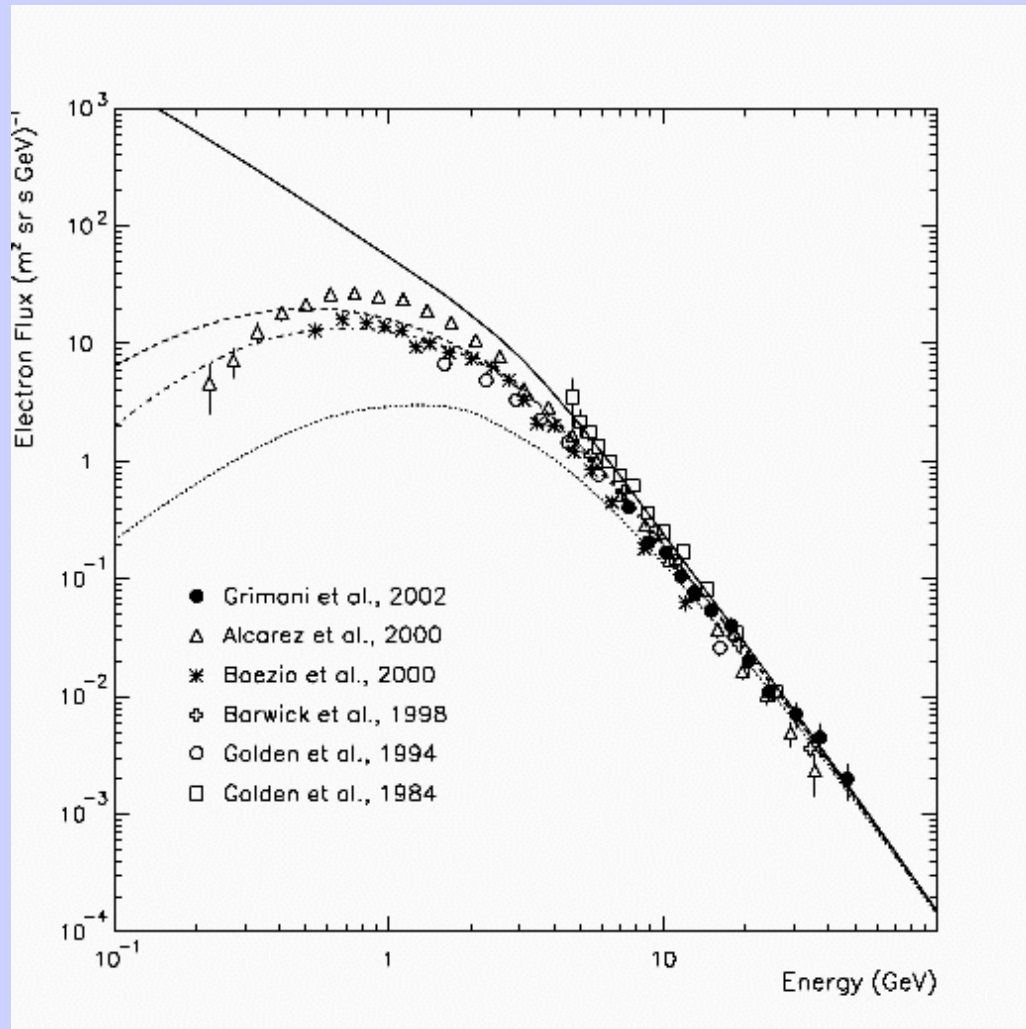
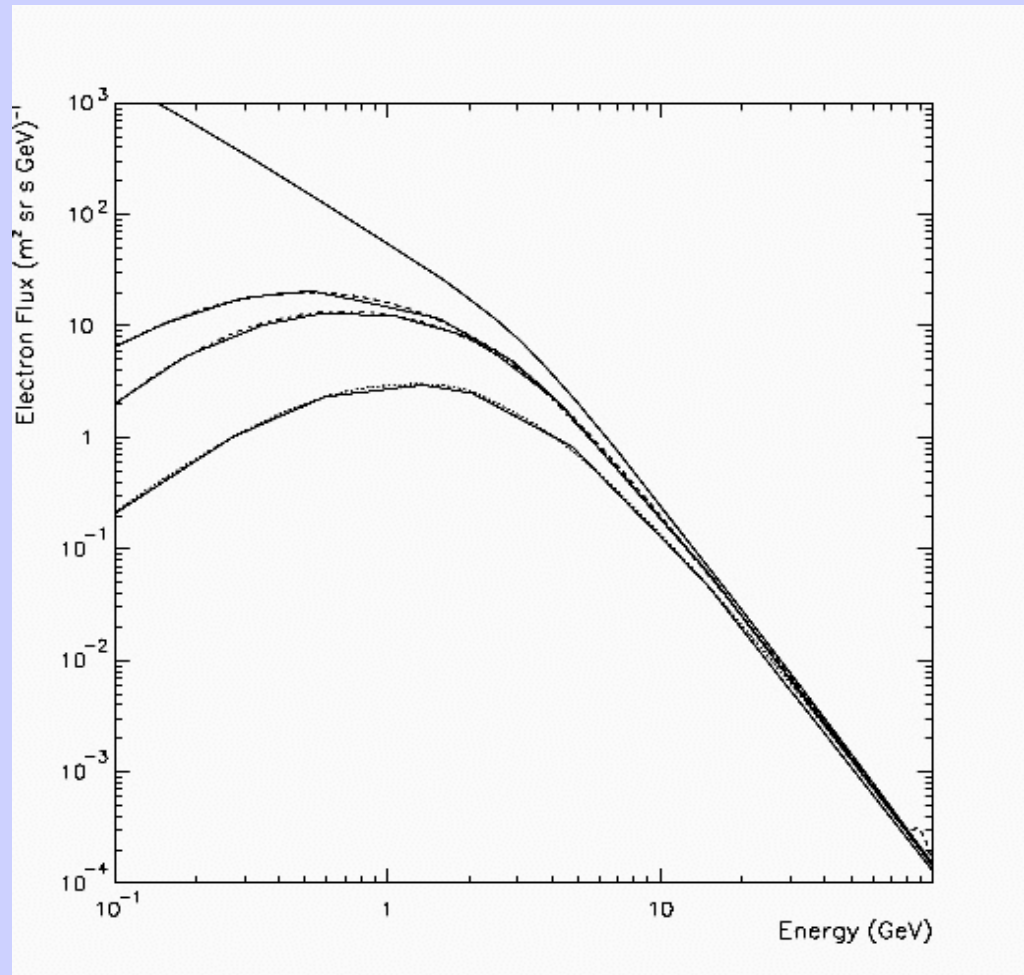


Figure from Grimani C. et al., A&A, 2002, 392, 287

Electron Flux Straight Line Interpolation



LISA SPACECRAFT SIMPLIFIED GEOMETRY (FLUKA MONTE CARLO PROGRAM)

- Test mass side: **4.6 cm**
- Test mass material: **gold**
- Test mass depth : **89 g/cm²**

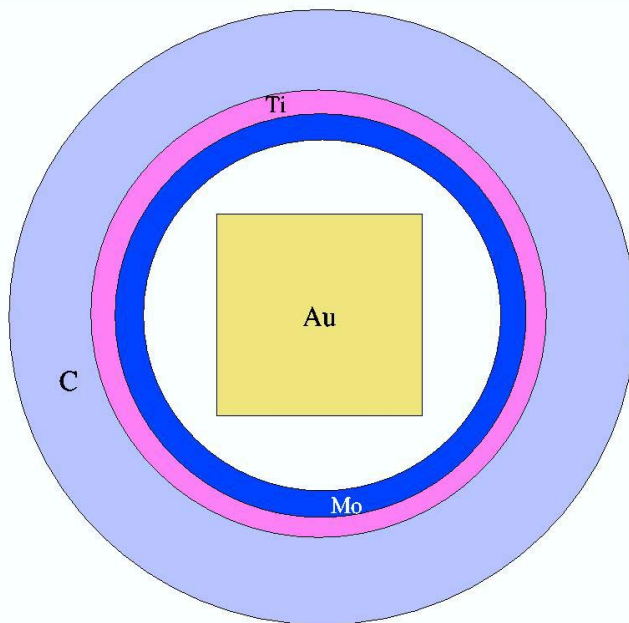


Figure from Vocca et al., Class. Quant. Grav. ,201, S665, 2004

| Material | Density (g cm ⁻³) | Thickness (cm) | Grammage (g cm ⁻²) |
|----------|----------------------------------|-------------------|-----------------------------------|
| C | 2.10 | 2.0 | 4.2 |
| Ti | 4.54 | 0.5 | 2.3 |
| Mo | 10.28 | 0.6 | 6.2 |
| Au | 19.32 | 4.6 | 88.9 |

FLUKA Monte Carlo Characteristics

Particle transport

| | Secondary particles | Primary particles |
|-------------------------------|------------------------------------|-------------------------------------|
| Charged hadrons | 1 keV ÷ 20 TeV | 100 keV ÷ 20 TeV |
| Neutrons | Thermal ÷ 20 TeV | Thermal ÷ 20 TeV |
| Muons | 1 keV ÷ 1 PeV | 100 keV ÷ 1 PeV |
| Electrons (low-Z) (high-Z) | 1 keV ÷ 1 PeV 150 keV ÷ 100 TeV | 70 keV ÷ 1 PeV 150 keV ÷ 100 TeV |
| Photons | 1 keV ÷ 1 PeV | 7 keV ÷ 1 PeV |

Net and Effective Charge Rate Definition

Test mass current: $I = e \lambda_{\text{NET}}$

Force noise proportional to: $S_Q^{1/2} = S_I^{1/2} / \omega$

Current shot noise: $S_I^{1/2} = \sqrt{2 e^2 \lambda_{\text{EFF}}}$

Where:

$$\lambda_{\text{NET}} = \sum_{j=-\infty}^{+\infty} j \lambda_j \quad \text{and} \quad \lambda_{\text{EFF}} = \sum_{j=-\infty}^{+\infty} j^2 \lambda_j$$

CHARGE RATE ON THE TEST MASSES - HEAVY NUCLEI

Solar minimum

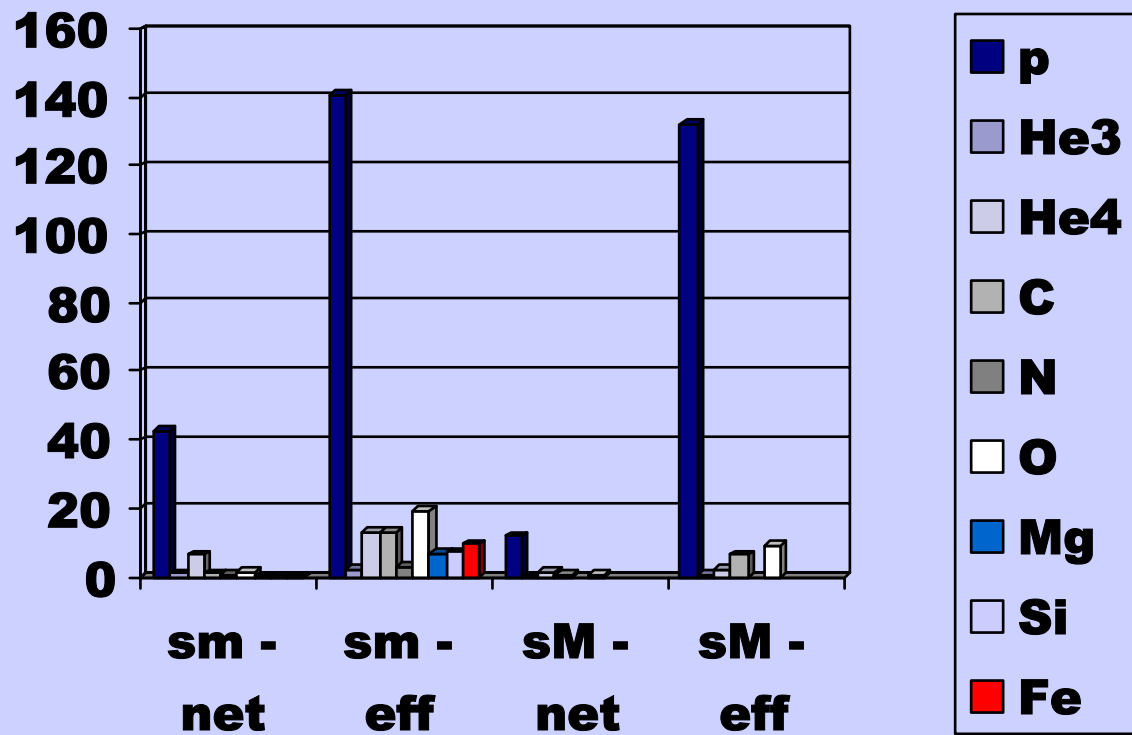
| Element | Z | Absolute charge rate (e ⁺ /s) | Effective charge rate (e/s) |
|-----------------|----|--|-----------------------------|
| ³ He | 2 | 0.97 | 1.97 |
| ⁴ He | 2 | 6.46 | 12.95 |
| C | 6 | 1.03 | 13.10 |
| N | 7 | 0.45 | 3.00 |
| O | 8 | 1.34 | 19.40 |
| Mg | 12 | 0.26 | 6.81 |
| Si | 14 | 0.20 | 7.12 |
| Fe | 26 | 0.15 | 9.64 |

CHARGE RATE ON THE TEST MASSES - HEAVY NUCLEI

Solar maximum

| Element | Z | Absolute charge rate (e ⁺ /s) | Effective charge rate (e/s) |
|-----------------|----|--|-----------------------------|
| ³ He | 2 | 0.17 | 0.37 |
| ⁴ He | 2 | 1.13 | 2.17 |
| C | 6 | 0.37 | 6.40 |
| N | 7 | <<0.1 | |
| O | 8 | 0.37 | 9.00 |
| Mg | 12 | <<0.1 | |
| Si | 14 | <<0.1 | |
| Fe | 26 | <<0.1 | |

| | Absolute charge rate (e ⁻ /s) | Effective charge rate (e/s) |
|-----------------------------|--|-----------------------------------|
| Electron (solar minimum) | 0.59 | 9.01 |
| Electron (solar maximum) | 0.34 (+) | 5.71 |



CONCLUSIONS

- * Heavy nuclei release about 25% of the charge generated by protons in the LISA test masses at solar minimum
- * Heavy nuclei release about 15% of the charge generated by protons in the LISA test masses at solar maximum
- * Electrons release about 1% of the charge generated by protons at both solar minimum and maximum