



LISA Symposium

An Acquisition Control for LISA

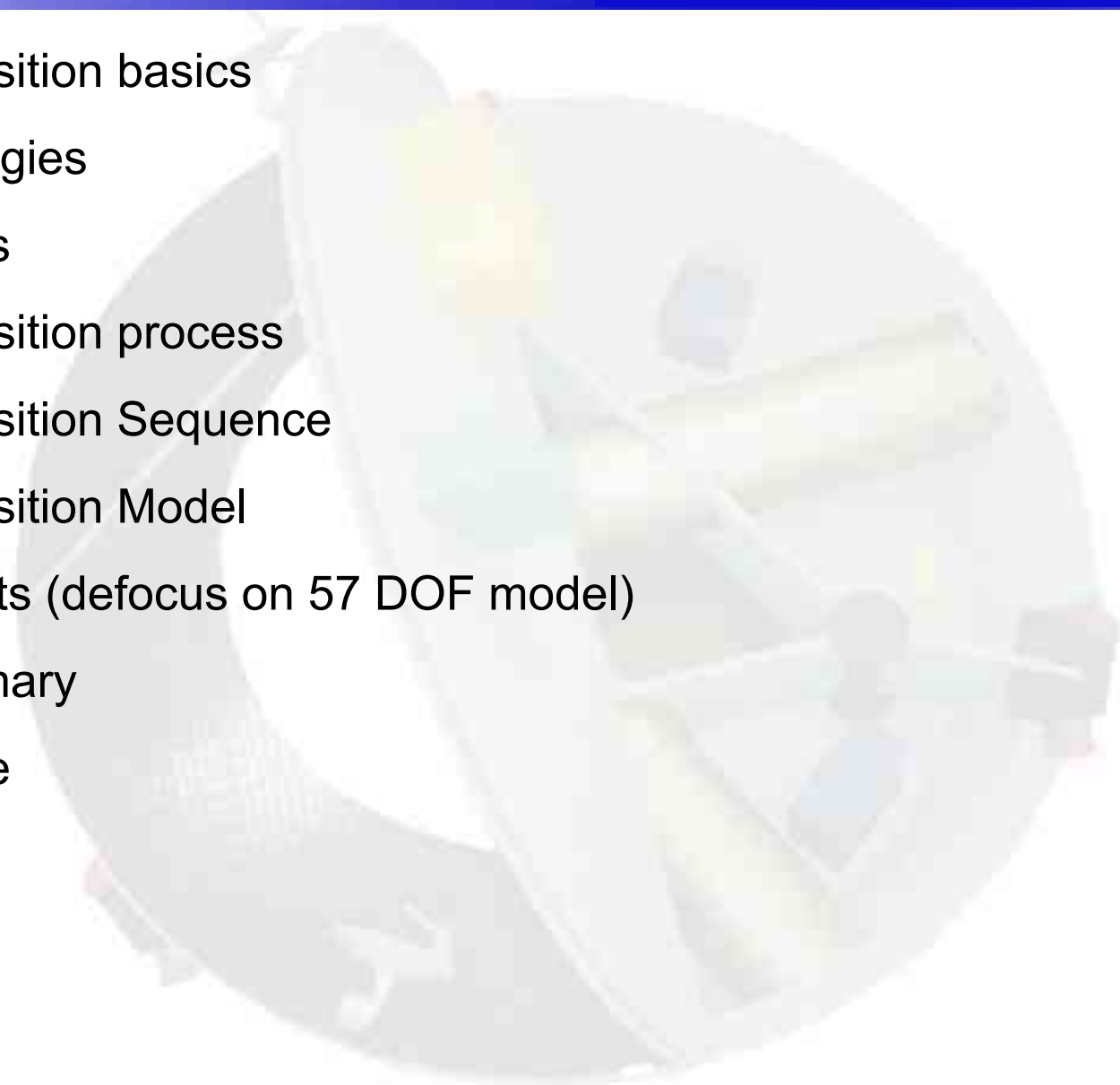
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- 🌀 Acquisition basics
- 🌀 Strategies
- 🌀 Issues
- 🌀 Acquisition process
- 🌀 Acquisition Sequence
- 🌀 Acquisition Model
- 🌀 Results (defocus on 57 DOF model)
- 🌀 Summary
- 🌀 Future



Constellation acquisition

- Not to be confused with “acquisition” as in buying stuff
- Defined as the process of bringing the three LISA spacecraft from “zero” to all six links phased locked and optimized (ready for science)
- Re-acquisition is a simpler (and faster) process with much smaller alignment and bias uncertainties

Process

- Spacecraft star tracker to telescope alignment calibrated with natural stars
- “Playlist” and ephemeris uplinked to all three spacecraft
- Spacecraft attitude (by star tracker) and point ahead prescribed
- For each link....autonomous process, but monitored by ground
 - Pointing acquisition (three strategies: defocus, scan, super-CCD)
 - Phase acquisition (scan frequency and lock to incoming)
- **Tune up constellation**
 - Verify inter-SC clock, comm. and ranging
 - Phase: Gain tuning for laser and offset/arm-lock loops
 - Pointing: Brightness and “flat-spot” search and optimization
 - Engineering Mode checkout, calibrations, and trial data runs

Defocus

- Spoil beam to cover uncertainty cone
 - 4 μ rad beam spoiled 10 x (to 40 μ rad), 100x weaker= \sim 0.7 pW
- Receiver fixed stares until integration time (SNR) good enough
 - This sets time required for “gyro-mode” stability

Scan

- Scan over uncertainty cone (continuous or step/stare)
- Receiver fixed stares until integration time (SNR) good enough
 - This sets scan rate and time required for “gyro-mode stability”

Thru-telescope Star Tracker

- Natural star fixes with visible CCD through LISA telescopes
 - $< 1 \mu$ rad NEA for 0.05 deg FOV down to magnitude (m_v) of 16, for 1 sec integration
- Issue of “imaging quality” FOV with telescope prescription

Defocus

- Fiber positioner mechanism complexity is driving issue
- Requires pointing stability over the phase-lock-up time
- Can be mitigated by both ends defocus and co-re-focusing

Scan

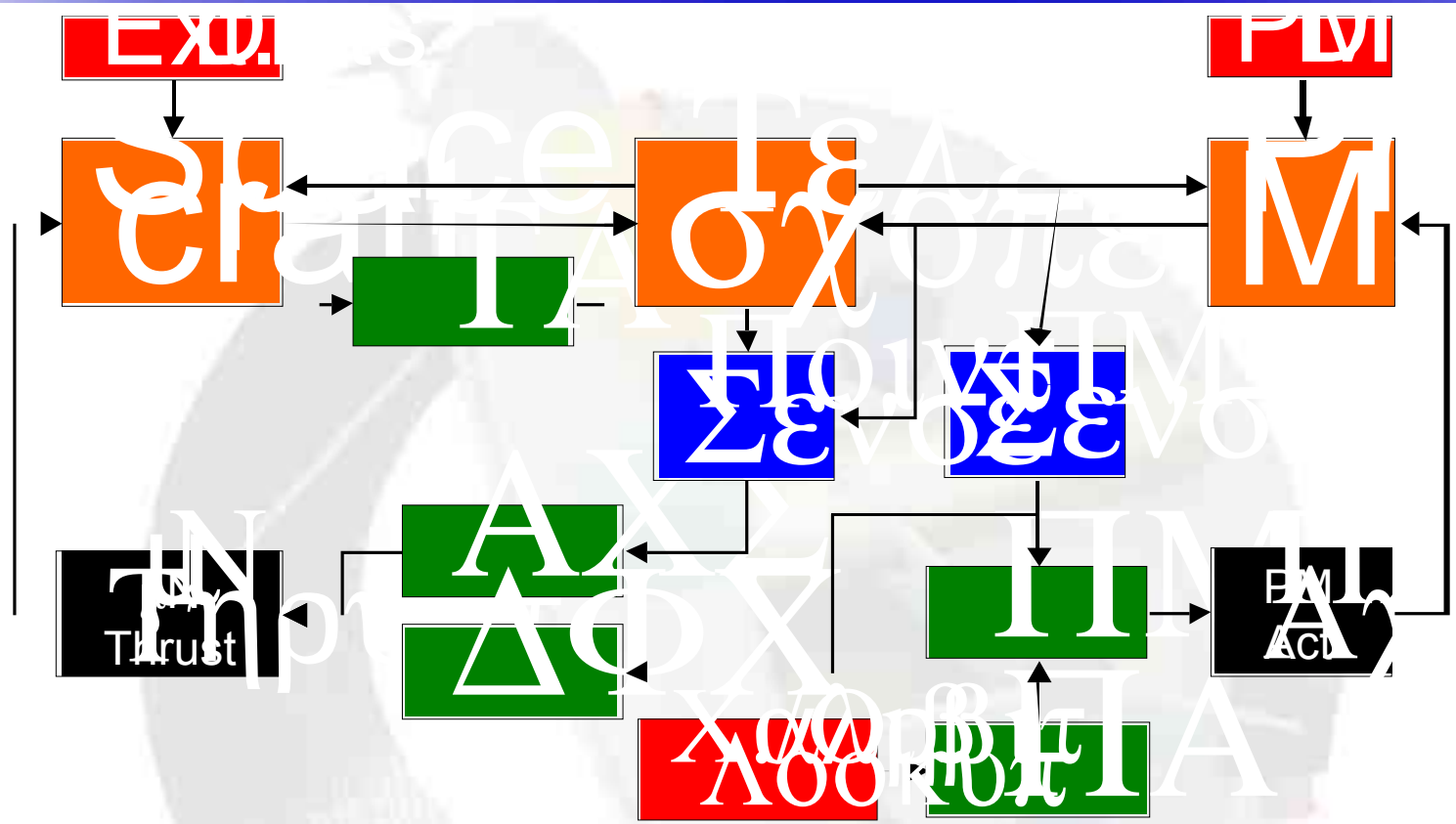
- Pointing stability over scan time is driving issue
- Can be mitigated with “gyro-mode” operation of GRSs
- Scan can always be done, even if another strategy is chosen as prime

Thru-telescope star tracker

- FOV of telescope is driving issue
- Requires good alignment knowledge on bench [or searching]
- Eliminates requirement to have high performance SC star trackers

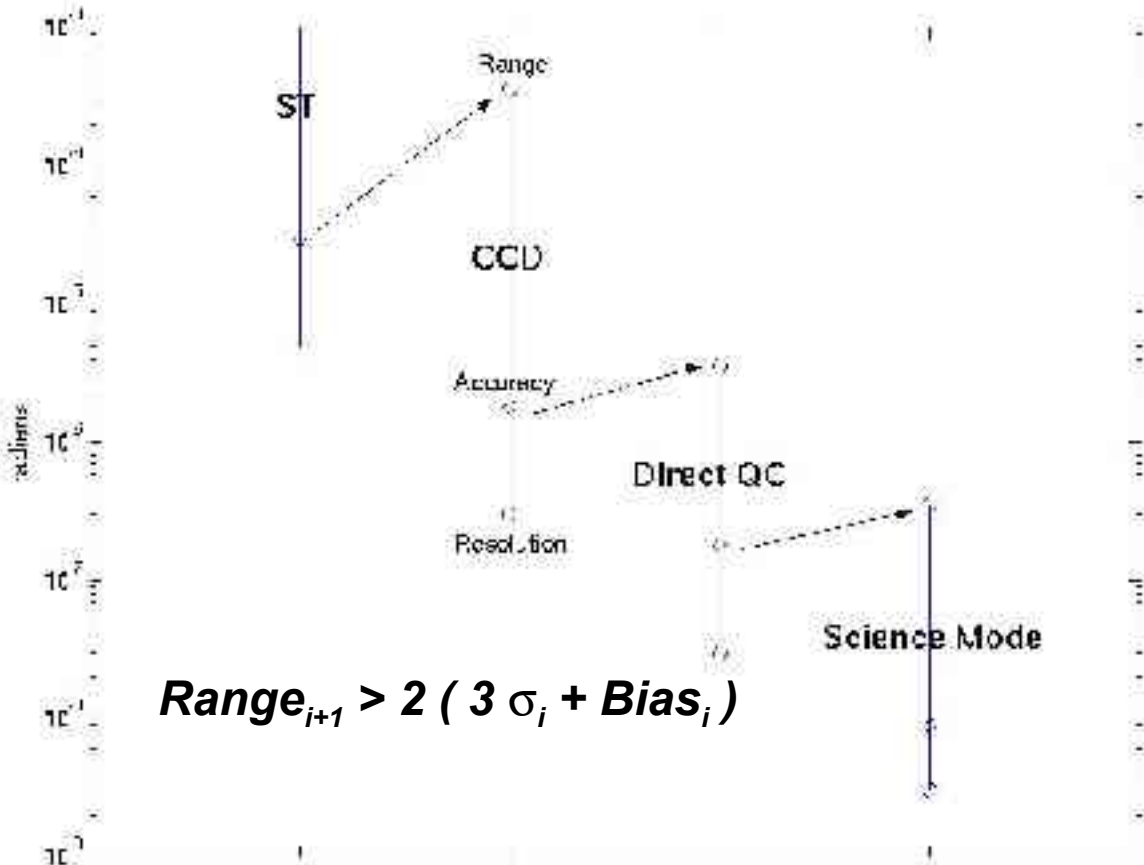


Disturbance Reduction System

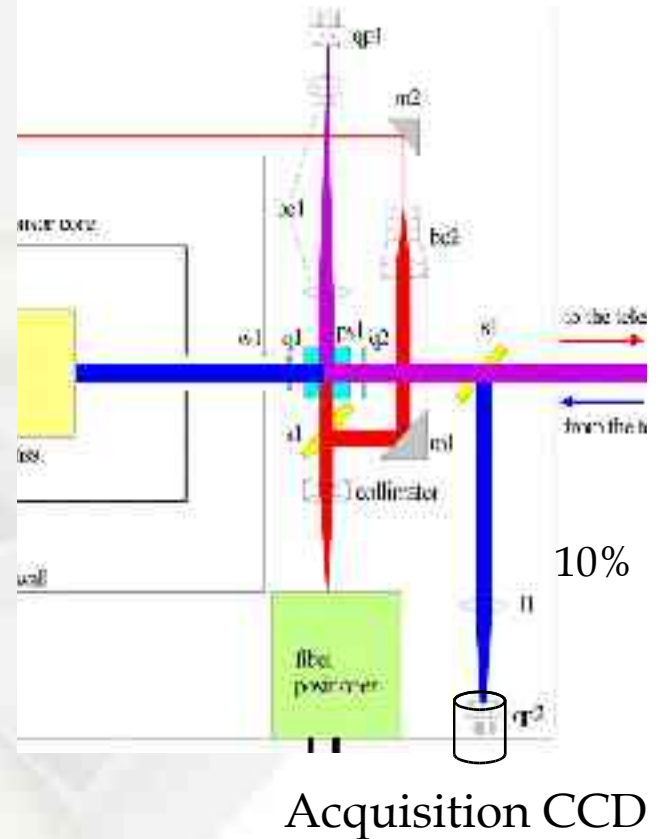


DFC=Drag Free Control System
 TA=Telescope Articulation
 ACS=Attitude Control System
 PM=Proof Mass Suspension
 PA=Point Ahead & Acquisition

Sensor	Resolution (1 Hz effective bandwidth)	Range	Alignment Bias to Next Sensor
Star Tracker	5 microrad	Full sky	20 microrad
CCD	0.3 microrad	360 microrad	0.9 microrad
Direct QC	30 nanorad	3.6 microrad	90 nanorad
Science Mode	3 nanorad	360 nanorad	N/A



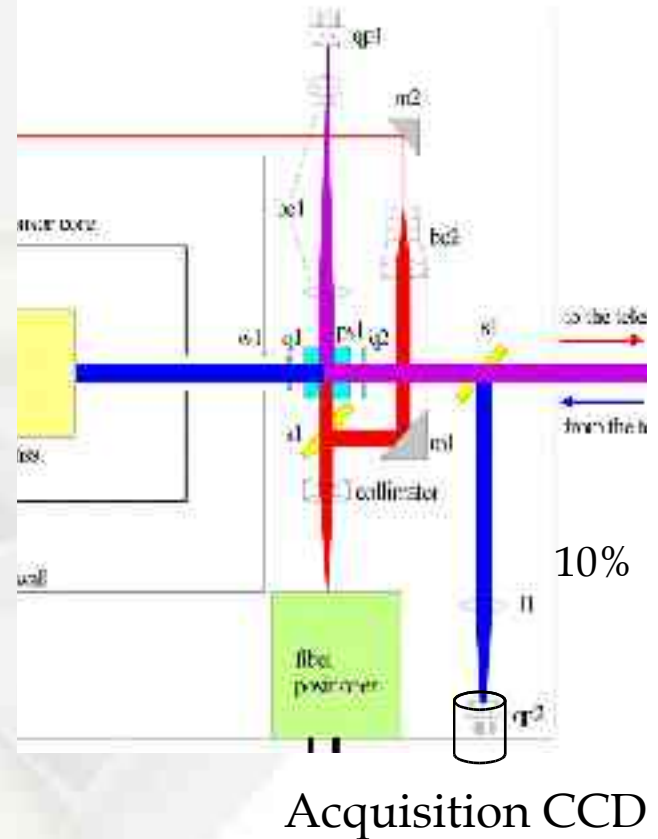
- Initiated from ground command and done autonomously
 - Spacecraft point at each other based on ephemeris information and star tracker measurements (Tracker mode)
 - The gravitational sensors are placed in accelerometer mode
 - Both spacecraft switch to the gyro mode: A Kalman filter that blends attitude error info from the star trackers with the rotational acceleration measurements from the gravitational sensors (Gyro mode)
 - The transmitting S/C spoils its outgoing beam
 - Looks for the incoming beam (from other S/C) in its acquisition CCD (CCD Mode): local laser is off
 - Centers on and switches to the Quad Cell detector (Direct QC Mode)



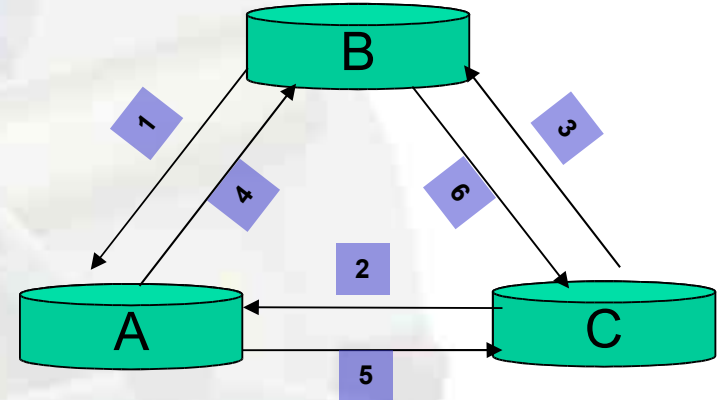
- The CCD mode may be bypassed if the performance of the Gyro mode allows
- Centers on the Quad cell and turns the local laser on for Heterodyned waverfront sensing
- The S/C is put pack in the Gyro mode during the 300 sec it may take for locking the beams

 Biases are calibrated out at each step making re-acquisition simpler

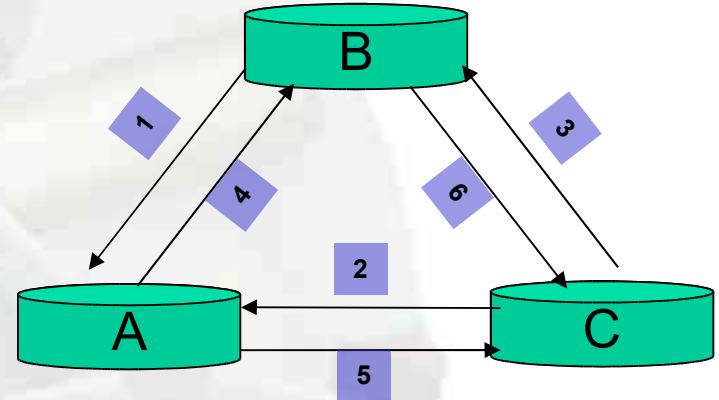
- Star tracker biases can be calibrated prior to the acquisition sequence
- Alternatively, a bias estimator may be implemented
- The estimator requires the adjacent arm to be in one of enhanced modes



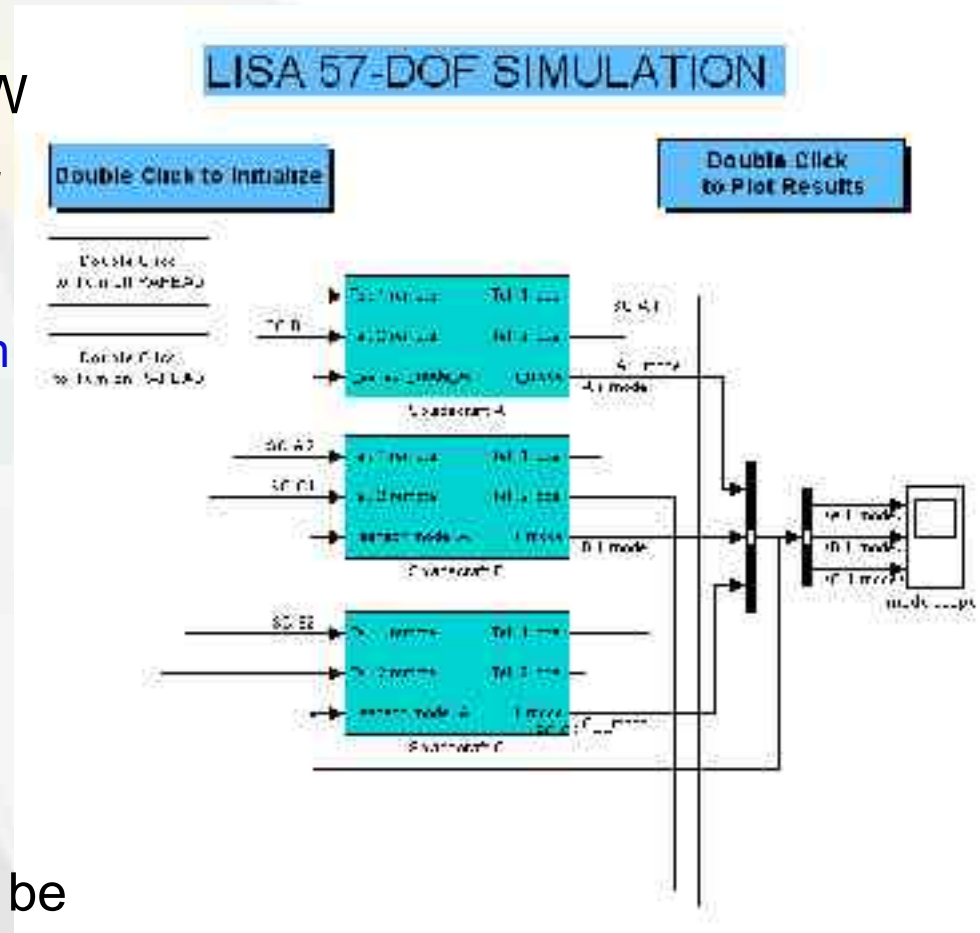
- Links 1 thru 3 are established first using the described process
- When establishing link 4 the local laser in S/C B must be initially turned off: **link 1 will be temporarily lost**
- Enhanced attitude knowledge is possible since link 2 is still active
- Current strategy is to use the sensor data from link 2 along with the attitude error estimate from the Kalman filter (Enhanced Gyro Mode)
- An alternative approach is to leave the attitude loop in an open-loop mode with thrusters firing at the last mean torque levels.
- Links 5 and 6 follow the same approach.



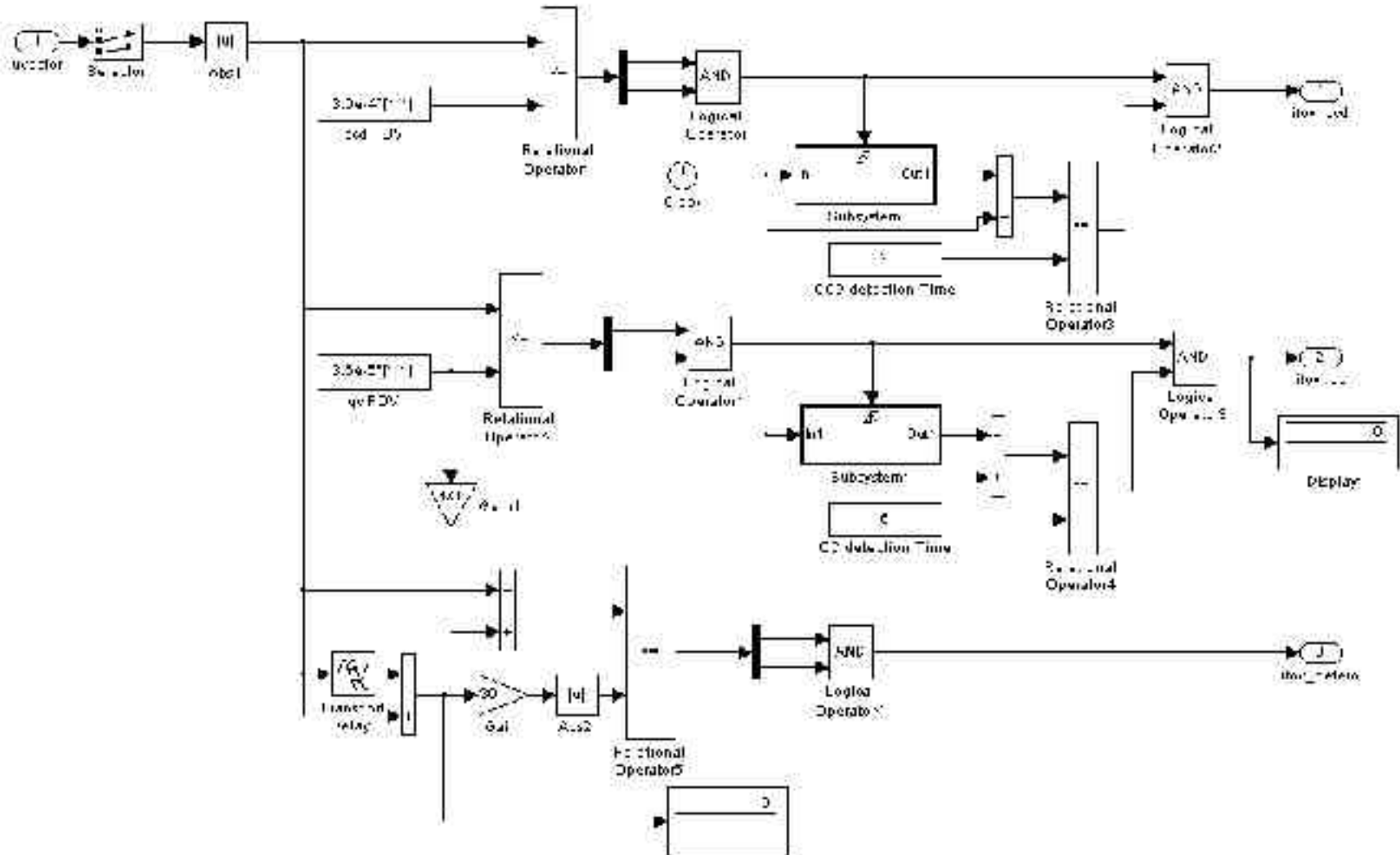
- Three cases considered for attitude determination
- No available fine attitude sensor
Measurements:
 - Star tracker + GRS accelerometer data (Kalman Filter)
 - Open-loop telescope articulation
- One link of fine sensor data available:
 - Fine sensor data used with the Kalman filter providing attitude error about the LOS
 - Open-loop telescope articulation
- Two links of fine sensor data available:
 - Fine sensor data is exclusively used to compute attitude and articulation errors



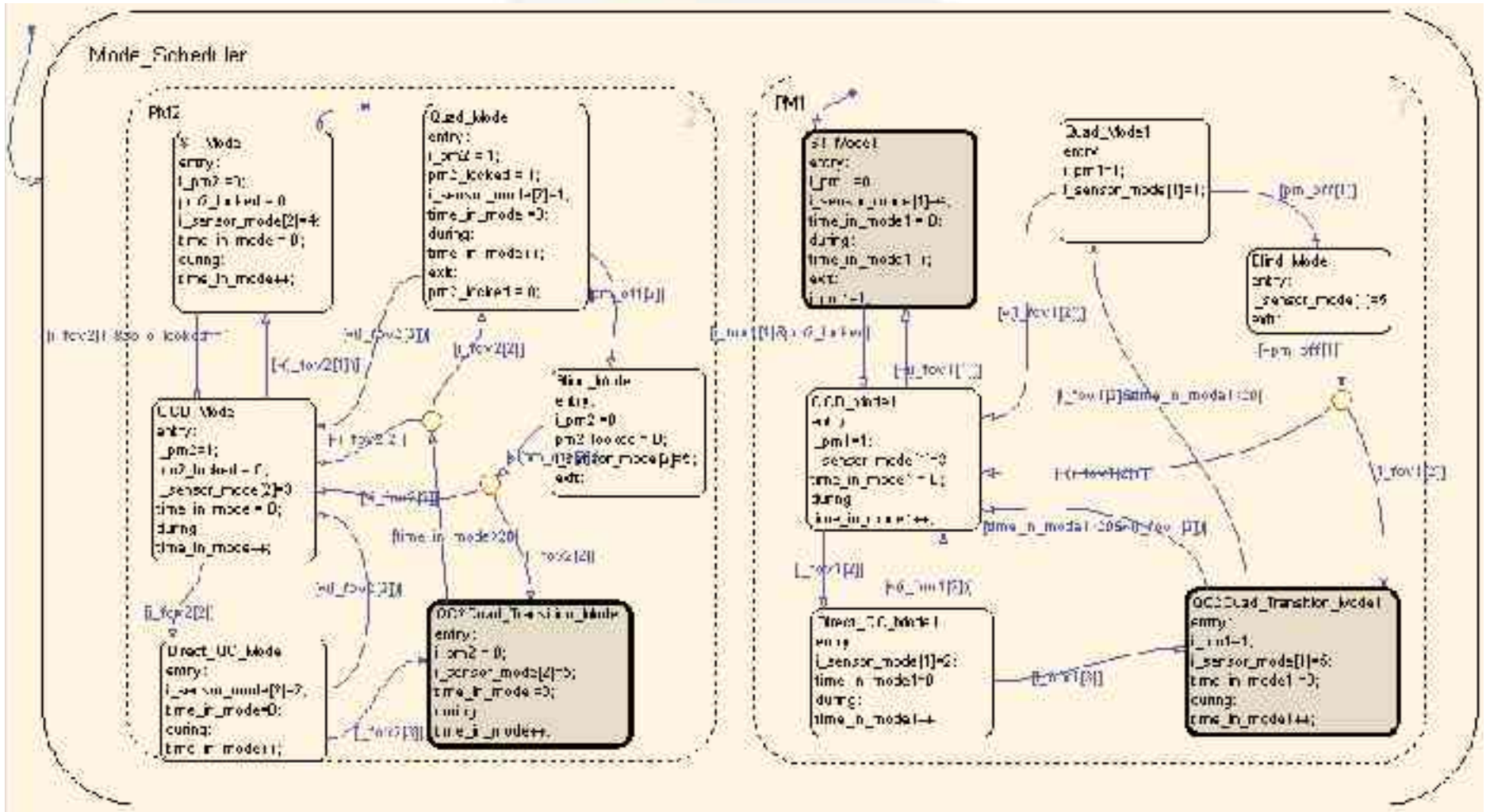
- Supports LISA acquisition control analysis
- Uses SIMULINK and STATEFLOW
- Include full rigid-body dynamics of the LISA formation
 - Same DRS control system in each S/C
 - Natural orbits from design optimization
 - Point-ahead compensation
- Currently supports the defocus strategy
- Scanning and other strategies will be supported in the near future



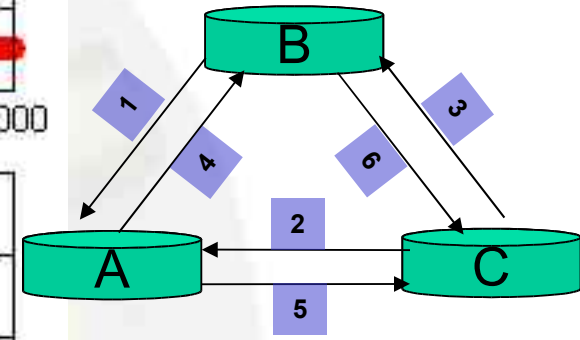
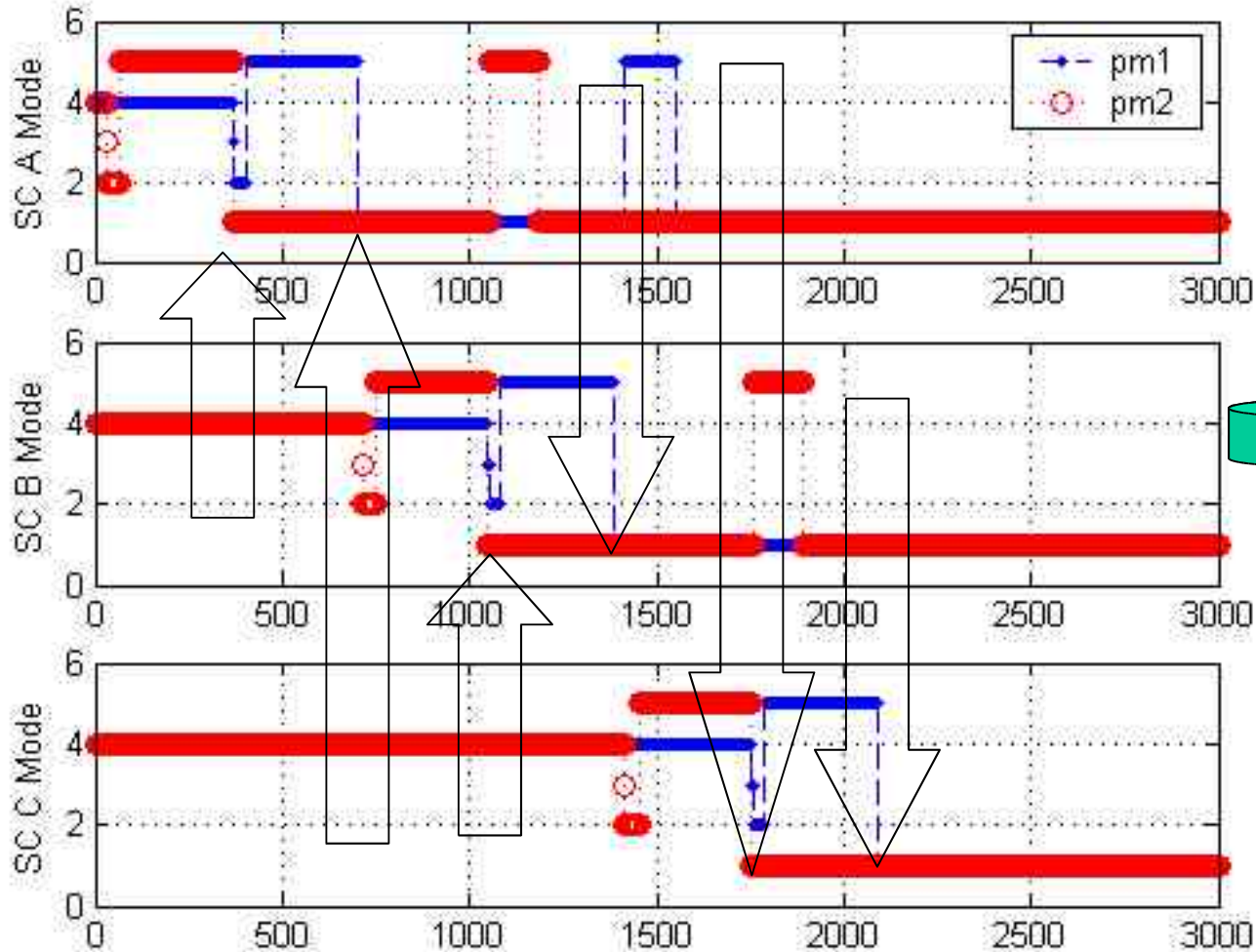
- Autonomous sensor selection and transition based on active and expected field of view



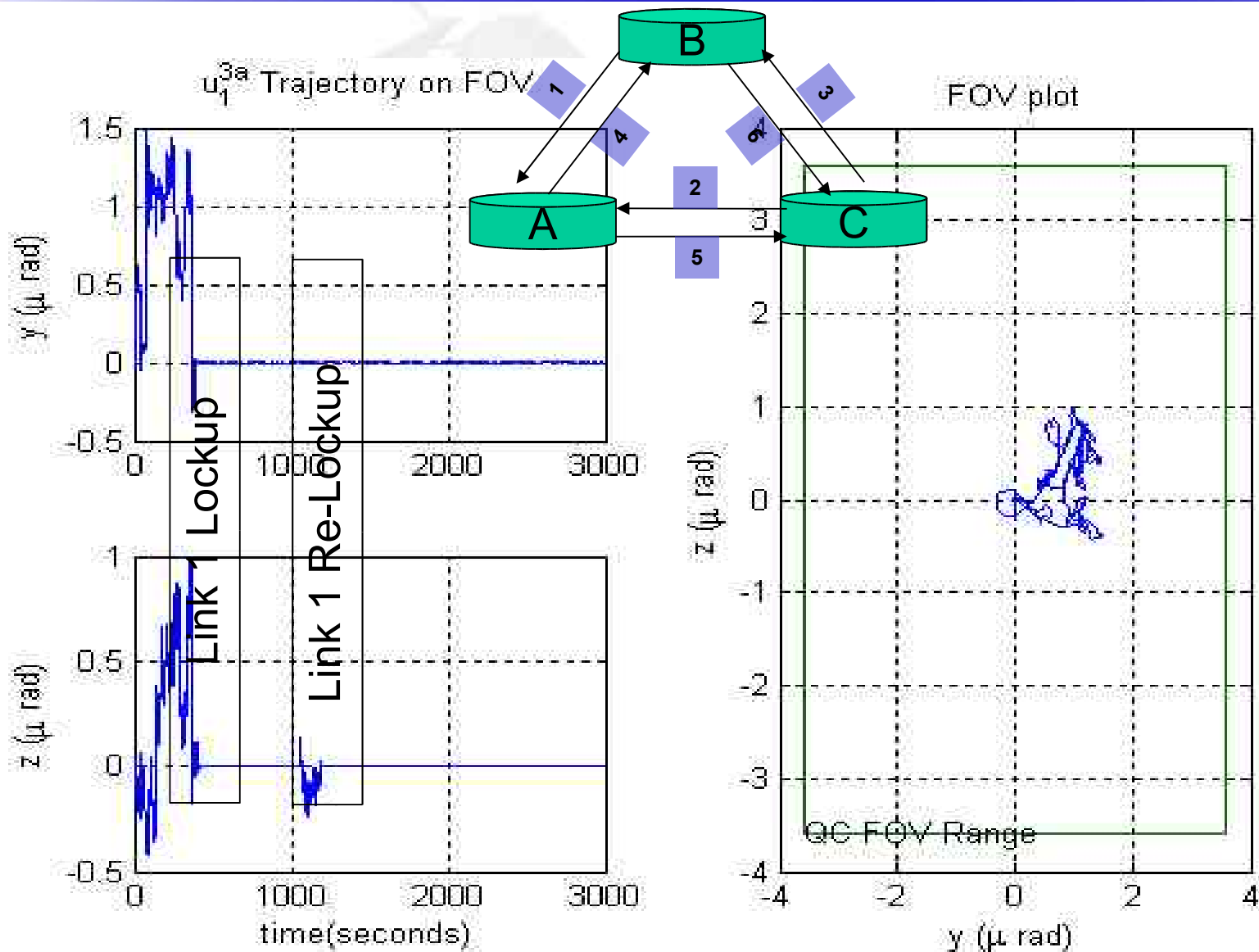
- STATEFLOW is ideal to model the transition logic between the sensors as the incoming light comes into and stays in the FOV

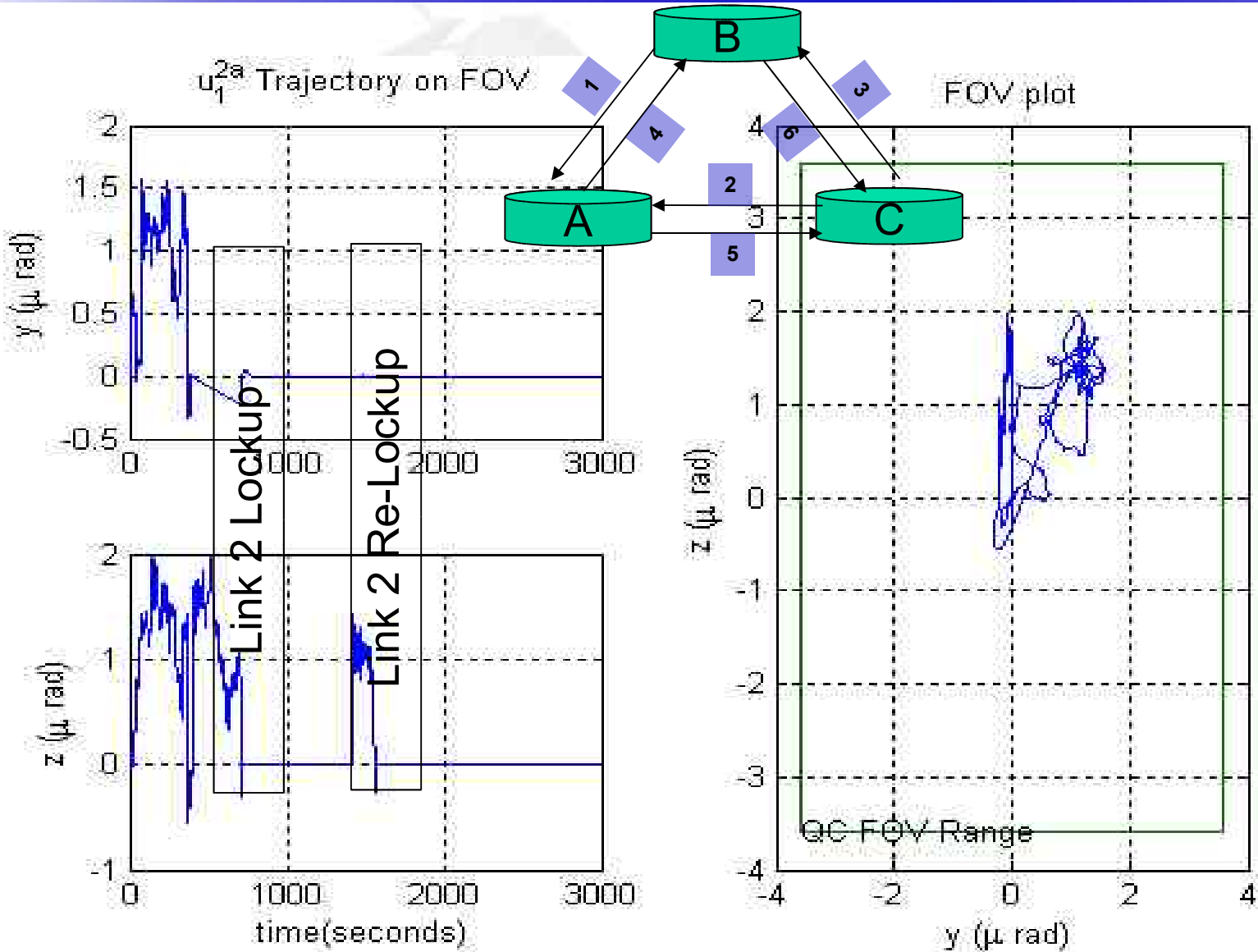


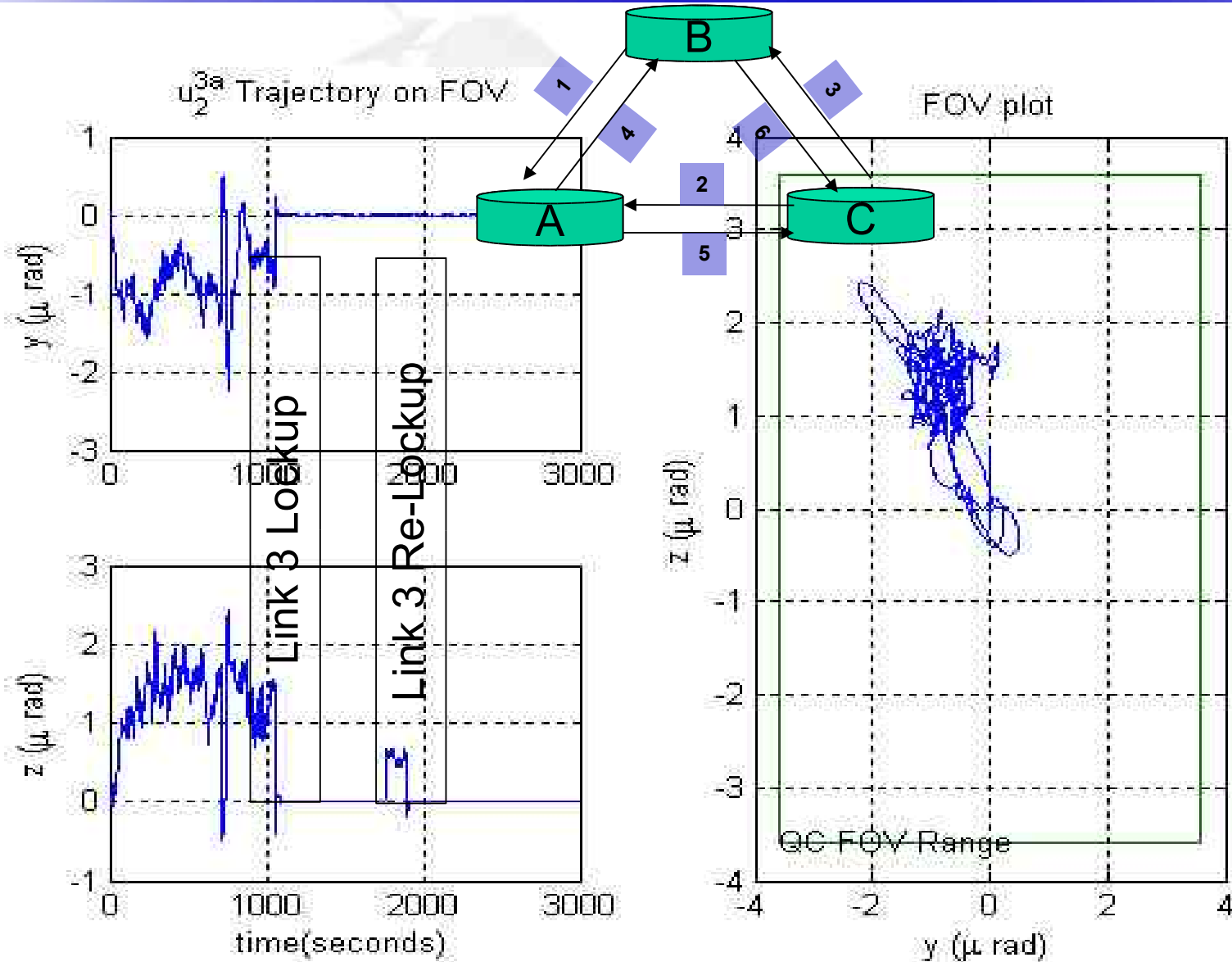
Link: 1 2 3 4 5 6

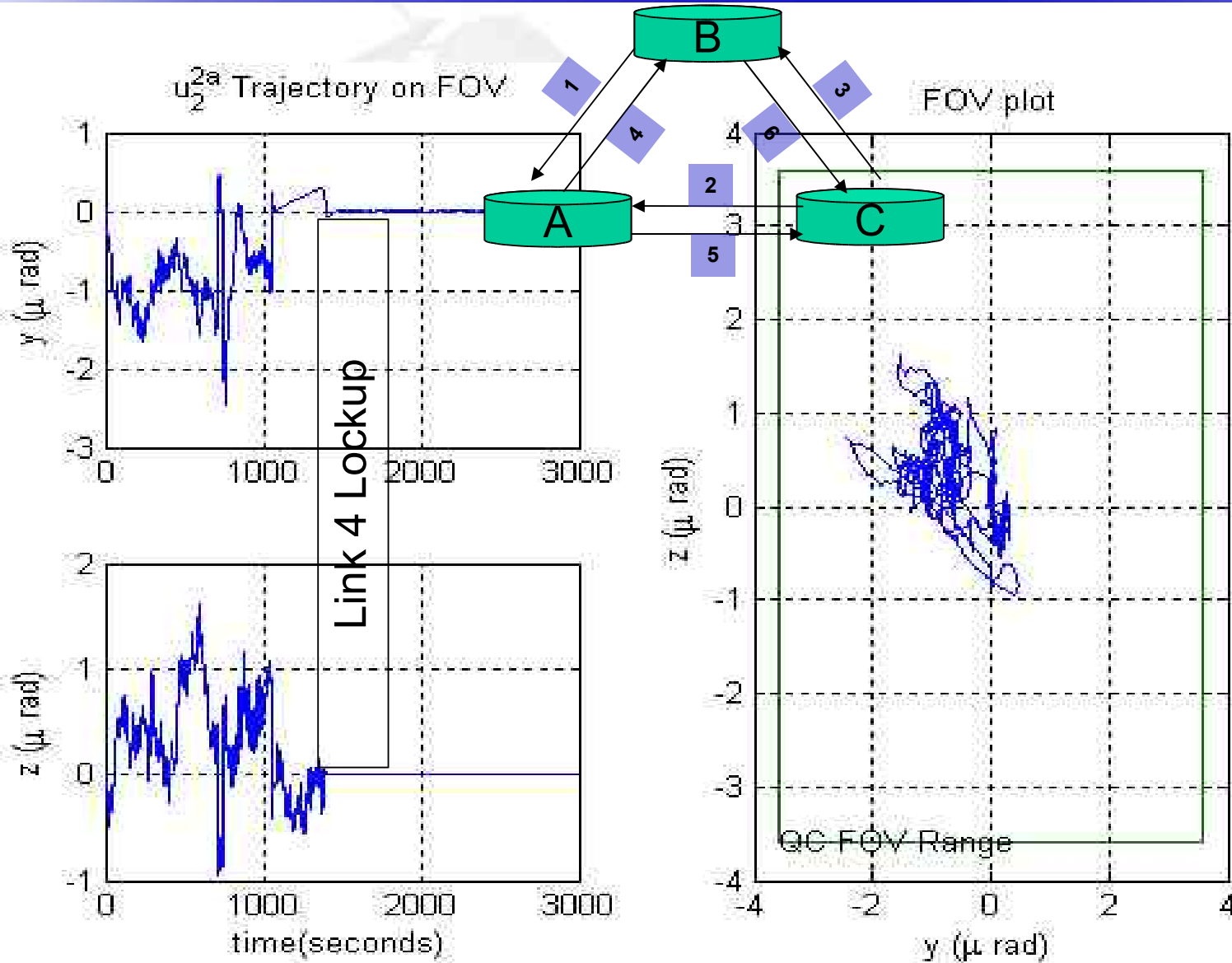


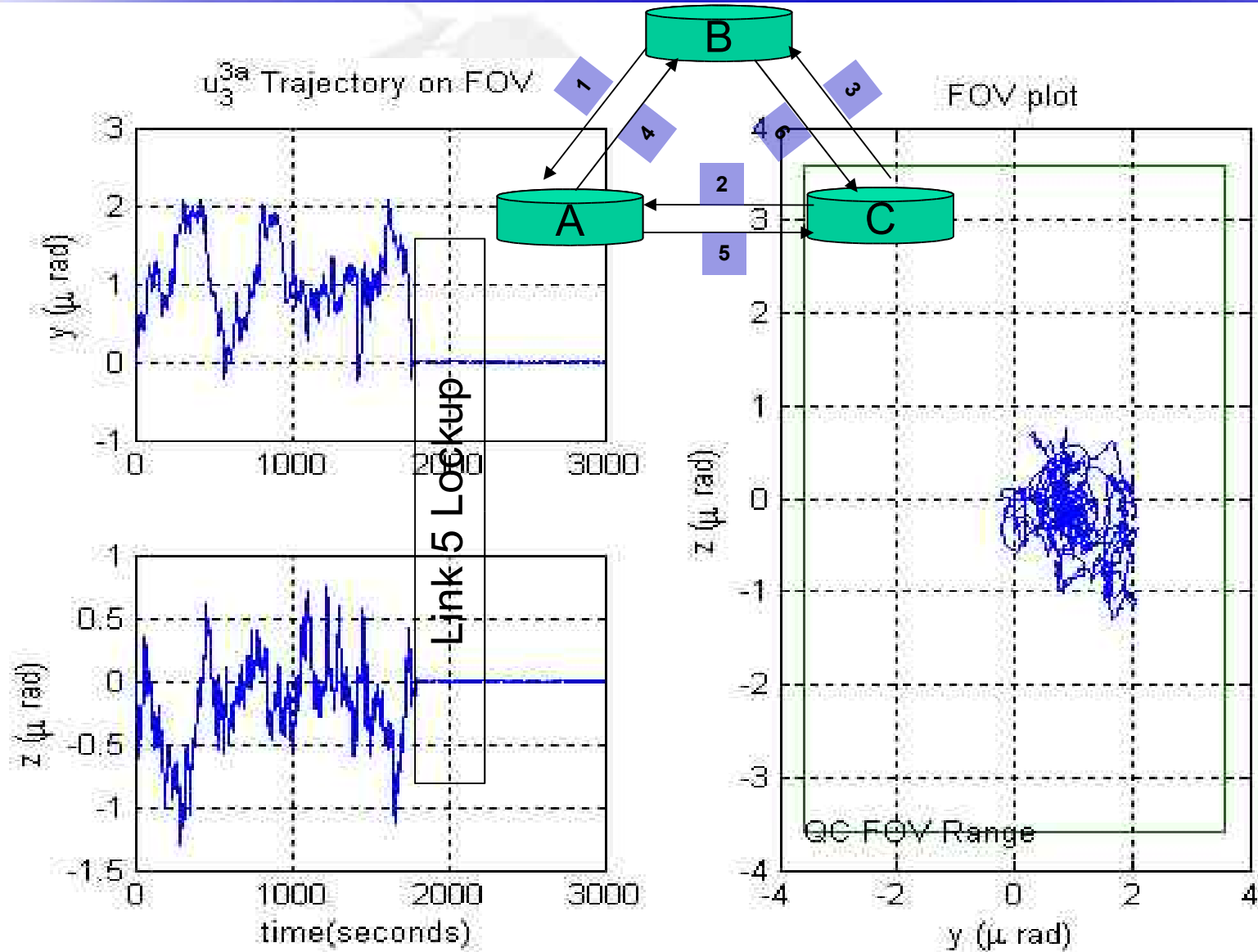
Mode 1: QC/Heterodyne
 Mode 2: QC/Direct
 Mode 3: CCD
 Mode 4: ST/Gyro
 Mode 5: Gyro/Enhanced*
 * Incoming laser information cannot be used.

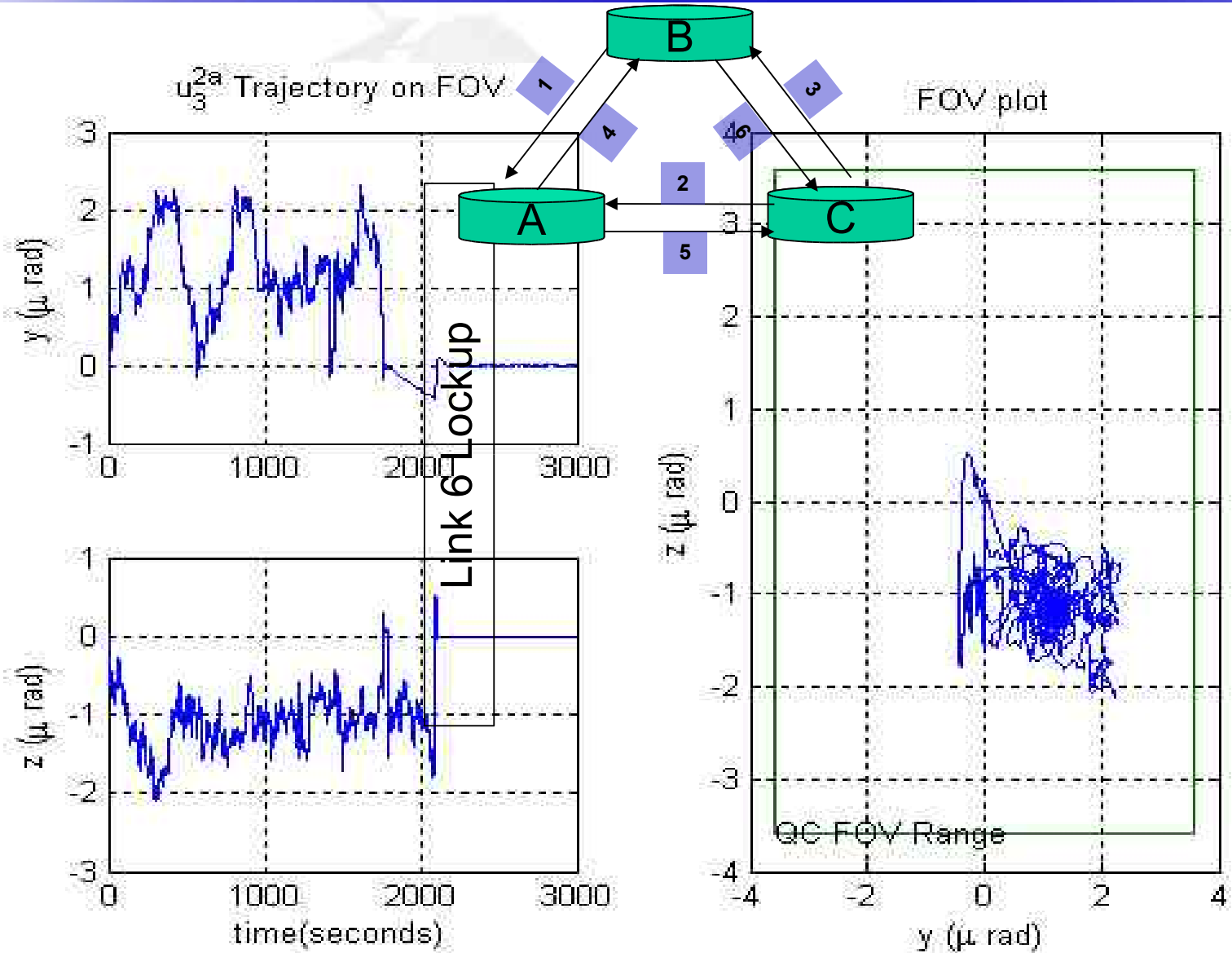










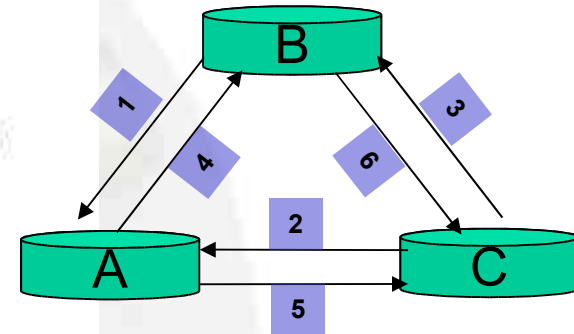
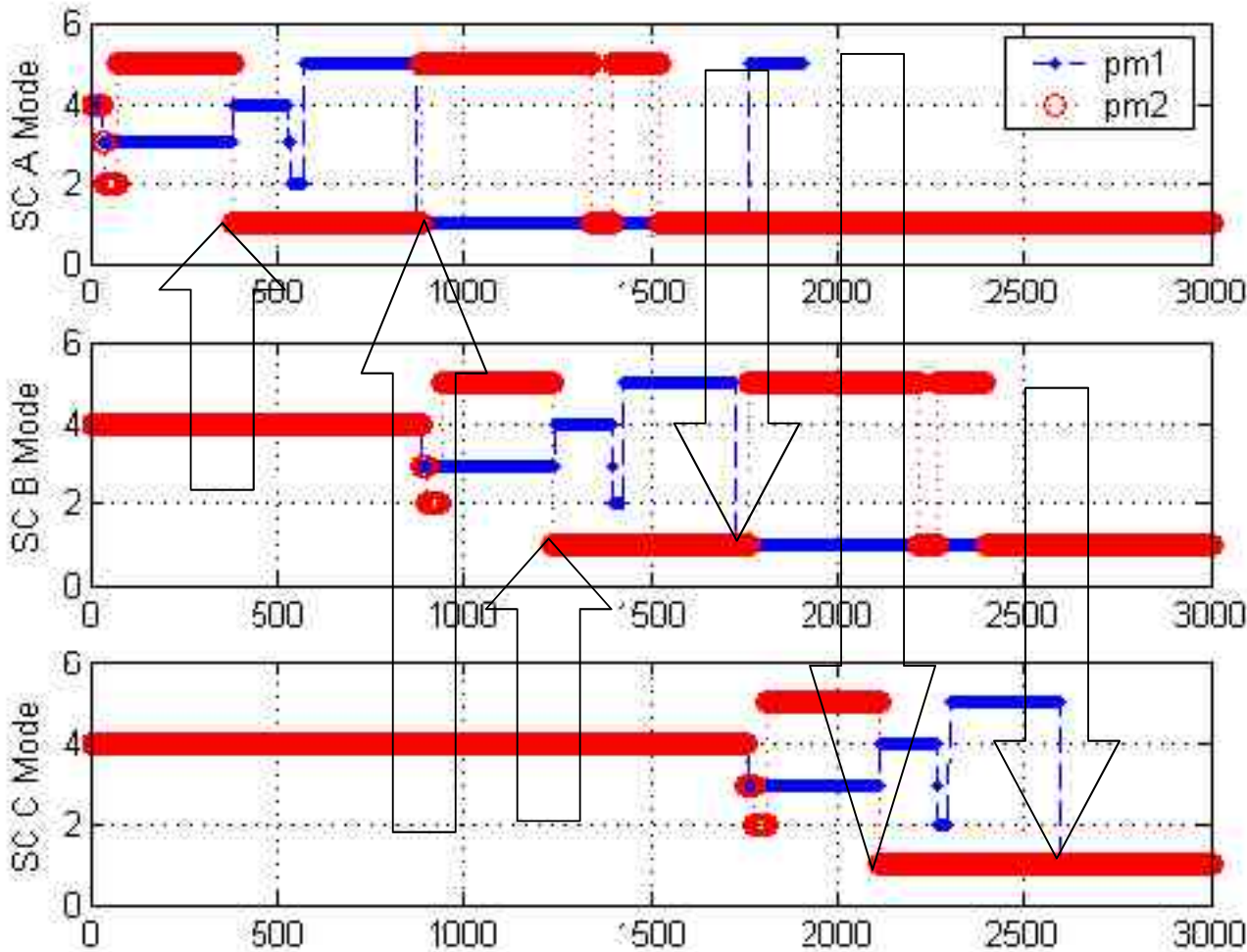




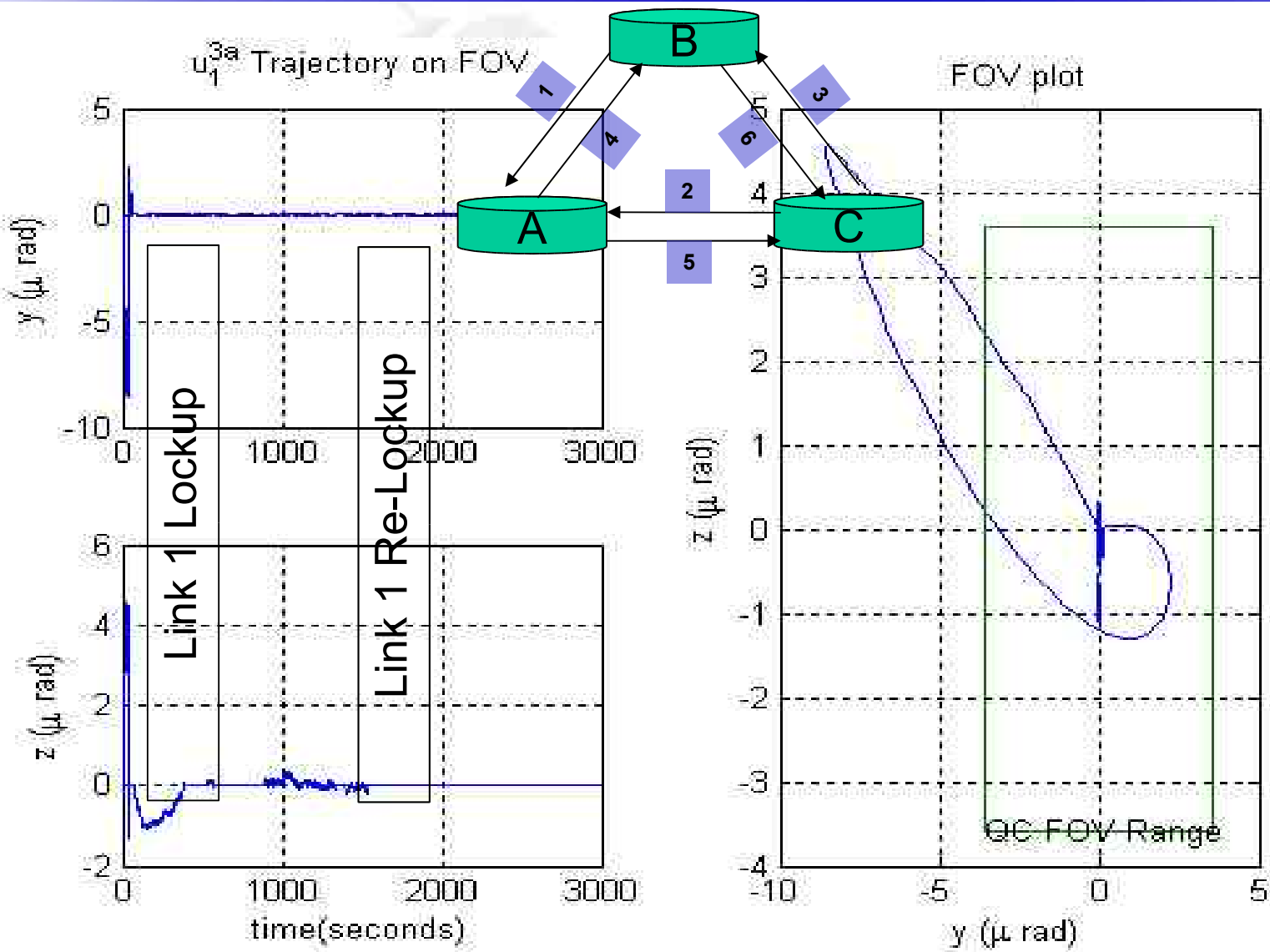
Simulation: Mode Transitions with Bias Estimation

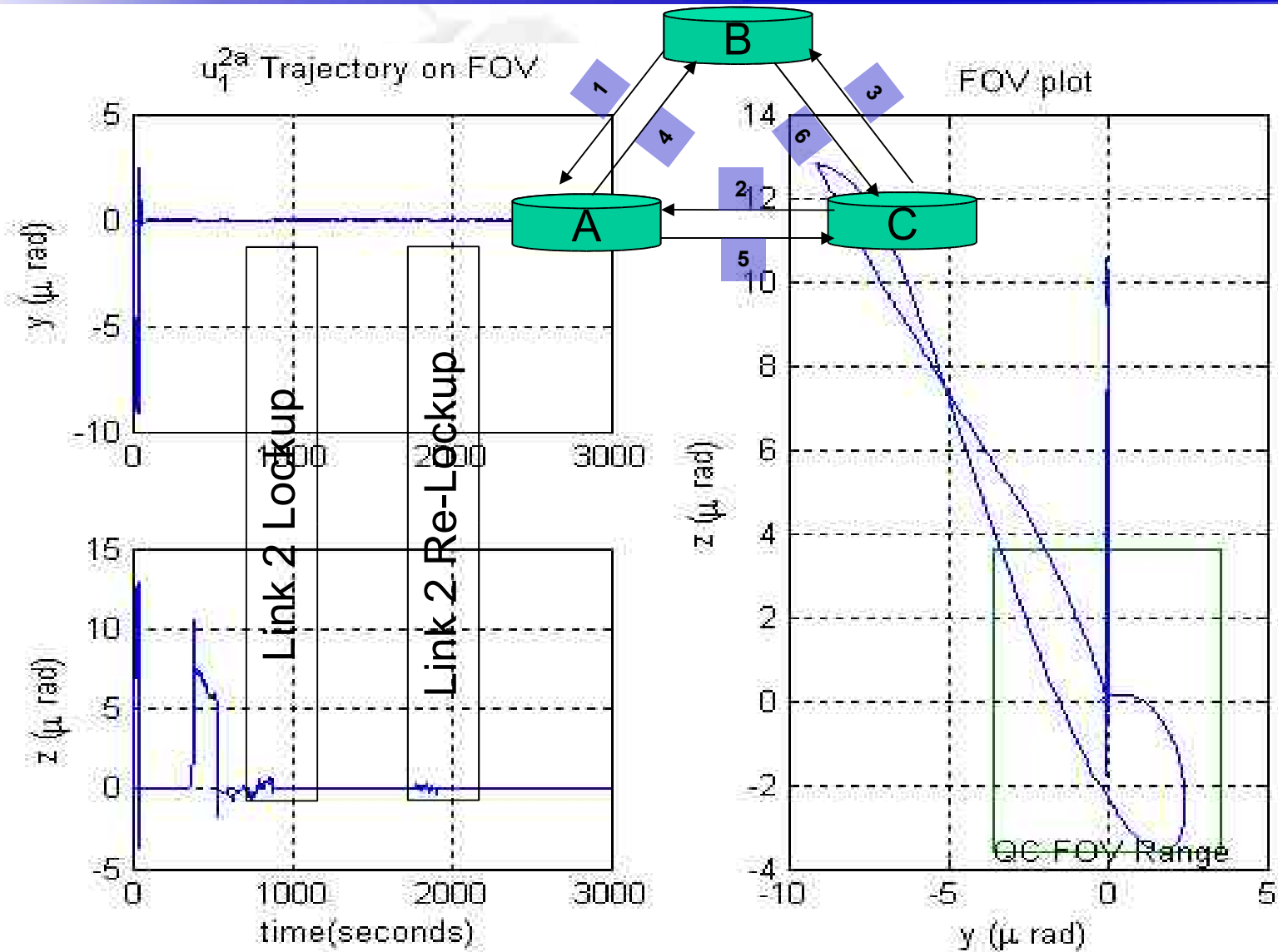


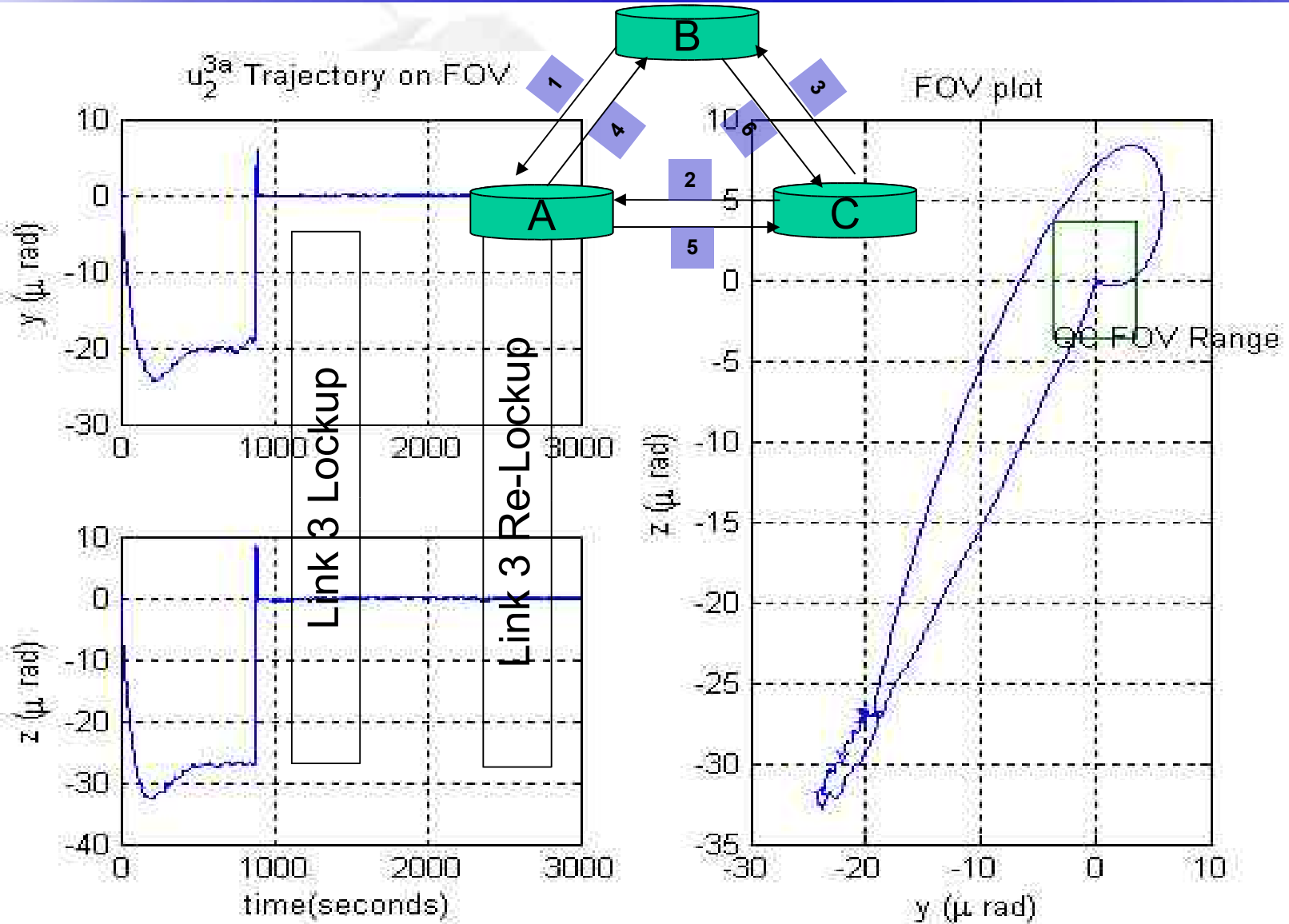
Link: 1 2 3 4 5 6

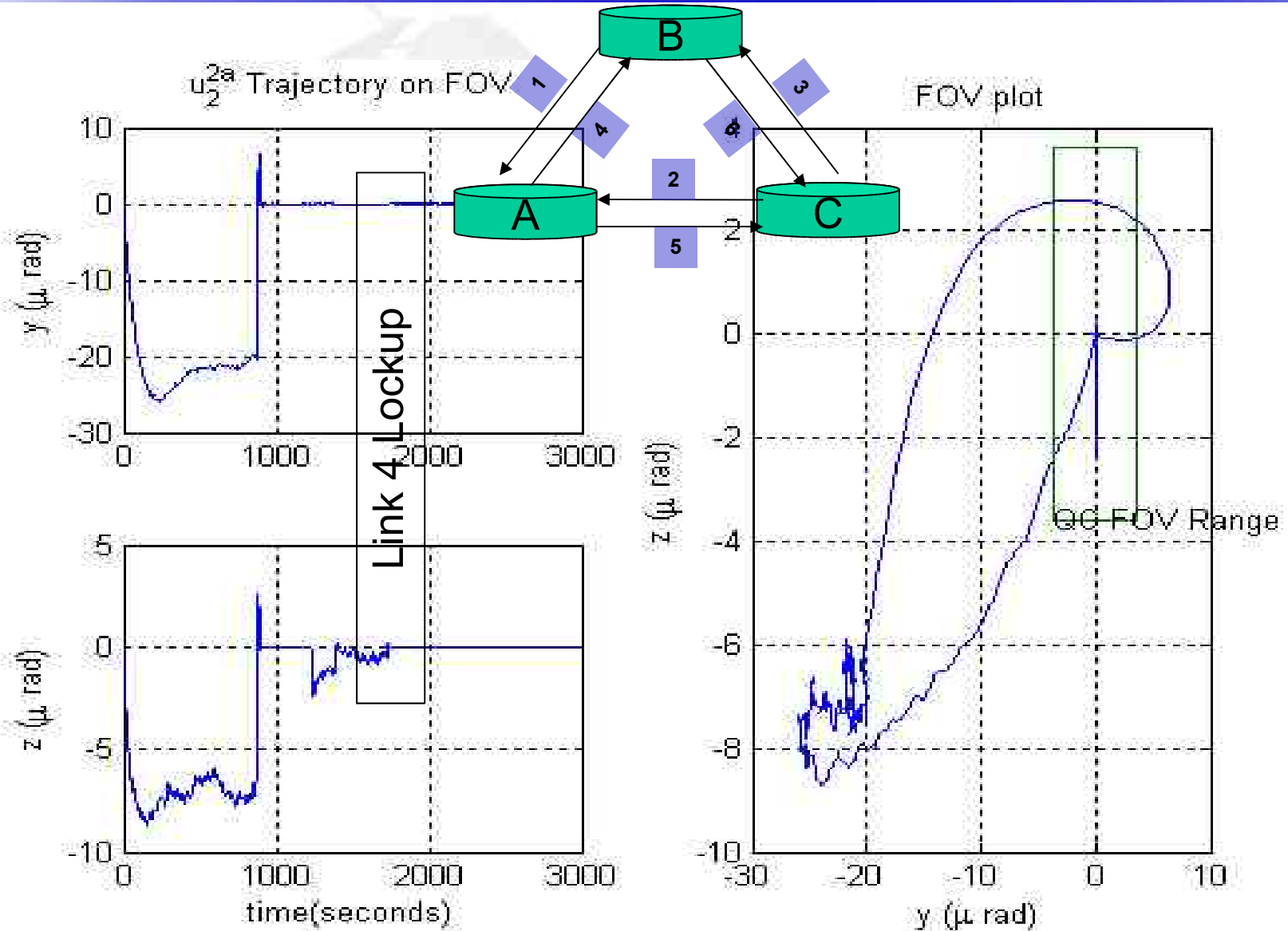


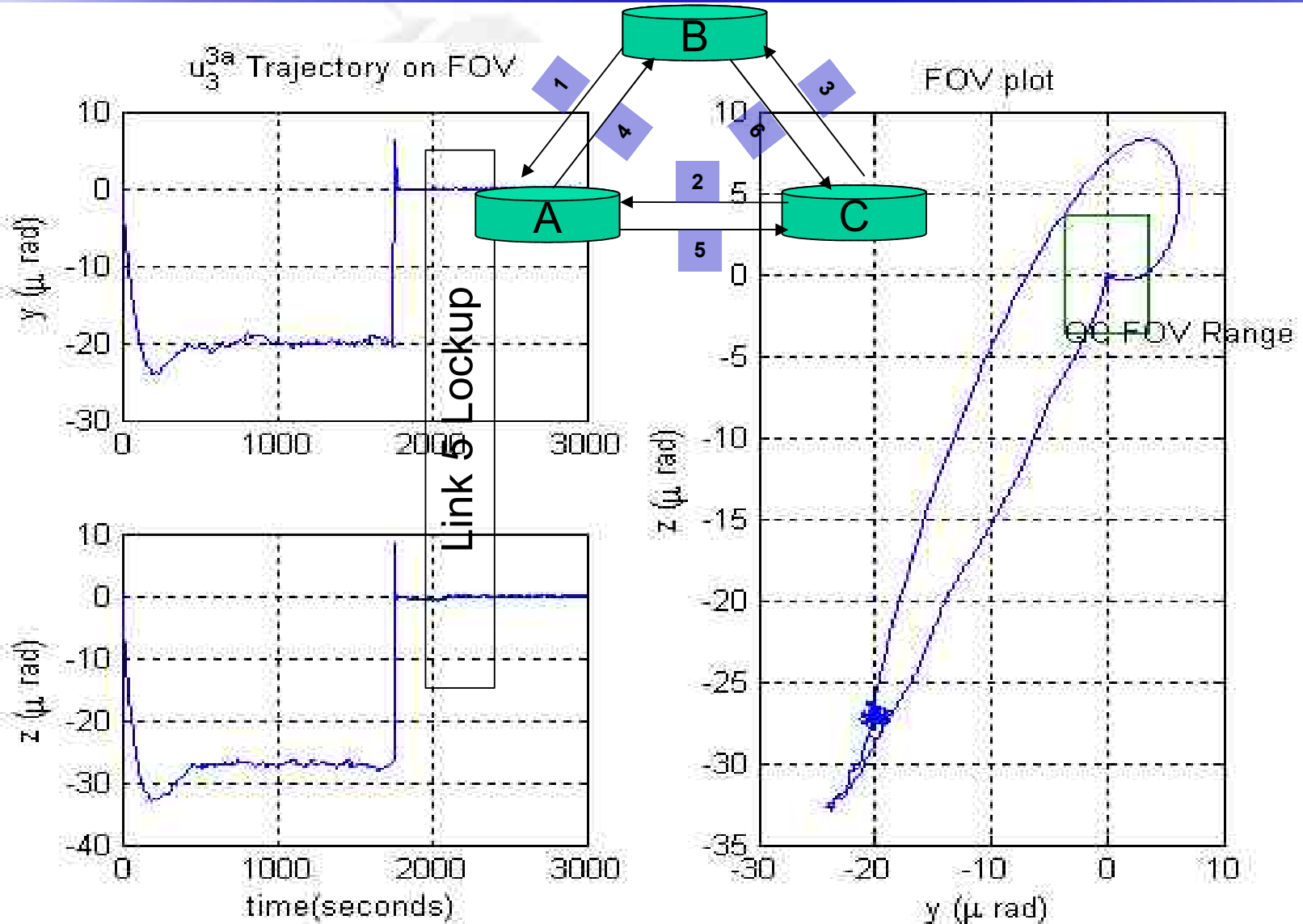
- Mode 1: QC/Heterodyne
 - Mode 2: QC/Direct
 - Mode 3: CCD
 - Mode 4: ST/Gyro
 - Mode 5: Gyro/Enhanced*
- * Incoming laser information cannot be used.

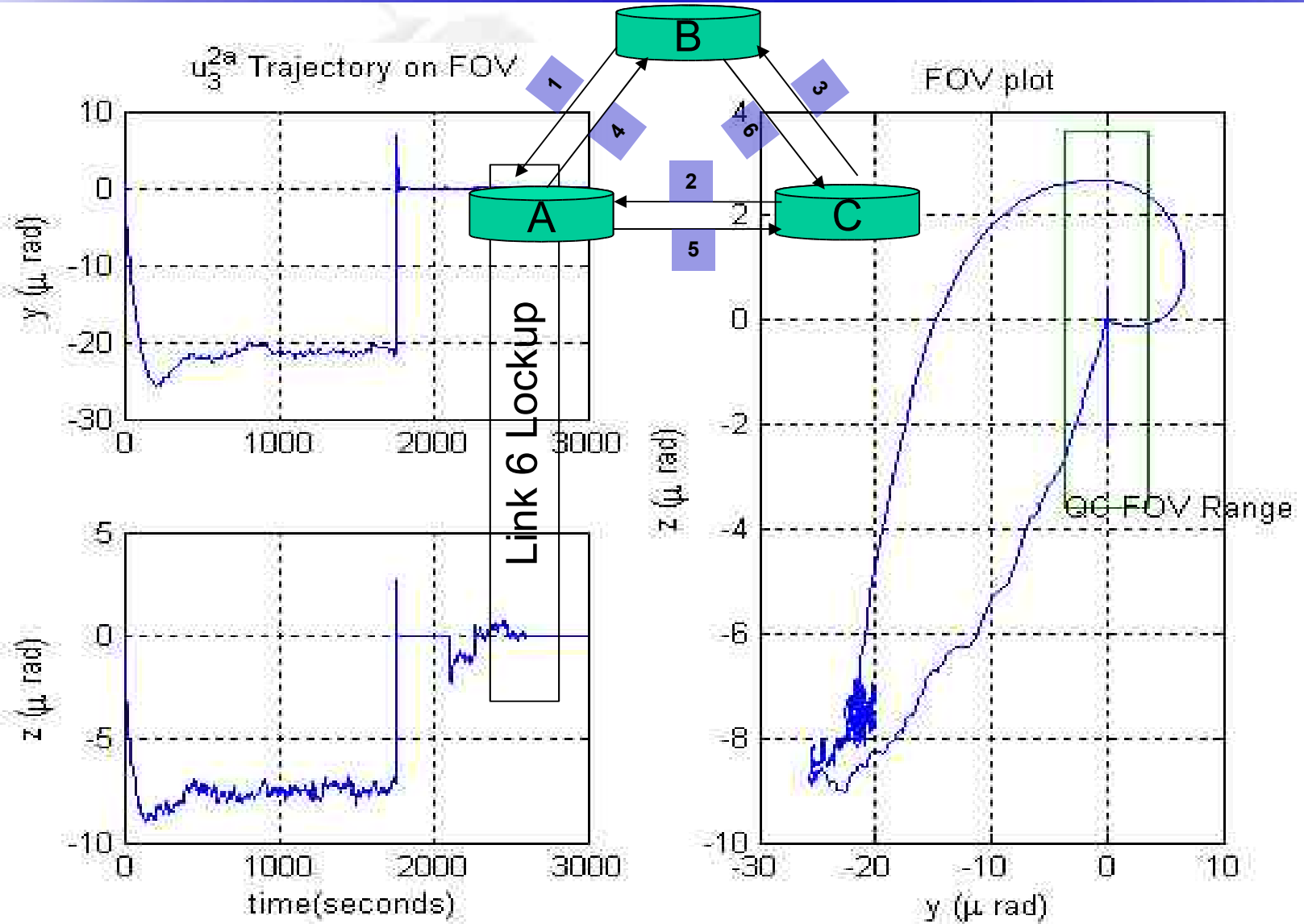












- An observatory laser acquisition strategy based on defocusing has been developed
 - Includes all six links
 - Requires an order to the sequence
 - Uses a Gyro mode for enhanced pointing stability
 - Star tracker bias estimation/calibration can be done a priori or with the acquisition algorithm
- A 57-dof Acquisition model has been developed
 - SIMULINK for time-domain analysis
 - STATEFLOW for switching logic
- Simulation results confirm the feasibility of this acquisition strategy

- Trade of all three strategies need completion
 - Defocus: stray light, mechanism complexity
 - Scan: gyro-mode stability, stray light
 - Thru-telescope star tracker: confirm feasibility
- To complete this, several engineering questions must be answered
 - Gyro-mode stability performance
 - Acquisition CCD performance with stray light
 - Quad cell mode (incoherent) performance
 - Proof mass alignment
 - Star tracker to CCD alignment
 - CCD to quad photodiode alignment
 - Local laser to quad photodiode alignment