



LISA Symposium 2004, Noordwijk, The Netherlands, July 12-15, 2004

Radiation reaction and Gravitational waveform from EMRI



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Problem solved?

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*Extreme Mass Ratio Binary:
Radiation reaction and gravitational waveform*

1: Introduction: Problem solved?



By a recent effort, we come to have the self-force induced by a linear perturbation.

$$\frac{D}{d\tau} V^\alpha = F^\alpha$$

So, what? Is it really available in constructing LISA templates?
We must know the validity of this calculation.



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For mass-ratio

$$\mu \approx 10^{-5}$$

dynamical time:

$$T_{\text{dynamical}} \approx 10 - 100 \text{ sec}$$

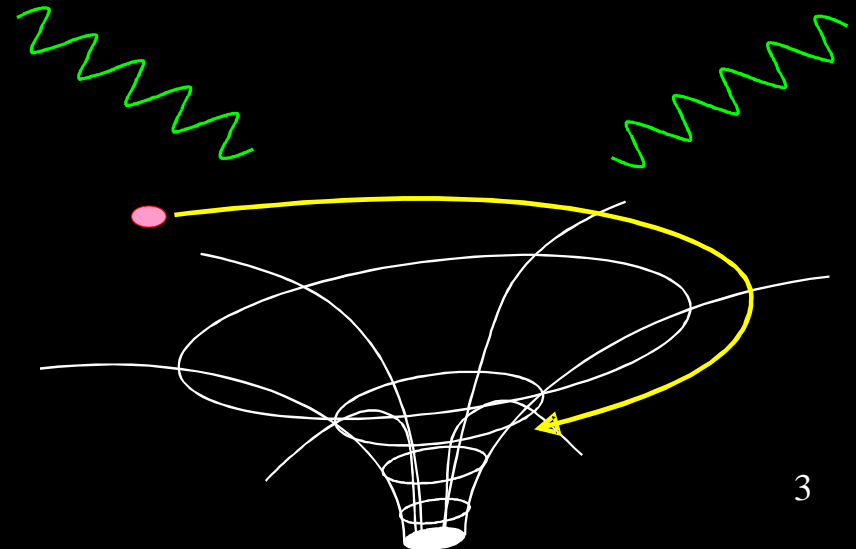
de-phasing time:

$$T_{\text{de-phase}} \approx v^{-2.5} \mu^{-0.5} T_{\text{dynamical}} \approx 1 \text{ day} - 1 \text{ week}$$

radiation reaction time:

$$T_{\text{reaction}} \approx v^{-5} \mu^{-1} T_{\text{dynamical}} \approx 1 - 10 \text{ years}$$

We need the orbital evolution of the radiation reaction time.





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Orbit Energetics:

Suppose we calculate the energy loss by a linear perturbation, the 2nd order perturbation acts as an error.

$$\dot{E} = \dot{E}^{(1)} + \dot{E}^{(2)} + \dots$$

$$(E)_{Error} \approx \dot{E}^{(2)} t$$

The linear approximation is valid enough.

$$T_{error} \approx v^{-7} \mu^{-2} T_{dynamical} \approx 10^7 \text{ years}$$

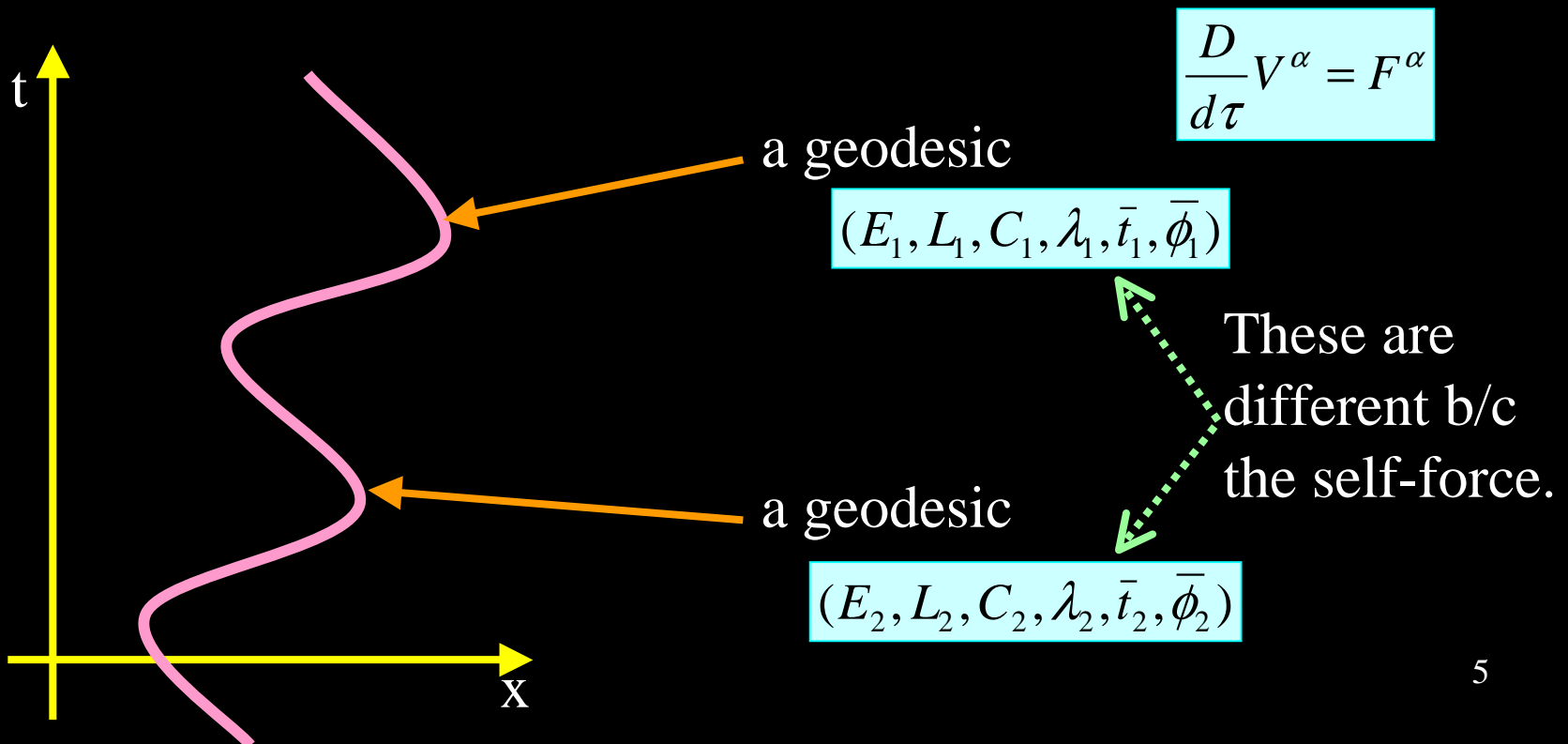
This type of estimate was widely done in PN formalism.

This estimate does not apply in the self-force calculation in general.

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2: Orbital Evolution & Gauge

We consider an orbital evolution by the adiabatic picture and discuss the gauge dependence of the resulting orbit.

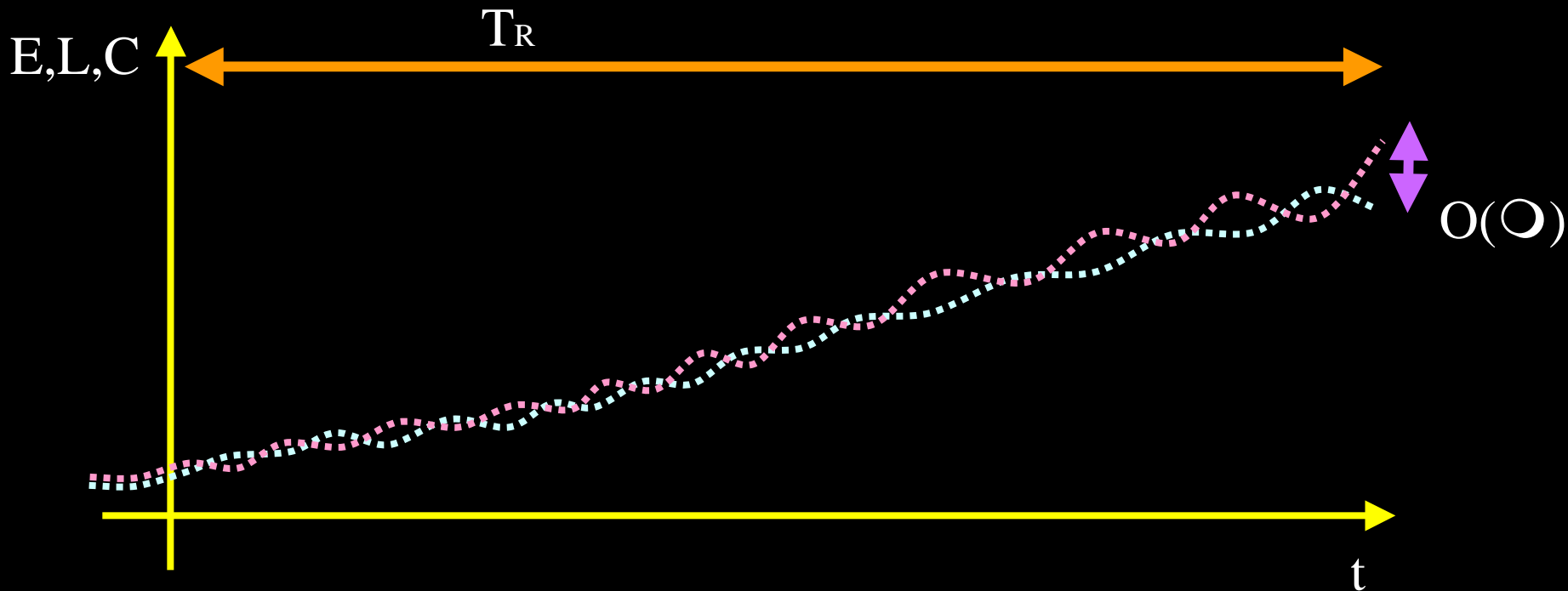




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We compare the orbital evolution by the self-force calculated by different gauges.

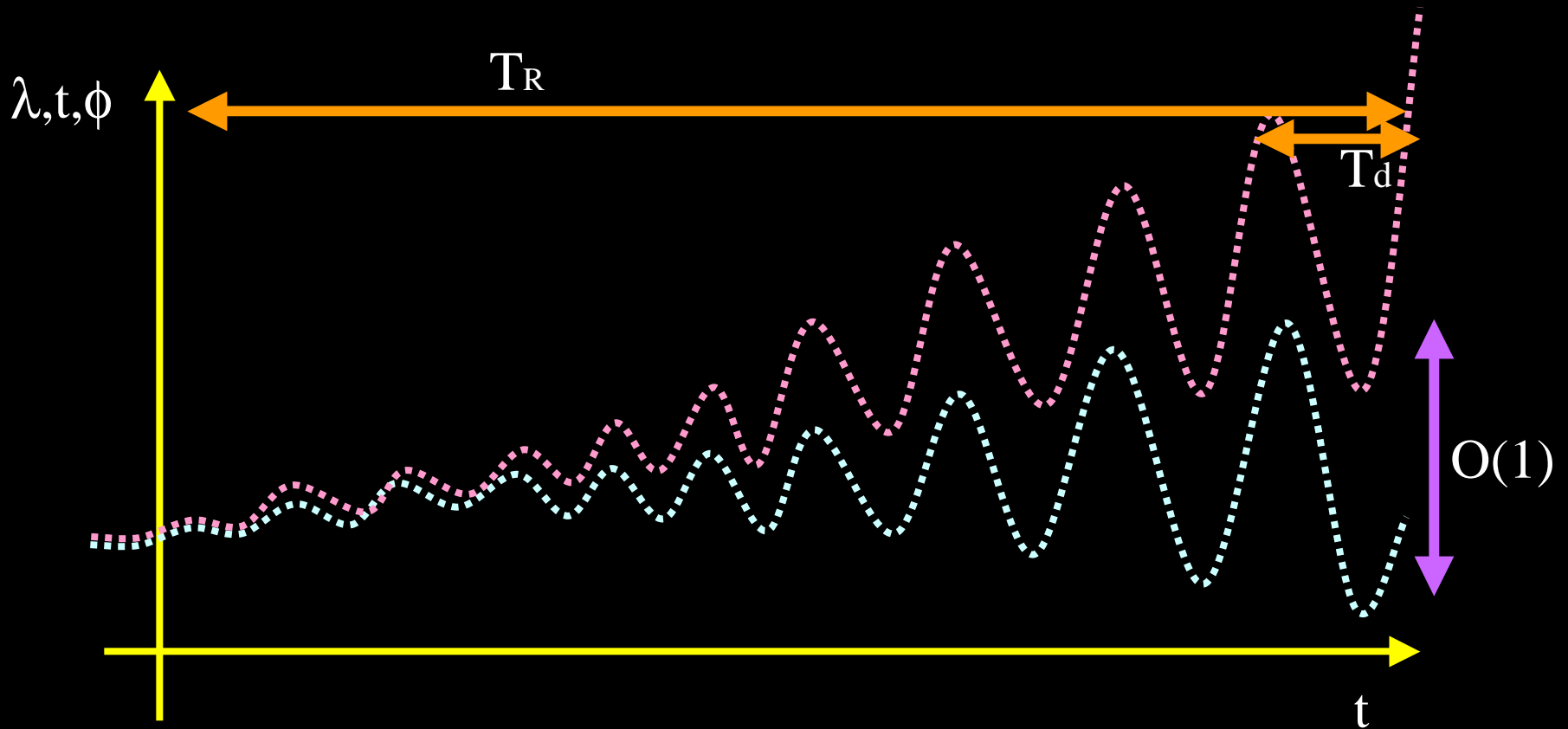


We have a same prediction of the orbital evolution.



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We have different predictions on the orbital evolution.
The adiabatic approximation is broken down.



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Different results of different gauge conditions suggest that we may not have a correct prediction at the radiation reaction time scale by the self-force calculation.

What is going on?

The gauge is growing!

$$g_{\alpha\beta} = g_{\alpha\beta}^{BH} + h_{\alpha\beta}^{(1)} + h_{\alpha\beta}^{(2)} + \dots$$

$$h_{\alpha\beta}^{(1)} \approx \mu \times t$$

$$h_{\alpha\beta}^{(2)} \approx \mu^2 \times t^2$$

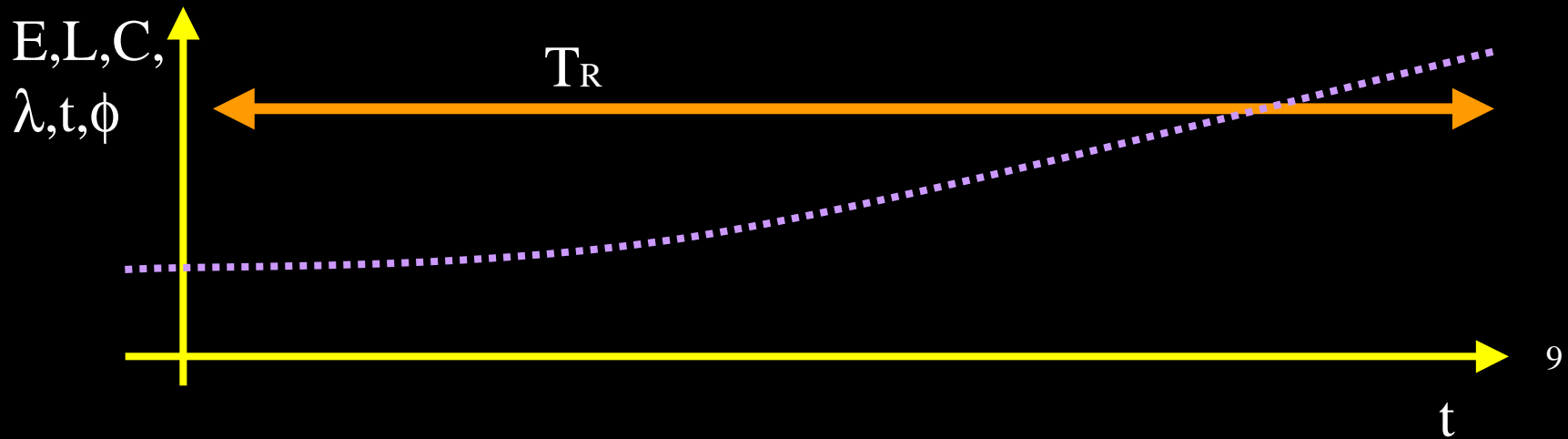
The system becomes non-perturbative at the radiation reaction time scale. We need to study 2nd order perturbation to find an optimal gauge.



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3: Radiation Reaction Gauge

We consider the self-force problem in a non-perturbative manner, and we find a special gauge condition (Adiabatic gauge). In this gauge, the orbital equation is written in an gauge invariant manner. (Improved Radiation Reaction Formula)





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Summary of this formula;

- 1) The estimate of the orbital energetics applies, hence, we can predict the orbital evolution beyond the radiation reaction time scale.
- 2) This is a non-trivial extension of PN radiation reaction formula used for a quasi-circular orbit, which was successfully applied to Nobel-awarded observation of the Hulse-Taylor binary.
- 3) This fully describes the orbital motion, including the orbital phasing.
- 4) The required computation is minimal. The regularization of the self-force is naturally done.



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$$\frac{d}{d\tau} E^i = F^i(E)$$

$$E^i = E^i(z, \nu)$$



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4: Waveform calculation

The self-force calculation:

$10^2(10^6)$ the self-force for a given geodesic
x

10^4 evolution for 1 year
x

10^{11} number of templates
x

10^{-6} ? extrapolation(?)
sec

We need 100 years to generate templates!

Apology: I almost finished my coding, however, at the end stage of coding, I found the problem on the validity of the self-force, and, I concentrated on this instead of finishing up the program.



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Improved Radiation Reaction Formula:

1 Radiation reaction for a given geodesic
 \times
 10 evolution for 1 year
 \times
 10^{11} number of templates
 \times
 $10^{-6?}$ extrapolation(?)

We need 1 month to generate templates



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5: Conclusion

We found a class of gauges with which we can predict an orbital evolution necessary for LISA project.

Making an efficient program to generate templates is in progress.