

Algebraic Approach to Time-Delay Data Analysis for Orbiting LISA

K. Rajesh Nayak and J-Y. Vinet

**Dept. Artemis
Observatoire de la Côte d'Azur,
Nice.**

TDI- Approach to suppress Laser Noise!

J.W. Armstrong, F.B. Estabrook and M. Tinto, *Astrophys. J* **527**, 814(1999).

- Elegant Approach to cancel the overwhelming ($\frac{\Delta v}{v_0} = 10^{-14} Hz / \sqrt{Hz}$) laser noise in the space based gravitational wave detectors.

TDI- Approach to suppress Laser Noise!

J.W. Armstrong, F.B. Estabrook and M. Tinto, *Astrophys. J* **527**, 814(1999).

- Elegant Approach to cancel the overwhelming ($\frac{\Delta v}{v_0} = 10^{-14} Hz / \sqrt{Hz}$) laser noise in the space based gravitational wave detectors.
- Laser beams are combined off-line by introducing the suitable time delays corresponding to the arm length to simulate the interferometer.

TDI- Approach to suppress Laser Noise!

J.W. Armstrong, F.B. Estabrook and M. Tinto, *Astrophys. J* **527**, 814(1999).

- Elegant Approach to cancel the overwhelming ($\frac{\Delta v}{v_0} = 10^{-14} Hz / \sqrt{Hz}$) laser noise in the space based gravitational wave detectors.
- Laser beams are combined off-line by introducing the suitable time delays corresponding to the arm length to simulate the interferometer.

F.B. Estabrook, M. Tinto and J.W. Armstrong, *Phys. Rev.* **D62**, 042002(2000).

Algebraic Approach to TDI Analysis

S. V. Dhurandhar, K. R. Nayak and J-Y. Vinet, *Phys. Rev. D* **65**, 102002(2002).

- Systematic approach based on computational commutative algebra.
- All the noises cancellation data combination can be generated from a set of 4 generators.
- This formalism can easily be extend to alternative geometry or the follow on missions.
- In the present work, we extend the formalism to include the motion of LISA(Sagnac effect)

Sagnac Effect in LISA

N. J. Cornish and R. W Hellings, *Class. Quantum Grav.* **20**, 4851(2003).

D. A. Shaddock, *Phys. Rev. D* **69**, 022001 (2004).

- The Rigid rotation of LISA triangle with period one year about its own axis results in Sagnac Phase. ($\simeq 4\vec{\Omega} \cdot \vec{A}$)

Sagnac Effect in LISA

N. J. Cornish and R. W Hellings, *Class. Quantum Grav.* **20**, 4851(2003).

D. A. Shaddock, *Phys. Rev. D* **69**, 022001 (2004).

- The Rigid rotation of LISA triangle with period one year about its own axis results in Sagnac Phase. ($\simeq 4\vec{\Omega} \cdot \vec{A}$)
- As a result up and down light travel time for the laser beam along same arm is different(by about few Km)

Sagnac Effect in LISA

N. J. Cornish and R. W Hellings, *Class. Quantum Grav.* **20**, 4851(2003).

D. A. Shaddock, *Phys. Rev. D* **69**, 022001 (2004).

- The Rigid rotation of LISA triangle with period one year about its own axis results in Sagnac Phase. ($\simeq 4\vec{\Omega} \cdot \vec{A}$)
- As a result up and down light travel time for the laser beam along same arm is different(by about few Km)
- This results in large residual laser noise, Hence set of new TDI solution were proposed.

- We apply the algebraic approach by taking the up and down path lengths are different for same arm and obtain a set of generators.

Delay operators used in TDI are defined as,

$$E_{ij}A(t) = A(t - L_{ij})$$

in the present case $L_{ij} \neq L_{ji}$ With this extra constraint, we follow the formalism given for stationary case.

- Earlier results assume a simple module in which LISA rotates only about its own axis!!
- In reality the motion of LISA is much more complex and the our study shows that the main for Sagnac effect comes form orbital motion.

Sagnac effect in LISA system

Metric in LISA frame, $O(\Omega)$

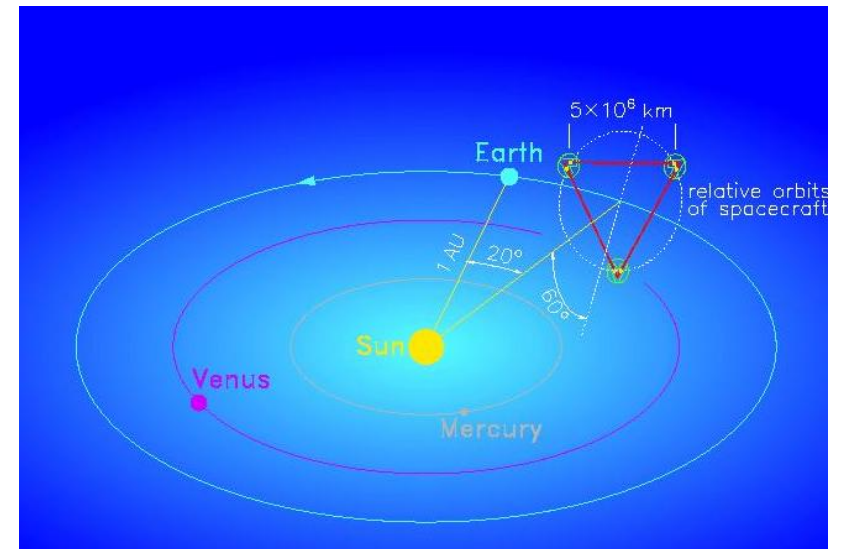
$$g_{00} = 1 + O(\Omega^2)$$

$$g_{01} = R_{\odot}\Omega \cos \Omega t - \frac{\Omega}{2}y - \frac{\sqrt{3}}{2}\Omega z \cos \Omega t,$$

$$g_{02} = R_{\odot}\Omega \sin \Omega t + \frac{\Omega}{2}x - \frac{\sqrt{3}}{2}\Omega z \sin \Omega t,$$

$$g_{03} = \frac{\sqrt{3}}{2}\Omega (x \cos \Omega t + y \sin \Omega t),$$

$$g_{11} = -1 = g_{22} = g_{33}$$



Metric is time dependent!

Sagnac Phase for LISA Beams

$$R_{\odot} \simeq 150 \times 10^6 \text{Km}$$

$$L \simeq 5 \times 10^6 \text{Km}$$

$$\Omega \simeq 10^{-7} \text{sec}^{-1}$$

$$\Delta\phi_{orbit}(t) / \Delta\phi_L \simeq R/L$$

$$\Delta\tau_{U1} = L \left(1 + \overbrace{\Omega R_{\odot} \sin \Omega t} - \underbrace{\frac{\Omega L}{4\sqrt{3}}} \right),$$

$$\Delta\tau_{U2} = L \left(1 + \overbrace{\Omega R_{\odot} \sin \left(\Omega t - \frac{\pi}{3} \right)} - \underbrace{\frac{\Omega L}{4\sqrt{3}}} \right),$$

$$\Delta\tau_{U3} = L \left(1 - \overbrace{\Omega R_{\odot} \sin \left(\Omega t + \frac{\pi}{3} \right)} - \underbrace{\frac{\Omega L}{4\sqrt{3}}} \right),$$

$$\Delta\tau_{V1} = L \left(1 - \overbrace{\Omega R_{\odot} \sin \left(\Omega t - \frac{\pi}{3} \right)} + \underbrace{\frac{\Omega L}{4\sqrt{3}}} \right),$$

$$\Delta\tau_{V2} = L \left(1 + \overbrace{\Omega R_{\odot} \sin \left(\Omega t + \frac{\pi}{3} \right)} + \underbrace{\frac{\Omega L}{4\sqrt{3}}} \right),$$

$$\Delta\tau_{V3} = L \left(1 - \overbrace{\Omega R_{\odot} \sin \Omega t} + \underbrace{\frac{\Omega L}{4\sqrt{3}}} \right).$$

The Generators for the Module

We apply the formalism similar to the stationary case and obtain the generators,

$$d^{(1)} = (E_{13}(1 - E_{23}E_{32}), E_{21}E_{13} - E_{23}, 1 - E_{23}E_{32}, 0, 0, 1 - E_{21}E_{13}E_{32}),$$

$$d^{(2)} = (E_{23}(1 - E_{13}E_{31}), 0, E_{21} - E_{31}E_{23}, E_{23} - E_{21}E_{13}, 0, E_{21}(1 - E_{13}E_{31})),$$

$$d^{(3)} = (0, 1 - E_{31}E_{13}, E_{32} - E_{12}E_{31}, E_{12} - E_{13}E_{32}, 1 - E_{13}E_{31}, 0),$$

$$d^{(4)} = (E_{12} - E_{13}E_{32}, E_{12}E_{21} - 1, E_{32}(E_{21}E_{12} - 1), 0, E_{21}E_{13}E_{32} - 1, 0),$$

$$d^{(5)} = (E_{23}E_{32} - 1, E_{31}E_{23} - E_{21}, 0, E_{23}E_{32} - 1, 0, E_{21}E_{32} - E_{31}),$$

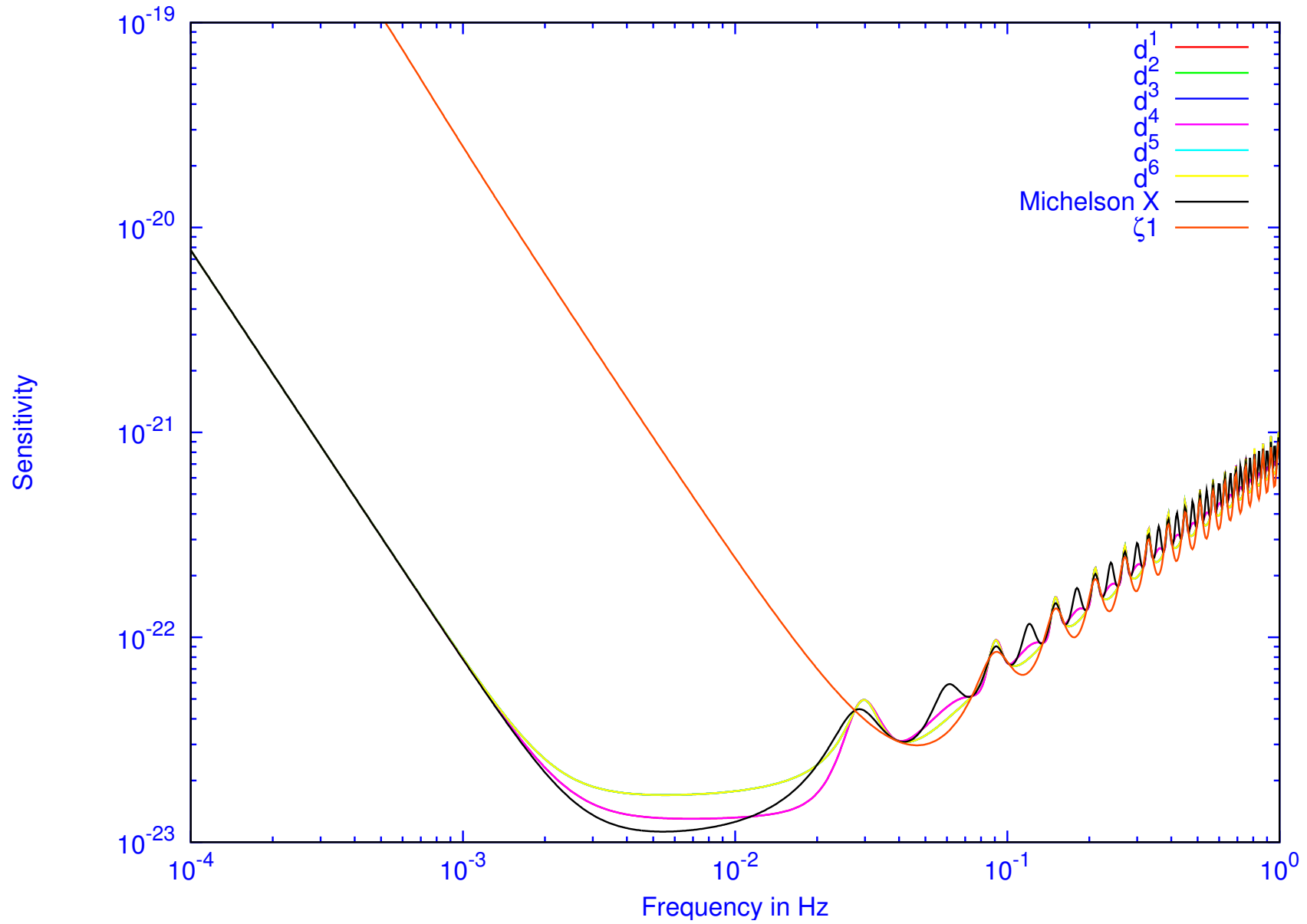
$$d^{(6)} = (E_{13} - E_{12}E_{23}, 0, 1 - E_{12}E_{21}, 0, E_{23} - E_{21}E_{13}, 1 - E_{12}E_{21}).$$

The Sagnac and Symmetric Sagnac solutions do not form a generating set. Any noise cancellation data combination X can be expressed as,

$$X = \sum_{I=1}^6 \alpha_{(I)} d^{(I)}$$

where $\alpha_{(I)}$ are polynomial coefficients in six variable E_{ij} .

The Sensitivity Curves



Conclusions

- The contribution from Sagnac effect is much larger than earlier predicted.

Conclusions

- The contribution from Sagnac effect is much larger than earlier predicted.
- In this case also laser noise cancellation solutions form the first module of syzygies over polynomial ring in six time delays E_{ij} .

Conclusions

- The contribution from Sagnac effect is much larger than earlier predicted.
- In this case also laser noise cancellation solutions form the first module of syzygies over polynomial ring in six time delays E_{ij} .
- The analysis can be extended to cancel optical bench noise in a straight forward manner.

Conclusions

- The contribution from Sagnac effect is much larger than earlier predicted.
- In this case also laser noise cancellation solutions form the first module of syzygies over polynomial ring in six time delays E_{ij} .
- The analysis can be extended to cancel optical bench noise in a straight forward manner.
- The generators are non-trivial in this case. The Sagnac and Symmetric Sagnac combinations do not form a generating set as in the stationary case.
- The formalism is useful as set of all solution can be parameterised in terms of coefficients of generators.

Applications(Stationary case)

- ★ **Maximisation of SNR:** Since each noise cancellation solution has different signal response, we can determine the data combination with maximum SNR

Applications(Stationary case)

- ★ **Maximisation of SNR:** Since each noise cancellation solution has different signal response, we can determine the data combination with maximum SNR
- ★ For sources without any prior information: there are both upper and lower bound for SNR and Network observable can be formed from un-correlated data combinations

T. A. Prince, M. Tinto, S. L. Larson and J. W. Armstrong, *Phys. Rev. D* **66**, 122002 (2002).
K. R. Nayak, A. Pai, S. V. Dhurandhar and J-Y. Vinet, *Class. Quantum Grav.* **20**, 1217(2003).

Applications(Stationary case)

- ★ **Maximisation of SNR:** Since each noise cancellation solution has different signal response, we can determine the data combination with maximum SNR
- ★ For sources without any prior information: there are both upper and lower bound for SNR and Network observable can be formed from un-correlated data combinations
T. A. Prince, M. Tinto, S. L. Larson and J. W. Armstrong, *Phys. Rev. D* **66**, 122002 (2002).
K. R. Nayak, A. Pai, S. V. Dhurandhar and J-Y. Vinet, *Class. Quantum Grav.* **20**, 1217(2003).
- ★ For a source in given direction: One can get the optimal SNR same as the above case but the lower bound is zero!
K. R. Nayak, S. V. Dhurandhar, A. Pai, and J-Y. Vinet, *Phys. Rev. D* **68**, 122001 (2003).