Density Map(s) at z~2 in COSMOS/UltraVISTA

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Motivation

- ★ Measure local density for environmental studies, morphology/SFR-density etc...
- ★ My primary interest: galaxy size vs. density (Zirm+ 2012, Papovich+ 2012, S. Patel's and other talks today)
- \star Use full information provided by photo-zs
- ★ Identify clusters/groups

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Photometric redshifts

- ★ Excellent template sets for fitting (incl. linear combinations of templates)
- ★ Bayesian priors
- \star z and redshift prob. distribution: P(z)
- ★ EAZY (Brammer, van Dokkum & Coppi 2008)
- ★ Apply to the large COSMOS/UltraVISTA field uV data now public: <u>http://goo.gl/BgVUv</u>

Input Catalog

- ★ K-band selected, includes BVrizYJHK
- ★ photo-zs derived with EAZY
- ★ systematic zeropoint offsets removed iteratively
- **★** star/galaxy separation using NIR colors

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An Exercise

 ★ Make no quality cut on photo-zs, use all P(z)s => more tracers
 ★ Is the P(z) itself statistically robust?
 ★ 100 MC realizations of photo-z catalog

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Monte Carlo P(z) realizations



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Sums of P(z)

- ★ Simple sum on a (RA, Dec, z) grid
- * Naturally weights well-determined photo-zs
- ★ Could also homogenize P(z)'s to account for precision variation among galaxy types (e.g., Quadri & Williams 2010). Perhaps not needed?

Cropped Density Map (2.0 < z < 2.3)



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What's Next?

★ Compare to other density estimators

- ★ Link to lower redshift density maps (e.g., Kovacs et al. 2010, 2011)
- * Correlate densities with properties of galaxies, pairs of galaxies (PhD student Allison Man)
- ★ Identify new groups/clusters

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Talking Points

- Utilize robust photo-zs as "anchors"
 via cross-correlation, akin to using spec-zs to trace
 structures (Newman 2008, Matthews & Newman 2010)
 also, spec-zs of course...
- \star Comparison to simulations
- Alter the prior(s)? include auxiliary data not well-modeled by photo-z templates (e.g., MIPS 24um, IRAC bump) iterate in response to density map?

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