# Proto-Groups at l. $8<z<3$ in zCOSMOS-deep 

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## zCOSMOS-deep

- BzK and ugr selected + cut in $B$ and $K$ band (star-forming galaxies at $z>1$ ) (Lilly et al. 2012, in prep.)
- 70\% sampled area covering $0.6 \times 0.62 \mathrm{deg}$ Full area: $0.92 \times 0.91 \mathrm{deg}$
- 3502 objects with reliable redshifts in the range $1.8<z<3, \sim 70 \%$ success rate

cf. zCOSMOS-bright, $120 \mathrm{~km} / \mathrm{s}$, 18 '000 objects to $z \sim 1$, about 500 groups with $\mathfrak{N}>2$ (Lilly et al. 2009, Knobel et al. 2012)


## Group finder



## Group finder

- FOF-method
- Calibrated with mocks
- $\mathrm{dr}=500 \mathrm{kpc}$ and $\mathrm{dv}=700 \mathrm{~km} / \mathrm{s}$
- Requiring $N>2$



## 42 candidate groups, most with 3 members

## 12 mock catalogs resembling zCOSMOS-deep <br> (Kitzbichler \&White 2007)

$\Rightarrow$ same velocity error
$\Rightarrow$ Cut in B and K to mimic $N(z)$-distribution

## Mocks

12 mock catalogs resembling
zCOSMOS-deep
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$\Rightarrow$ same velocity error
$\Rightarrow$ Cut in B and K to mimic $N(z)$-distribution

applied group-finder \& got consistent numbers (up to 20\%)

## What are these structures in the mocks?

- Only $0.2 \%$ fully assembled at epoch of observation
- $\sim 10 \%$ at least partially assembled


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## Which of todays haloes would be detected?



## (Photo-z) galaxies associated <br> with these structures

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There is an excess of photo-z objects around the candidate groups...

# (Photo-z) galaxies associated with these structures 




There is an excess of photo-z objects around the candidate
...but there is no evidence of colour differentiation groups...

## (Photo-z) galaxies associated with these structures

- In SDSS and zCOSMOS-bright (to z~1) all of the environmental differentiation is in the satellite population (Peng et al. 2010 and 2011, Knobel et al. 20I2, Kovac et al. 2012 in prep.)
- If still true at $\mathrm{z} \sim 2$ and if the detected structures are not yet assembled, then we would not expect any colour-differentiation
- 42 (spectroscopic) proto-groups at z~2. Also confirmed by photo-z sample
- Most of them are probably not assembled yet, will however do so by $z=0$.
- With the full sample $\sim 70 \%$ of todays high mass haloes are detectable, we catalogue ~35\%

