

Planetary mass brown dwarfs in young clusters

Catarina Alves de Oliveira

Jérôme Bouvier, Estelle Moraux, Hervé Bouy, Gaspard Duchêne, Thomas Maschberger
Péter Ábráham, Gábor Marton, Christophe Pinte



European Space Agency

How do stars & brown dwarfs form?

Initial Mass Function

Wide-field IR surveys

Complete census of young clusters

Accretion

Discs

Multiplicity

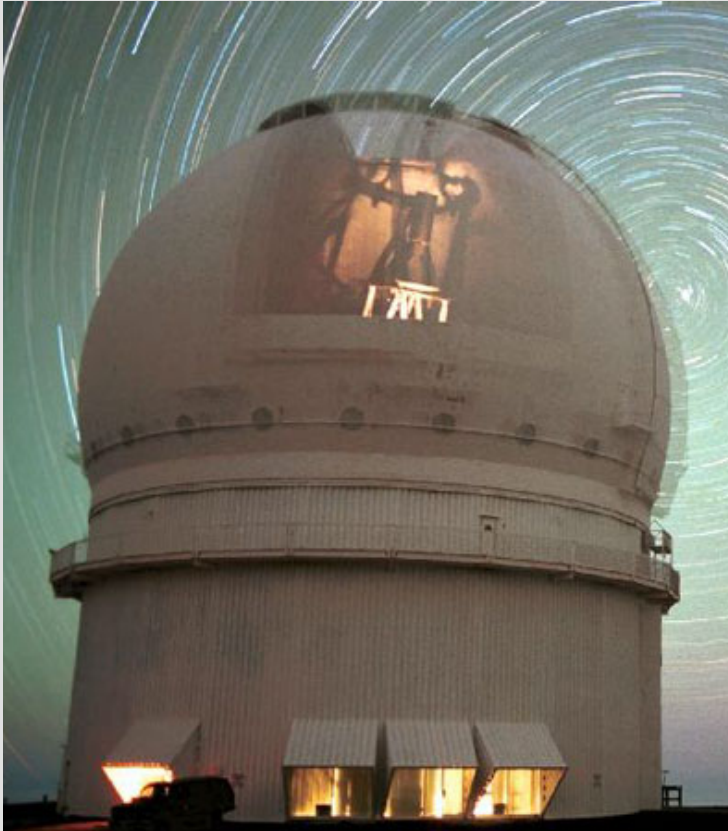
Kinematics

X-rays

Where can planets form?

CFHT survey

Science case



Credit: J.-C. Cuillandre, CFHT

CFHT, 3.6m telescope, Mauna Kea

WIRCam, near-IR camera (YJHKs)

MegaCam, optical camera (z)

Survey of nearby clusters with
deep exposures over large fields:

IC 348

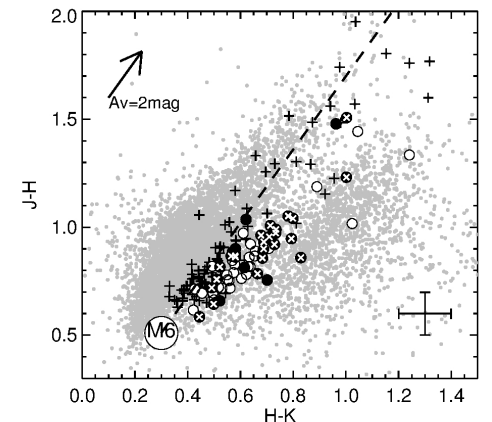
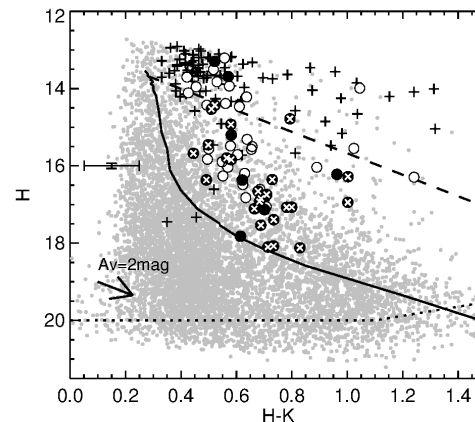
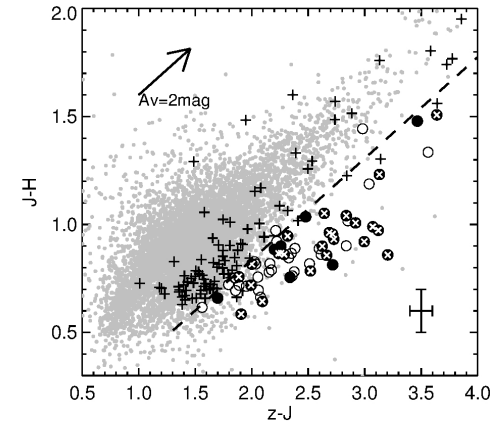
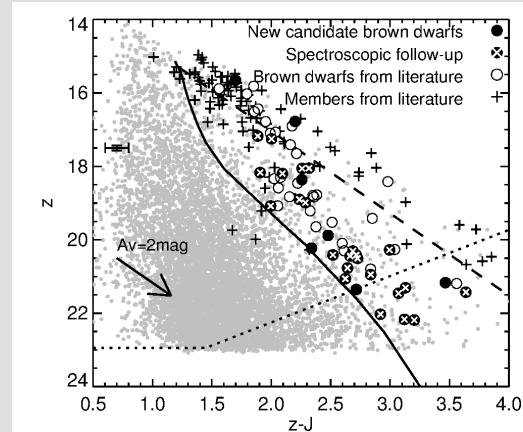
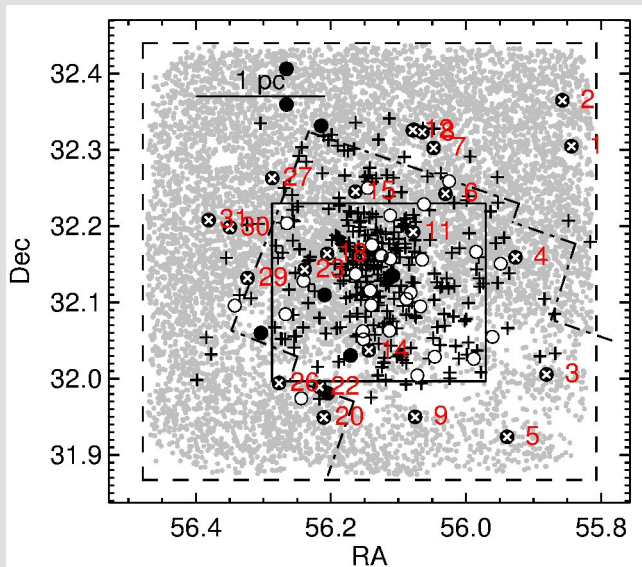
Rho Oph

CFHT survey

Observations

2 Myr
316 pc
345 members (30 brown dwarfs)

CFHT-photometric survey:
31 substellar candidates

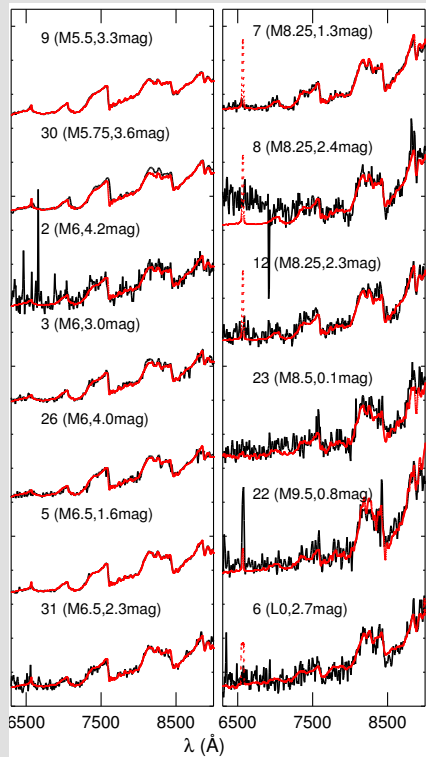


IC 348

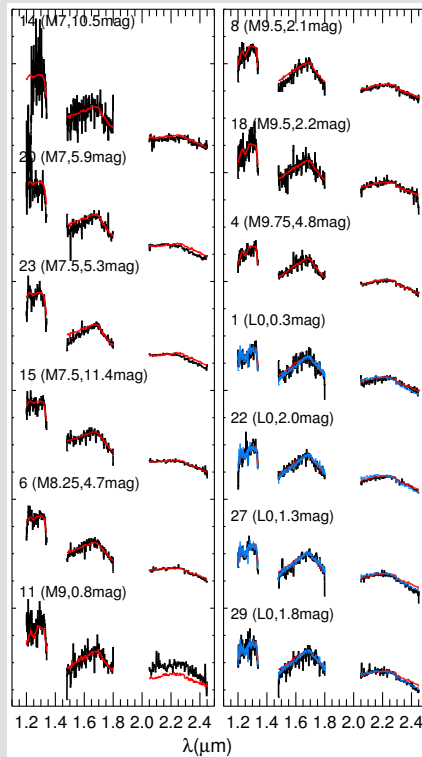
Photometry

Alves de Oliveira et al. 2013 (A&A, 549, 123A)

Osiris/GTC



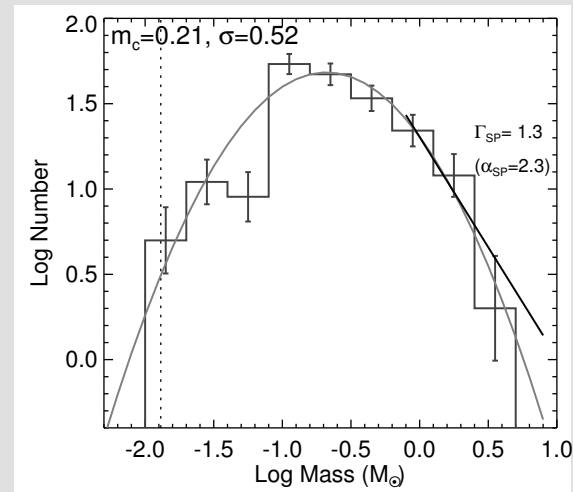
GNIRS/GEMINI



Confirmed spectroscopically 16 new members, including **13 brown dwarfs**

5 new members with L0 spectral type (according to models $\sim 13M_{Jup}$)

No indication for IMF variations

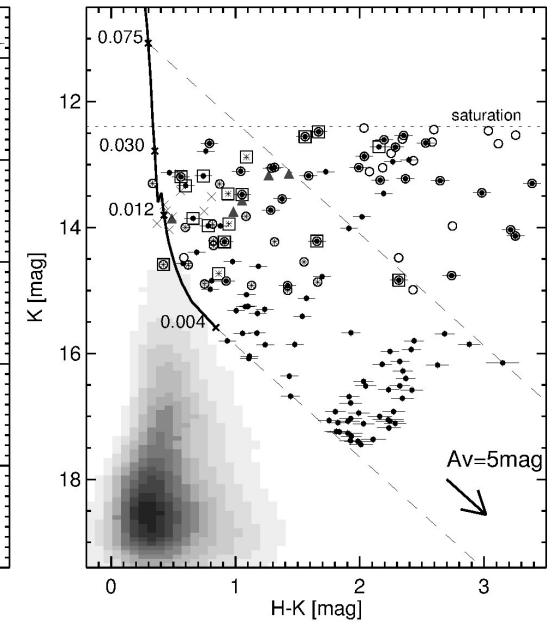
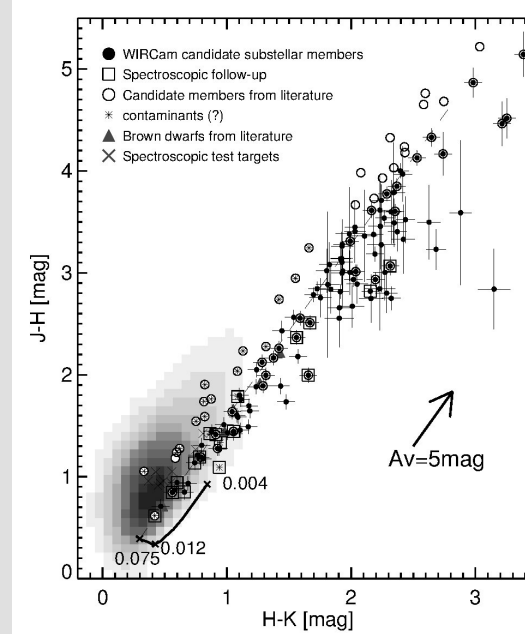
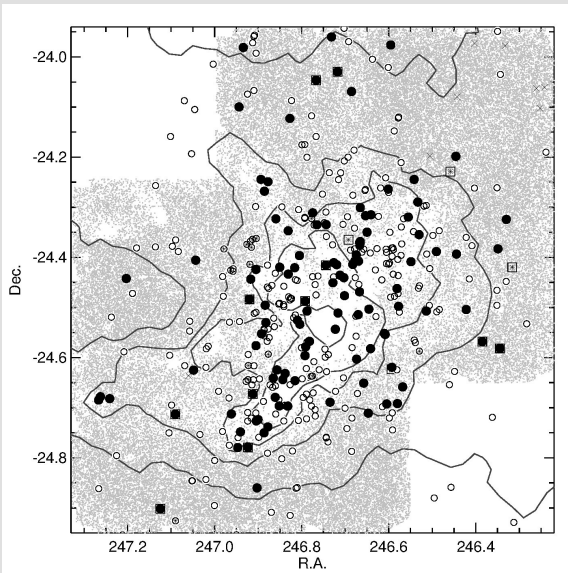


IC 348

Spectroscopy

1 Myr
120 pc
~300 members (11 brown dwarfs)

CFHT-photometric survey:
110 substellar candidates

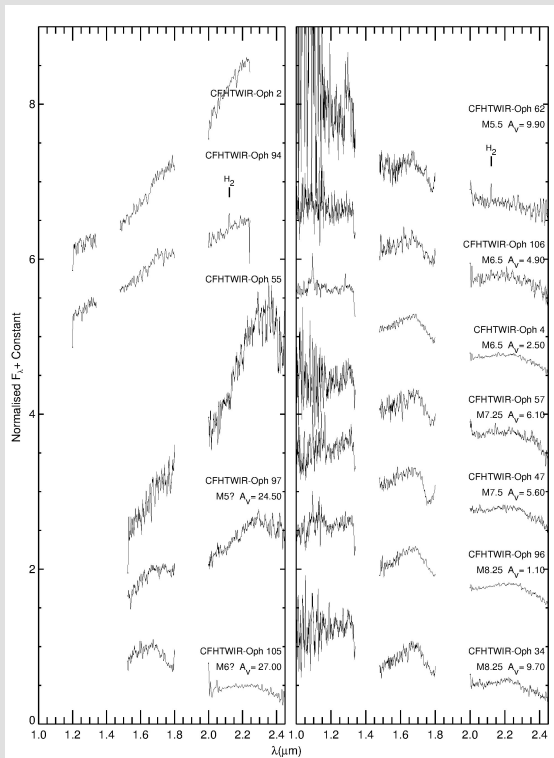


Rho Oph

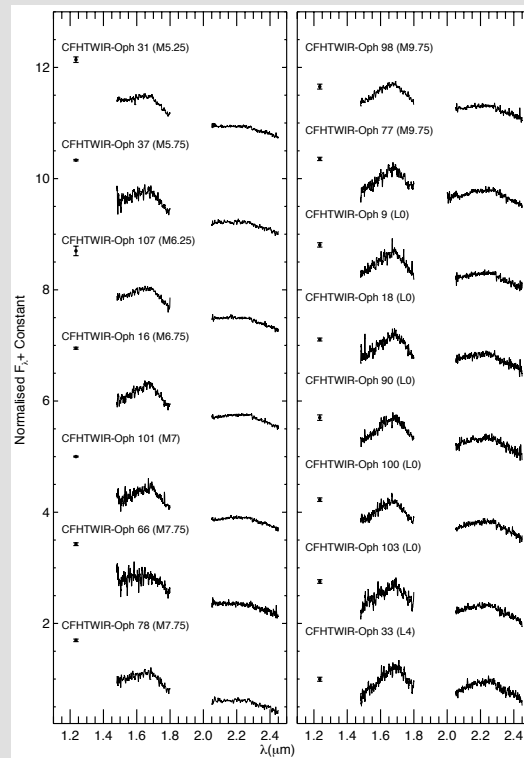
Photometry

Alves de Oliveira et al. 2010 (A&A, 515, 75A)
 Alves de Oliveira et al. 2012 (A&A, 539, 151A)

SofI / NTT



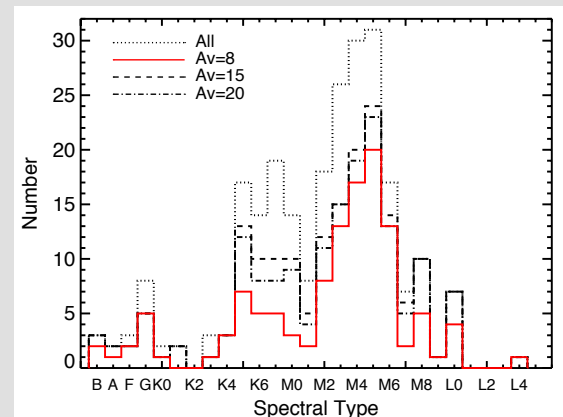
ISAAC / VLT



Confirmed spectroscopically 27 new members, including **20 brown dwarfs**

4 new L0 members and 1 L4 (according to models $\sim 5\text{-}9M_{\text{Jup}}$)

No indication for IMF variations

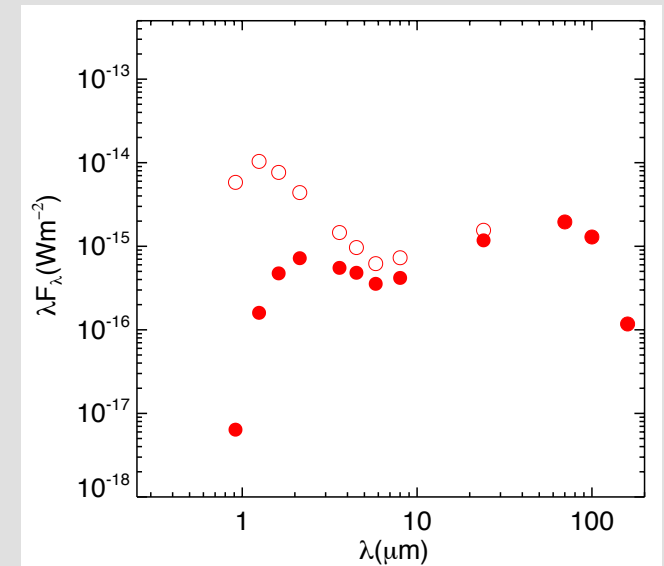
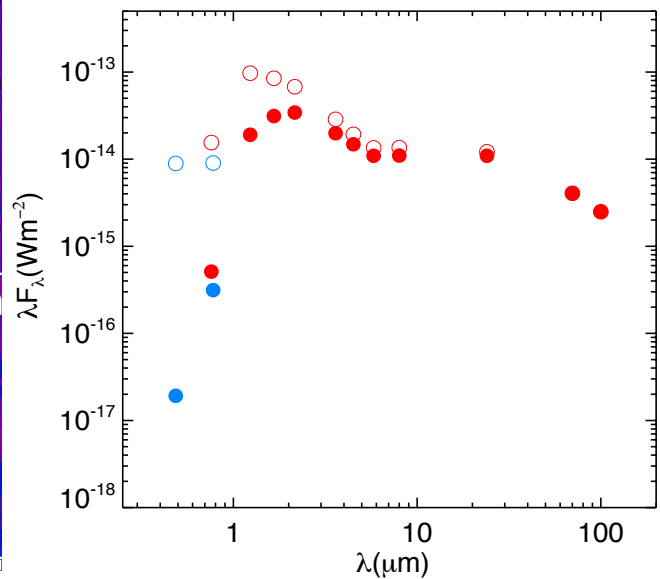
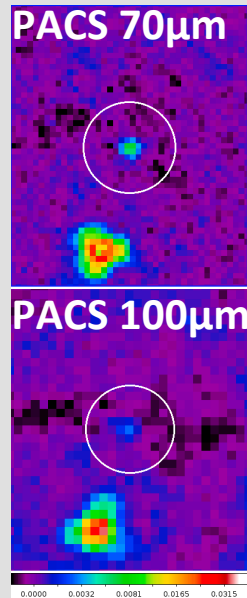


Rho Oph

Spectroscopy

Alves de Oliveira et al., *in prep*: **Herschel** survey of brown dwarf disks in Rho Oph

- Detect 12 brown dwarf disks ($M6 < \text{SpT} < M8.5$) at $70\mu\text{m}$ and $100\mu\text{m}$ (5 at $160\mu\text{m}$)



Rho Oph

Disks with *Herschel*

CFHT photometric survey of IC348 and Rho Oph uncovered several candidate brown dwarfs, and spectroscopic follow-up confirmed 13 new brown dwarfs in IC 348, and 20 in Rho Oph

Discovery of first L0 dwarfs in IC 348 ($\sim 13M_{Jup}$) and Rho Oph (4 L0, 1 L4, 4-9 M_{Jup})

No evidence for variation of the mass function in both clusters (down to $\sim 13M_{Jup}$ in IC348, and $\sim 4M_{Jup}$ in Rho Oph)

New census represents fundamental step for further studying properties of young brown dwarfs, e.g. disks in Rho Oph

Conclusions
