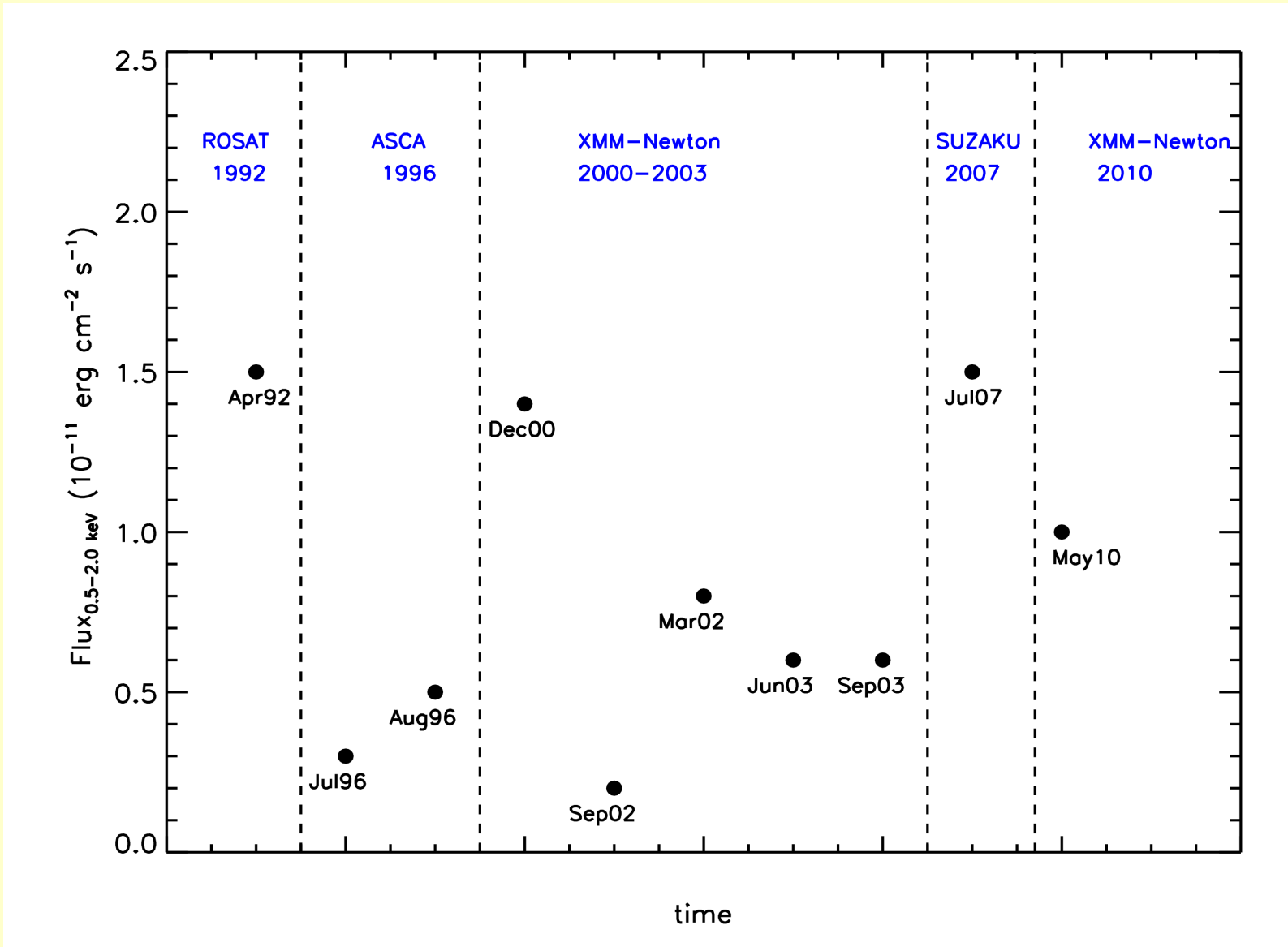
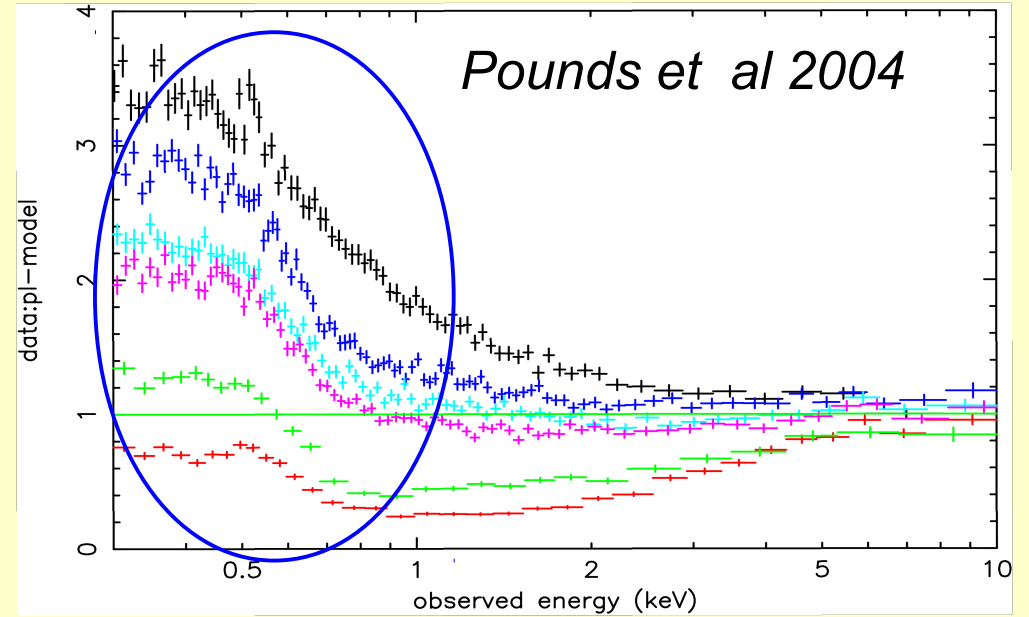
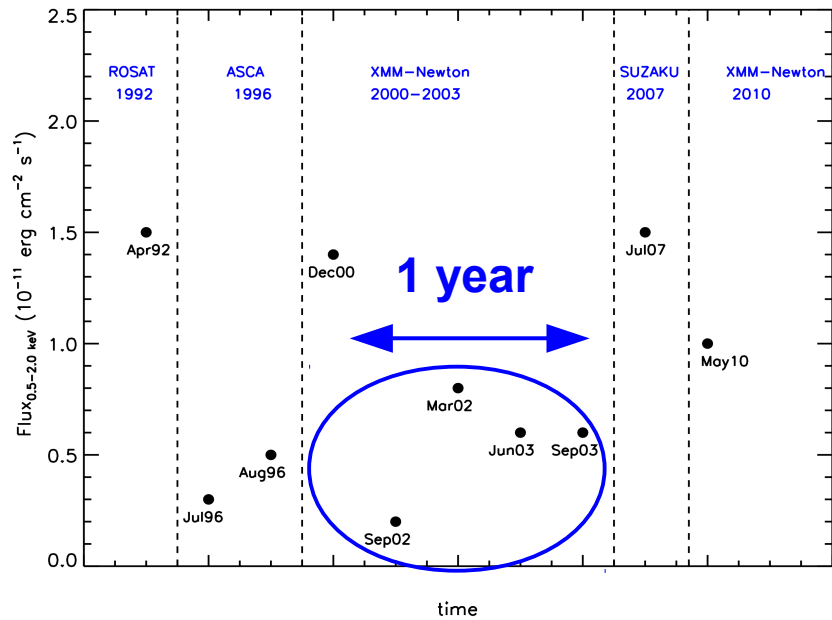


***X-ray/UV observation
of
1H0419-577***

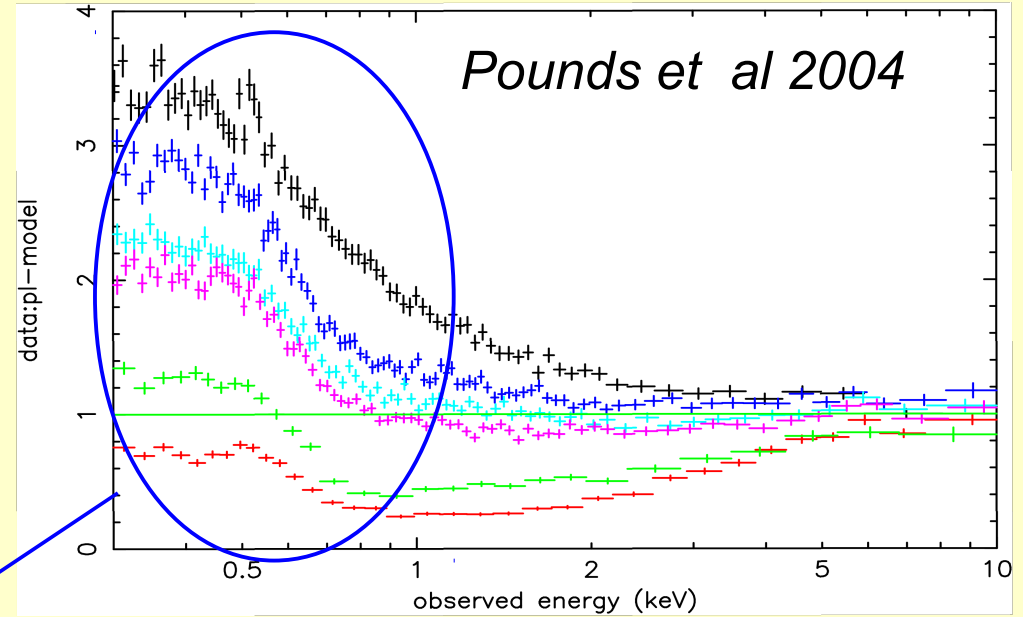
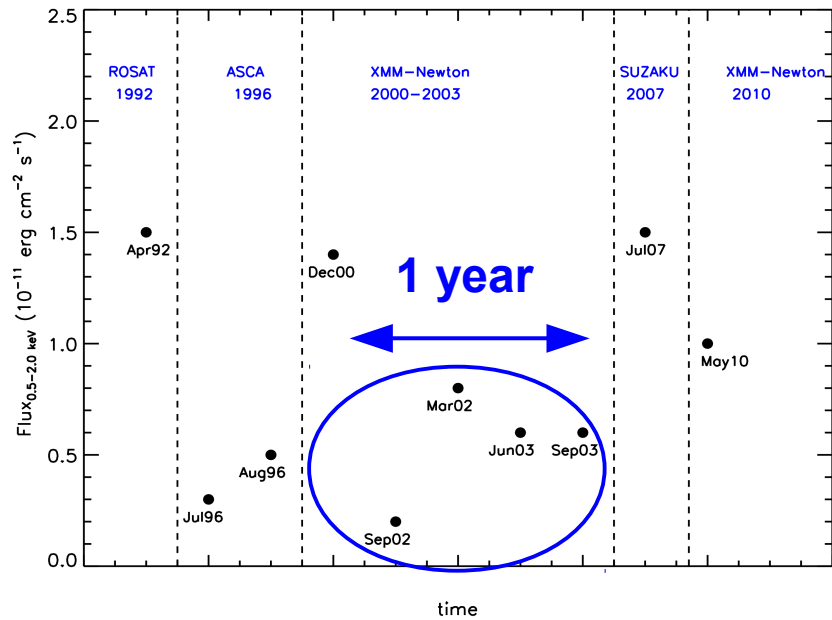
1H0419-577: a peculiar Seyfert 1



The X-ray variability

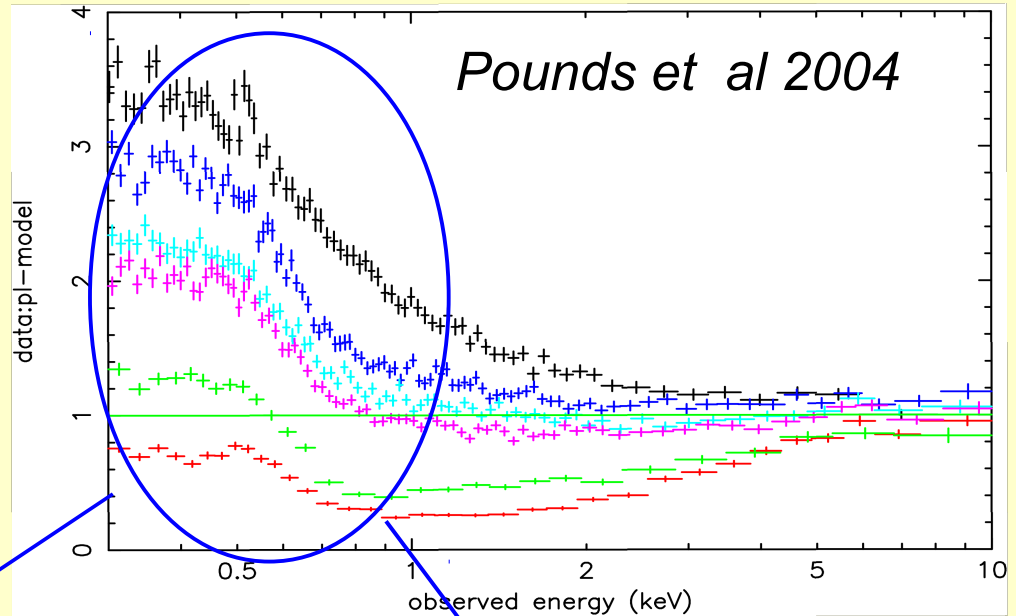
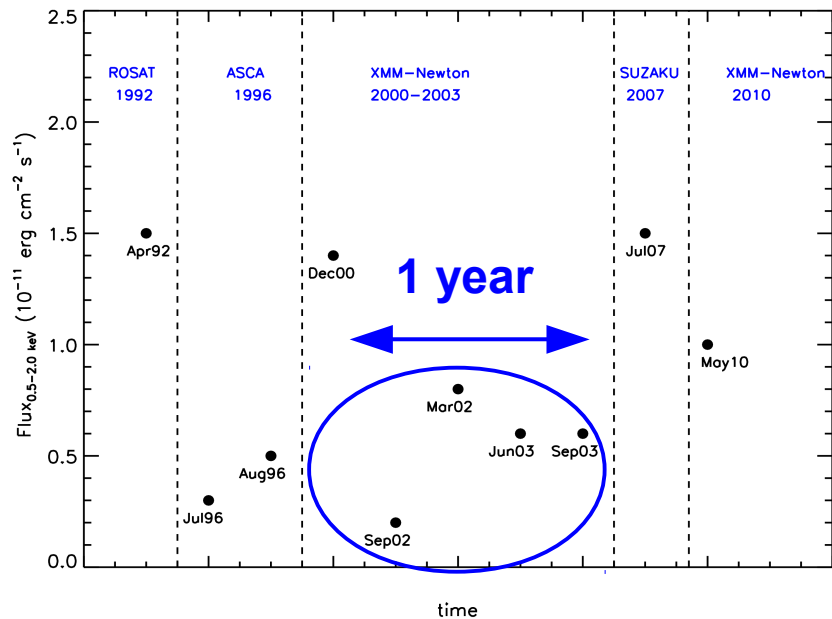


The X-ray variability



**Absorption based models
(e.g. Pounds 2004,
Turner 2009)**

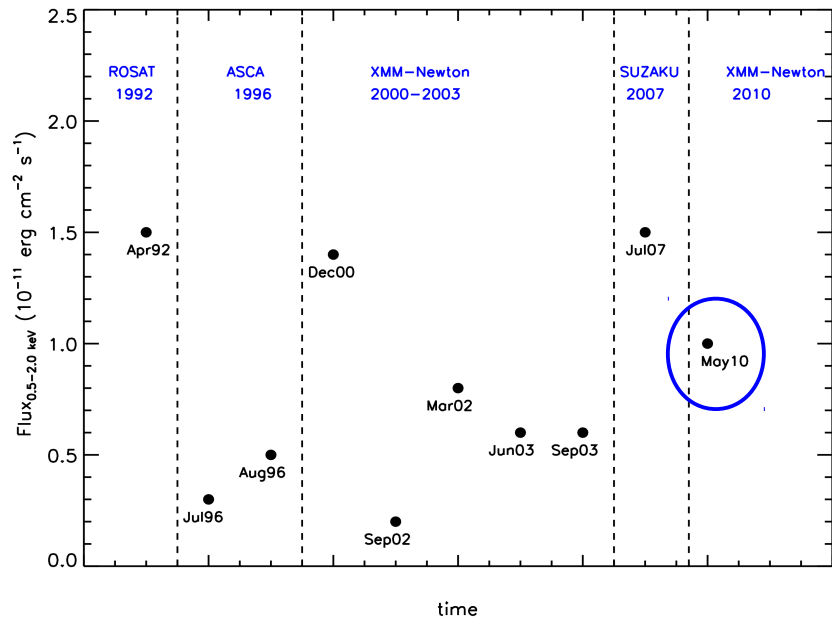
The X-ray variability



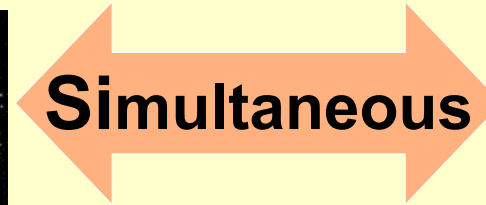
**Absorption based models
(e.g. Pounds 2004,
Turner 2009)**

**Reflection model
(e.g. Fabian 2005)**

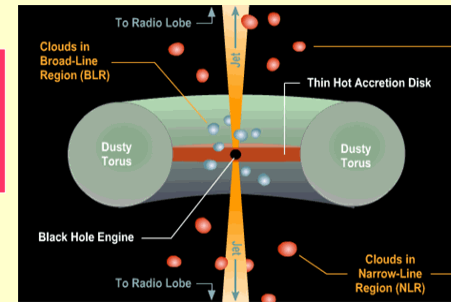
An X-ray/UV campaign



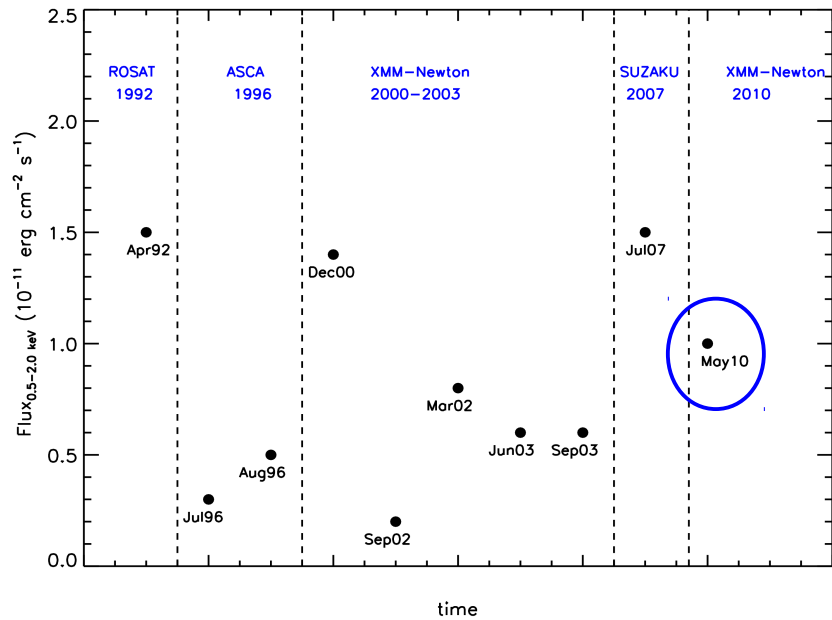
**XMM-Newton
167 ks**



**HST/COS
19 ks**



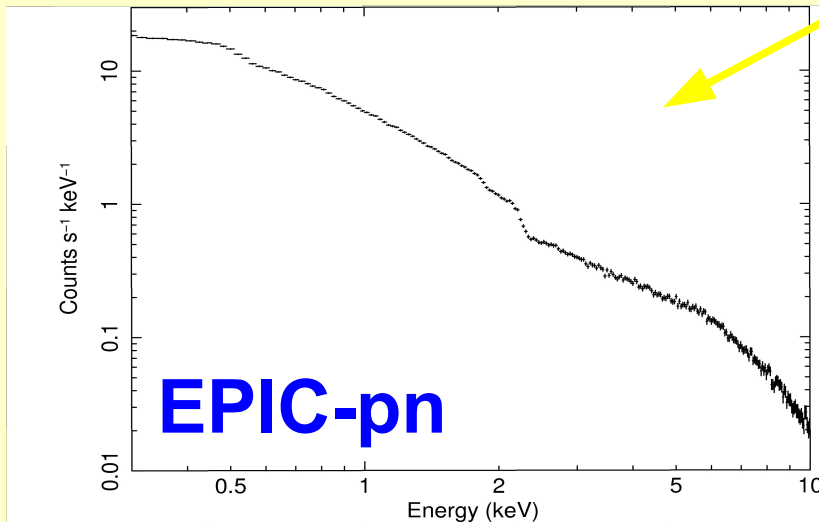
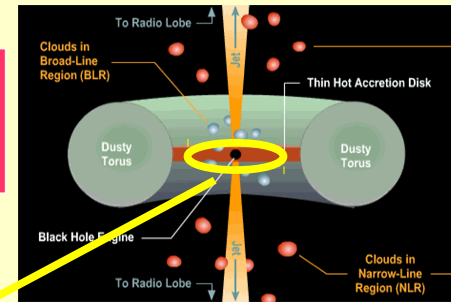
An X-ray/UV campaign

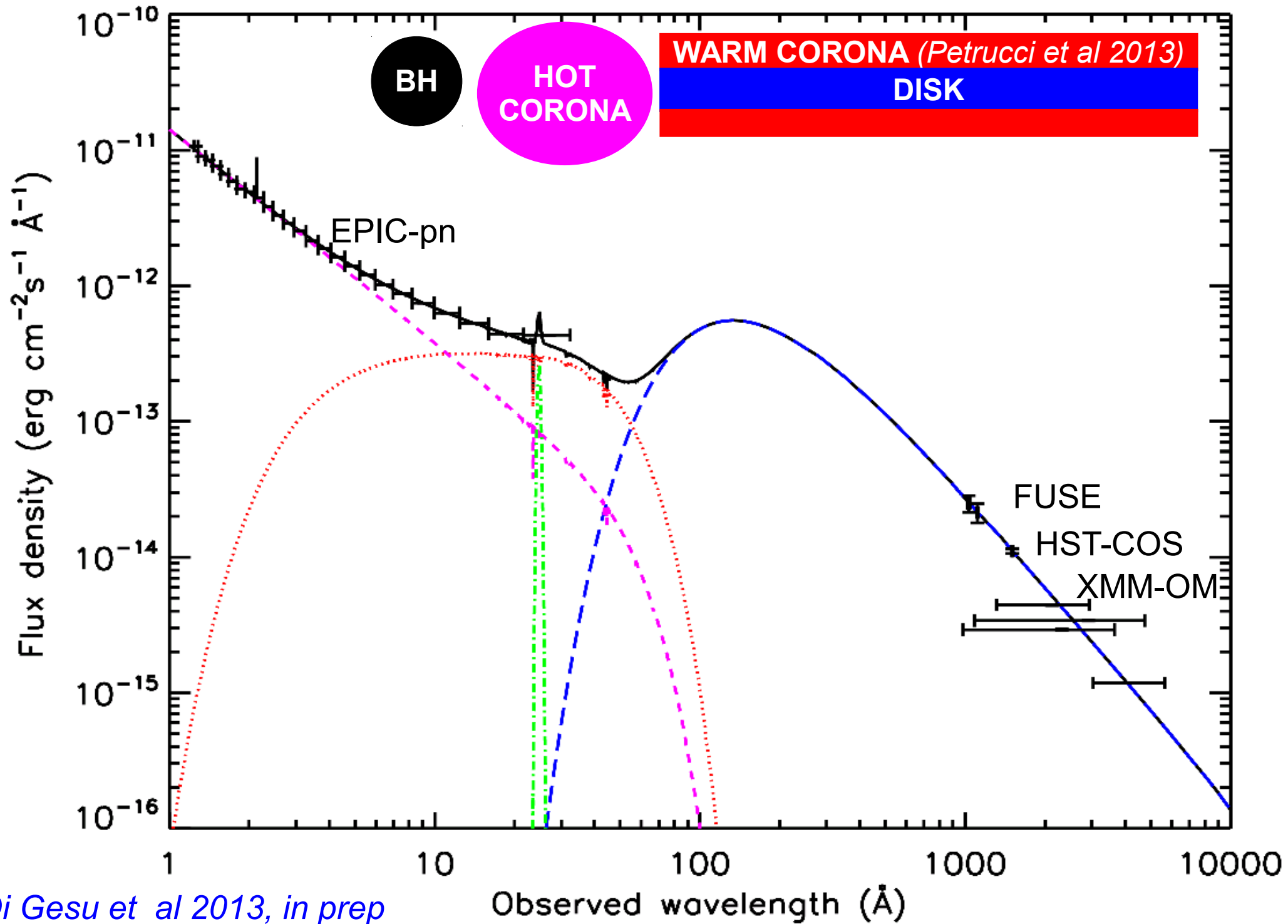


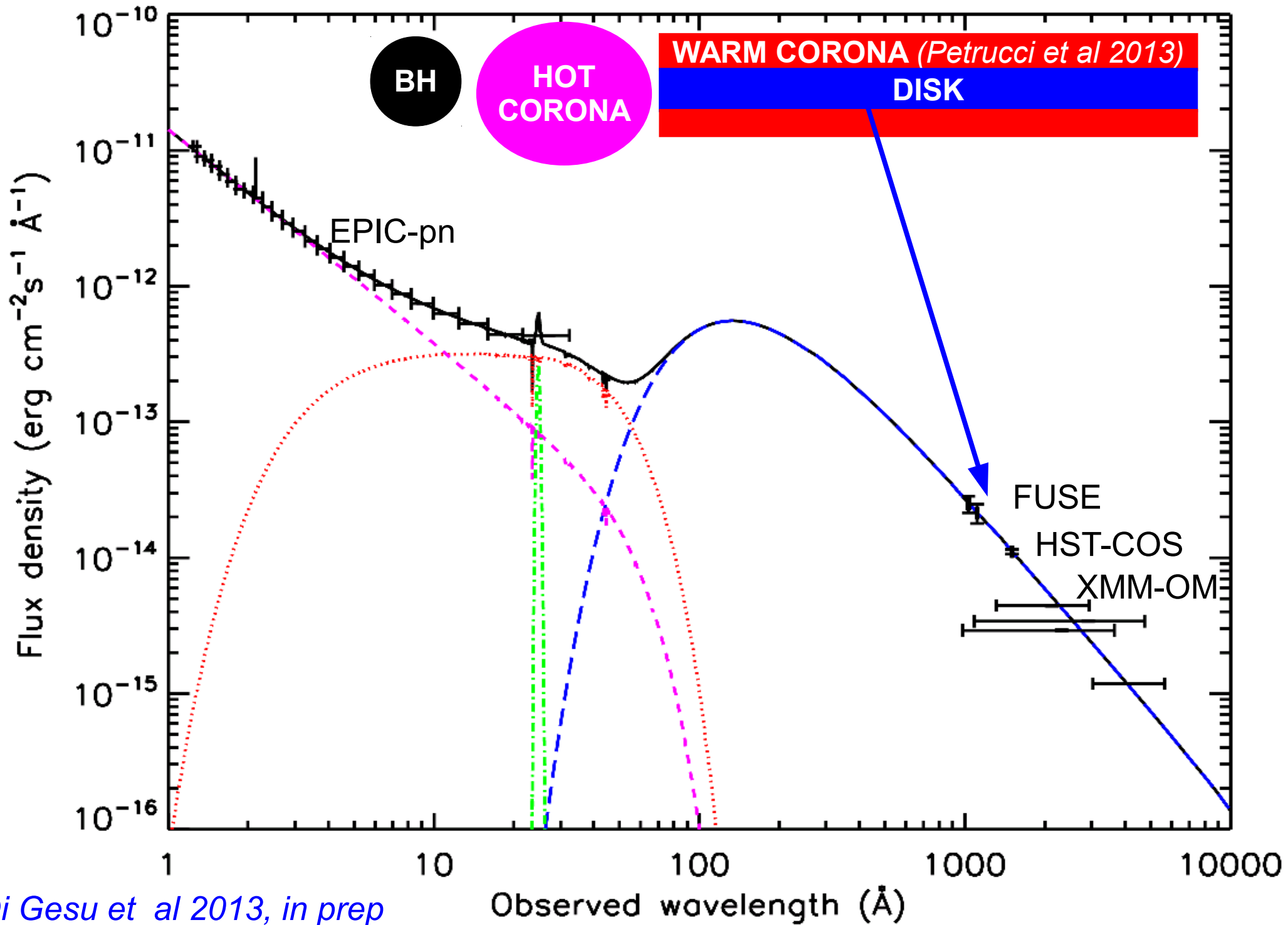
**XMM-Newton
167 ks**

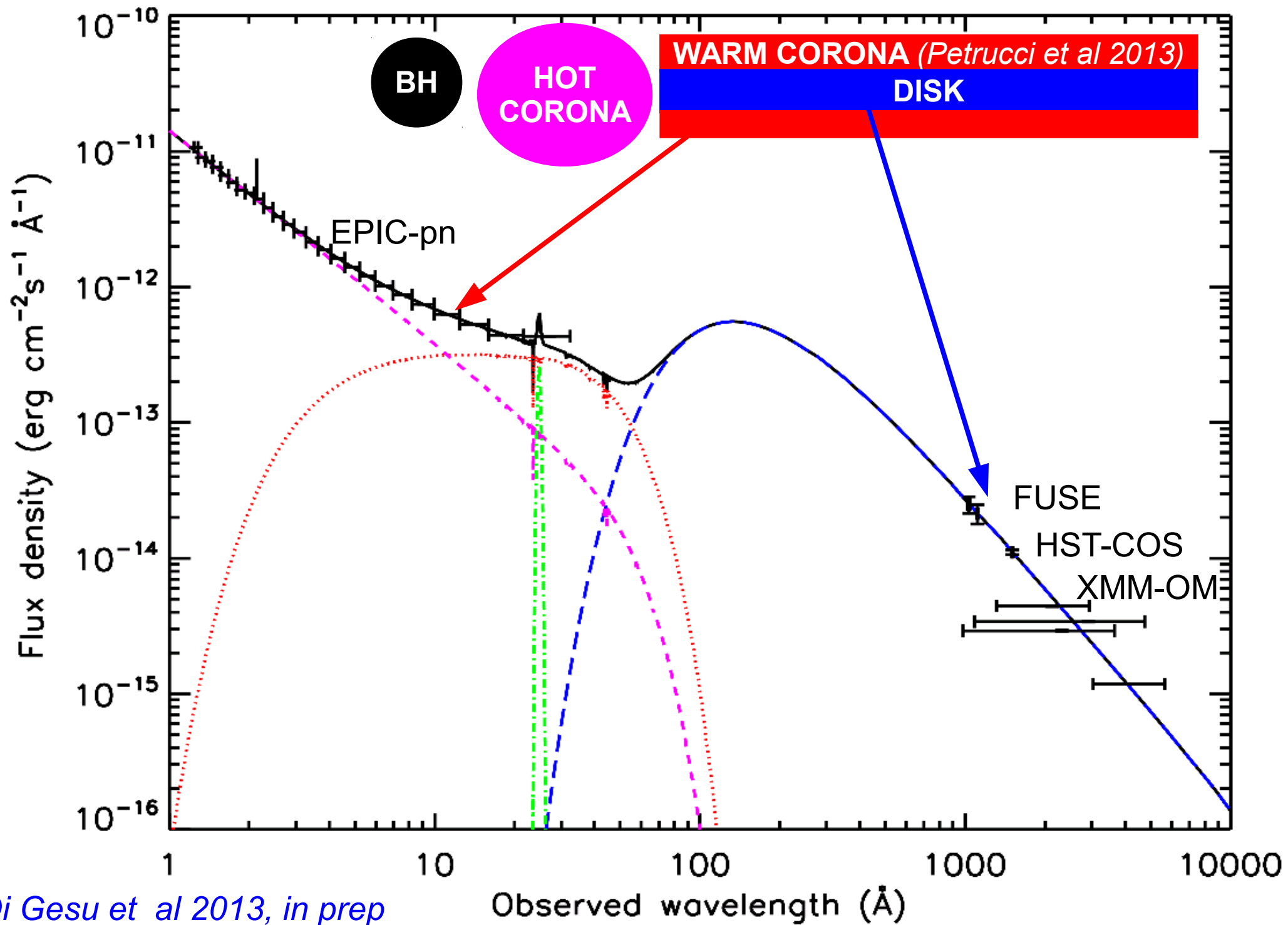


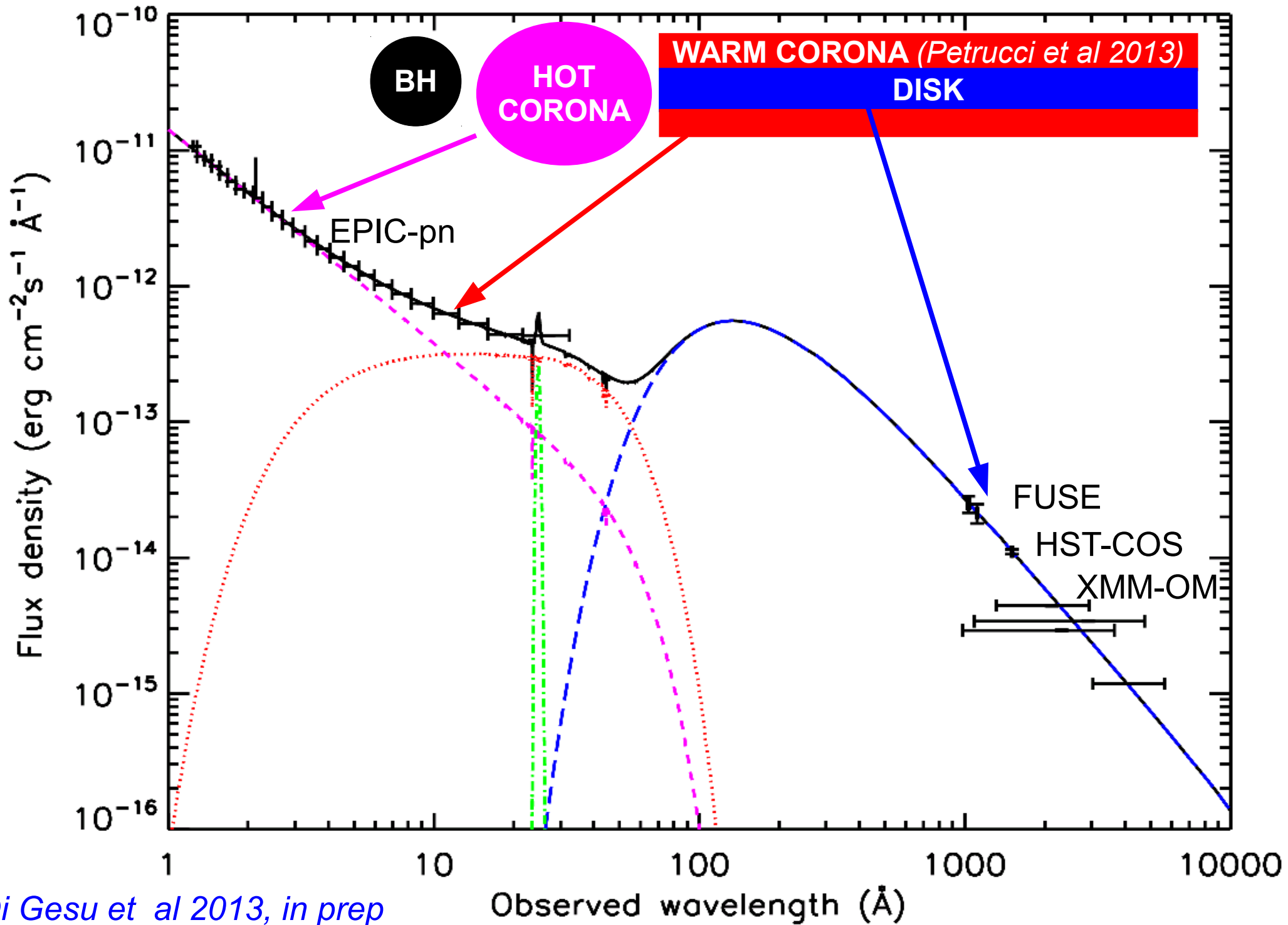
**HST/COS
19 ks**



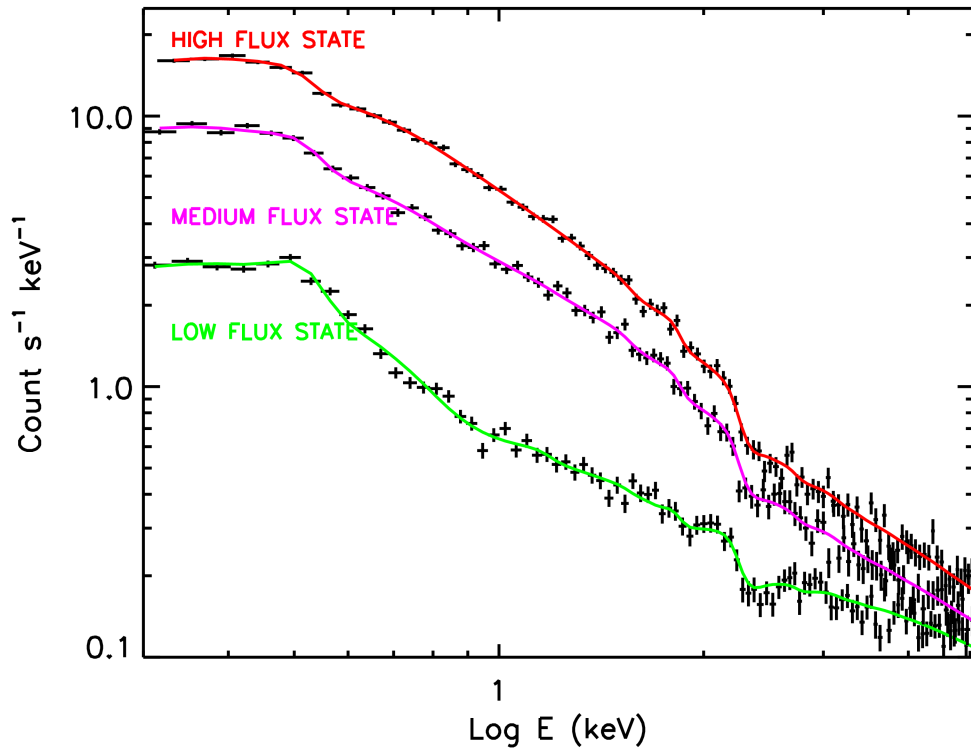






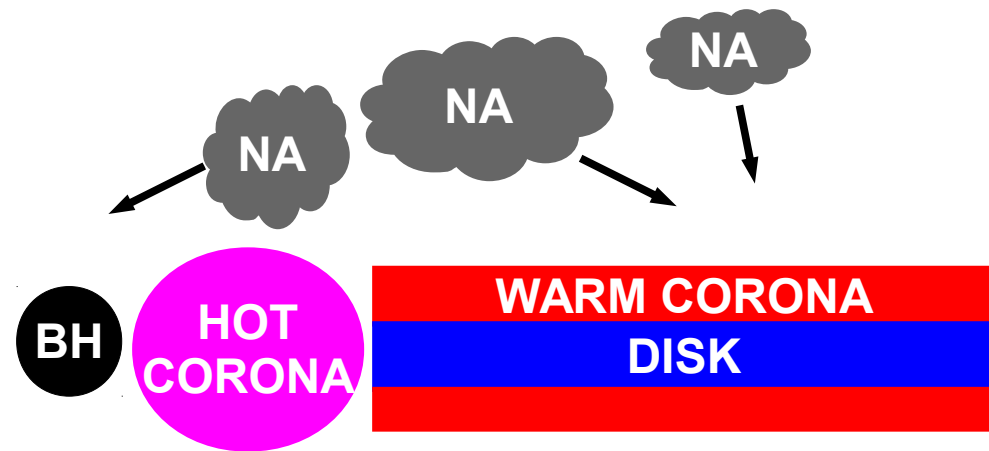


Fitting the variability:

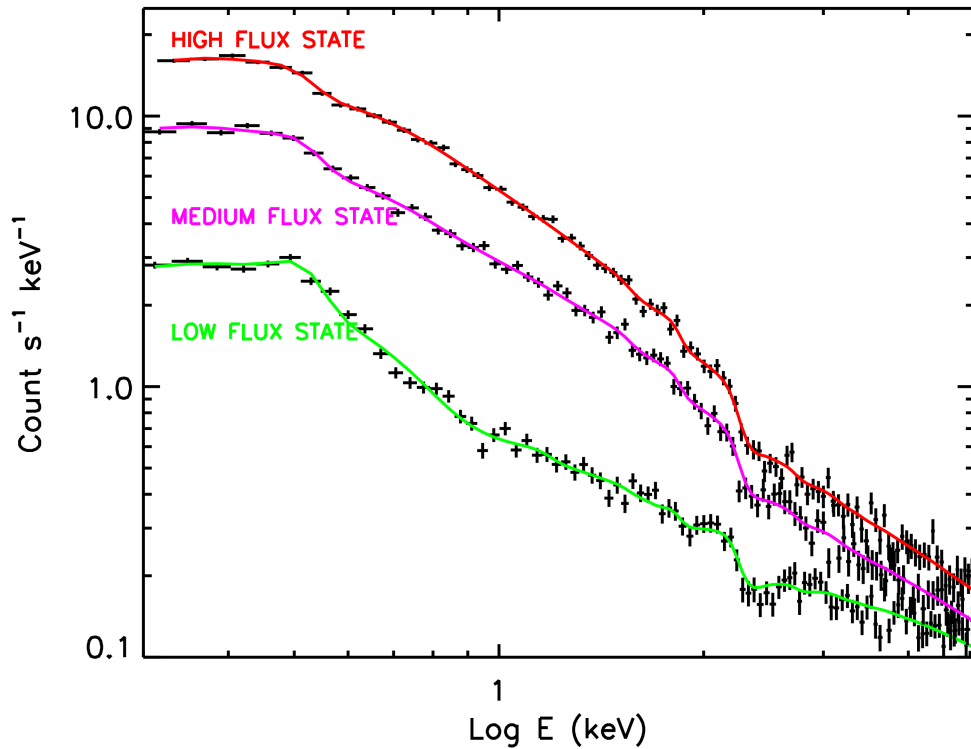


HIGH STATE: $\text{Log}N_{\text{H}} \leq 20 \text{ cm}^2$
MED STATE: $\text{Log}N_{\text{H}} \approx 22.0; C_{\text{V}} \approx 49\%$

LOW STATE: $\text{Log}N_{\text{H}} \approx 22.3 C_{\text{V}} \approx 94$

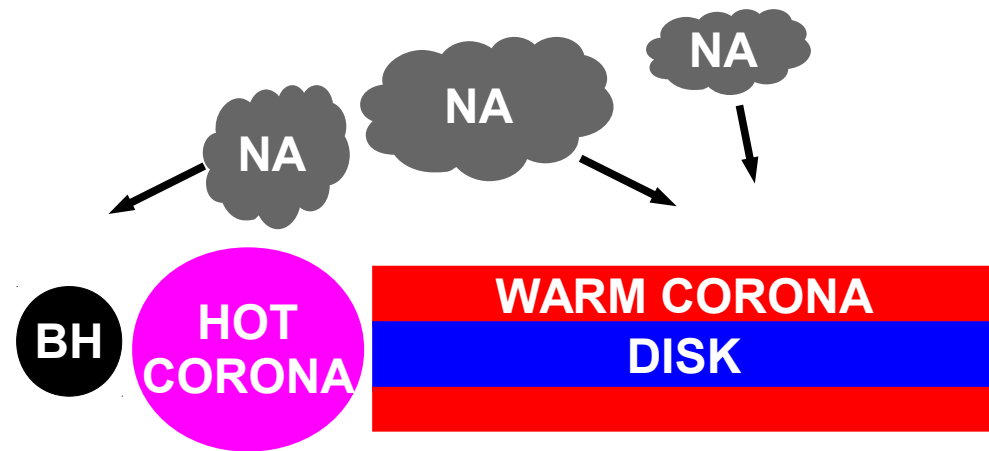


Fitting the variability:



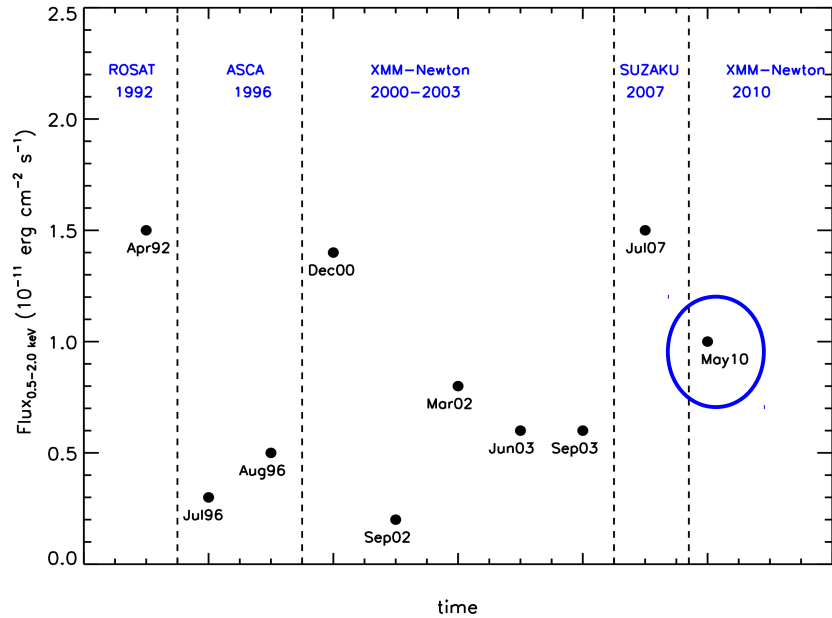
HIGH STATE: $\text{Log}N_{\text{H}} \leq 20 \text{ cm}^2$
MED STATE: $\text{Log}N_{\text{H}} \approx 22.0; C_{\text{V}} \approx 49\%$

LOW STATE: $\text{Log}N_{\text{H}} \approx 22.3 C_{\text{V}} \approx 94$

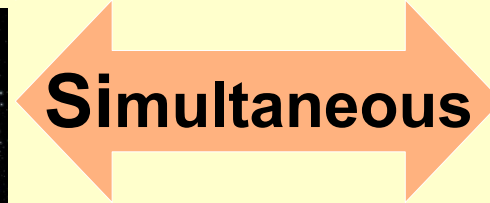


→ The obscuration explains almost all the variability.

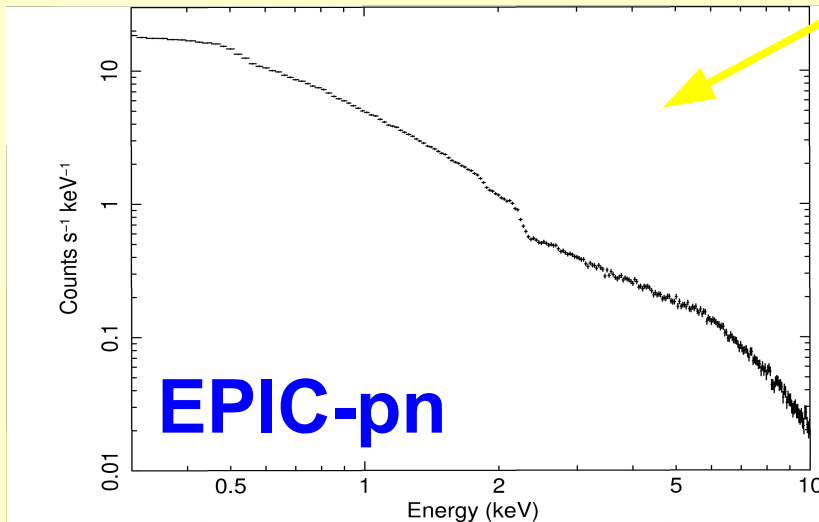
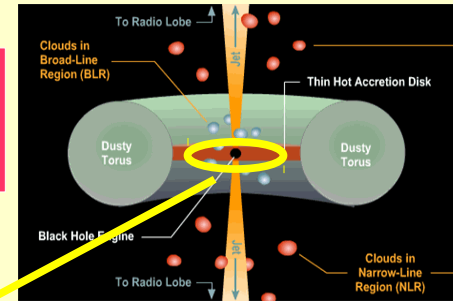
An X-ray UV campaign



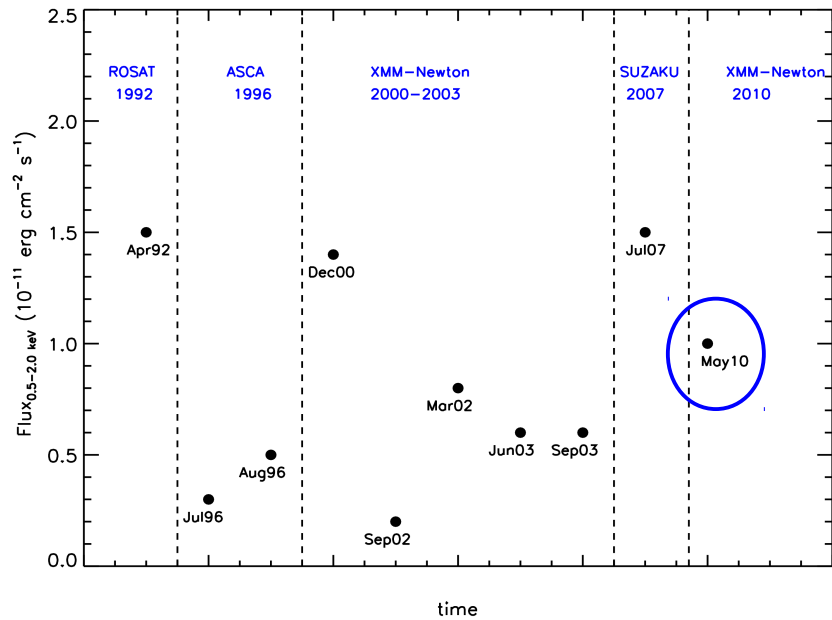
**XMM-Newton
167 ks**



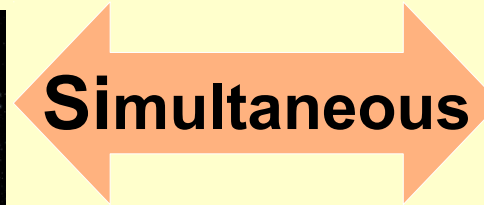
**HST/COS
19 ks**



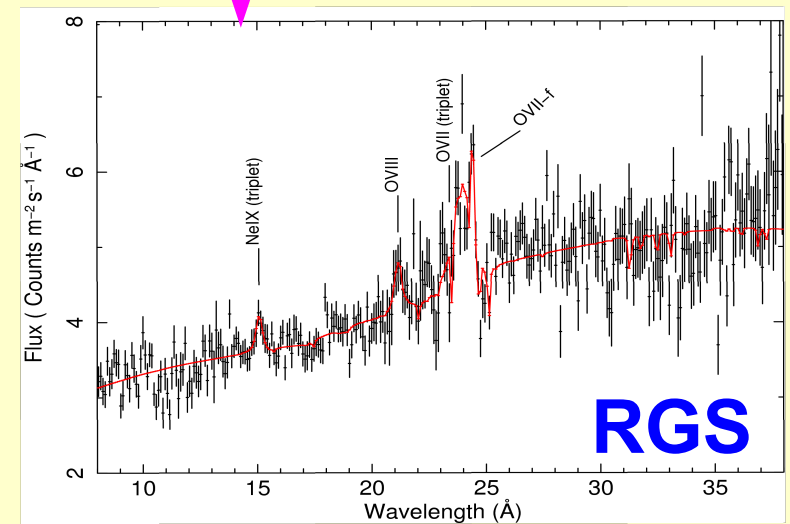
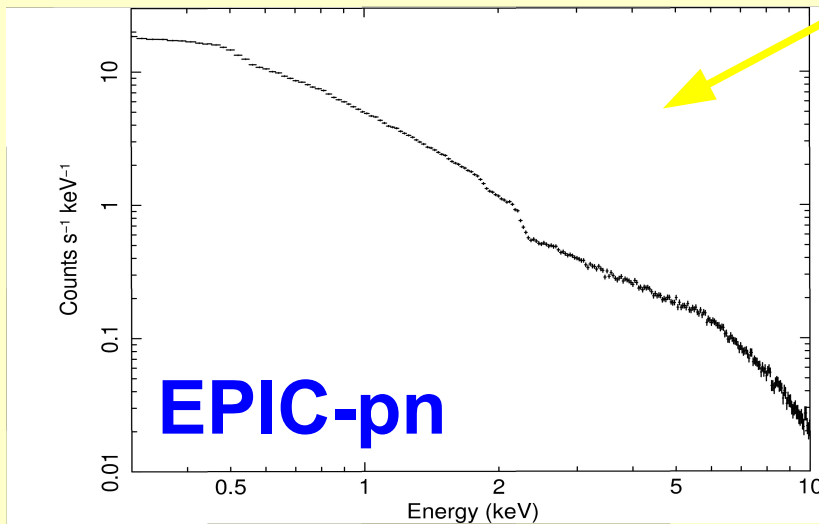
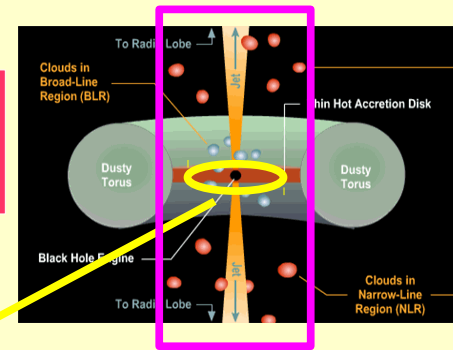
An X-ray UV campaign



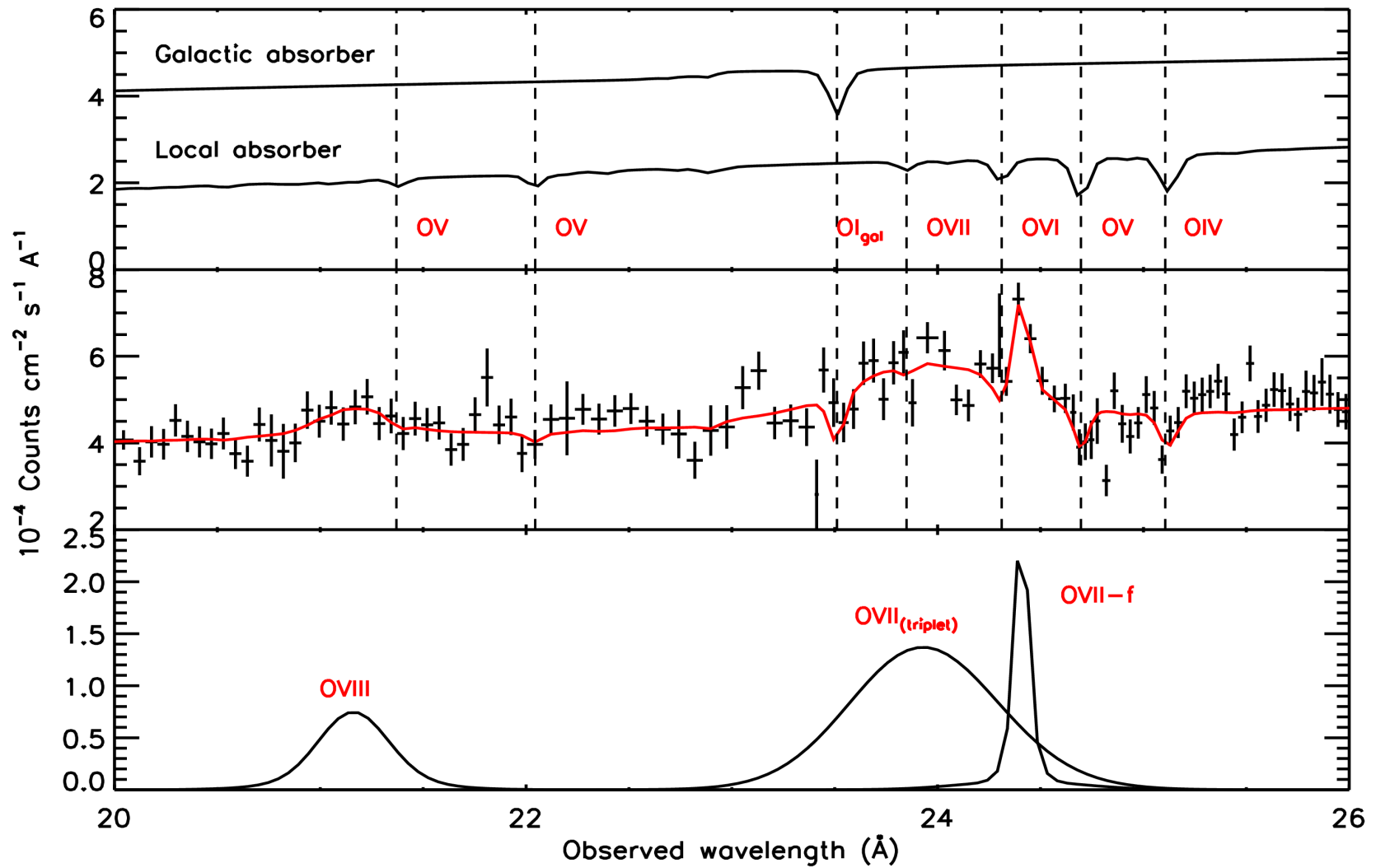
**XMM-Newton
167 ks**



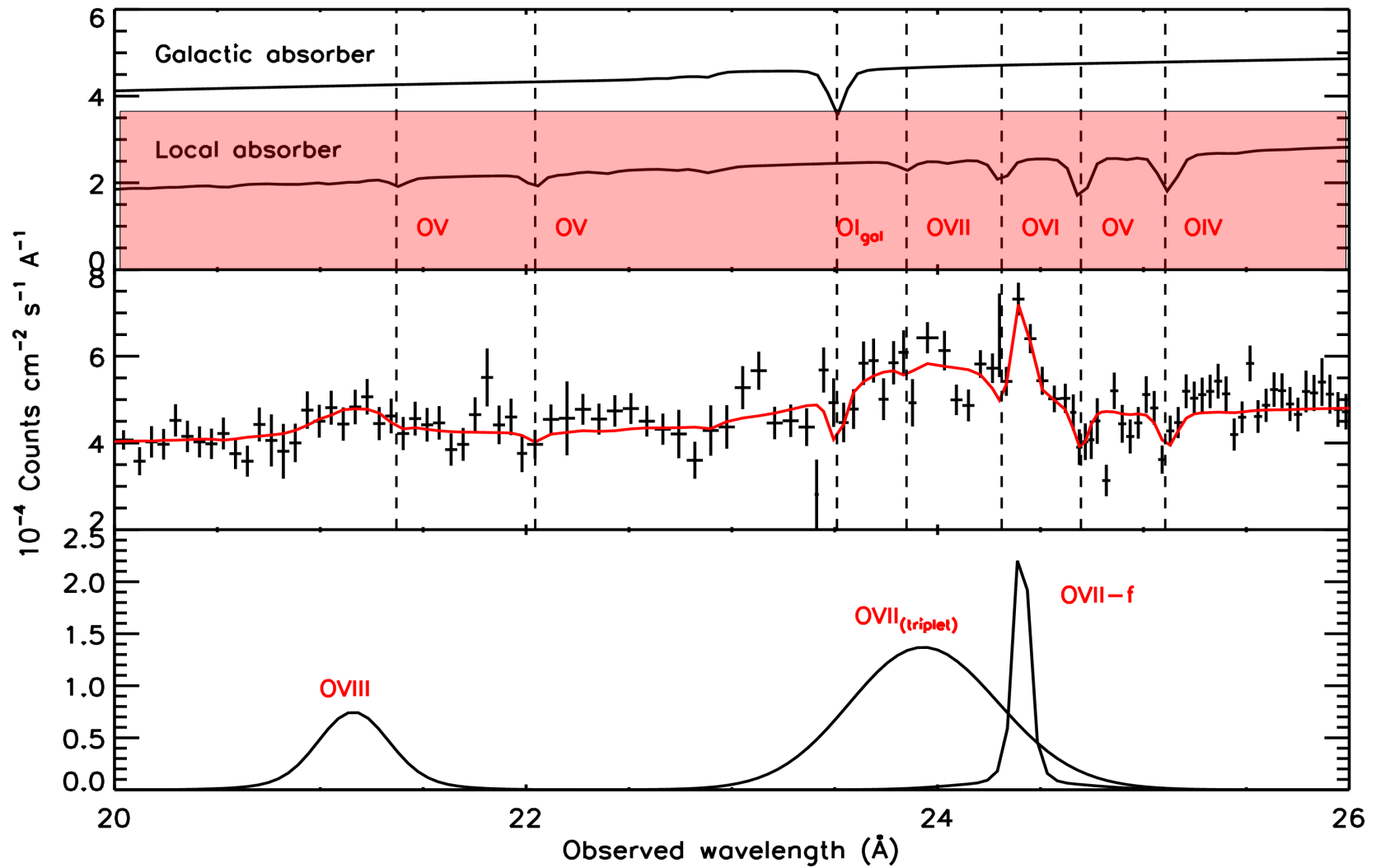
**HST/COS
19 ks**



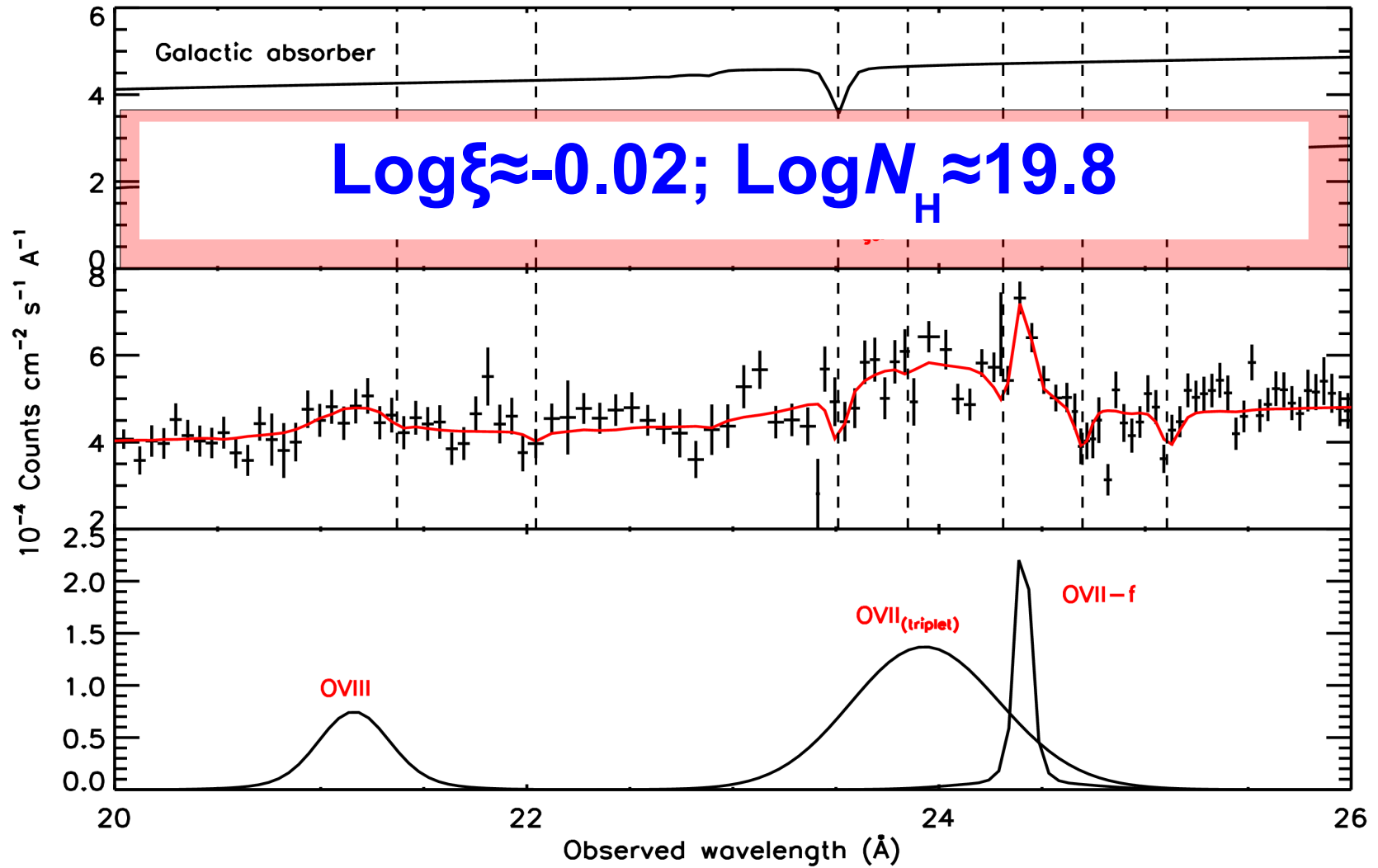
The RGS spectrum



The RGS spectrum



The RGS spectrum



The absorbing system

X-rays

$$n_{\text{H}} < 25 \text{ cm}^{-3}$$
$$\text{Log } N_{\text{H}} \approx 19.8$$
$$\text{Log } \xi \approx -0.02$$

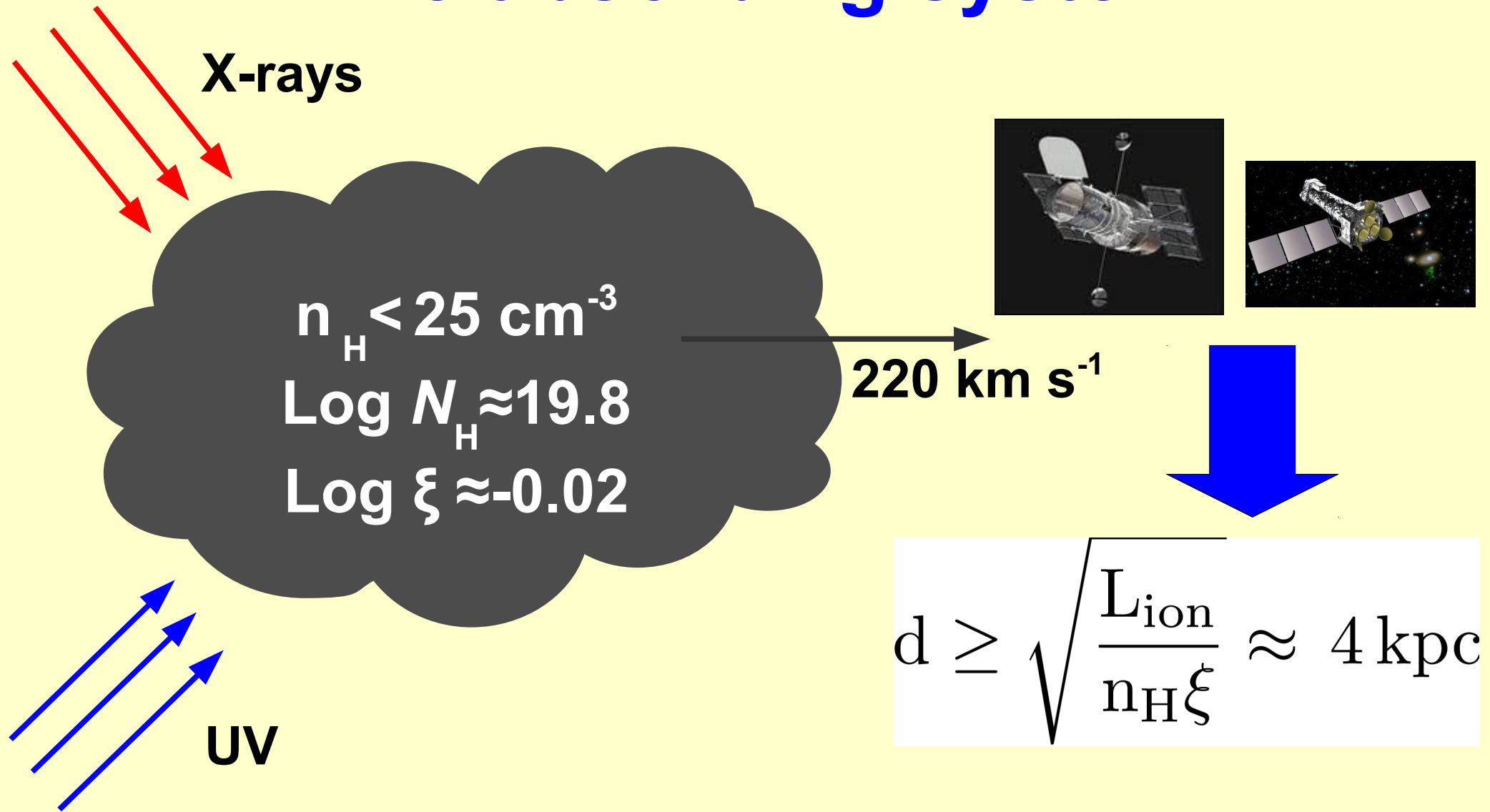
220 km s⁻¹



UV

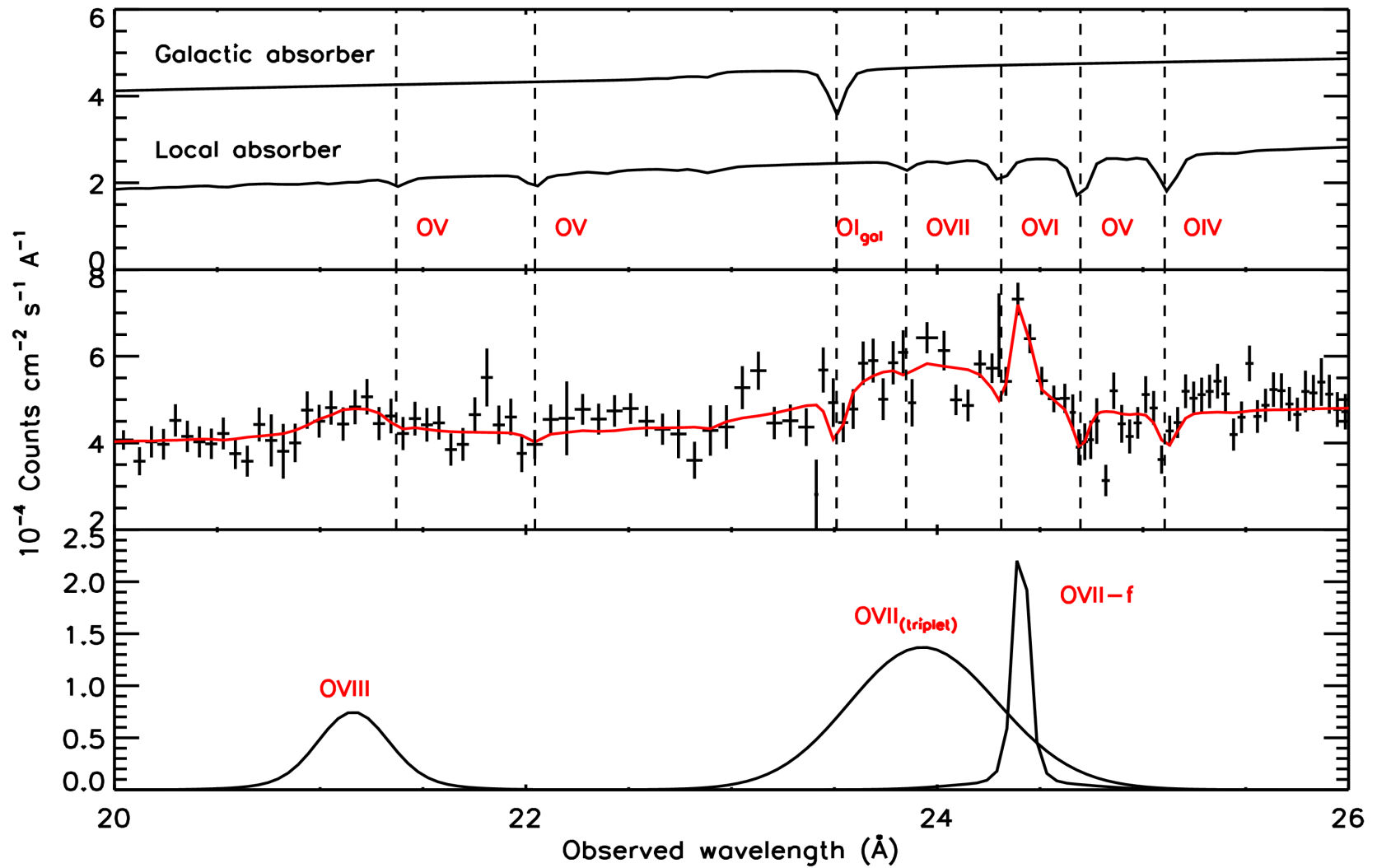
$$d \geq \sqrt{\frac{L_{\text{ion}}}{n_{\text{H}} \xi}} \approx 4 \text{ kpc}$$

The absorbing system

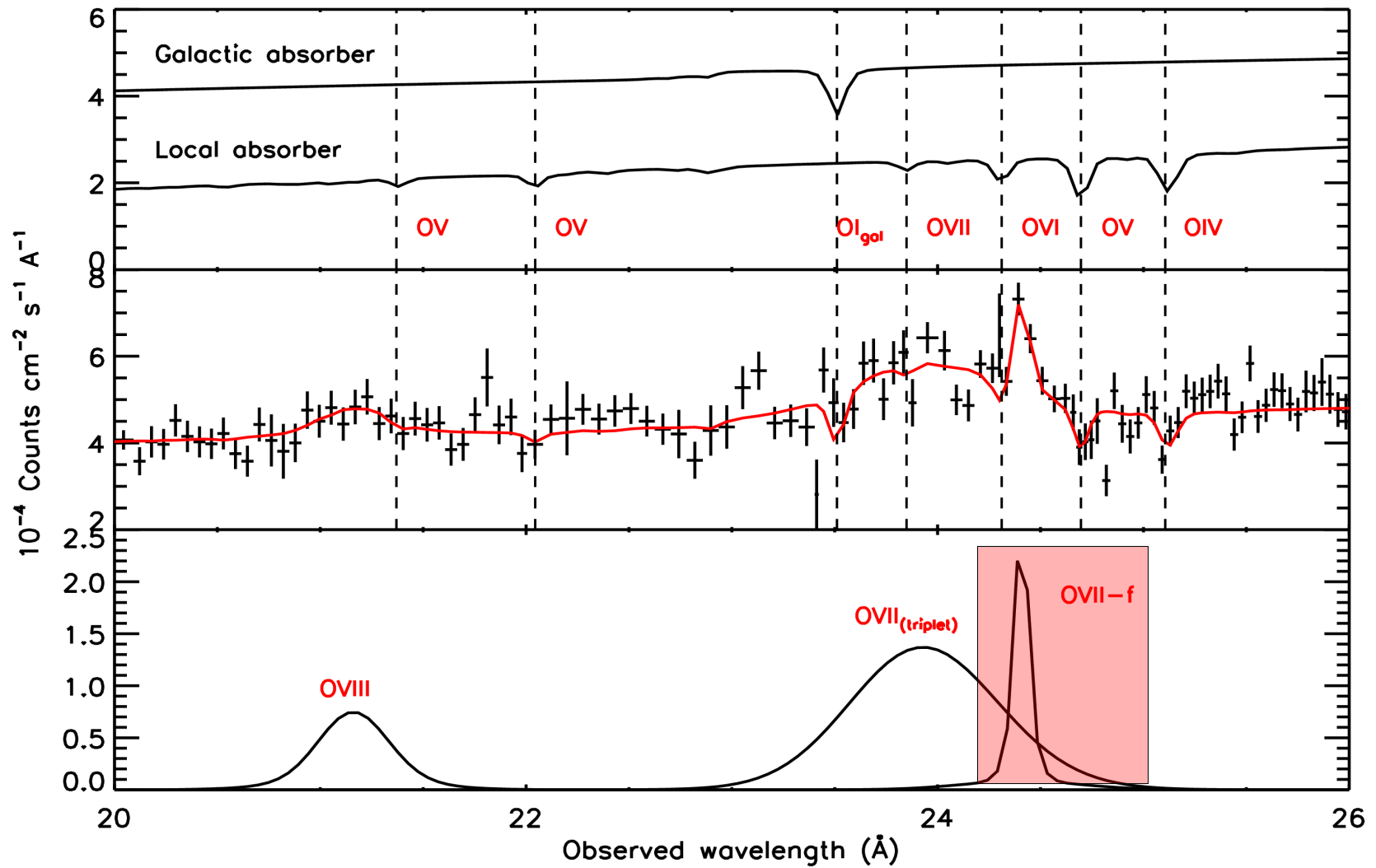


→ First example of a galactic scale X-ray absorber.

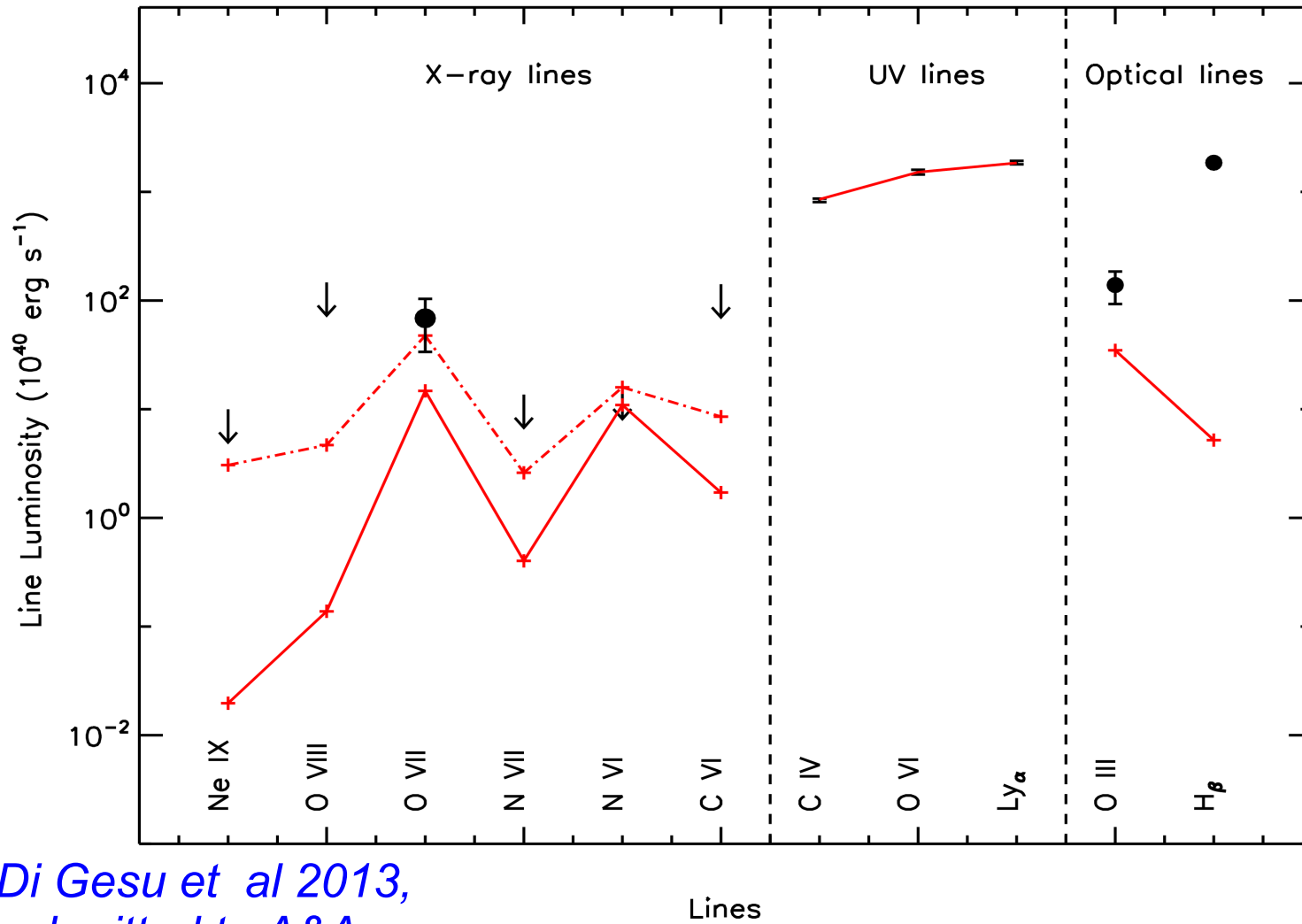
The RGS spectrum



The RGS spectrum

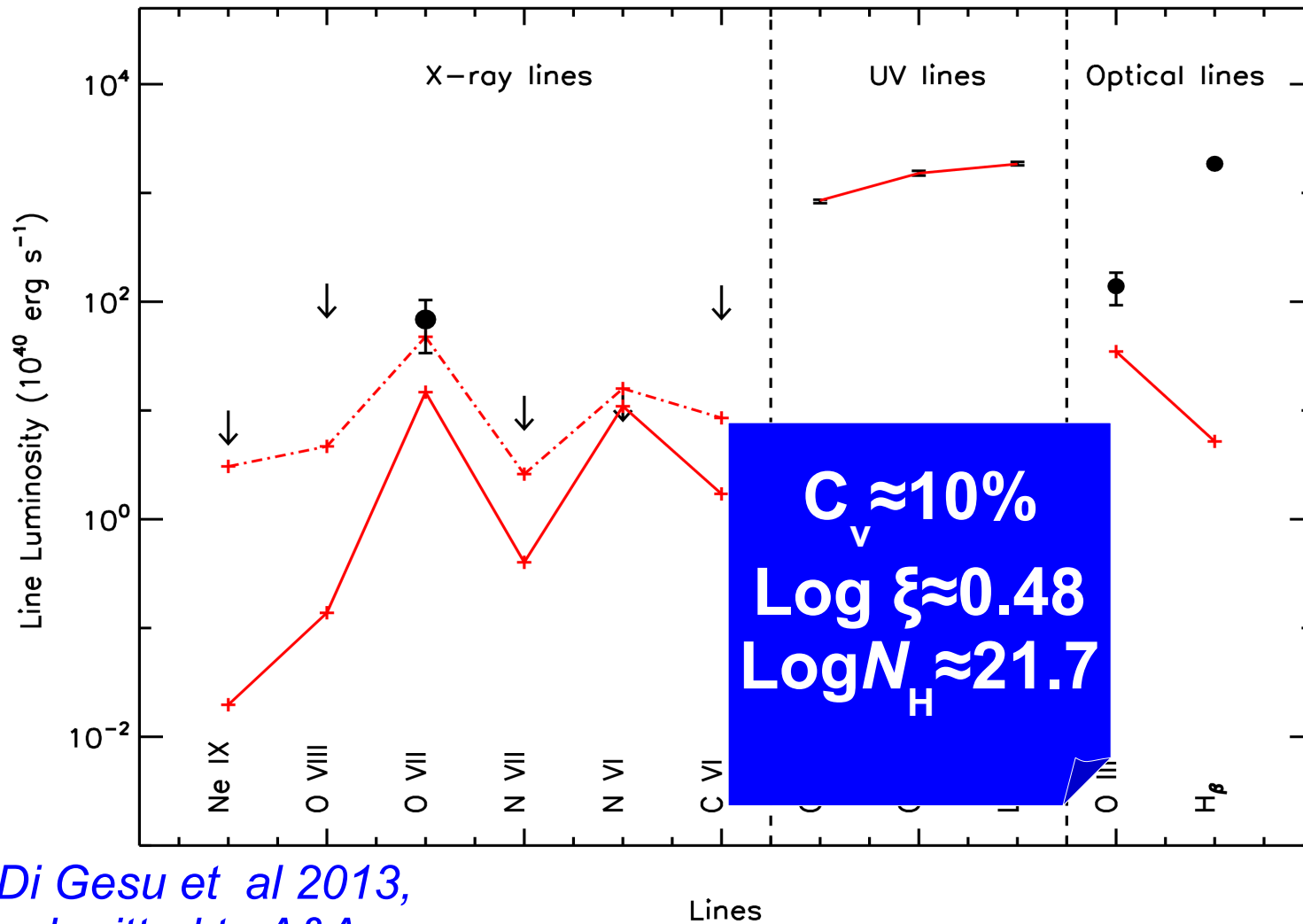


The NL emitting system



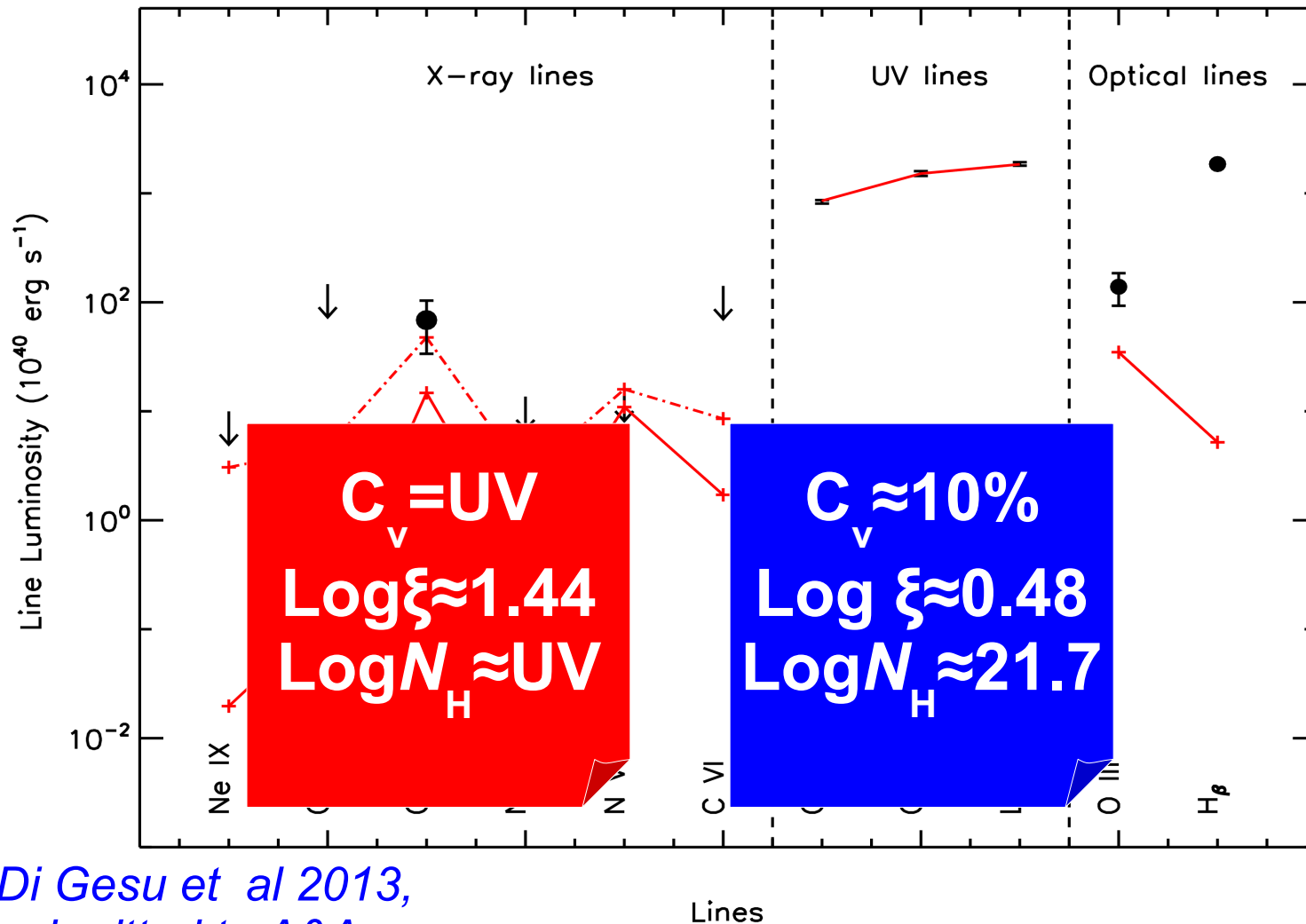
*Di Gesu et al 2013,
submitted to A&A*

The NL emitting system



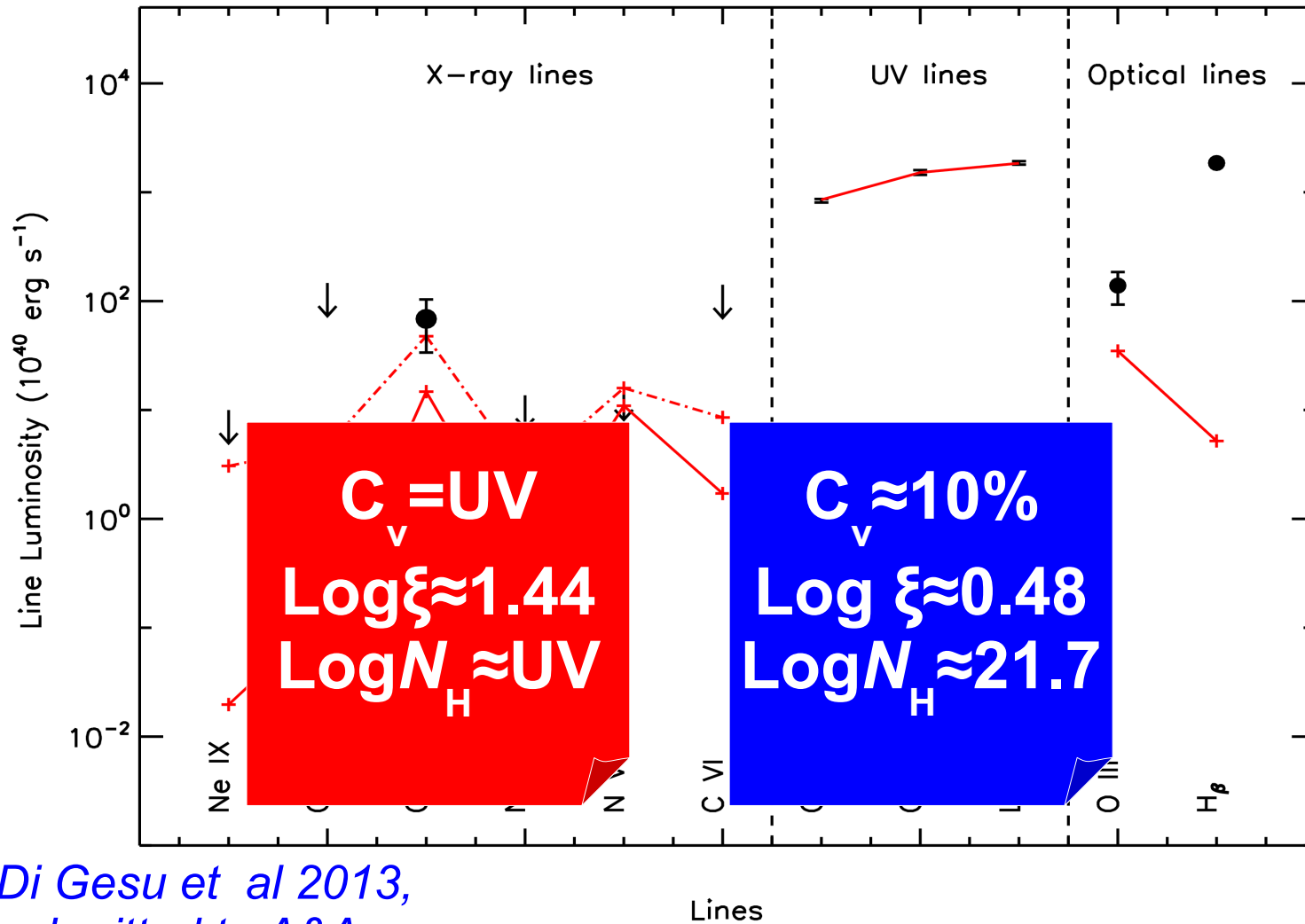
*Di Gesu et al 2013,
submitted to A&A*

The NL emitting system



*Di Gesu et al 2013,
submitted to A&A*

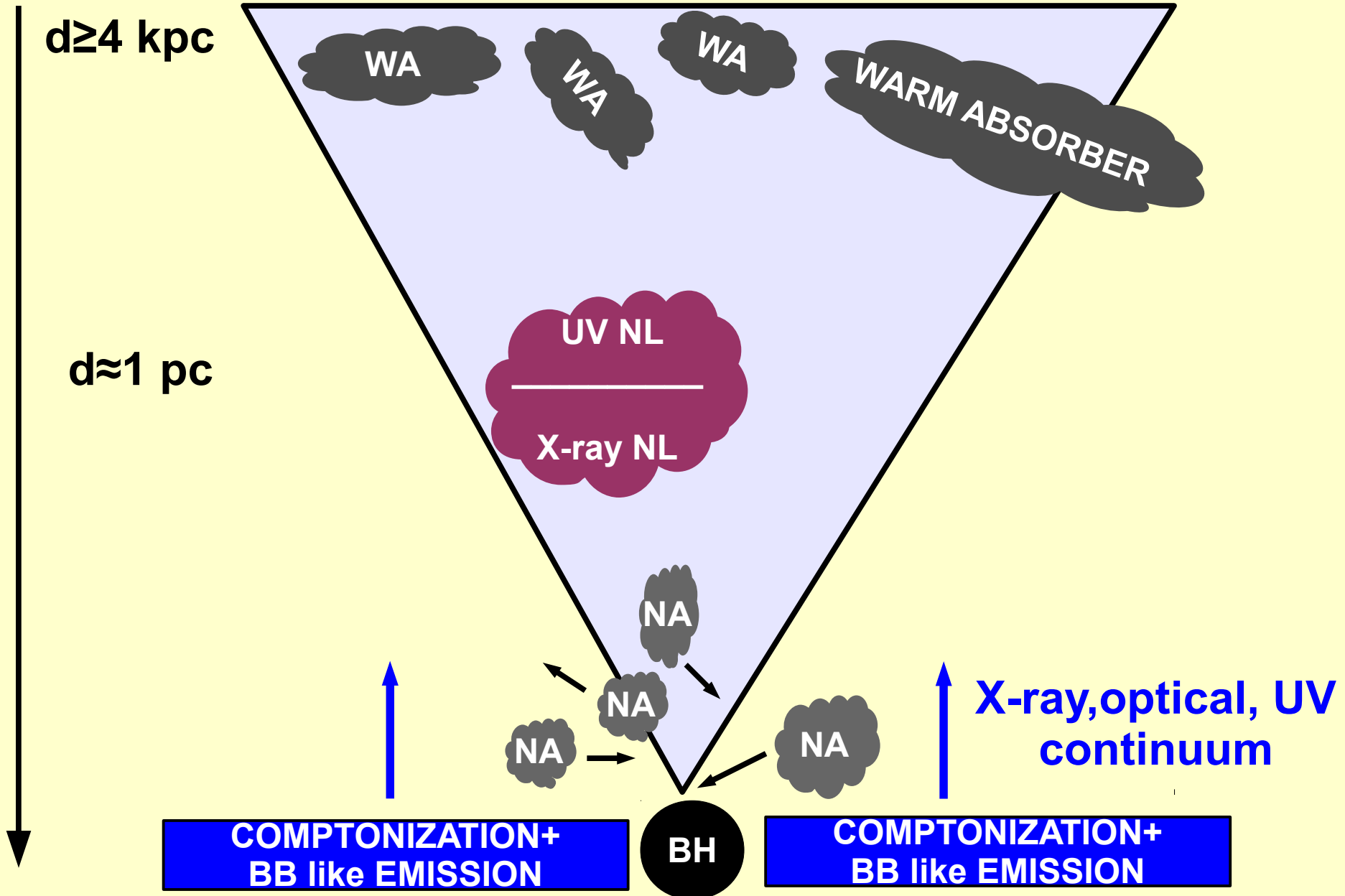
The NL emitting system



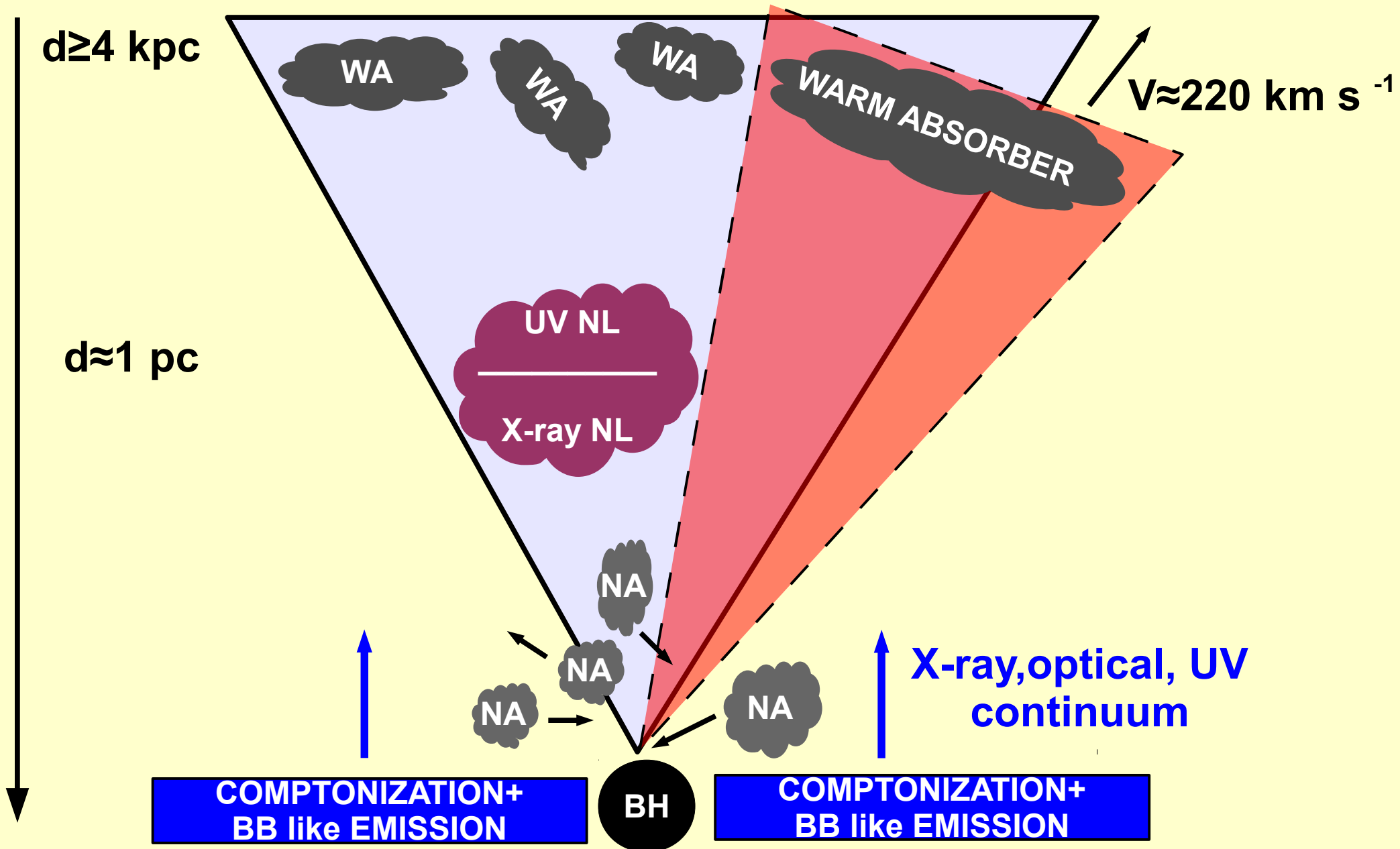
*Di Gesu et al 2013,
submitted to A&A*

→ UV and X-ray emitters are adjacent layers of the gas.

A global picture



A global picture



Summary

- **Present and past X-ray observation are well fitted by Comptonization plus neutral obscuration.**
- **Line emitting region stratified in ionization.**
- **Galactic scale outflow: first X-ray absorber detected so far away from the nucleus.**