

# Detailed analysis of Kepler-10: Synergy between asteroseismology and exoplanet research

Kepler-10b (animation: NASA)



CoRot-7b  
(Figure: ESO/L. Calçada)

Hans Kjeldsen



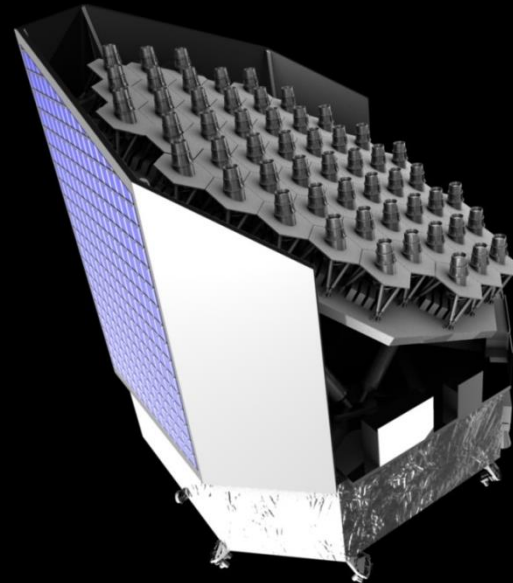
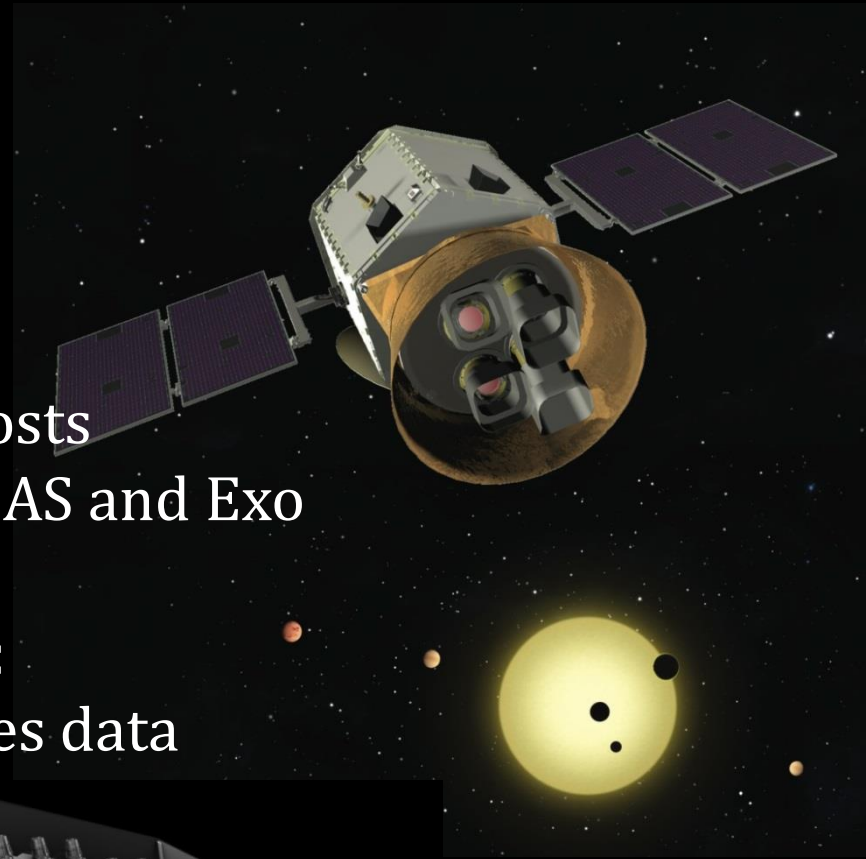
SAC  
STELLAR ASTROPHYSICS CENTRE

# PLATO and TESS

- CoRoT: Separate AS and Exo
- Kepler: AS for selected planet hosts
- PLATO and TESS: Simultaneous AS and Exo

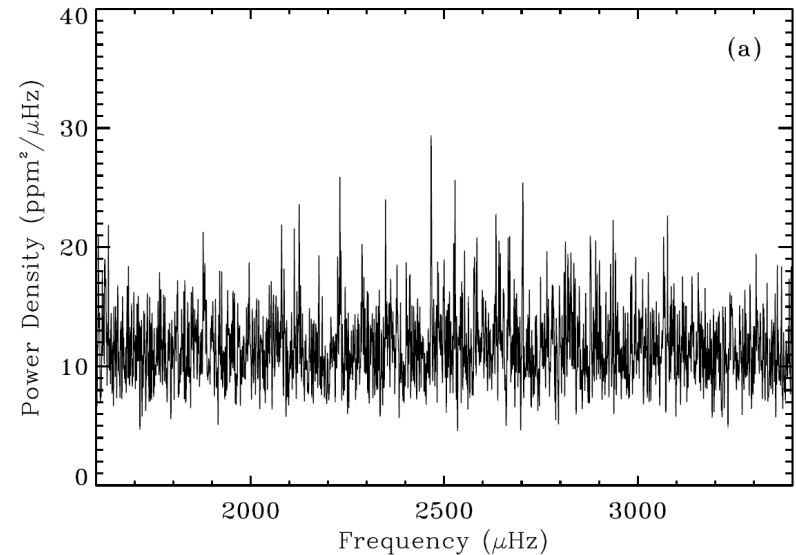
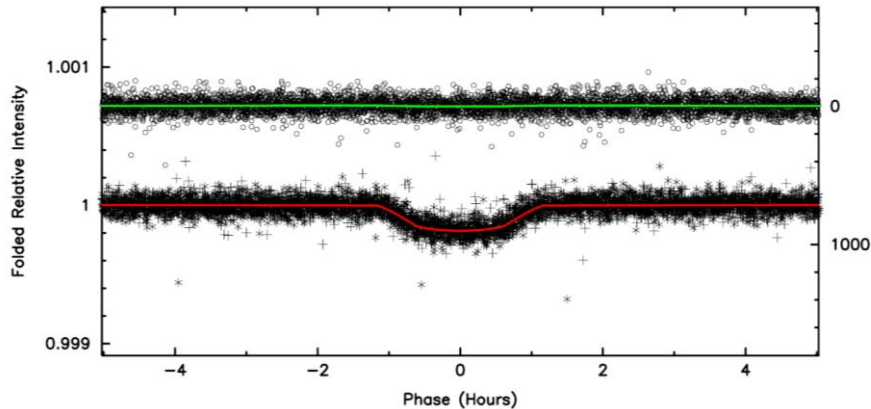
Accurate properties of exoplanets:

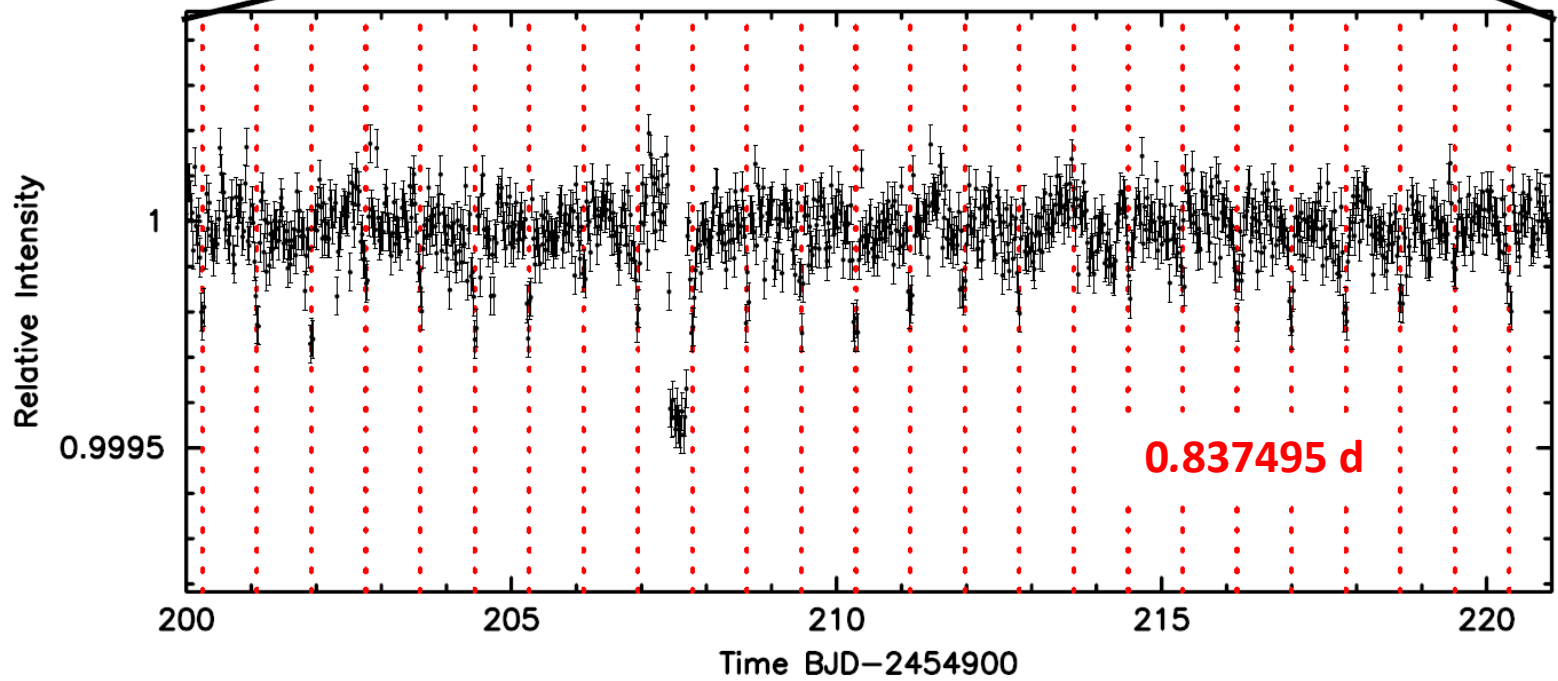
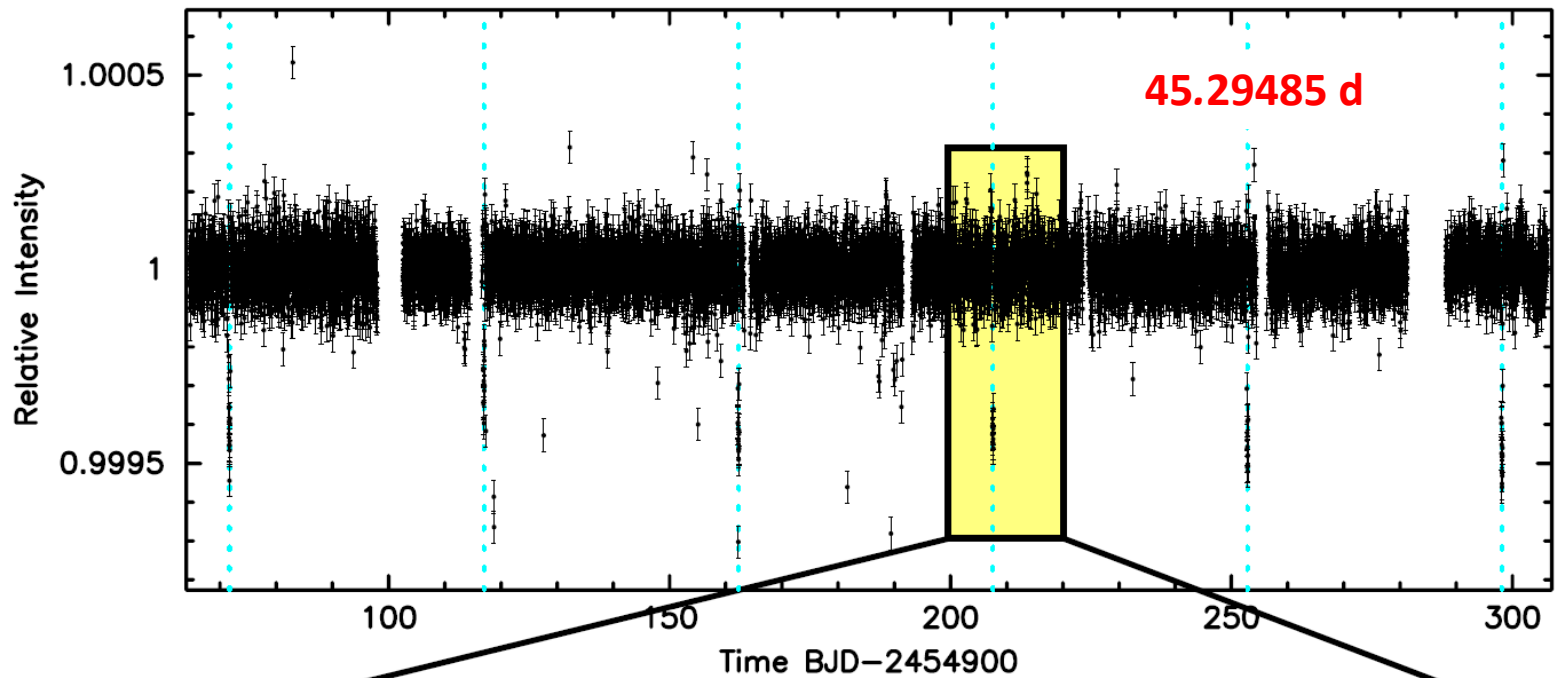
- SNR and length of the time series data
- Kepler-10:
  - Batalha et al. 2011: **275d**
  - All data to date: **850 d**

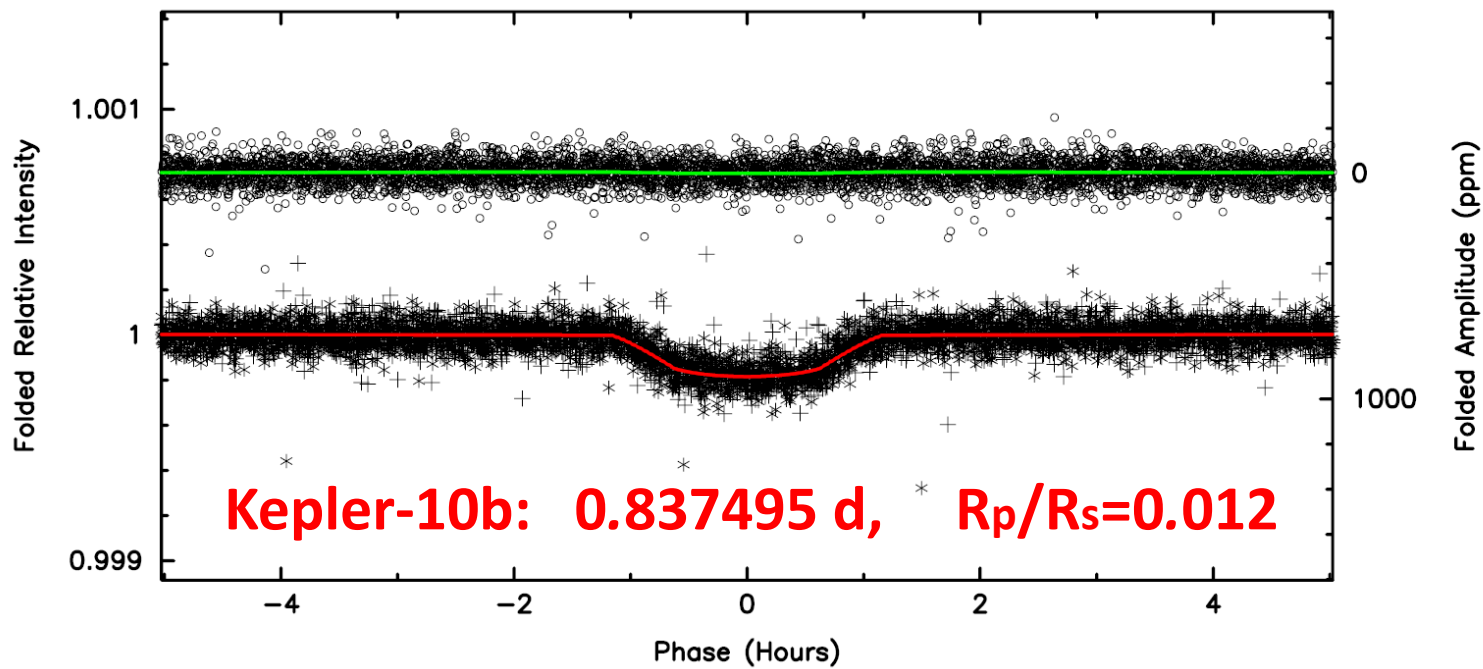
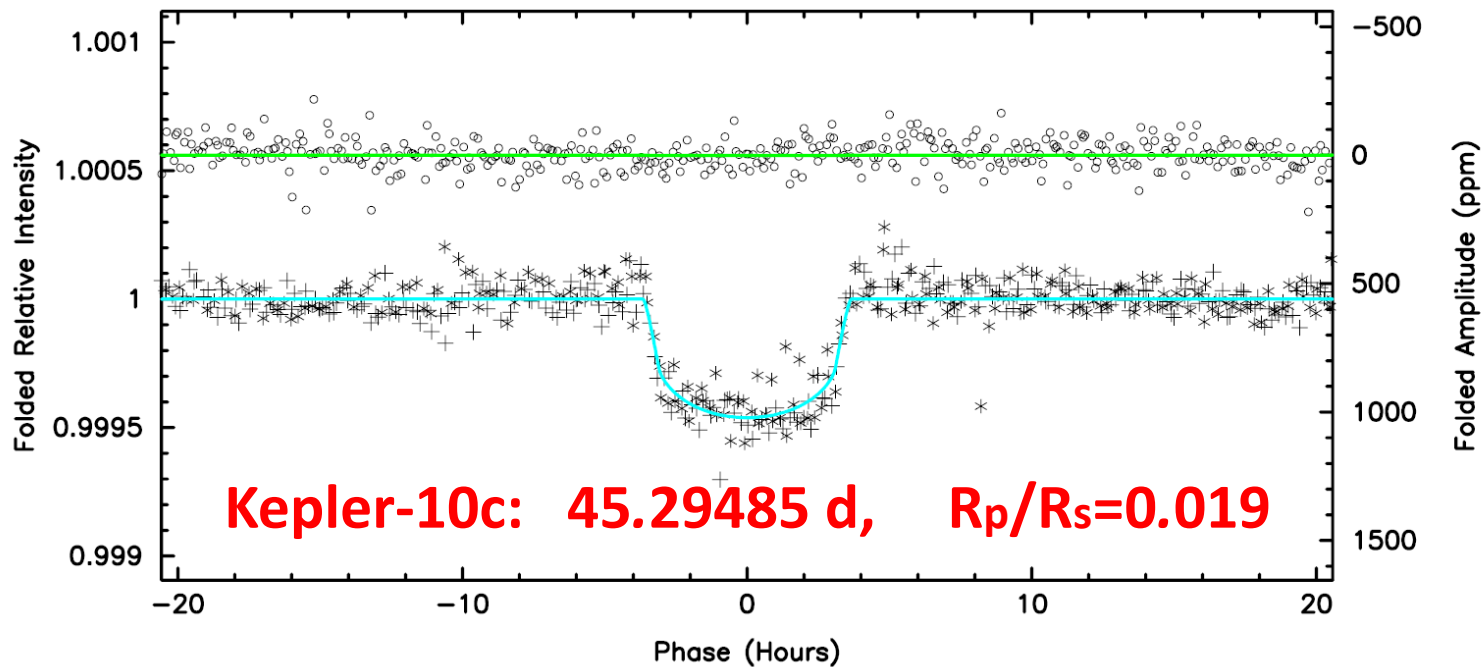


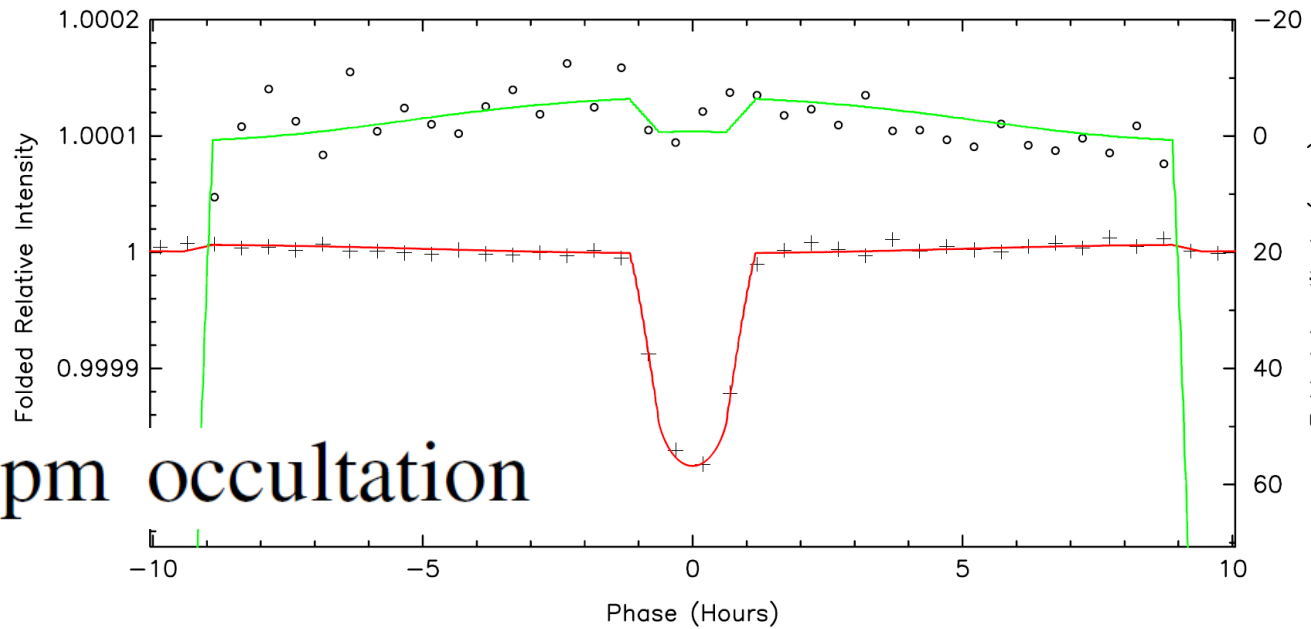
## KEPLER'S FIRST ROCKY PLANET: KEPLER-10b\*

NATALIE M. BATALHA<sup>1</sup>, WILLIAM J. BORUCKI<sup>2</sup>, STEPHEN T. BRYSON<sup>2</sup>, LARS A. BUCHHAVE<sup>3</sup>, DOUGLAS A. CALDWELL<sup>4</sup>, JØRGEN CHRISTENSEN-DALSGAARD<sup>5,6</sup>, DAVID CIARDI<sup>7</sup>, EDWARD W. DUNHAM<sup>8</sup>, FRANCOIS FRESSIN<sup>3</sup>, THOMAS N. GAUTIER III<sup>9</sup>, RONALD L. GILLILAND<sup>10</sup>, MICHAEL R. HAAS<sup>2</sup>, STEVE B. HOWELL<sup>11</sup>, JON M. JENKINS<sup>4</sup>, HANS KJELDSSEN<sup>5</sup>, DAVID G. KOCH<sup>2</sup>, DAVID W. LATHAM<sup>3</sup>, JACK J. LISSAUER<sup>2</sup>, GEOFFREY W. MARCY<sup>12</sup>, JASON F. ROWE<sup>2</sup>, DIMITAR D. SASSELOV<sup>3</sup>, SARA SEAGER<sup>13</sup>, JASON H. STEFFEN<sup>14</sup>, GUILLERMO TORRES<sup>3</sup>, GIBOR S. BASRI<sup>12</sup>, TIMOTHY M. BROWN<sup>15</sup>, DAVID CHARBONNEAU<sup>3</sup>, JESSIE CHRISTIANSEN<sup>2</sup>, BRUCE CLARKE<sup>4</sup>, WILLIAM D. COCHRAN<sup>16</sup>, ANDREA DUPREE<sup>3</sup>, DANIEL C. FABRYCKY<sup>3</sup>, DEBRA FISCHER<sup>17</sup>, ERIC B. FORD<sup>18</sup>, JONATHAN FORTNEY<sup>19</sup>, FORREST R. GIROUARD<sup>20</sup>, MATTHEW J. HOLMAN<sup>3</sup>, JOHN JOHNSON<sup>21</sup>, HOWARD ISAACSON<sup>12</sup>, TODD C. KLAUS<sup>20</sup>, PAVEL MACHALEK<sup>4</sup>, ALTHEA V. MOOREHEAD<sup>18</sup>, ROBERT C. MOREHEAD<sup>18</sup>, DARIN RAGOZZINE<sup>3</sup>, PETER TENENBAUM<sup>4</sup>, JOSEPH TWICKEN<sup>4</sup>, SAMUEL QUINN<sup>3</sup>, JEFFREY VANCLEVE<sup>4</sup>, LUCIANNE M. WALKOWICZ<sup>12</sup>, WILLIAM F. WELSH<sup>22</sup>, EDNA DEVORE<sup>4</sup>, AND ALAN GOULD<sup>23</sup>

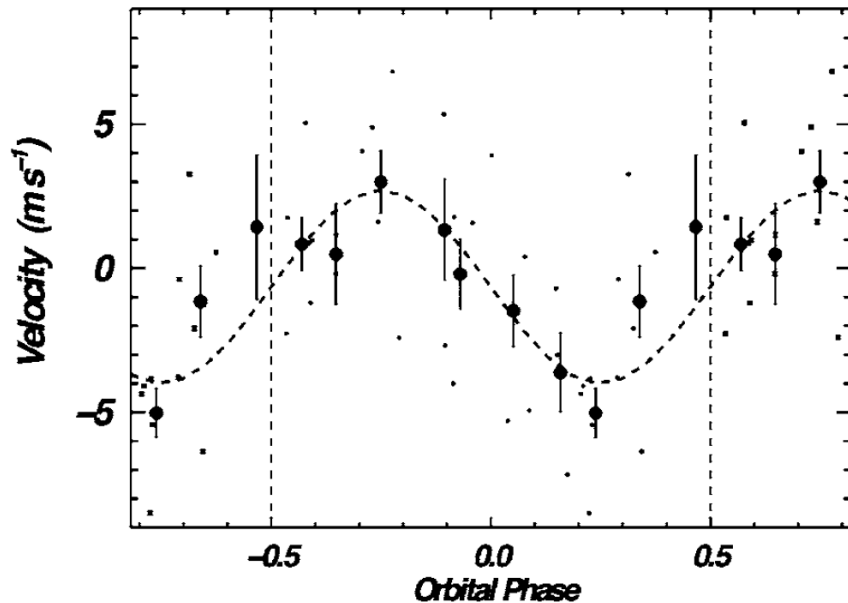






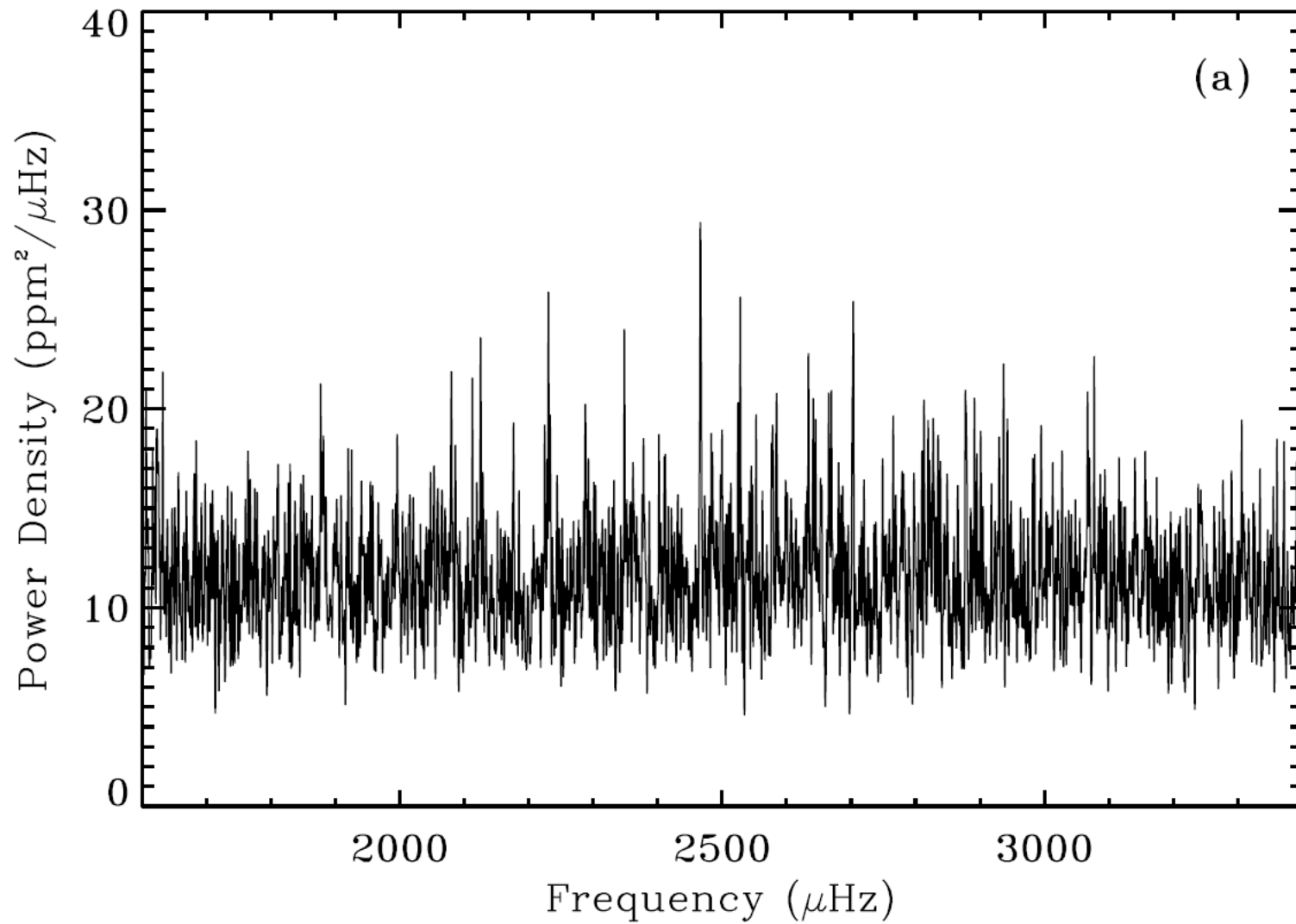


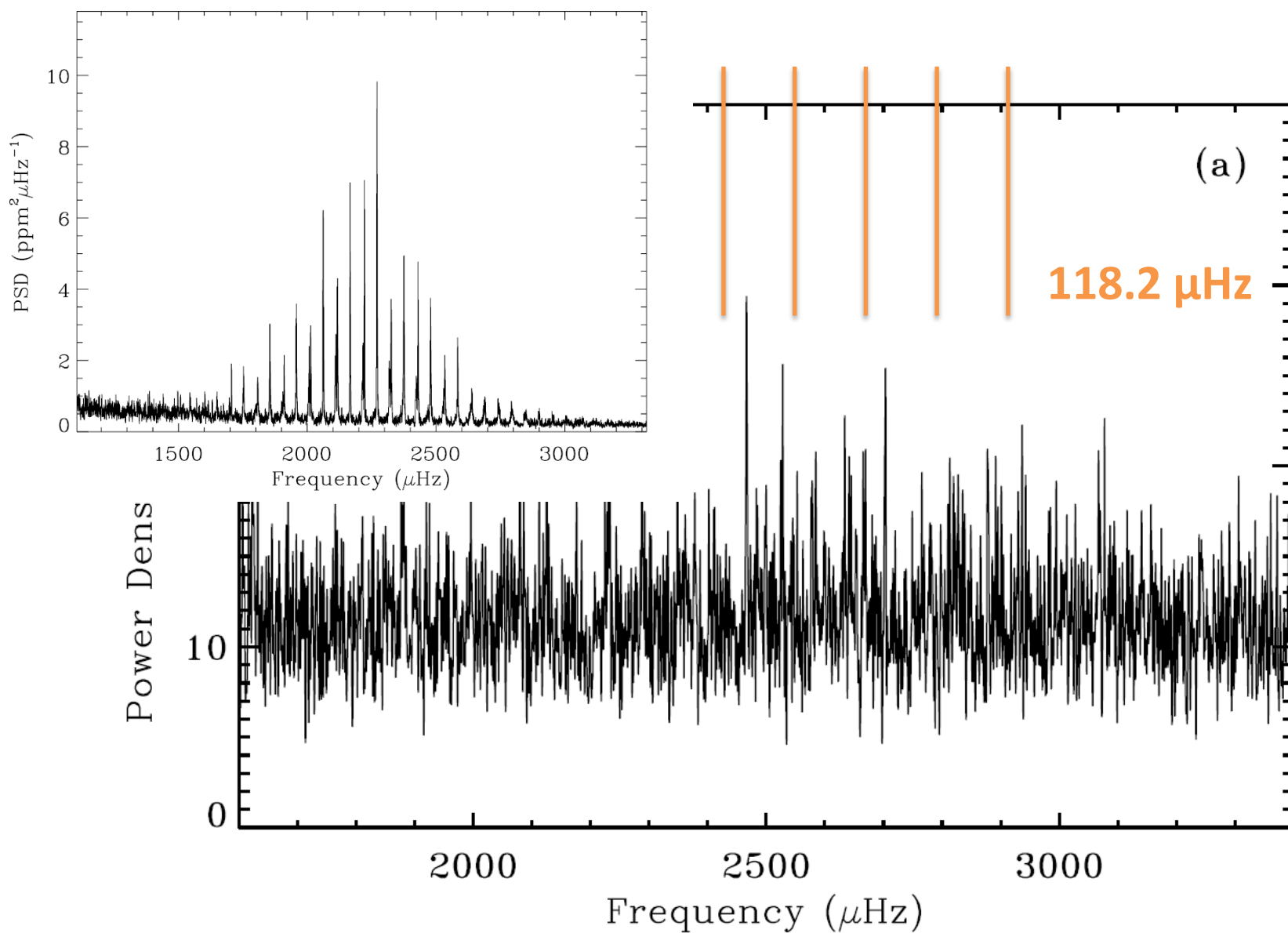
**Kepler-10b: 0.837495 d,  $R_p/R_s=0.012$**



Mass,  $M_P (M_{\oplus})$

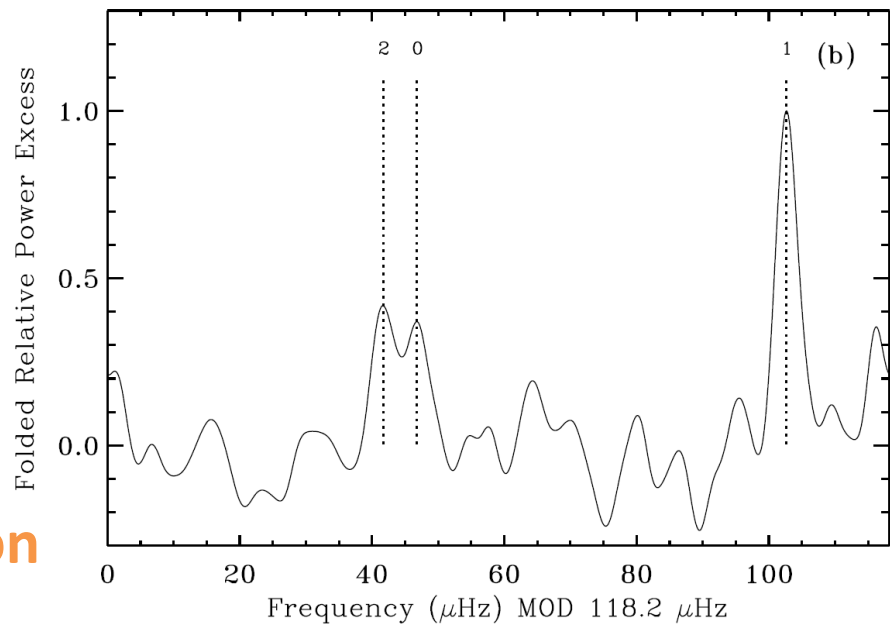
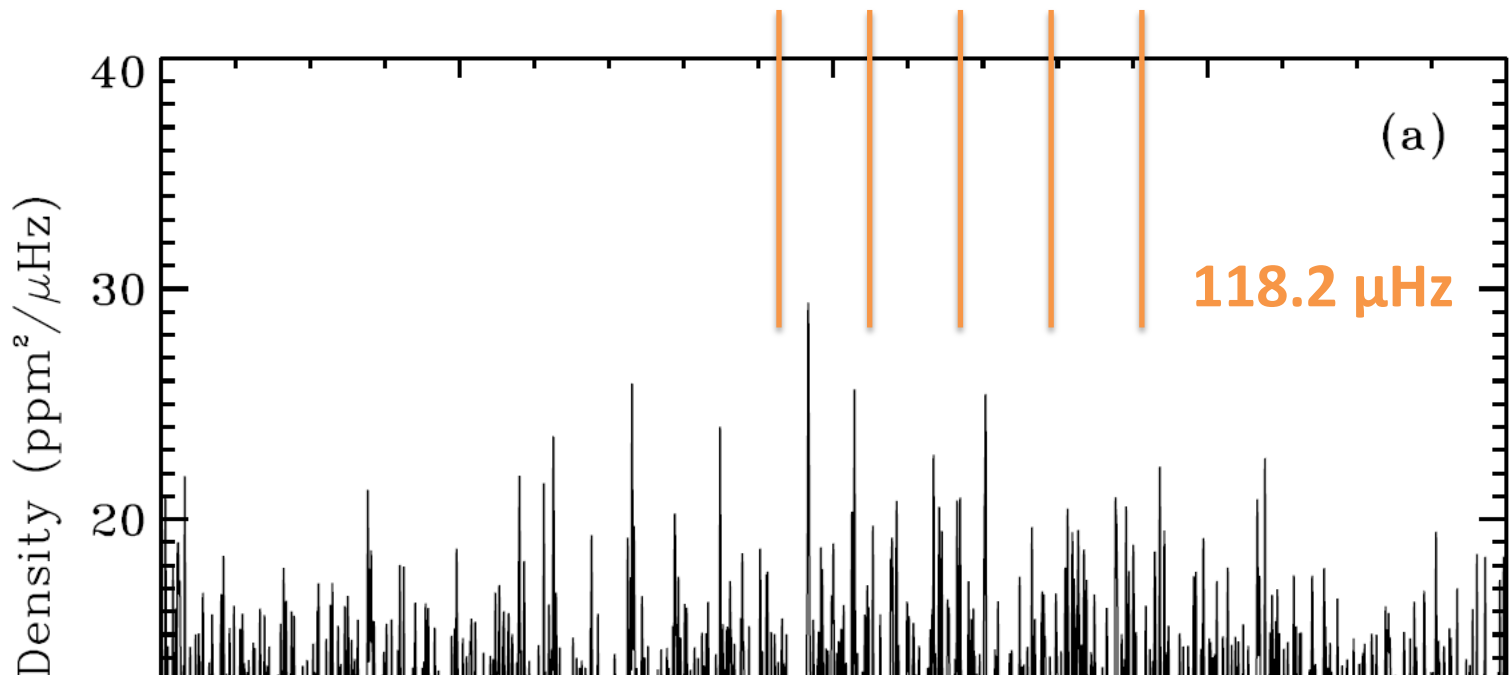
$$4.56^{+1.17}_{-1.29}$$





**Large frequency separation**

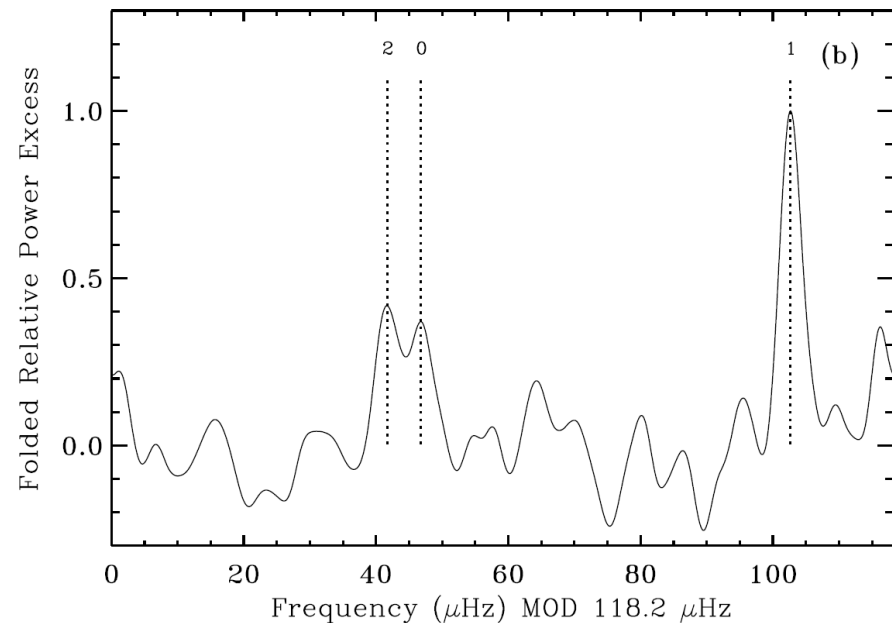




Small frequency separation

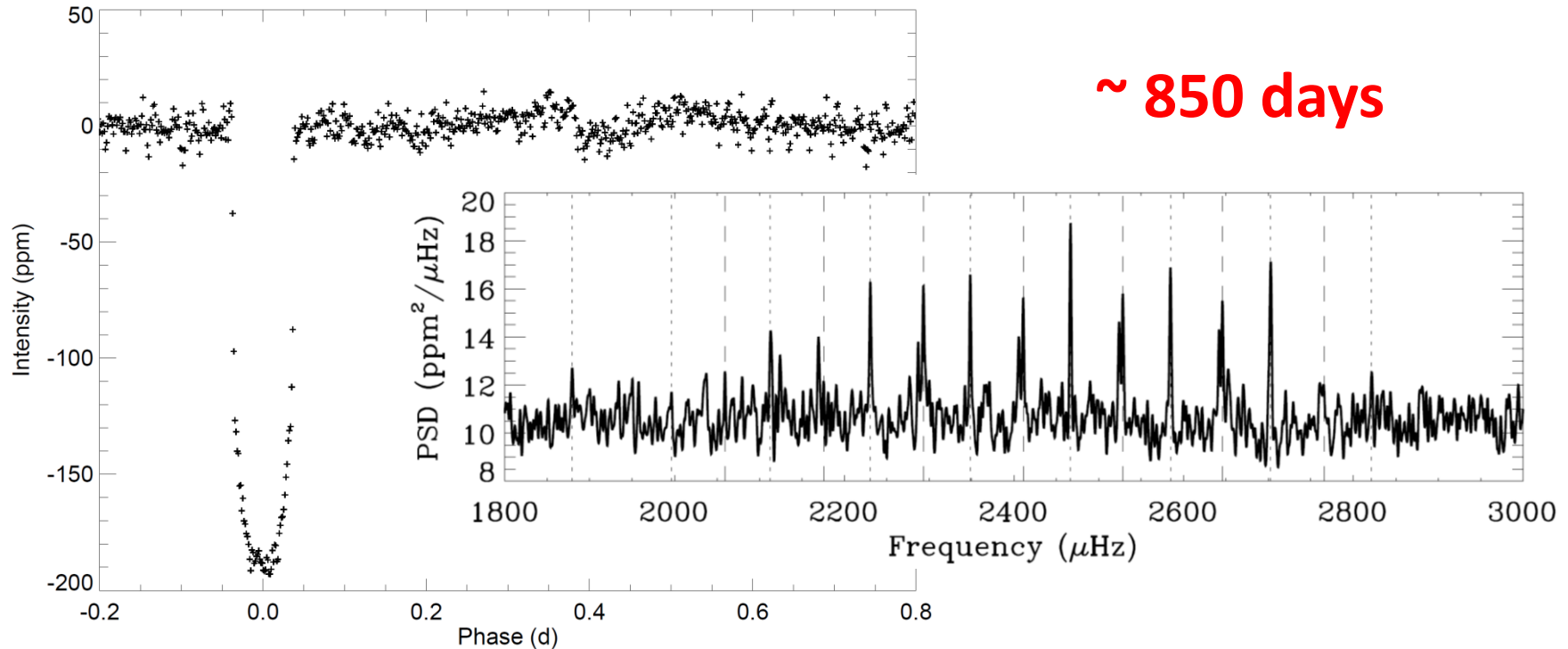
Mass (Msun)	$0.995 \pm 0.060$	( 6%)
Radius (Rsun)	$1.056 \pm 0.021$	( 2%)
Age (Gyr)	$11.9 \pm 4.5$	(38%)

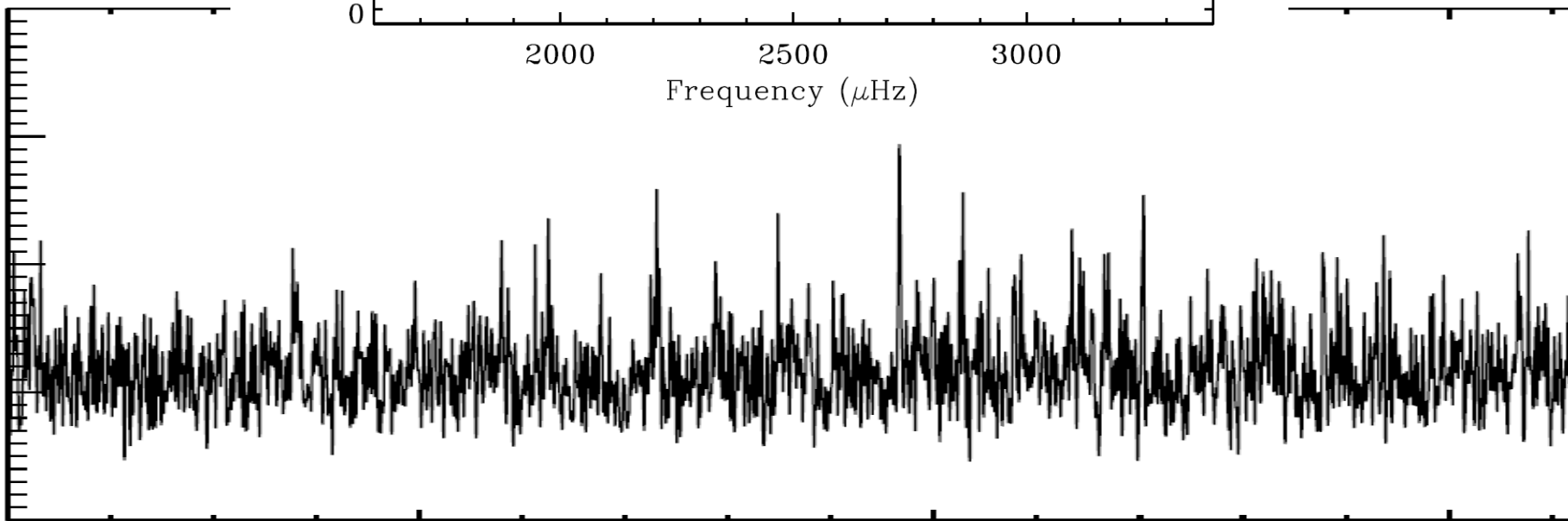
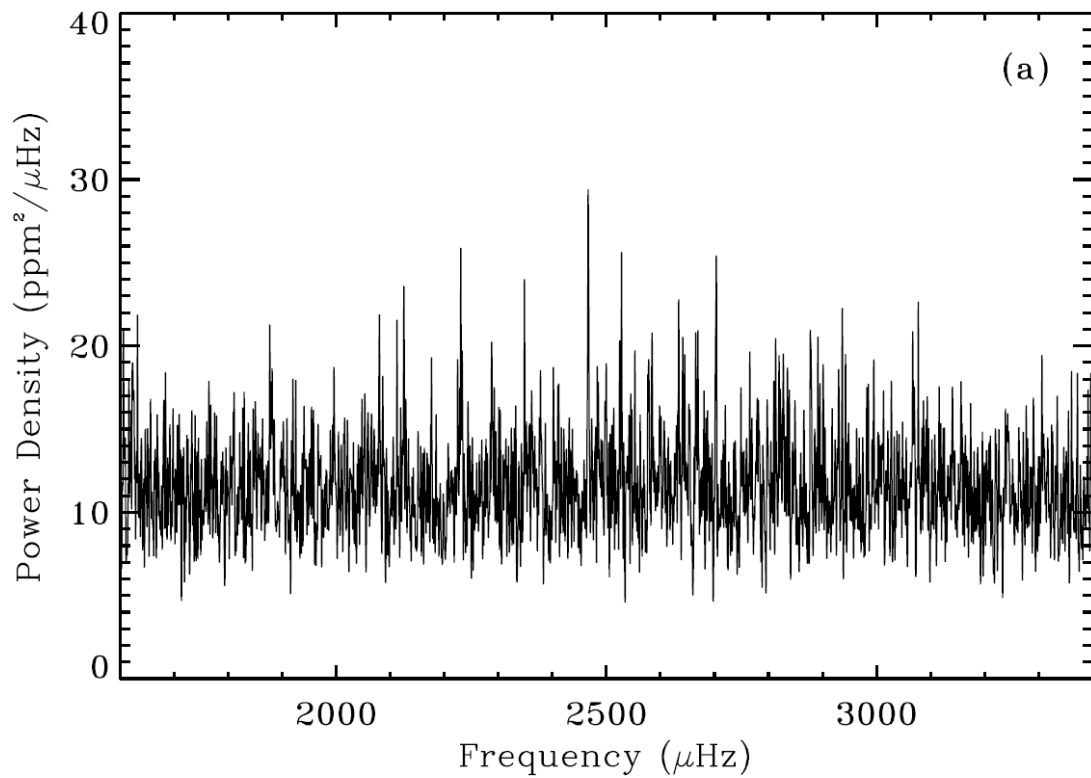
- **Batalha et al. 2011**

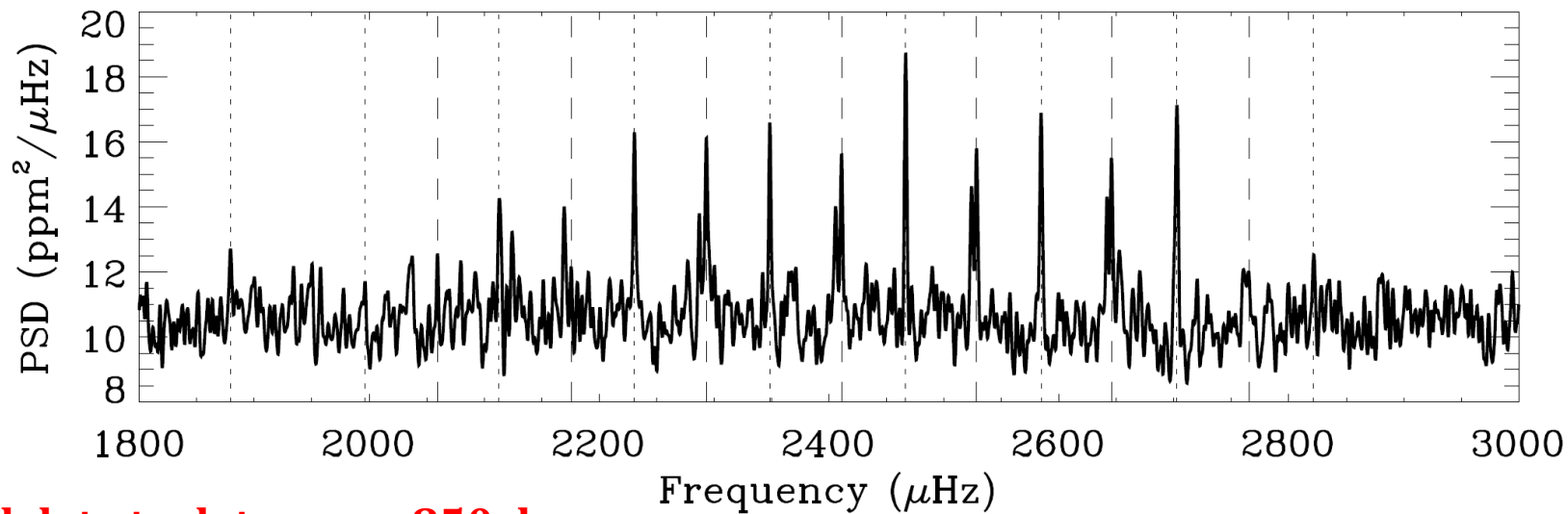


# Analysis of more than two years of data...

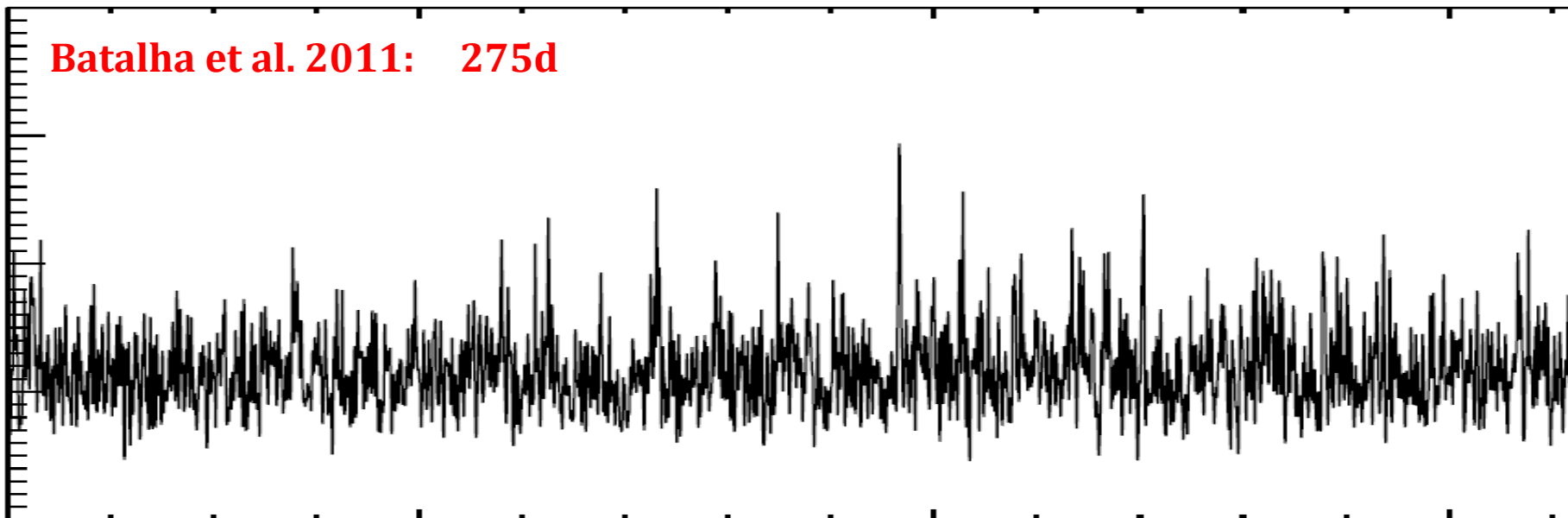
Alexandra Fogtman-Schulz, Brian Hinrup, Vincent Van Eylen, Jørgen Christensen-Dalsgaard, Hans Kjeldsen, Victor Silva Aguirre and Brandon Tingley  
Stellar Astrophysics Centre, Department of Physics and Astronomy, Aarhus University, Denmark.

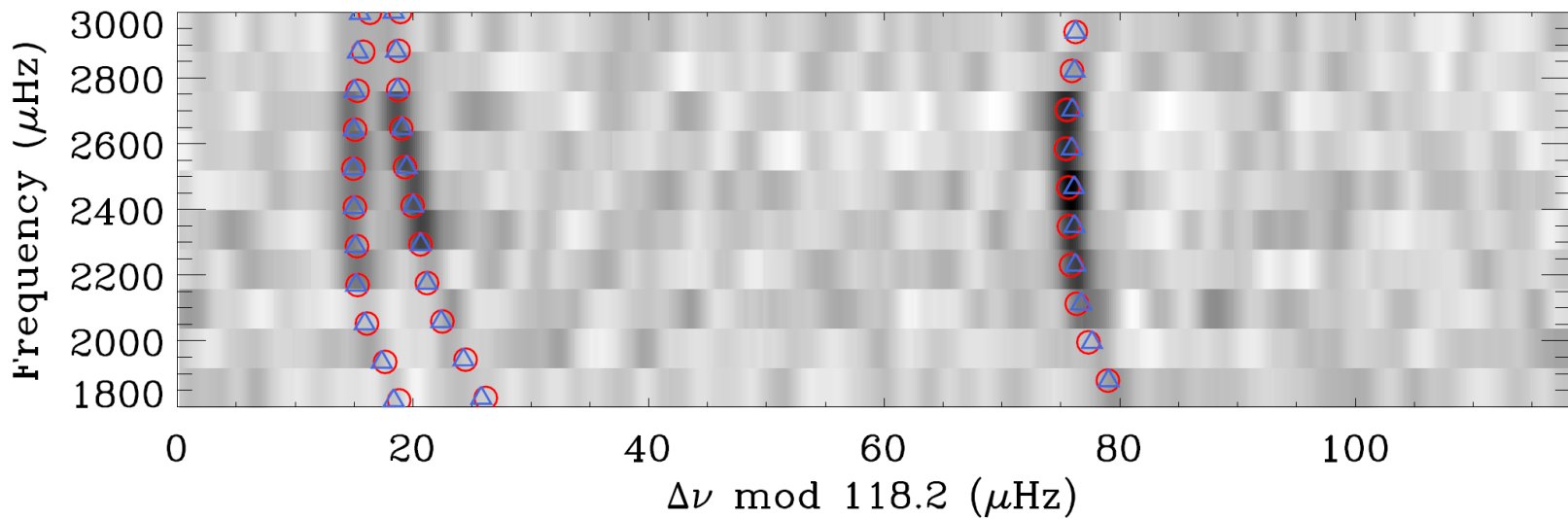
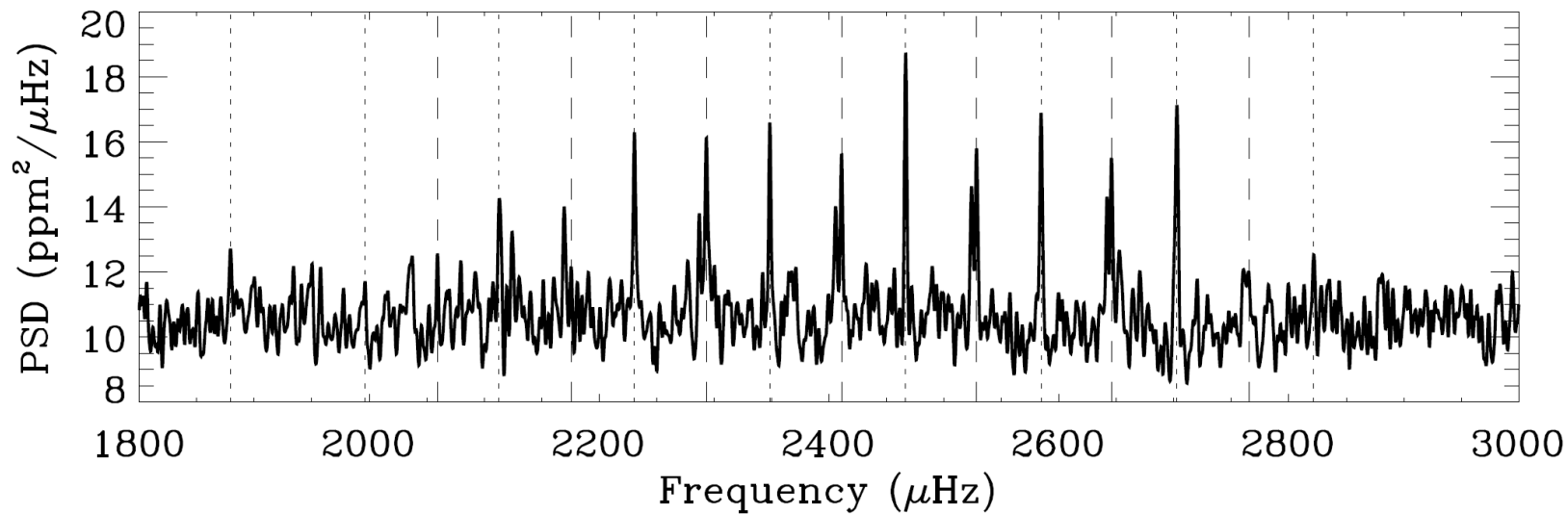






**All data to date: 850 d**



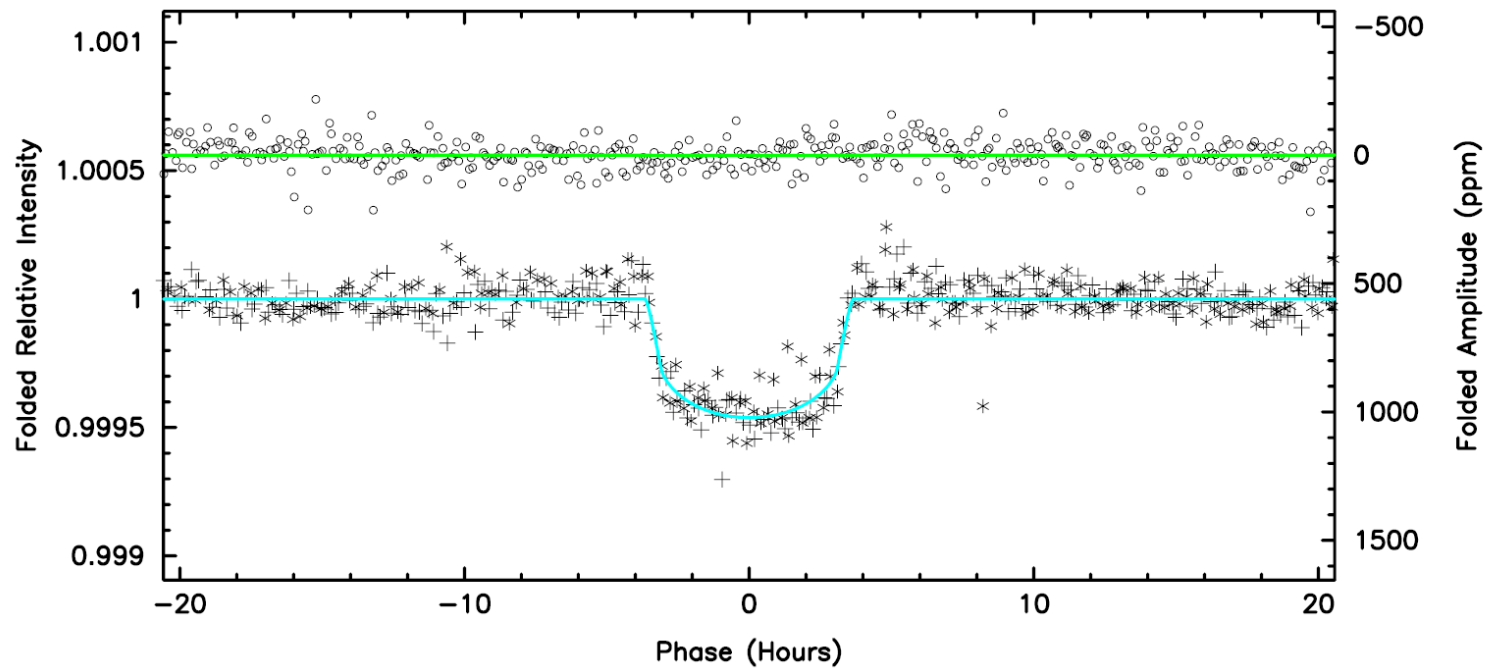


Mass (Msun)	$0.995 \pm 0.060$
Radius (Rsun)	$1.056 \pm 0.021$
Age (Gyr)	$11.9 \pm 4.5$

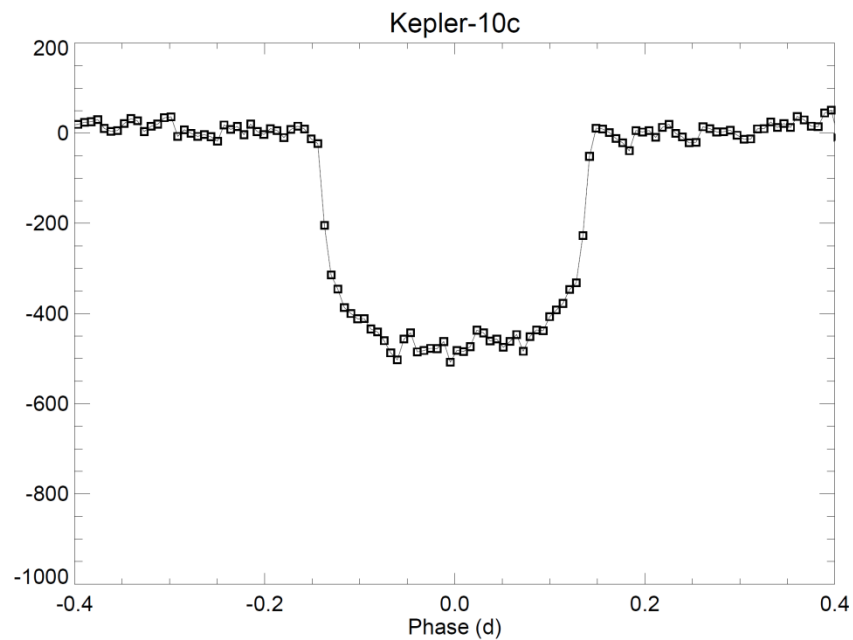
**- Batalha et al. 2011**

Mass (Msun)	$0.913 \pm 0.022$
Radius (Rsun)	$1.065 \pm 0.009$
Age (Gyr)	$10.4 \pm 1.4$

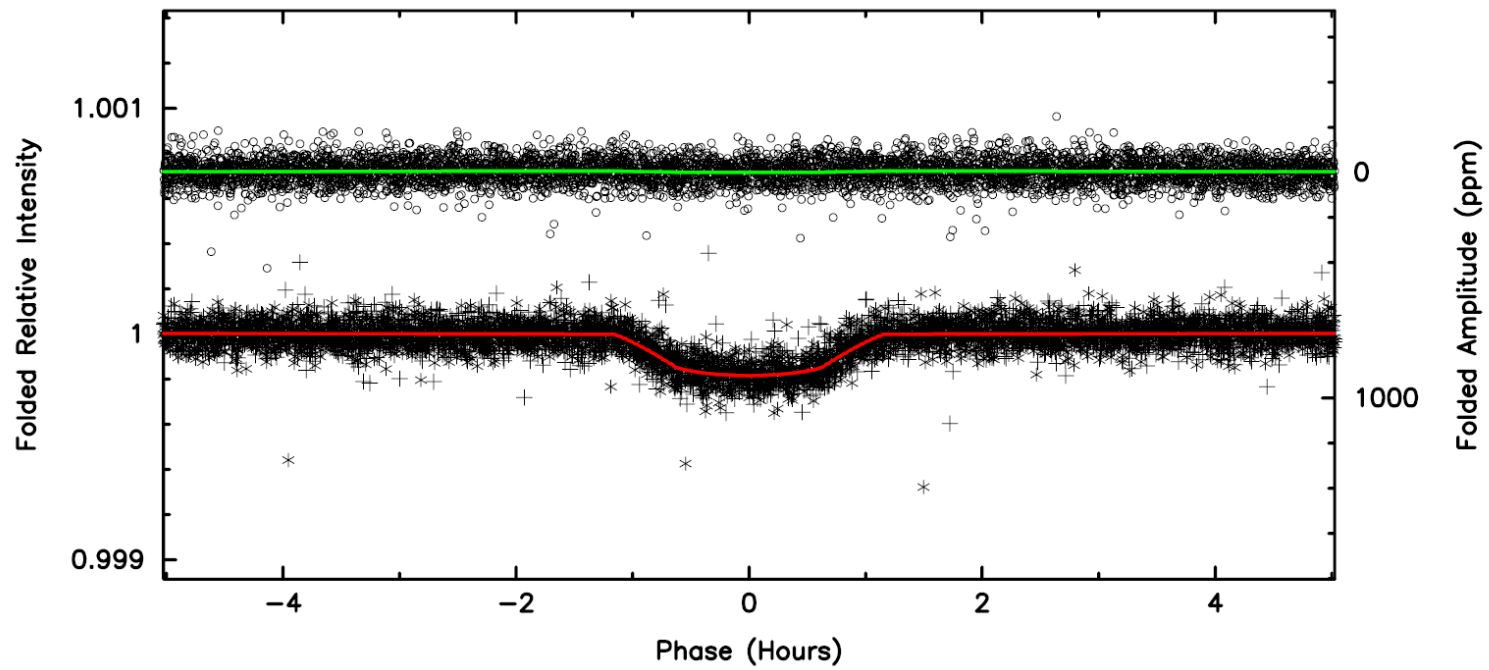
**- All data to date**



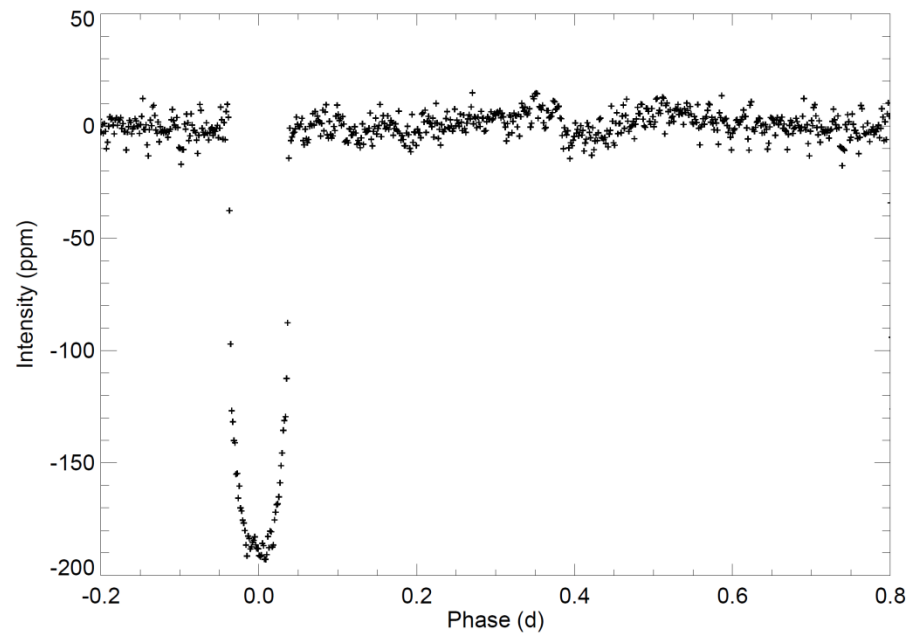
# Kepler-10c

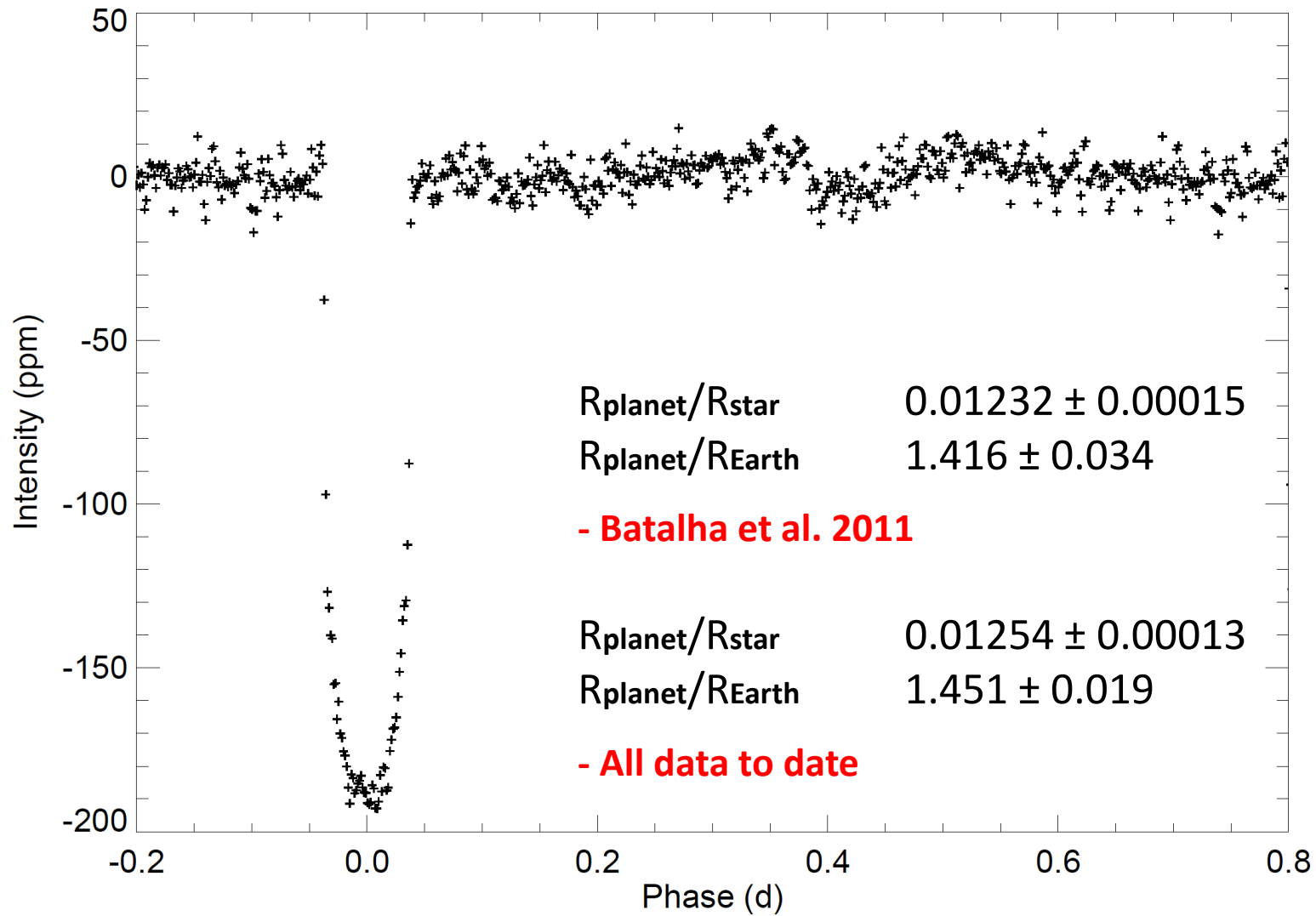




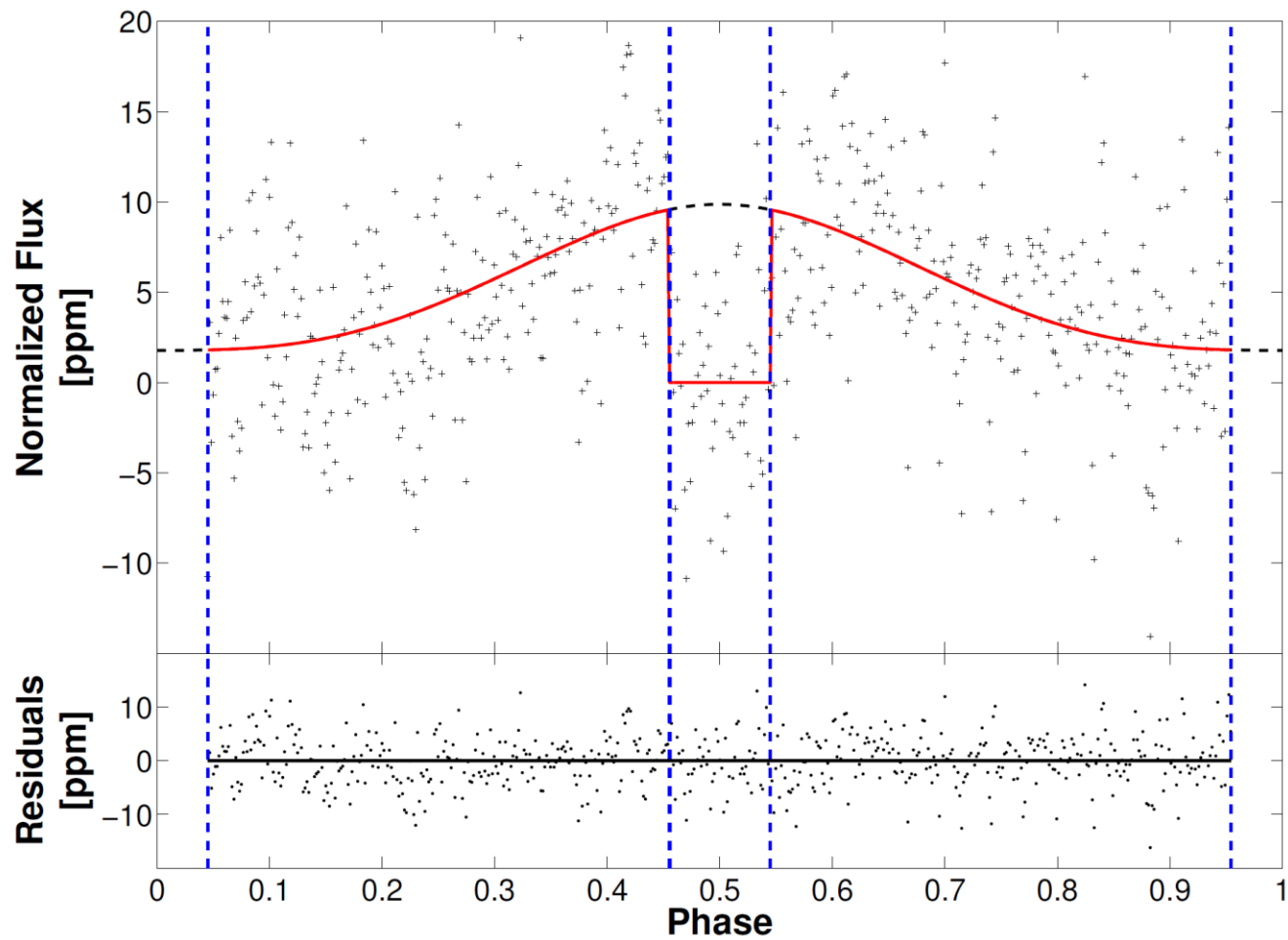


# Kepler-10b

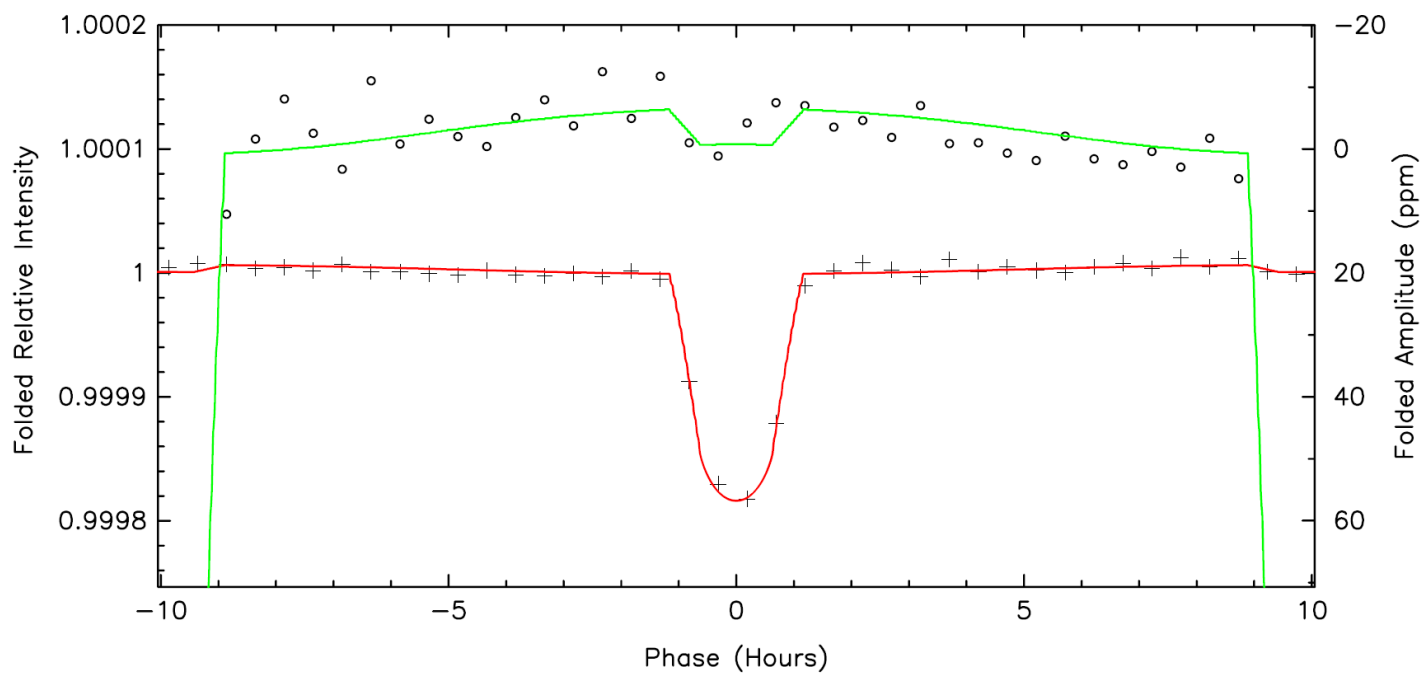
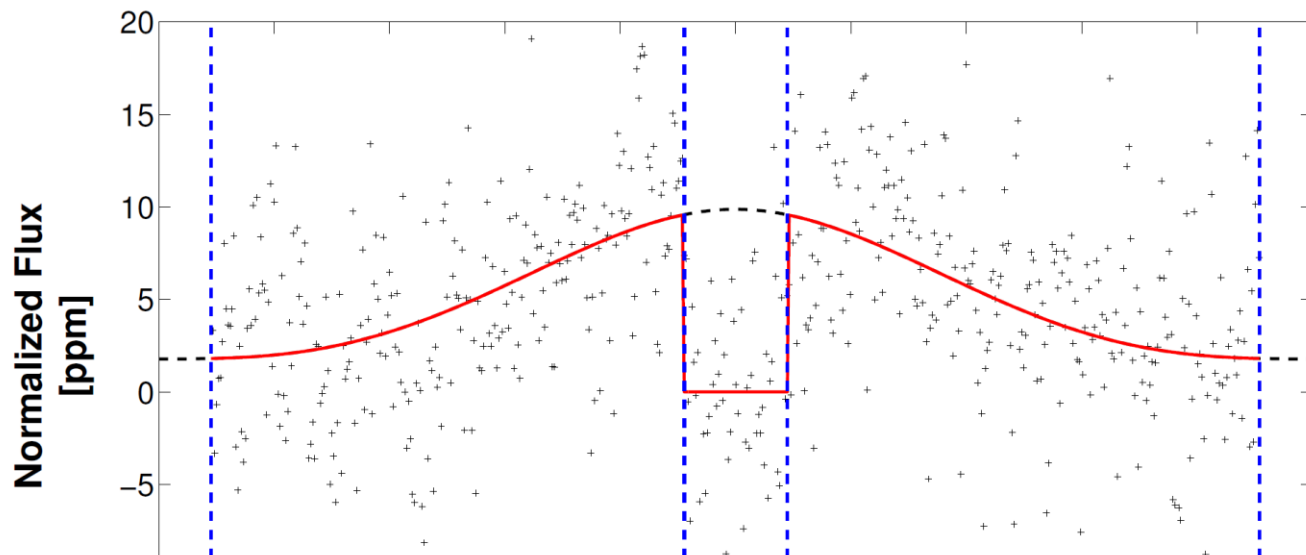




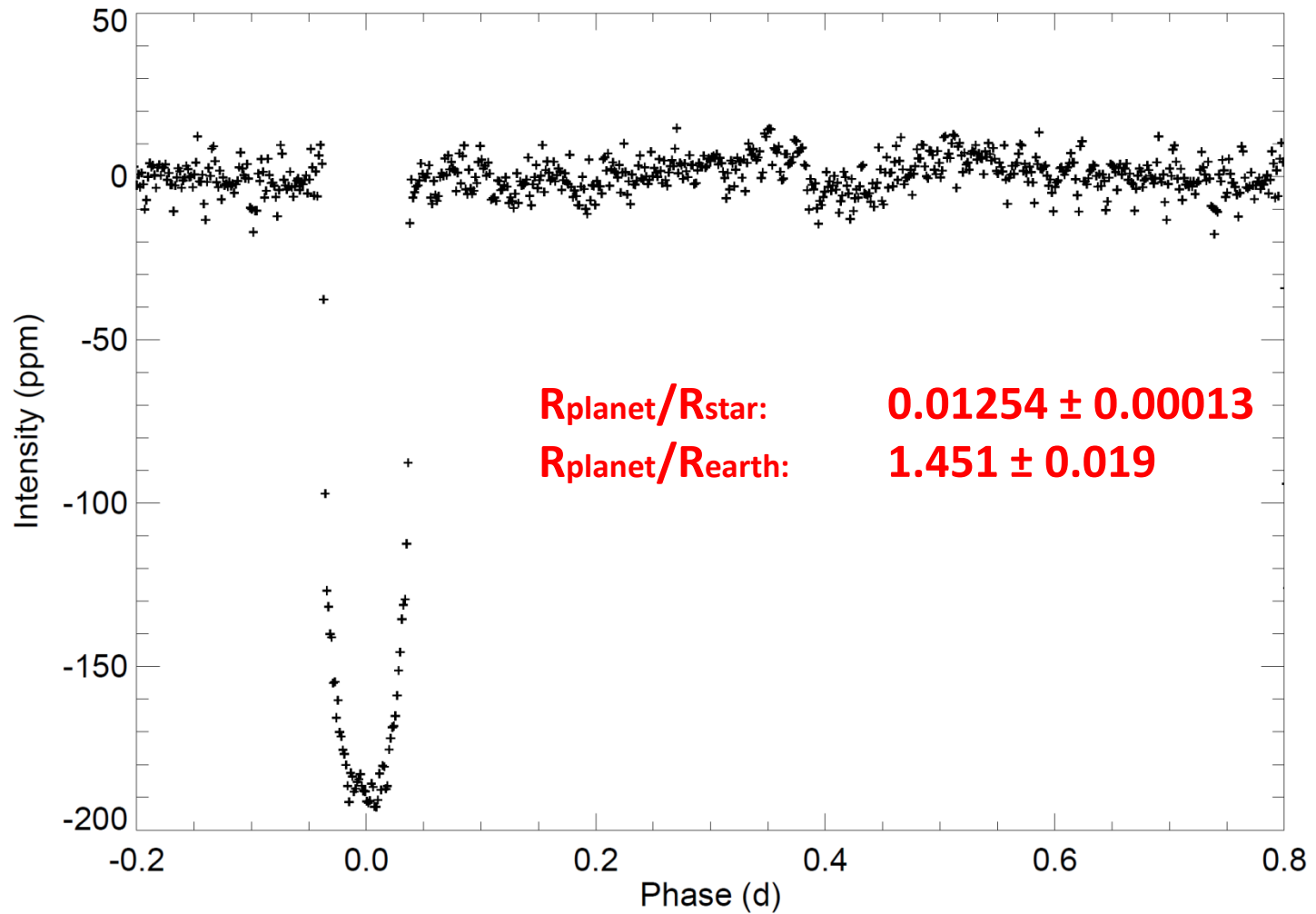
$\delta_{\text{occultation}}$  (ppm)  $9.9 \pm 1.0$



$\delta_{\text{occultation}}$  (ppm)  $9.9 \pm 1.0$



# Are planet properties at the 1% level accurate or only precise?



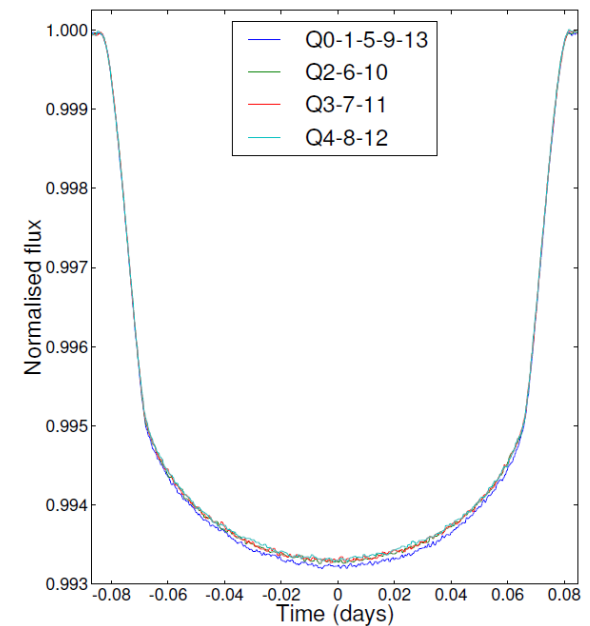
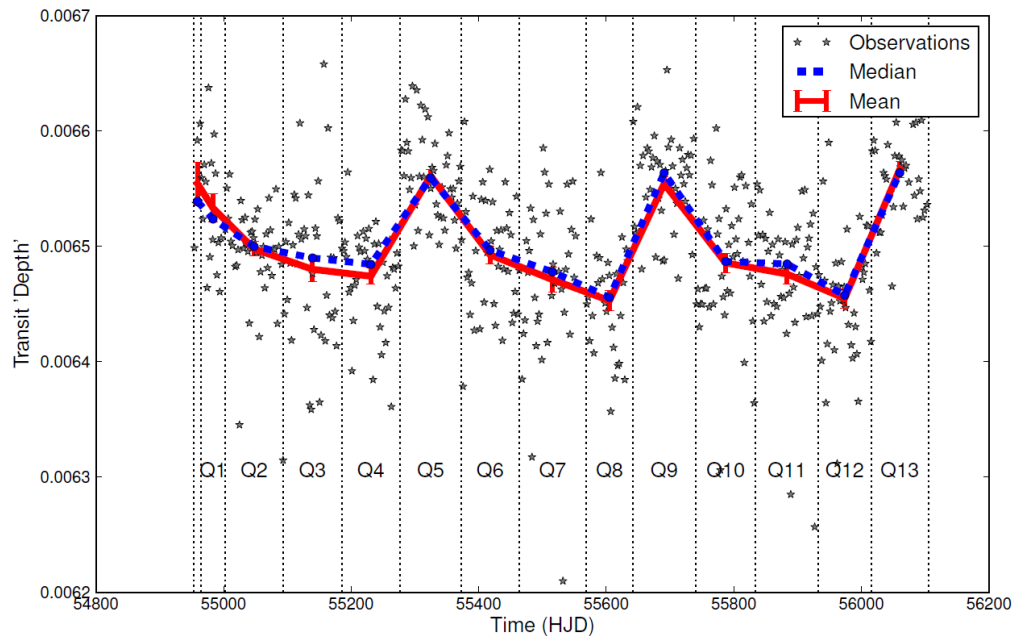
# Are planet properties at the 1% level accurate or only precise?

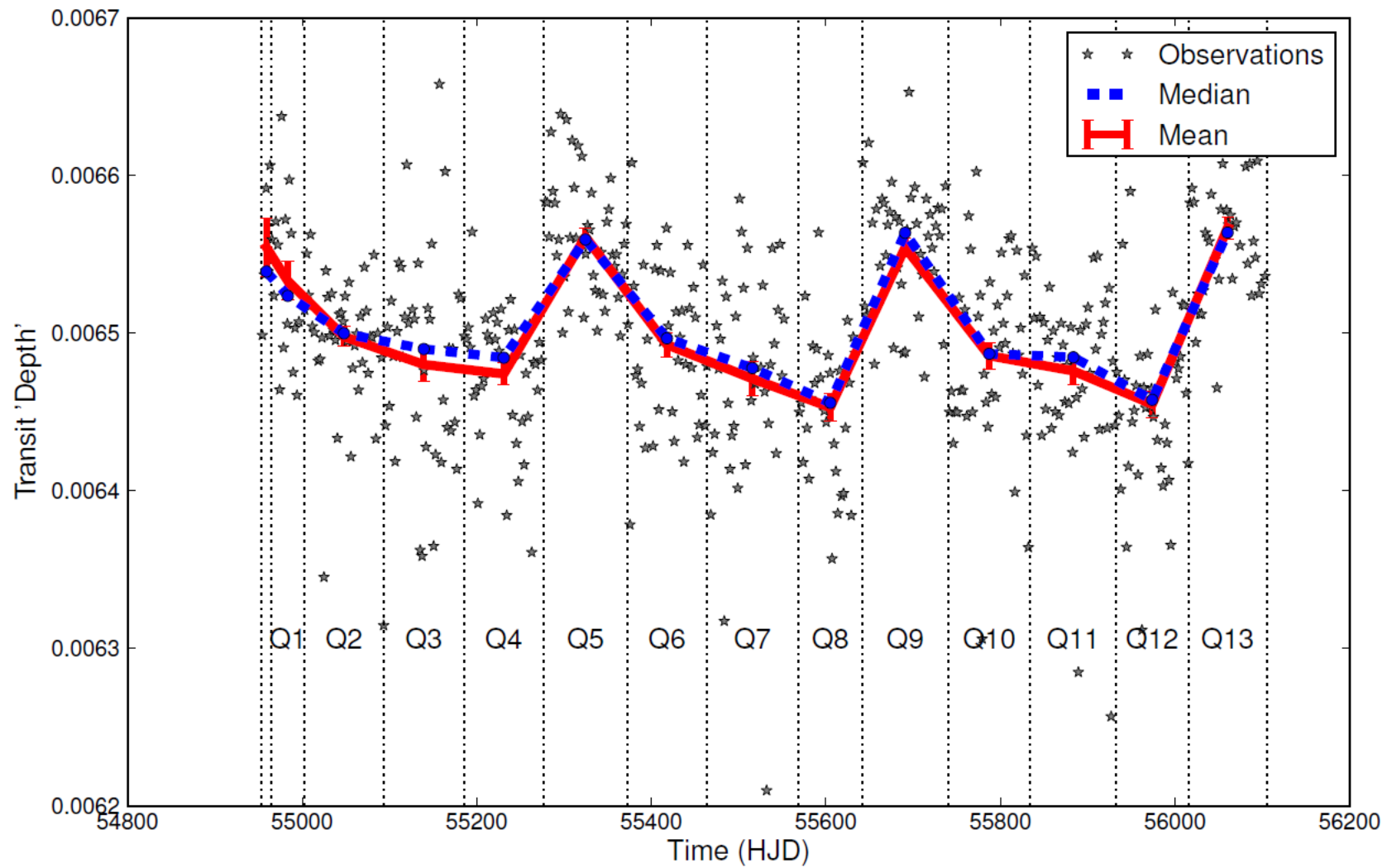
arXiv:1307.6959v1 [astro-ph.EP] 26 Jul 2013

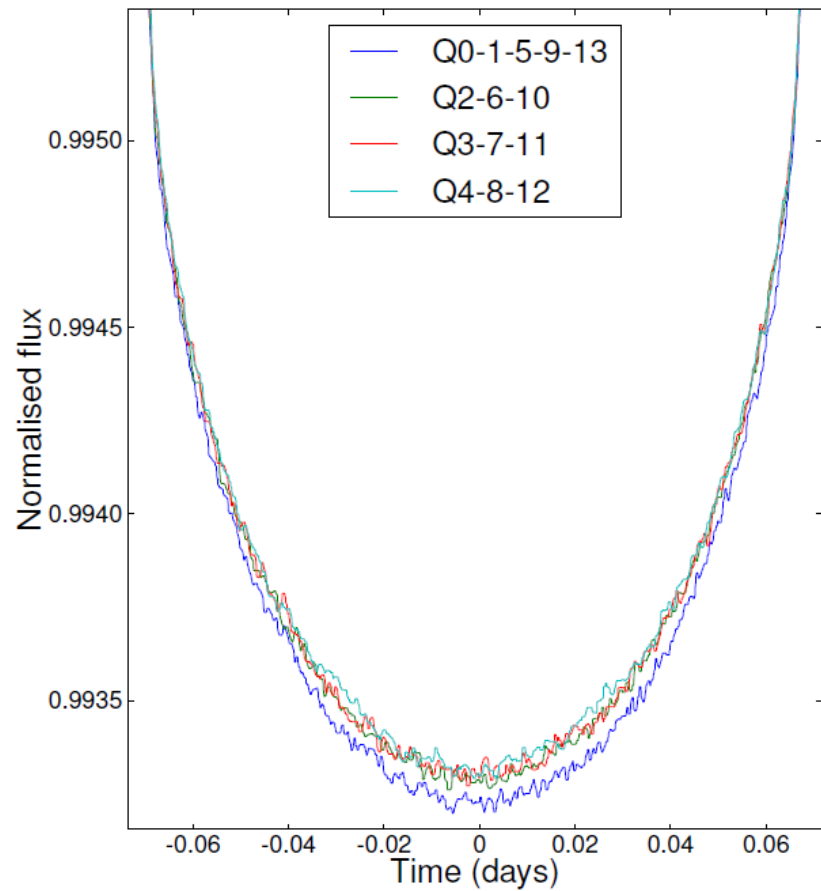
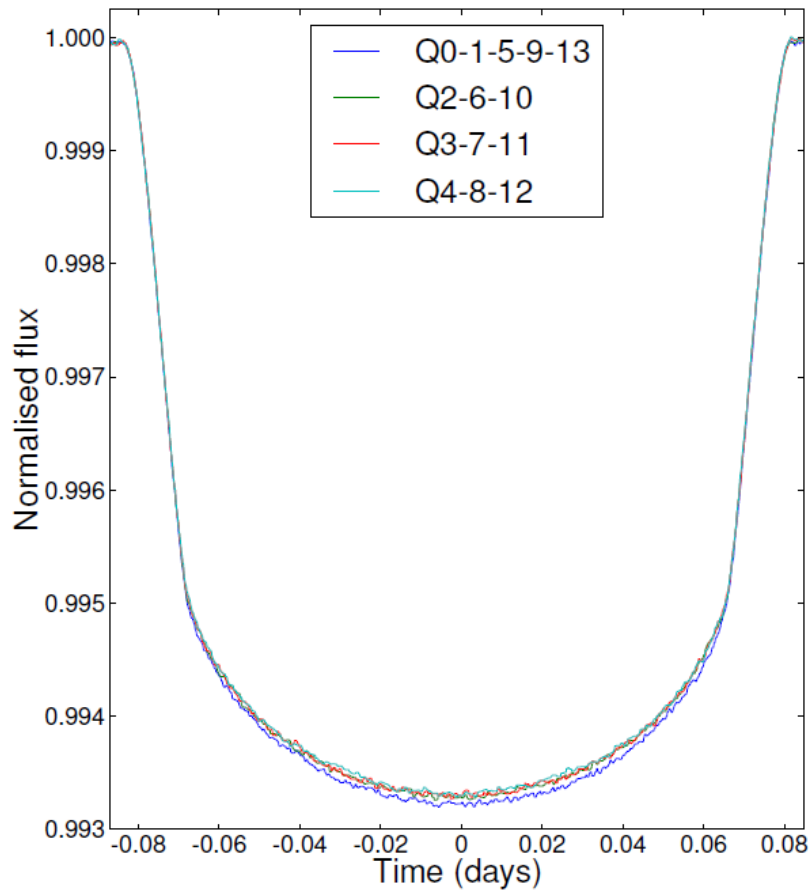
## Investigation of systematic effects in *Kepler* data: Seasonal variations in the light curve of HAT-P-7b

V. Van Eylen, M. Lindholm Nielsen, B. Hinrup, B. Tingley and H. Kjeldsen

*Stellar Astrophysics Centre, Department of Physics and Astronomy, Aarhus University, Ny Munkegade 120, DK-8000 Aarhus C, Denmark.*







Season	Module	Channel	'Depth' (ppm)	Rel. 'depth'	'Depth' diff. (%)	$R_p/R_\star$
1	17	58	$6557.2 \pm 3.5$	$1.00869 \pm 0.00054$	$0.869 \pm 0.054$	$0.077618^{+0.000073}_{-0.000073}$
2	19	66	$6491.8 \pm 4.5$	$0.99863 \pm 0.00068$	$-0.137 \pm 0.068$	$0.077355^{+0.000035}_{-0.000035}$
3	9	26	$6475.9 \pm 6.0$	$0.99618 \pm 0.00091$	$-0.382 \pm 0.091$	$0.077330^{+0.000056}_{-0.000059}$
4	7	18	$6461.1 \pm 4.7$	$0.99392 \pm 0.00071$	$-0.608 \pm 0.071$	$0.077229^{+0.000044}_{-0.000043}$



## Kepler-10:

Mass ( $M_{\text{sun}}$ )	$0.913 \pm 0.022$	(2.4%)
Radius ( $R_{\text{sun}}$ )	$1.065 \pm 0.009$	(0.85%)
Age (Gyr)	$10.4 \pm 1.4$	(13%)

## Kepler-10b:

$R_{\text{planet}}/R_{\text{star}}$	$0.01254 \pm 0.00013$	(1.0%)
$R_{\text{planet}}/R_{\text{Earth}}$	$1.451 \pm 0.019$	(1.3%)

**The key is to extend the length of the time series**