The ALMA Observing Tool

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EUROPEAN ARC ALMA Regional Centre

What is the ALMA OT?

- * JAVA-based application, can be run on any OS that has a recent version of JAVA (currently JAVA 6 and 7 supported) installed
- * Two flavours: web-start (recommended; updates done automatically) or tarball (more robust)
- * The ALMA OT is the one-stop shop for
 - Preparing ALMA Observing Proposals (Phase I)
 - ✓ Submitting proposals to the ALMA archive
 - ✓ Generating Scheduling Blocks (Phase II)
 - ✓ Submitting Phase II material to the archive
- Designed to allow users without in-depth technical knowledge of sub-mm/interferometry concepts to prepare high-quality proposals



Cycle2Test2

Phase I - Concepts

- OT software can be freely downloaded from the ALMA Science Portal. Then run locally, although intermittent internet connection required (e.g. database access, submission)
- * PIs prepare ALMA proposals containing
 - ✓ General information (title, abstract, PI and co-Is etc)
 - ✓ Science Case (4 page PDF)
 - ✓ One or several Science Goals (SGs contain technical information)
- * All PIs and Co-Is must be registered with the ALMA Science Portal
- * All proposals must pass a validation check before submission

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Project - Observing Tool for ALMA, version Cycle2Test2

Phase I: Science Proposal

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Phase I: Science Proposal

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- 2. Create a new proposal by either:
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Phase I – General info

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Phase I – General info

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Phase I – Science Goal(s)

- A Science Goal defines the scientific requirements of an observation
 - * Field setup (several point sources with offset pointings OR 1 mosaic)
 - * One Spectral setup (frequencies, spectral resolution...)
 - * Calibration requirements
 - * One set of Control & Performance parameters (spatial resolution and LAS, sensitivity required, timing constraints...)
- The correlator setup, array configuration(s), time required to achieve requested sensitivity, observing sequence etc. are then all defined by the OT based on the user input
- * Technical Justification to be included in each SG by PI

Phase I – Science Goal(s)

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Phase I – Spectral Setup

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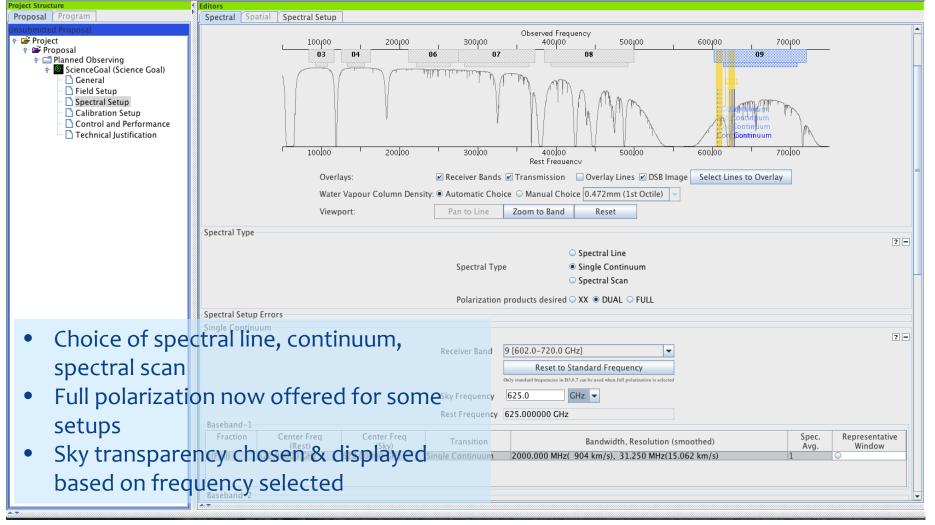
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Perspective 1

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The Spectral Line Picker

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Spectral Setup – Spectral Line

ALMA Observing Tool (Cycle2)

Project - Observing Tool for ALMA, version Cycle2Test2

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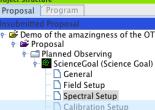
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Spectral Spatial Spectral Setup Visualisation

In the table below, it is possible to define up to 16 spectral windows, 4 per baseband as long as the total Fraction per baseband is no more than 1. Each baseband is 2GHz wide and can be separately configured i.e. each spectral window can have a different bandwidth and resolution. Note that for bands 3, 4, 6, 7 and 8, it is not possible to put 3 basebands in one sideband and the fourth one in the other.

Left/right click to zoom in/out, grab sliding bar to pan Note: Moving LO1 here is for experimentation only – actual setup determined by the windows

- Complicated spectral setups with up to 16 spectral windows with different resolution now allowed
- Can be set up with the spectral line picker + manually, and are Vapour Column Density: Automatic Choice Manual Choice 1.262mm (4th Octile) visualised in the graphical display and table

Non-valid spectral setups immediately give an error Spectral Line message

260,00 270,00 280. $180 \times = 0.6-5$ 260.00 270.00 280,00 Rest Frequency Receiver Bands Transmission Overlay Lines DSB Image Select Lines to Overlay Pan to Line Zoom to Band Reset ? -Spectral Line Spectral Type Single Continuum Spectral Scan ? -Center Free Spec epresentative Transition Bandwidth, Resolution (smoothed) (Rest) (Sky) Avg. Window 226.34036 GHz 225.37126 GHz CO v=2 2-1 117.188 MHz(156 km/s), 122.070 kHz(0.162 km/s) 227.00456 GHz 117.188 MHz(155 km/s), 122.070 kHz(0.162 km/s) 226.03262 GHz Si13CC 10(2.9)-9(2. Add Delete Select Lines to Observe in Baseband-1...

Spectral Setup – Spectral Scan

🗯 ALMA Observing Tool (Cycle2)

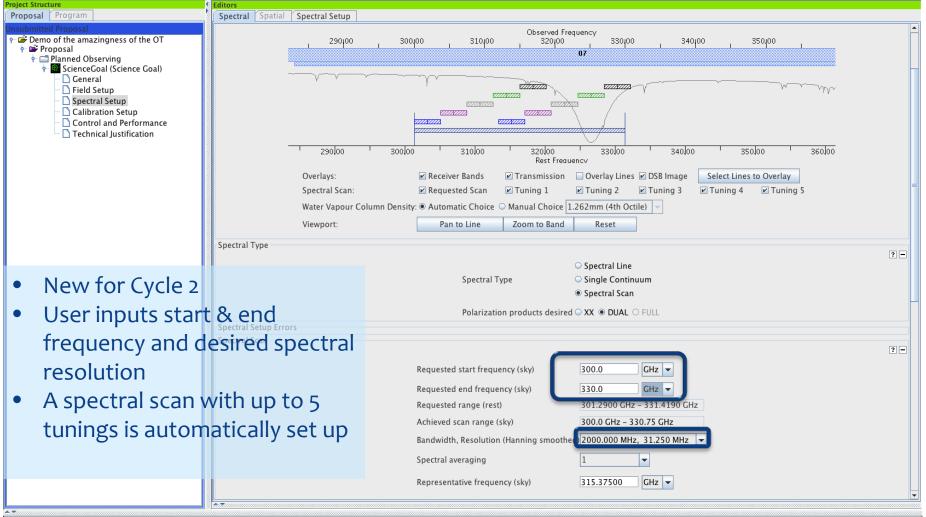
Project - Observing Tool for ALMA, version Cycle2Test2

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Phase I – Calibration Setup

MAX ALMA Observing Tool (Cycle2)

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Proposal Program

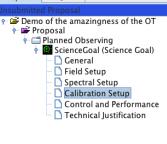
Project - Observing Tool for ALMA, version Cycle2Test2

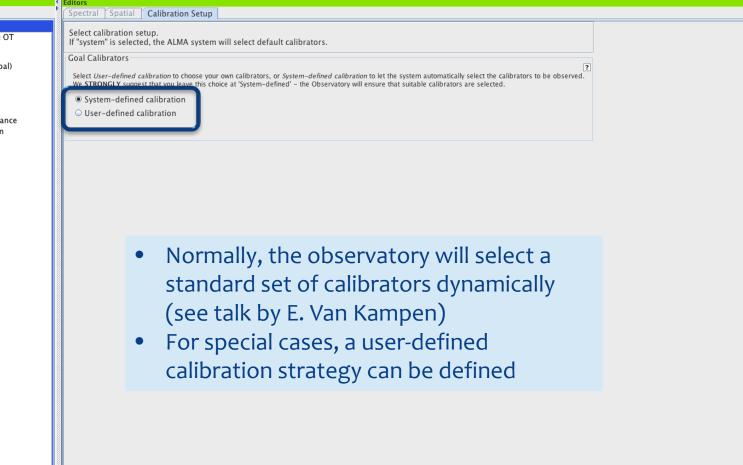
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ALMA Observing Tool (Cycle2)	Project - Observing Tool for ALMA, version Cycle2Test2
ile <u>E</u> dit <u>V</u> iew <u>T</u> ool <u>S</u> earch <u>H</u> elp	Persp
roject Structure Proposal Program	Editors Spectral Spatial Control and Performance
nsubmitted Proposal	These parameters are used to control various aspects of the observations, including the required antenna configurations and integration times.
P Demo of the amazingness of the OT P Proposal	Control and Performance
Planned Observing	Configuration Information
– 🗋 General – 🗋 Field Setup	Antenna Beamsize (1.2 * λ / D) 12m 27.438 arcsec 7m 47.036 arcsec
 Spectral Setup Calibration Setup 	Number of Antennas 12m 34 7m 9 TP 2
Control and Performance	Most extended 12m configuration Most compact 12m configuration Longest baseline (L _{max}) 1.50789000000002 km 165.641 m
Technical Justification	Synthesized beamsize (λ/L _{max}) 0.182 arcsec 1.656 arcsec
	Shortest baseline (L _{min}) 40.611 m 14.189 m
	Maximum recoverable scale (0.6λ/L _{min}) 4.054 arcsec 11.602 arcsec
	Desired Performance
	Desired Angular Resolution 1.00000 arcsec
	Largest Angular Structure in source O Point Source O Extended So 10.00000 arcsec 💌
	Desired sensitivity per pointing 0.10000 Jy v equivalent to 2.40725 K v
	Bandwidth used for Sensitivity RepresentativeWindowResolution V Frequency Width 0.122070 MHz
User inputs	
largest angu	Science goal integration time estimate Time Estimate
• •	is more uniterrequired due to d,v coverage issues: (indicible justified) or res lo inc
requested o	over a certain bandwidth ves 🛛 🔊
Choices are	guided by the configuration
information	n that is updated depending on
the capabili	ities of the array
the capabili	des of the array

Project Observing Tool for ALMA	version (velo)Test2	
Project - Observing Tool for ALMA,	Version Cycle2 Lest2	Perspective 1
tial Control and Performance		
rformance Information Immsize (1.2 * λ / D) 12m 27.438 arcsec Antennas 12m 34 Most extended 12m configura eline (L _{max}) 1.507890000000002 km beamsize (λ/L _{max}) 0.182 arcsec seline (L _{mix}) 40.611 m ACA Ne ACA use is not recommended The ALMA Observing Tool suggests this observation the largest recoverable angular scale of the 12m arra interview	7m 47.036 arcsec 7m 9 TP 2 ation Most compact 12m configuration 165.641 m 1.656 arcsec 14 189 m accessity Estimator does not need the ACA because ay configuration(s) exceeds the largest angular scale in source.	2
tegration time estimate dthe scenes quired due to u,v coverage issues? (must be justified) O	Time Estimate	
	ial Control and Performance ial Control and Performance ers are used to control various aspects of the observation formance Information imsize (1.2 * λ / D) 12m 27.438 arcsec intennas 12m 34 Most extended 12m configur. eline (L _{max}) 1.507890000000002 km beamsize (λ/L _{max}) 0.182 arcsec eline (Louc) 40 611 m ACA Ne ACA use is not recommended The ALMA Observing Tool suggests this observation the largest recoverable angular scale of the 12m arra int int complementary ACA Observations? complementary ACA Observations? </th <th>ial Control and Performance ers are used to control various aspects of the observations, including the required antenna configurations and integration times. formance Information msize (1.2 * λ/D) 12m 27.438 arcsec 7m 40 7m 9 TP Most extended 12m configuration Most compact 12m configuration Beamsize (N/Lmax) 0.182 arcsec 1.5078 9000000002 km 165.641 m beamsize (N/Lmax) 0.182 arcsec 1.656 arcsec eline (L_max) 40.611 m 14.189 m 40.611 m 14.189 m 40.611 m 14.189 m 40.611 m 14.189 m 67 ACA use is not recommended The ALMA Observing Tool suggests this observation <i>does not need</i> the ACA because The Hargest recoverable angular scale of the 12m array configuration(s) exceeds the largest angular scale in source. 0K Cancel of or Sensitivity RepresentativeWindowResolution ▼ Frequency Width 0.122070 MHz complementative ACA Observations? Yes No Suggest Time Estimate of the Sensitivity Representative WindowResolution ▼ Frequency Wid</th>	ial Control and Performance ers are used to control various aspects of the observations, including the required antenna configurations and integration times. formance Information msize (1.2 * λ/D) 12m 27.438 arcsec 7m 40 7m 9 TP Most extended 12m configuration Most compact 12m configuration Beamsize (N/Lmax) 0.182 arcsec 1.5078 9000000002 km 165.641 m beamsize (N/Lmax) 0.182 arcsec 1.656 arcsec eline (L_max) 40.611 m 14.189 m 40.611 m 14.189 m 40.611 m 14.189 m 40.611 m 14.189 m 67 ACA use is not recommended The ALMA Observing Tool suggests this observation <i>does not need</i> the ACA because The Hargest recoverable angular scale of the 12m array configuration(s) exceeds the largest angular scale in source. 0K Cancel of or Sensitivity RepresentativeWindowResolution ▼ Frequency Width 0.122070 MHz complementative ACA Observations? Yes No Suggest Time Estimate of the Sensitivity Representative WindowResolution ▼ Frequency Wid

ALMA Observing Tool (Cycle2) 🎖 👗 📿 🕘 🕴 💷 🚸 🔤 💽 (Charged) Mon 15:25 Suzanna Randall 🔍 Project - Observing Tool for ALMA, version Cycle2Test2 File Edit View Tool Search Help Perspective 1 1 1 2 2 4 🛋 🕞 🔚 🗐 🕸 🔅 🕾 🖂 📟 🖬 🖌 🗘 🌖 📀 Proposal Program Spectral Spatial Control and Performance These parameters are used to control various aspects of the observations, including the required antenna configurations and integration times. 🕆 🖨 Demo of the amazingness of the OT 👇 🖼 Proposal Control and Performance ← □ Planned Observing
 ? ALMA OT - Information ScienceGoal (Science Goal) Configuration Informat General Estimated time Antenna Beamsize (Field Setup Spectral Setup Number of Antenna TP 2 Requested sensitivity 100.0000 mJy Calibration Setup Bandwidth used for sensitivity 0.122 MHz tion Control and Performance Longest baseline (L Representative frequency (sky, first source) 225.37 GHz Technical Justification Precipitable water vapour (all sources) 1.262mm (4th Octile) Synthesized beams Shortest baseline (L ALMA 12m Array - 34 antennas 10.00 s Time on source per pointing (first source) Maximum recoveral Total number of pointings (all sources) 1 Total time on source 10.00 s Desired Performance Total time on calibrators 13.55 min Desired Angular Resol Total overheads 398.00 s Total 12m array time (inc. calibration & overheads) 20.35 min Largest Angular Struct arcsec 💌 **Calibration Breakdown** Desired sensitivity per 2 10725 К 🔻 Estimated number of tunings required 1 1.68 min 1 x SidebandRatio Bandwidth used for Se equency Width 0.122070 MHz 3 x Pointing 54.00 s 1 x Amplitude (inc. AtmosphericCal) 3.27 min Do you request comple 1 x Bandpass (inc. AtmosphericCal) 5.77 min Science goal integratic 1 x Phase (inc. AtmosphericCal) 1.27 min stimate 40.00 s 1 x Atmospheric Additional calibration overheads 6.53 min Is more time required Are the observations t Estimated total time for science goal 20.35 min OK

- OT gives an estimate of the observing time required to reach the requested sensitivity
- Overheads and calibrations included

ALMA Observing Tool (Cycle2)		🎖 👗 📿 Ð 🕴 🚥 🚸 些 💽 (Charged) Mon 15:33 Suzanna Randall 🔍
	Project - Observing Tool for ALMA, version Cycle2Te	
Eile Edit View Tool Search Help		Perspective 1
Proposal Program	Spectral Spatial Control and Performance	
Insubmitted Proposal	Shortest baseline (L _{min}) 40.611 m 14.1	89 m
P Demo of the amazingness of the O	Maximum recoverable scale (0.6λ/L _{min}) 4.054 arcsec 11.6	i02 arcsec
Planned Observing ScienceGoal (Science Goal)	Desired Performance	
- 🗋 General	Desired Angular Resolution 1.00000 arcsec	
– 🗋 Field Setup – 🗋 Spectral Setup		
Calibration Setup	Largest Angular Structure in source OPoint Source Extended S	o 10.00000 arcsec -
Control and Performan Technical Justification	Desired sensitivity per pointing 0.10000 Jy 👻 equiv	alent to 2.40725 K 💌
	Bandwidth used for Sensitivity RepresentativeWindowResolut	ion 🔻 Frequency Width 0.122070 MHz
	Do you request complementary ACA Observations? O Yes No	Suggest
	Science goal integration time estimate	Time Estimate
	Is more time required due to u,v coverage issues? (must be justified) • Yes O Nc Enter total time e	stimate 0.00000 min 👻
	Are the observations time-constrained?	ecific Dates O Multiple Epochs Continuous Monitoring
	Please specify the arrangment of visits for your Monitoring specified : 1	2
	observation.	Monitoring Constraints 32:06.510Z with a monitoring length of 3.0 h
	Monitoring can either be for a specific date or at an arbitrary date.	
• Time constr	ained observations now handled by	=
	do not enter a date it will be assumed that this visit	
OT	can be run at any time.	Add Delete
_		Add Time Range
Can specify	several possible time windows,	
multiple visi	ts, continuous monitoring (length	
· · · · · · · · · · · · · · · · · · ·	iming information time estimate based on	
	ig overndes time estimate based on	
sensitivity)		
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 Coordinated 	observations possible in the future	

Phase I – Technical Justification

ALMA Observing Tool (Cycle2)

Project - Observing Tool for ALMA, version Cycle2Test2

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Proposal Program	Spectral Spatial Technical Justification				
Unsubmitted Proposal	Enter a Technical Justification for this Scien	ce Goal, paying special attention to the parameter	s reproduced below.		^
🕂 🖼 Proposal	B revant science parameters				
🛉 🚍 Planned Observing 🛉 🚭 ScienceGoal (Science Goal)	Sensitivity	0.10 Jy Angular Resolution	? - 1.00 arcsec		
🚽 🗋 General	Bandwidth for sensitivity	122.07 kHz Largest angular scale	10.00 arcsec		
– 🗋 Field Setup – 🗋 Spectral Setup	Representative frequency	225.37 GHz ACA	No		
 Calibration Setup 	Expected source properties				
 Control and Performance Technical Justification 	Continuum:		? -		
	Aggregate bandwidth	3.81 GHz Sensitivity	77.53 mJy		
	Peak flux density	2.00 mJy SNR	0.78		
	Line:				
	Peak flux density	0.50 Jy SNR	5.00		
	Line width	20.00 km/s Resolutions per FWHM	985		
	Oynamic Range (cont. peak/line rms)	0.03			
			? -		
	Field setup:	 Input and exp 	ected perfor	mance parameters	=
	Spectral Setup:	echoed back t			
	Calibration:	echoed back t	ouser		
	Control and Parameters:	 Non-standard 	choices are l	isted and must be	
	* Time Constrained	Non Standard	choices are i	isted and must be	
	* High Data rate (> 6 MB /s)	fully justified			
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	This proposal is the te is definitely feasible.	chnically most amazing proposal in history and			
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Phase I – Validation & Submission

🗯 ALMA Observing Tool (Cycle2)			🎖 👗 📿 🤨 🕴 💷 🚸 💷 💽 (Charged) Mon 15:58 Suzanna	a Randall 🔍
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<u>File Edit View T</u> ool <u>S</u> earch <u>H</u> elp				Perspective 1
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Project Structure	Editors			
	Spectral Spatial Technical Justificatio	n		
	Spectral Spatial Technical Justification Enter a Technical Justification for this Scie Relevant science parameters Sensitivity Bandwidth for sensitivity Representative frequency Expected source properties Continuum: Aggregate bandwidth Peak flux density Line: Peak flux density Line width Dynamic Range (cont. peak/line rms) Non-standard choices Field setup: Spectral Setup: Calibration: Control and Parameters: * Time Constrained * High Data rate (> 6 MB /s) * High Data rate (> 6 MB /s)	ence Goal, paying special attention to the parameters rep 0.10 Jy Angular Resolution 122.07 kHz Largest angular scale 225.37 GHz ACA 3.81 GHz Sensitivity 2.00 mJy SNR 2.050 Jy SN	<pre> 1.00 arcsec 10.00 arcsec No P 77.53 mJy 0.78 project at any time to get npleteness/technical problem performed upon submission posals that have validation </pre>	
	1 error, 1 warning No document found – you must add a S 12M array data rate is 32.17 MB/s whice		Suggestion he proposal node in the Proposal tab and add your document essive data rate requires scientific justification	

Beyond Phase I

- * Re-submission of projects possible up until the deadline
- All projects submitted by the deadline are ranked based on their scientific merit
- From Cycle 2 the ranking will follow the ABCD grades as in place for ESO proposals
- Technical feasibility is evaluated if not technically feasible, the proposal is rejected
- Proposals judged likely to be executed pass to Phase II they then appear in the Program tab of the OT
- Scheduling Blocks (SBs) are generated semi-automatically by a member of staff using the OT, and are then approved (often after iteration & fine-tuning) by the PI
- * After PI approval, the Phase II SBs are submitted to the archive by ALMA staff and set to Ready in the Project Tracker software

Phase II – Scheduling Blocks

- For each Science Goal an ObsUnitSet containing one or more SBs is created simply by pressing a button, then some checks & tweaks are done manually
- * Normally one SB per array (12-m, ACA, TP)
- An SB is the smallest executable unit and contains all the technical details of the science observations as well as associated calibrations
- * SBs normally run for 1-1.5 hours and are executed repeatedly until the time on source (calculated by OT based on sensitivity or specified by user for monitoring) has been reached

Phase II – Scheduling Blocks

ALMA Observing Tool (Early Science)

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Spatially extended [CII] in a z=4.8 SMG - Observing Tool for ALMA, version Cycle1Phasell

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Spectral Spatial Group 1 : Calibrators
Observing Group Name Calibrators

Proposa Program Control and Performance LESS J033229.4-275619-SI LESS J033229.4-275619[12m Array SB] Group 1 : Calibrators 14 Targets Titan Amplitude (Amplitude) 1337 Pointing for Amplitude (Titan) (Pointing) 0402 Bandpass (Science) (Pointing) B0402 Phase (Pointing) 0402 Bandpass (Science) (Bandpass) B0402 Phase (Phase) Titan Amplitude (Atmospheric) 80402 Bandpass (Science) (Atmospheric) B0402 Phase (Atmospheric) [R] LESS J033229.4-275619 Primary: (Science) LESS J033229.4-275619 Primary: (Atmospheric) Ceres Amplitude (Amplitude) 30507 Pointing for Amplitude (Ceres) (Pointing) Ceres Amplitude (Atmospheric) Resources 9 8 Field Sources Primary: LESS J033229.4-275619 Amplitude Titan Pointing for Amplitude (Titan) J1337 Bandpass (Science) B0402 Phase B0402 Amplitude Ceres Pointing for Amplitude (Ceres) B0507 Amplitude Titan 6 Instrument Setup CII_red Science setup[12m Array SB] (4 BBCs) Band 3 Pointing setup[12m Array SB] (4 BBCs) TDM Atm Cal setup[12m Array SB] (4 BBCs) CII_red Science setup[12m Array SB] (4 BBCs) Band 3 Pointing setup[12m Array SB] (4 BBCs) TDM Atm Cal setup[12m Array SB] (4 BBCs) - 🛃 9 Observing Parameters LESS J033229.4-275619 Params PhaseCalParameters PointingCalParameters PointingCalParameters AmplitudeCalParameters AmplitudeCalParameters AtmosphericCalParameters AtmosphericCalParameters BandpassCalParameters

		All Ava	allable largets		
Source Name	RA	DEC	Spectral Spec	Rest Freq	Purpose
Titan	00:00:00	00:00:00	CII_red Science set	329.5 GHz	AmpCal
J1337	13:37:39	-12:57:24	Band 3 Pointing se	98.0 GHz	PntCal
B0402	04:03:53	-36:05:01	Band 3 Pointing se	98.0 GHz	PntCal
B0402	04:03:53	-36:05:01	Band 3 Pointing se	98.0 GHz	PntCal
B0402	04:03:53	-36:05:01	CII_red Science set	329.5 GHz	BndCal
B0402	04:03:53	-36:05:01	CII_red Science set	329.5 GHz	PhsCal
Titan	00:00:00	00:00:00	TDM Atm Cal setup	329.5 GHz	AtmCal
B0402	04:03:53	-36:05:01	TDM Atm Cal setup	329.5 GHz	AtmCal
B0402	04:03:53	-36:05:01	TDM Atm Cal setup	329.5 GHz	AtmCal
LESS J0332	03:32:29	-27:56:19	CII_red Science set	329.5 GHz	Science
LESS J0332	03:32:29	-27:56:19	TDM Atm Cal setup	329.5 GHz	AtmCal
Ceres	00:00:00	00:00:00	CII_red Science set	329.5 GHz	AmpCal
B0507	05:10:02	18:00:41	Band 3 Pointing se	98.0 GHz	PntCal
Ceres	00:00:00	00:00:00	TDM Atm Cal setup	329.5 GHz	AtmCal

All Assollable Terrar

Observing Group Targets

	Index	Source Name	RA	DEC	Spectral Spec	Rest Freq	Purpose
	1	J1337	13:37:39	-12:57:24	Band 3 Pointing s	98.0 GHz	PntCal
	2	Titan	00:00:00	00:00:00	TDM Atm Cal setup	329.5 GHz	AtmCal
	3	Titan	00:00:00	00:00:00	CII_red Science se	329.5 GHz	AmpCal
	4	B0402	04:03:53	-36:05:01	Band 3 Pointing s	98.0 GHz	PntCal
	5	B0402	04:03:53	-36:05:01	TDM Atm Cal setup	329.5 GHz	AtmCal
	6	B0402	04:03:53	-36:05:01	CII_red Science se	329.5 GHz	BndCal
	7	B0507	05:10:02	18:00:41	Band 3 Pointing s	98.0 GHz	PntCal
	8	Ceres	00:00:00	00:00:00	CII_red Science se	329.5 GHz	AmpCal
_	9	Ceres	00:00:00	00:00:00	TDM Atm Cal setup	329.5 GHz	AtmCal

Delete ↑ ↓

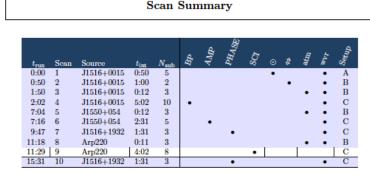
 SBs contain a rather complicated set of parameters fixing the technical details of the observations that are then parsed to scripts executing the observations at the telescope

The Observing Script Simulator

- The presentation of SBs in the OT can be confusing to novice users; in particular the observing sequence and potential calibrators cannot be easily understood
- The Observing Script Simulator simulates what will happen at the telescope assuming a given execution date
- For each SB, it outputs the scan sequence and a time-breakdown

OSS Summary of OSS-453-0-v1.txt

BP=Bandpass Cal, **AMP**=Amplitude Cal, **PHASE**=Phase Cal, **SCI**=Science Target, \odot =Pointing, \Leftrightarrow =Sideband Ratio, **atm**=T_{sys}, **wvr**=Water Vapour Radiometry



Setups

A B6 Pointing Setup

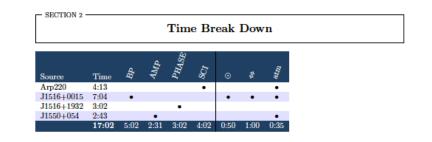
B Tsys Setup #1

C Manual window Science setup

Science Setup:

SECTION :

Band 6 SPWs 284.6, 286.2, 296.8, 298.4 GHz



The Observing Script Simulator

- The OSS also queries the ALMA calibrator data base and shows which calibrators are likely to be selected at run-time
- In the future, the OSS will be incorporated into the OT to make things more user-friendly
- Eventually, PIs will be able to create, submit and set to Ready their own SBs

- SECTION 3 -

Calibrator Query

3.1 Bandpass Calibrator

The OSS performs a search of possible bandpass calibrators within 45.0 deg from the science target. A SNR of 50.0 per channel must be reached within 3.0 to 15.0 min, which conforms to a limiting flux density of 718 mJy (90 mJy if smoothing to 64ch in FDM). The minimum elevation is set to 56.0 deg.

The OSS query returned the following targets (the final choice is marked in bold):

Source	Ang.Sep.	Flux Density	Ref.Frequency
J1512-0905	33.0°	1.8 Jy	284.6 GHz
J1516+0015	23.7°	0.53 Jy	284.6 GHz
J1513-1012	34.1°	0.52 Jy	284.6 GHz

3.2 Phase Calibrator

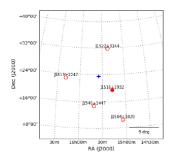
The OSS performs a search of possible phase calibrators within 15.0 deg from the science target. A SNR of 15.0 over the full bandwidth must be reached within 0.5 to 2.0 min, which conforms to a limiting flux density of 18.4 mJy.

The OSS query returned the following targets (the final choice is marked in **bold**):

Source	Ang.Sep.	Flux Density	Ref.Frequency
J1516+1932	5.8°	0.45 Jy	284.6 GHz
J1540+1447	8.8°	0.4 Jy	284.6 GHz
J1504+1029	14.9°	0.36 Jy	284.6 GHz
J1619+2247	10.2°	0.12 Jy	284.6 GHz
J1522 + 3144	8.7°	0.1 Jy	284.6 GHz

3.3 Amplitude Calibrator

source	size	baselines*			
Pallas	0.2"	3 / 6			
Ceres	0.5"	1/6			
Titan	0.8"	1/6			
Callisto	1.2"	1/6			
Ganymede	1.3"	1/6			
Neptune	2.2"	1/6			
Uranus	3.3"	1/6			
Mars	3.9"	1/6			
Venus	9.7"	0/6			
Jupiter	34.9"	0/6			
* fraction of baselines that are short					
enough to not resolve out the source					



No suitable Solar System Object found. Choosing Grid Source J1550+054 instead.

Conclusion

The OT is constantly evolving – many new capabilities still to come

- A huge effort was made to make this tool as self-explanatory and user-friendly as possible
- * There is extensive documentation available
- * Any suggestions for improvement / comments are welcome!